Inverse Bremsstrahlung Absorption


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Inverse bremsstrahlung absorption was measured based on transmission through a finite-length plasma that was thoroughly characterized using spatially resolved Thomson scattering. Expected absorption was then calculated using the diagnosed plasma conditions while varying the absorption model components. To match data, it is necessary to account for (i) the Langdon effect; (ii) laser-frequency (rather than plasma-frequency) dependence in the Coulomb logarithm, as is typical of bremsstrahlung theories but not transport theories; and (iii) a correction due to ion screening. Radiation-hydrodynamic simulations of inertial confinement fusion implosions have to date used a Coulomb logarithm from the transport literature and no screening correction. We anticipate that updating the model for collisional absorption will substantially revise our understanding of laser-target coupling for such implosions.

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In long-pulse laser-plasma experiments, inverse bremsstrahlung (IB) absorption is typically the dominant mechanism coupling laser energy to the plasma, and is often the only mechanism taken into account in models. This is true for the conventional approaches to every branch of inertial confinement fusion. Despite its importance, there remains no consensus on how to calculate IB absorption. In particular, absorption rates are proportional to the Coulomb logarithm, also known as the Gaunt factor, which has myriad definitions [1–9]. Moreover, calculations often include a Langdon absorption-reduction factor which assumes that laser-heated electron distribution functions (EDFs) become super-Gaussian [10]. Although recent measurements have confirmed the existence of such EDFs [11,12], attempts to validate the absorption-reduction factor itself have been fewer.

In this Letter, we present measurements of IB absorption and compare to predictions using various models. The precision of both the absorption measurements and the spatially resolved plasma conditions that serve as inputs to the model predictions enables us to discriminate between theories. We find that accounting for the Langdon effect is essential, even for moderate intensities and low-Z ions. We also find that the maximum impact parameter in the Coulomb logarithm for IB absorption should depend on the laser frequency rather than the plasma frequency, which is consistent with most bremsstrahlung-specific models [1–5,7] but different from Coulomb logarithms used for transport processes such as electron-ion temperature relaxation or heat conductivity [6,8]. Finally, even for the very underdense plasmas studied here \( n_e/n_c < 0.04 \), where \( n_c \) is the electron density normalized by the critical density \( n_\text{c} \), it is necessary to include a correction for ion screening, and we propose a suitable form.

The measurements were obtained on the OMEGA laser system at the University of Rochester’s Laboratory for Laser Energetics using the laser-plasma instability platform. While the setup for the absorption campaigns was described more fully in Ref. [13], we briefly summarize the essential ingredients here. A supersonic gas jet with a 1-mm-outlet-diameter nozzle emitted a narrow column of gas at target chamber center. The gas was ionized and preheated by thirteen 351-nm beams at \( t = 0 \) ns, each of which had a 500-ps-square pulse shape. The energy in each beam was a primary variable and ranged from 35 to 190 J per beam. A 527-nm probe beam with energy ranging from 1 to 4 J in a 100-ps Gaussian picket pulse was then injected from the P9 port into the plasma. Thomson-scattered light from that probe beam was collected using the imaging Thomson scattering system [14], which spatially resolved the plasma conditions over a distance of about 1.5 mm. Other campaign variables included probe timing (0.3 ns—during the heater beams, or 0.6 ns—after the heater beams, where the timing refers to the point at which the laser power reaches 2% of the peak power); gas composition (H\(_2\), CH\(_4\), or a mixture of 45% N\(_2\) and 55% H\(_2\)); and gas jet backing pressure to vary the peak plasma density (0.0037 < \( n_e/n_\text{c} < 0.036 \)).

The transmission, \( T \), of the probe beam was determined by comparing an upstream measurement (of an uncoated wedged pickoff located in the P9 beam just before the target chamber) to a downstream measurement after target chamber center using the P9 transmitted beam diagnostic.
(TBD) [15]. The ratio of these detectors was calibrated each day using a pair of shots through vacuum with an otherwise identical setup, so that data shots differed only by the presence of the absorbing plasma. Special care was taken to remain below threshold for potentially confounding sources of loss such as backscatter (monitored and maintained below a noise floor of $\approx 0.1\%$) and beam spray [16], as well as possible absorption-enhancement mechanisms like the return-current instability [17]. This was achieved by minimizing the probe intensity (while maintaining good Thomson-scattering signals) and peak plasma density (while still getting measurable absorption), and by selecting gases with a high hydrogen fraction to ensure strong ion-acoustic-wave damping. With competing losses kept small, the absorption $\kappa$ was simply $1 - T$, with an estimated uncertainty of $(+0.07\%, -0.12\%)$ [13].

Other details of the setup varied slightly between three configurations. In config. No. 1 [Fig. 1(a)], all beams were pointed 1 mm above the exit face of the gas jet nozzle. The preheating beams were conditioned with full smoothing by spectral dispersion (SSD) and elliptical E-IDI-300 distributed phase plates (DPPs) oriented to complement the SSD such that each beam’s profile was approximately round with a FWHM diameter just over 200 $\mu$m. The probe beam was conditioned with a CircSG100 DPP to have a 97-um-FWHM diameter, and the Thomson scattering system field-of-view (TSS FOV) was centered around target chamber center. In config. No. 2, SG5-850 DPPs (714-um-FWHM) conditioned the preheating beams without SSD to heat the entire plasma more uniformly, all beams were pointed 1.5 mm from the gas jet to improve beam clearance and reduce noise on the ion-feature Thomson scattering, and the TSS FOV was biased 300 $\mu$m toward the probe entrance to better capture the low-density tail on one side of the plasma. Cylindrical axisymmetry of the gas or plasma density about the gas jet axis allowed us to mirror those measurements onto the probe exit side while introducing minimal error. The probe was conditioned with a CircSG200 DPP to have a 165-um-FWHM diameter. Finally, config. No. 3 was identical to config. No. 2 except for the use of a refractive (rather than reflective) Thomson-scattering collection telescope; its larger (> 4 mm) FOV eliminated the need for the axisymmetric assumption.

A sample characterized plasma profile is shown in Fig. 1(b). For the shots that included mid-Z ions, the ionization state was determined using FLYCHK [18]. Note that the plasmas were finite in length and measured almost in their entirety, so the absorption calculations are well constrained.

We can then use the measured plasma conditions to predict how much absorption should occur as the probe beam propagates through the plasma. Including corrections for both the Langdon effect ($f_L$) and ion screening ($f_{sc}$), the absorption rate (with units m$^{-1}$) is

$$\kappa = \nu e_i \frac{n_e/n_c}{c\sqrt{1 - n_e/n_c}} f_L f_{sc},$$  

where $\nu e_i = 2.91 \times 10^{-12} T_e^{-3/2} \sum_i (Z_i^2 n_i \ln \Lambda_{IB,i})$ is the electron-ion collision rate including the summation over all ion species. (SI units are used throughout this Letter with the exception of temperatures in eV.)

The Coulomb logarithm generally arises from the need to bound an integral in the classical derivation for the electron-ion collision rate using maximum and minimum impact parameters $b_{\text{max}}$ and $b_{\text{min}}$, but one or both are ostensibly constrained depending on the derivation. While the following list is by no means exhaustive, it aims to include some of the better known theories while also spanning some key differences across all theories.

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Sommerfeld-Maue [1, 19]: 

$$\kappa = \frac{\pi}{\sqrt{3}} \langle g_{ff} \rangle \int \left( \frac{2}{e^2} \right)^{5/2} \frac{\nu e_i}{\omega} \max \left( \frac{Z e_i}{4\pi e_i T_e}, \frac{h}{2\sqrt{3}m_e T_e} \right)$$  

Oster [2]: 

$$\kappa = \ln \left( \frac{2}{\nu e_i} \right)^{5/2} \frac{\nu e_i}{\omega} \max \left( \frac{Z e_i}{4\pi e_i T_e}, \frac{h}{2\sqrt{3}m_e T_e} \right)$$  

Lee-More [6]: 

$$\kappa = \ln \left( \frac{\lambda_{D,e}}{12\pi e_i T_e}, \frac{h}{2\sqrt{3}m_e T_e} \right)$$  

FIG. 1. An example experimental setup (config. No. 1) is shown in (a). An example of the measured plasma conditions is shown in (b). For density and temperature, solid markers denote points that were explicitly measured by Thomson scattering. Open markers indicate extrapolations based on the measured profile.
Sommerfeld’s seminal result—a function of the EDF $f_e(v)$, laser frequency $\omega$, and $Z$—is an exact quantum-mechanical solution for the bremsstrahlung emitted in a single binary electron-ion collision, but it is in terms of complicated hypergeometric functions that can be computationally prohibitive to evaluate [1]. However, there have been many approximations to Sommerfeld over the ensuing century [19–24]. To estimate Sommerfeld in our analysis, we use Pradler’s Eq. (12) for the thermally averaged Gaunt factor, which in turn uses their Eq. (10) for the fitting formula to the velocity-dependent Gaunt factor [19]. Oster [Eq. (3)] was also derived from a binary-collision approach, but additional assumptions (straight-line approximation to the hyperbolic trajectory) facilitated a much simpler formula [2]. In the leading order-unity term, $\gamma = 0.577$ is Euler’s constant. It is otherwise written in terms of a maximum impact parameter $v_i/\omega$ where $v_i = \sqrt{T_e/m_i}$ is the electron thermal velocity, and a minimum impact parameter that is either classical (the distance of closest approach) or quantum-mechanical (related to the thermal de Broglie wavelength). In the quantum limit, it simplifies considerably to $\ln(4T_e/\hbar \omega e)$, which is found elsewhere in the literature [7,25,26]. Though not shown, the well-known formula for the high-frequency limit (when $\omega \gg \omega_p$, where $\omega_p$ is the plasma frequency) from Dawson-Oberman [3], later revised by Johnston [5] and adopted by the NRL plasma formulary [27], has the same format as Oster but lacks the order-unity terms to reflect the indeterminacy of $b_{\text{min}}$ in that derivation, although $b_{\text{max}}$ was rigorously derived. They could instead have chosen minimum impact parameters in order to match the results of Oster. The justification offered in the bremsstrahlung literature for this common $b_{\text{max}}$ is that the collision time must be short compared to the period of the wave, otherwise the heating is rendered ineffective. (Note that—given the laser-frequency dependence—the Coulomb logarithm is expected to be specific to each beam in the case of overlapping beams with different frequencies.)

For transport processes such as electron thermal conductivity [6] or electron-ion temperature relaxation [8], however, it is typical to assume a Debye length for the maximum impact parameter, with Lee-More [Eq. (4)] allowing ions to participate in the shielding (giving $\lambda_{D,el} = \sqrt{(e_0 T_e T_i/n_e e(ZT_e + T_i))}$ whereas Dimonte-Daligault [Eq. (5)] concluded that ions do not participate (therefore $\lambda_{D,e} = \sqrt{(e_0 T_e/n_e e)} = v_i/\omega_p$). Although transport Coulomb logarithms are not necessarily relevant to IB absorption, the community has not maintained a clear distinction between the two, and as we will discuss later, Lee-More has long been used to compute IB absorption in the codes that simulate inertial confinement fusion experiments. The various theories also differ by other numerical factors inside the logarithm.

The Langdon absorption-reduction factor is

$$f_L = 1 - 0.553/[1 + (0.27/\alpha)^{0.75}],$$

where $\alpha = Z_{\text{eff}} v_o^2/v_i^2$, $Z_{\text{eff}} = \langle Z^2 \rangle / \langle Z \rangle$, and $v_o$ is the oscillatory velocity of electrons in the laser field [10]. The factor assumes that EDFs become increasingly non-Maxwellian at high laser intensity. Matte et al. later showed using Fokker-Planck simulations that the bulk electrons relevant to IB absorption conform to a super-Gaussian EDF with an exponent given by $m(\alpha) = 2 + 3/(1 + 1.66/\alpha^{-0.724})$, and they stated that such EDFs are consistent with the Langdon factor to within 1% for any $\alpha$ [28]. While we have consistently verified the Matte formula when we have sensitivity to the EDF with Thomson scattering [11,12], the absorption-reduction factor itself has not yet been directly validated. All of the Thomson-scattering analysis here assumes super-Gaussian EDFs with $\alpha$ determined by the instantaneous overlapped intensity, i.e., the probe only (at $t = 0.6$ ns) or an effective intensity summing the probe and the preheating beams [multiplied by $(2/3)^2$ in order to be “$2\omega$ equivalent” since $v_o \propto \omega^{-1}$ in the Langdon factor] if co-timed.

Finally, although the Coulomb logarithms based on binary collisions (Sommerfeld, Oster) neglected medium effects, and we can conclude from Dawson’s work that medium effects are not important in the high-frequency limit [3,29], it seems apparent from molecular dynamics (MD) simulations [9] and partial wave calculations of screened Coulomb potentials [30] that ion screening becomes important for densities of interest above $n_e \gtrsim 0.01 n_c$. To account for this, we define the screening correction factor

$$f_{\text{sc}} = \frac{\ln((2 T_e/m_o \sqrt{1+y^2}) y)}{\ln((2 T_e/m_o \sqrt{1+y^2}) 1)},$$

where $y = l_{ic}/b_{\text{max}} = l_{ic} \omega/v_i$ is the ratio of the effective screening length to the maximum impact parameter. This term is Rosznyai’s (simplified) Gaunt factor [31] normalized by the Gaunt factor derived in the Born approximation without screening, and is valid for $y \gtrsim 0.005$ and $\hbar \omega \ll T_e$, which applies to most laser-produced plasmas. Although the Debye length is the most common choice for the screening length in weakly coupled plasmas, we choose instead the mean interionic distance, or Wigner-Seitz radius, $l_{ic} = a_i = (3/4 \pi n_i)^{1/3}$ where $n_i$ is the average ion density, in order to match the data. This factor reproduces the differences between screened and unscreened Coulomb potentials for the weakly coupled examples in Ref. [30] very well. A qualitative explanation for the impact of screening in the case when $a_i < b_{\text{max}}$ is that electrons interact with more than one ion simultaneously when the ions are relatively closely

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FIG. 2. Ratios of predicted-to-measured absorption for twenty-two distinct OMEGA shots (a) without the Langdon factor or screening correction; (b) with Langdon but no screening; or (c) accounting for both the Langdon effect and ion screening, along with the uncertainty. Only Oster is shown here so that the uncertainty can be seen more clearly, but the agreement is comparable for Sommerfeld.

The differences between Sommerfeld and Oster are slight. When the quantum-mechanical minimum impact parameter was invoked (most cases of hydrogen), the two models were almost identical. Differences arose when the classical minimum impact parameter was invoked (all of the carbon and nitrogen collisions), suggesting close collisions that violate Oster’s straight-line approximation are more important in those cases. For direct-drive inertial confinement fusion coronal plasma conditions, the quantum $b_{\text{min}}$ is more relevant so Oster should suffice as a good approximation to Sommerfeld.

The average value of $f_{\text{sc}}$ was 0.89, so the ion-screening correction factor is the final ingredient that brings predictions into agreement with the measurements [Fig. 2(c)], with Oster being on average just 2% below the data, and Sommerfeld just 2% above. We therefore conclude that Sommerfeld, with corrections for the Langdon effect and ion-screening, is likely the best model for IB absorption, but Oster reasonably approximates Sommerfeld in most situations. Had we used $f_{\text{sc}}(\lambda_D,c)$ instead of $f_{\text{sc}}(a_i)$, screening would have been a much smaller (%-level) correction—not enough to match the data.

The implications for direct-drive inertial confinement fusion are significant. Accurately predicting time-dependent laser-target coupling has been a long-standing challenge. Early on, simulations were tuned using a variable flux limiter in the heat transport model. Later, when the nonlocal model removed the flux-limiter knob, a crossed-beam energy transfer (CBET) package was added to better match the data [34]. When discrepancies persisted, multipliers were added to the CBET model due to a legacy belief that laser-plasma instabilities are not well understood and difficult to model quantitatively [35]. However, focused experiments have shown that CBET can be calculated reliably when the plasma conditions are well known [11,36,37].

While the Langdon factor is generally used, radiation hydrodynamics have long defaulted to using Lee-More...
without any ion-screening correction for IB absorption [38–40]. Figure 3 shows the impact of the proposed model changes using the well-known shot 90288 (described in Ref. [41]) as an example. Switching only from Lee-More to Oster would modestly redistribute absorption from lower density to higher density, which could actually benefit the implosion. However, the ion-screening correction systematics, but we expect that the same mechanisms (and no additional mechanisms) remain in play for IB absorption up to critical density. We therefore believe that a revised model motivated by the results presented here will substantially modify our understanding of direct-drive inertial confinement fusion.

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