Intrinsic reflection asymmetry in atomic nuclei

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The experimental and theoretical evidence for intrinsic reflection-asymmetric shapes in nuclei is reviewed. The theoretical methods discussed cover a wide spectrum, from mean-field theory and its extensions to algebraic and cluster approaches. The experimental data for nuclear ground states and at low and high spin, cited as evidence for reflection asymmetry, are collected and categorized. The extensive data on electric dipole transition moments and their theoretical interpretation are surveyed, along with available data on electric octupole moments. The evidence for reflection-asymmetric molecular states in light nuclei is summarized. The application of reflection-asymmetric theories to descriptions of the fission barrier, bimodal fission, superdeformation, and hyperdeformations is reviewed, and some other perspectives in the wider context of nuclear physics are also given. [S0034-6861(96)00102-X]

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FIG. 1. The low-lying rotational spectra of ²²⁴Ra, compared with that of the H³⁵Cl molecule. The spectrum of ²²⁴Ra is taken from Poynter *et al.* (1989a). The rotational constants for the H³⁵Cl molecule are taken from Landolt-Börnstein (1974).

I. INTRODUCTION

The existence of nuclei with stable deformed shapes was realized early in the history of nuclear physics. The observation of large quadrupole moments led to the suggestion that some nuclei might have spheroidal shapes, which was confirmed by the observation of rotational band structures and measurements of their properties. For most deformed nuclei, a description as an axial- and reflection-symmetric spheroid is adequate to reproduce the band's spectroscopy. Because such a shape is symmetric under space inversion, all members of the rotational band will have the same parity; however, with the first observation of negative-parity states near the ground state by the Berkeley group in the 1950s (Asaro et al., 1953; Stephens et al., 1955), the possibility arose that some nuclei might have a shape asymmetric under reflection, such as a pear shape (Strutinsky, 1956; Lee and Inglis, 1957).

Extensive investigations into the structure of nuclei with low-lying negative-parity states has led to the conclusion that, while reflection-asymmetric shapes can play a role in the band structure, they are not as stable as the familiar quadrupole deformations. The situation is illustrated by the low-lying spectrum of a representative case, ²²⁴Ra. Its spectrum is compared to that of the linear molecule HCl in Fig. 1. The molecule's reflection asymmetry permits both parities in its rotational spectrum, perfectly interleaved according to the energy formula $E_I \sim J(J+1)$. The spectrum of ²²⁴Ra has both parities as well, but in two bands that are displaced from each other. Nevertheless, there are good reasons for describing such a system as a common band. Experimentally, not only the proximity of energies, but also the dipole transitions between the subbands are characteristic of an intrinsic shape having an electric dipole moment. The displacement of the two parities means that fluctuations of shape back to symmetry must also be significant. In light nuclei, the resonances observed in collisions and other reactions may also be interpreted in terms of "molecular states" of alpha particles and other light clusters, which lack reflection symmetry. Analogies have also been drawn between molecular and baryonic spectra, describing baryons as symmetric or asymmetric tops (Iachello, 1989).

This review addresses both experimental and theoretical progress in this area. On the experimental side, spectroscopy has been extended to very high spins, and indicative transitions, including dipole and octupole, can now be measured. On the theoretical side, following an early phase of phenomenological theory, we have moved to an era where stable shapes can be predicted and understood from self-consistent mean-field theory, and furthermore the softness of the reflectionasymmetric deformation that was apparent in Fig. 1 is described by a more sophisticated theory going beyond the mean field.

The organization of this review is as follows. Section II contains the definitions of parameters used to define the nuclear shape in terms of both mass and charge distribution. Section III summarizes in detail various theoretical approaches, some of which have been reviewed earlier (Leander, 1985a; Nazarewicz, 1985; Rohoziński, 1988; Åberg et al., 1990). Sections IV and V collect experimental data pertaining to energy levels and transitions, and ground-state properties, respectively; for previous reviews of experimental systematics, see Leander and Sheline (1984), Leander and Chen (1988), Jain et al. (1990), Ahmad and Butler (1993), and Sheline (1993a). Section VI explores the properties of the rotating nuclear reflection-asymmetric shape and discusses them in terms of experimental observations. Section VII examines theoretical models of the electric dipole transition moment, an important observable in this context; for an earlier review see Butler and Nazarewicz (1991). Section VIII presents a survey of molecular structures in nuclei in the sd and fp shell, and of theoretical methods used to describe these light systems. Section IX discusses how the concept of octupole mass deformation has been applied to the description of fission and (more recently) to super- and hyperdeformed nuclear states. Finally, Sec. X offers a broader perspective on the consequences of nuclear reflection asymmetry and on future avenues of investigation.

II. DEFINITIONS

The anisotropy of nuclear shape is described in terms of intrinsic moments and deformation parameters. Shape parametrizations and intrinsic moments specific to reflection-asymmetric nuclei (e.g., octupole deformed) are discussed below in Sec. II.A (isoscalar moments) and Sec. II.B (isovector electric dipole moment).

A. Reflection-asymmetric shapes

In many applications, the nuclear shape is parametrized in terms of a spherical harmonic (multipole) expansion. The spheroidal nuclear surface is defined by means of standard deformation parameters $\alpha_{\lambda\mu}$ describing the length of the radius vector pointing from the origin to the surface (Bohr, 1952; Hill and Wheeler, 1953):

$$R(\Omega) = c(\alpha) R_0 \left[1 + \sum_{\lambda=2}^{\lambda_{\max}} \sum_{\mu=-\lambda}^{+\lambda} \alpha_{\lambda\mu} Y^*_{\lambda\mu}(\Omega) \right], \qquad (1)$$

with $c(\alpha)$ being determined from the volumeconservation condition and $R_0=r_0A^{1/3}$. The requirement that the radius be real imposes the condition

$$(\alpha_{\lambda\mu})^* = (-)^{\mu} \alpha_{\lambda-\mu}.$$
⁽²⁾

The three dipole deformations, $\alpha_{1\pm 1}$ and α_{10} , are given by the constraint that fixes the center of mass (c.m.) at the origin of the body-fixed frame:

$$\int_{V} \mathbf{r} d^{3} \mathbf{r} = 0, \tag{3}$$

where V is the total volume enclosed by the surface defined in Eq. (1). For shapes axially symmetric with respect to the z axis, all deformation parameters with $\mu \neq 0$ vanish. The remaining deformation parameters $\alpha_{\lambda 0}$ are usually called β_{λ} :

$$\beta_{\lambda} \equiv \alpha_{\lambda 0}. \tag{4}$$

For a well-defined axial octupole minimum in the total energy, the intrinsic charge octupole moment is

$$\mathcal{Q}_{30,c} \equiv \int 2r^3 P_{30} \rho_c(\mathbf{r}) d^3 \mathbf{r}, \tag{5}$$

where $\rho_c(\mathbf{r})$ is the charge density. Assuming ρ_c =const inside the sharp surface of Eq. (1), $\mathcal{Q}_{30,c}$ can be related to the deformations β_{λ} by (Leander and Chen, 1988)

$$Q_{30,c} = \frac{3}{\sqrt{7\pi}} Z R_0^3 \bar{\beta}_3, \qquad (6)$$

where

$$\bar{\beta}_{3} = \beta_{3} + \frac{5}{\sqrt{4\pi}} \left(\frac{4\sqrt{5}}{15} \beta_{2} \beta_{3} + \frac{6}{11} \beta_{3} \beta_{4} + \frac{60\sqrt{7}}{91\sqrt{11}} \beta_{4} \beta_{5} + \cdots \right).$$
(7)

The mass octupole moment and corresponding deformations are defined in a similar way.

There exist several parametrizations of reflectionasymmetric shapes other than the α and β parametrizations discussed above. A compilation of other parametrizations is contained in the Appendix.



Quadrupole-octupole shapes

FIG. 2. Quadrupole-octupole shapes represented by multipole expansion, Eq. (1). In all cases, the same axial quadrupole deformation $\alpha_{20}=\beta_2=0.6$ is assumed. The four shapes correspond to octupole deformations with $\mu=0, 1, 2, \text{ and } 3 \ [\alpha_{30}=\beta_{30}; \alpha_{3-\mu}=(-1)^{\mu}\alpha_{3\mu}=\beta_{3\mu}/2; \beta_{3\mu}=0.35]$. (Courtesy of T. Misu.)

The number of deformation parameters that appear in the multipole expansion Eq. (1) grows rapidly with λ . For instance, the general quadrupole-plus-octupole shape is described by two quadrupole deformations (α_{20} and α_{22} , or β_2 and γ) and seven independent deformations $\alpha_{3\mu}$. Figure 2 displays four shapes resulting from the superposition of axial quadrupole and octupole deformations with μ =0, 1, 2, and 3.

A general parametrization of the combined quadrupole-octupole field, covering all possible shapes without double counting, was proposed by Rohoziński (1990). The basic requirements are that the parametrization obeys simple transformation rules under O_h (a group of 48 transformations changing the names and arrows of the axes), and have simple ranges for the parameters. After introducing the seven real Cartesian components $a_{3\mu}$ and $b_{3\mu}$ (Rohoziński *et al.*, 1982; Rohoziński, 1988),

$$\alpha_{30} = a_{30}, \ \alpha_{3\pm\mu} = \frac{(\pm 1)^{\mu}}{\sqrt{2}} (a_{3\mu} \pm ib_{3\mu}) \ (\mu = 1, 2, 3),$$
(8)

the latter can be expressed in terms of one "radial" coordinate β_3 ($\beta_3 \ge 0$) and six "angular" biharmonic coordinates ($\delta_0, \delta_1, \delta_2, \gamma_0, \gamma_1, \gamma_2$) describing both the octupole distortion and its orientation with respect to the intrinsic frame defined by the quadrupole tensor. The coordinates given by Eq. (8) can be expressed as

$$a_{30} = \beta_3 \cos \delta_0 \cos \delta_1 \cos \gamma_1,$$

$$a_{31} = \beta_3 \cos \delta_0 \sin \delta_1 \cos \delta_2 \cos \gamma_2,$$

$$a_{32} = \beta_3 \cos \delta_0 \cos \delta_1 \sin \gamma_1,$$

$$b_{31} = \beta_3 \cos \delta_0 \sin \delta_1 \sin \delta_2 \cos \gamma_3,$$

$$b_{32} = \beta_3 \sin \delta_0,$$

$$b_{33} = \beta_3 \cos \delta_0 \sin \delta_1 \sin \delta_2 \sin \gamma_3,$$

(9)

where $-\pi/2 \le \delta_0 \le \pi/2$, $0 \le \delta_i \le \pi/2$ (i=1,2), $-\pi/2 \le \gamma_1 \le \pi/2$, $-c - \pi/2 \le \gamma_2 \le \pi/2 - c$, $c - \pi/2 \le \gamma_3 \le \pi/2 + c$, and $\sin c = \sqrt{5/8}$, $\cos c = \sqrt{3/8}$ (Rohoziński, 1990).

Attempts to find a unique parametrization of the pure-octupole field defining an intrinsic frame (i.e., without involving the quadrupole field) turned out to be less successful (Hamamoto, Zhang, and Xie, *et al.*, 1991).

B. The electric dipole moment

The nuclear electric dipole moment is a measure of the shift between the center of charge and the center of mass of the nucleus. Assuming that nucleons are pointlike particles, the E1 moment is

$$\boldsymbol{D} = \sum_{i=1}^{Z} e_i(\boldsymbol{r}_{p,i} - \boldsymbol{R}_{\text{c.m.}}) = e(\boldsymbol{r}_p - Z\boldsymbol{R}_{\text{c.m.}}).$$
(10)

Neglecting the proton-neutron mass difference $(A\mathbf{R}_{c.m.}=\mathbf{r}_p+\mathbf{r}_n)$, **D** is equal to

$$\boldsymbol{D} = e \frac{N}{A} \boldsymbol{r}_p - e \frac{Z}{A} \boldsymbol{r}_n \,. \tag{11}$$

Equation (11) can be written alternatively as

$$\boldsymbol{D} = e \frac{ZN}{A} (\boldsymbol{r}_{p,\text{c.m.}} - \boldsymbol{r}_{n,\text{c.m.}}), \qquad (12)$$

where $\mathbf{r}_{p,\text{c.m.}} = \mathbf{r}_p / Z$ and $\mathbf{r}_{n,\text{c.m.}} = \mathbf{r}_n / N$ are the center-ofmass coordinates for protons and neutrons, respectively.

For reflection-symmetric systems, the nucleonic (proton and neutron) densities have three symmetry planes, so that $\langle \mathbf{r}_n \rangle = \langle \mathbf{r}_p \rangle = 0$, and hence $\langle \mathbf{D} \rangle = 0$ (the notaton $\langle \cdots \rangle$ denotes the expectation value in the intrinsic state). However, if density distributions are reflection asymmetric, then in general $\mathbf{r}_{p,c.m.} \neq \mathbf{r}_{n,c.m.}$, and a large static E1 moment may arise in the intrinsic frame. For an axially deformed system, having $\langle x \rangle = \langle y \rangle = 0$, the intrinsic dipole moment is aligned along the symmetry axis (z axis), and its value D_0 can be calculated directly from Eq. (11). In the most general case of triaxial and reflection-asymmetric density distributions, the intrinsic dipole moment is characterized by three spherical components, $D_{\pm 1}$ and D_0 .

III. THEORETICAL DESCRIPTIONS

In this section, we review the many theoretical approaches to reflection-asymmetric nuclear shapes. We begin in Sec. III.A with general arguments, before moving on to the reflection-asymmetric mean-field approach (Sec. III.B) and its extensions (Sec. III.C), algebraic approaches (Sec. III.D), cluster models (Sec.III.E), and, for completeness, vibrational models (Sec. III.F).

A. Microscopic origin of static octupole deformations

The mechanism responsible for the appearance of static deformations in nuclei arises from the degeneracy of eigenvalues of a single-particle Hamiltonian around the Fermi level, leading to instability with respect to shape vibrations (spontaneous symmetry breaking). This is the Jahn-Teller effect (Jahn and Teller, 1937; Reinhard and Otten, 1984). Stable reflection-asymmetric deformations in the body-fixed frame can be attributed to a parity-breaking odd-multipolarity interaction which couples intrinsic states of opposite parity.

To illustrate the transition from symmetric to reflection-asymmetric shapes, we use arguments based on the random-phase approximation (RPA) with a separable multipole interaction. A simple nuclear Hamiltonian representing nuclear vibrations, the pairing-plusmultipole Hamiltonian, can be written as (Lane, 1964; Soloviev, 1976):

$$H = \sum_{j} e_{j}c_{j}^{+}c_{j'} - \frac{1}{2}\sum_{\lambda} \kappa_{\lambda} \sum_{\mu=-\lambda}^{+\lambda} Q_{\lambda\mu}^{+} \cdot Q_{\lambda\mu} + H_{\text{pair}}, \quad (13)$$

where the first term on the right-hand side is the spherical shell-model potential, the second term represents a long-range separable multipole-multipole force generating the collective motion, H_{pair} is the pairing Hamiltonian, and *j* stands for the set of quantum numbers (n, ℓ, j) . In Eq. (13), $Q_{\lambda\mu}$ is the multipole operator

$$Q_{\lambda\mu}^{+} = \sum_{jj'} \langle j | f_{\lambda}(r) Y_{\lambda\mu}(\Omega) | j' \rangle c_{j}^{+} c_{j'}, \qquad (14)$$

where $f_{\lambda}(r)$ is the radial form factor [given, e.g., by the derivative of the average potential (Bohr and Mottelson, 1975)]. A coupling between single-particle states of opposite parity is produced by the octupole-octupole (λ =3) residual interaction. For the Hamiltonian [Eq. (13)] representing simple octupole vibrations, the excitations E_{oct} of the system can be computed by means of the RPA; they are solutions to the dispersion equation (Lane and Pendlebury, 1960; Veje, 1966; Soloviev, 1976; Ring and Schuck, 1980)

$$\sum_{j,j'} \frac{|\langle j \| Q_3 \| j' \rangle|^2 (u_j v_{j'} + u_{j'} v_j)^2 (E_j + E_{j'})}{(E_j + E_{j'})^2 - E_{\text{oct}}^2} = \frac{7}{\kappa_3}, \quad (15)$$

where $E_j = \sqrt{(e_j - \lambda)^2 + \Delta^2}$ are the quasiparticle energies, and u_j and v_j are the usual BCS occupation coefficients.

The lowest root of Eq. (15) represents the lowfrequency collective octupole vibration. If the coupling



FIG. 3. Total nuclear energy as a function of octupole deformation β_3 , for different values of octupole coupling κ_3 . The vibrational limit (a) corresponds to small values of κ_3 . Here the octupole stiffness $C_3 = (d^2 E/d\beta_3^2)_{\beta_3=0}$, is positive. At the point of instability (b) $\kappa_3 = \kappa_{3,crit}$ [Eq. (16)], $C_3 = 0$ and the system becomes unstable to octupole vibrations. For $\kappa_3 > \kappa_{3,crit}$ (c) the system is permanently deformed ($C_3 < 0$); the parity splitting results from tunneling between two degenerate minima at $\beta_3 = \pm \bar{\beta}_3$, separated by a barrier V_B . In the limit of rigid octupole deformation (d), i.e, an infinite barrier separating the minima, the parity splitting vanishes.

constant κ_3 is relatively small, the system behaves like a vibrator [Fig. 3(a)]. If the value of κ_3 is increased, vibrations become more and more collective and the vibrational frequency decreases. At the critical point, defined by

$$\kappa_{3,\text{crit}} = 7 \left\{ \sum_{j,j'} \frac{|\langle j \| Q_3 \| j' \rangle|^2 (u_j v_{j'} + u_{j'} v_j)^2}{E_j + E_{j'}} \right\}^{-1}, \quad (16)$$

the lowest solution of Eq. (15) has zero energy, i.e., the collective vibrational state becomes degenerate with the ground state [Fig. 3(b)]. At this point the RPA breaks down, the system becomes unstable against vibration induced by the octupole-octupole force (Thouless, 1961), and further increase of the coupling constant leads to permanent intrinsic octupole deformation [see Fig. 3(c)]. Pairing correlations, through the change in the energy denominator in Eq. (15) and the reduction of the numerator through the uv factor, tend to increase the critical value of κ_3 . One can thus say that pairing has a tendency to make the system less octupole deformed.

It is important to remember that actual nuclei are relatively small systems, and that finite-size effects, which manifest themselves through dynamical correlations (fluctuations), are crucial. The fluctuations wash out transitions from the "spherical" to the "deformed" phase and result in a smooth and continuous pattern of $E_{\rm oct}$ and octupole deformation β_3 , proportional to $\langle 0^+ \| Q_3 \| 3^- \rangle$ (see the discussion in Sec. III.C). Even for



FIG. 4. Nuclear spherical single-particle levels. The most important octupole couplings are indicated.

the best cases of nuclear stable octupole deformations, the octupole barriers separating the two degenerate minima are relatively small, i.e., the extreme limit of the rigid static octupole deformation shown in Fig. 3(d) is never approached.

As can be seen in Eq. (16), the necessary condition for the presence of low-energy octupole collectivity is the existence, near the Fermi level, of pairs of orbitals strongly coupled by the octupole interaction. As shown in Fig. 4, for normally deformed systems the condition for strong octupole coupling is satisfied for particle numbers associated with the maximum $\Delta N=1$ interaction between the intruder subshell (ℓ , j) and the normal-parity subshell (ℓ -3,j-3). The regions of nuclei with strong octupole correlations correspond to particle numbers near 34 ($g_{9/2} \leftrightarrow p_{3/2}$ coupling), 56 ($h_{11/2} \leftrightarrow d_{5/2}$ coupling), 88 ($i_{13/2} \leftrightarrow f_{7/2}$ coupling), and 134 ($j_{15/2} \leftrightarrow g_{9/2}$ coupling). That is, the tendency towards octupole deformation occurs just above closed shells.

B. Reflection-asymmetric mean-field approach

The concept of stable octupole deformation appears naturally in the unconstrained mean-field approach, in which the average potential, and consequently the total energy, is a functional of the nucleonic density (Ring and Schuck, 1980).

The average nuclear mean field, in which nucleons move as independent particles, can be obtained from the Hartree-Fock (HF) theory. The starting point is the general two-body Hamiltonian

$$H = \sum_{ij} t_{ij} c_i^+ c_j^+ + \frac{1}{4} \sum_{ijkl} \bar{v}_{ijkl} c_i^+ c_j^+ c_l c_k^-, \qquad (17)$$

where \bar{v}_{ijkl} is the antisymmetrized matrix element of a two-body effective (usually density-dependent) force. The *A*-body wave function is approximated by a Slater determinant whose orbitals are determined by minimization of the total energy. This variation leads to an eigenvalue problem which defines both the single-particle orbitals $\{\Phi_i, i=1, \ldots, A\}$ and the single-particle energies e_i :

$$(t+\Gamma)\Phi_i = e_i\Phi_i. \tag{18}$$

The resulting self-consistent HF potential,

$$\Gamma_{ij}[\rho] = \sum_{kl} \bar{v}_{ikjl} \rho_{lk}, \qquad (19)$$

depends on the orbitals through the density matrix ρ .

The pairing (particle-particle) components of the effective force can be approximated in the framework of the BCS theory, or they can be treated on the same footing as the particle-hole interactions through the Hartree-Fock-Bogolyubov (HFB) theory. The solution of the HF+BCS or HFB equations gives the binding energy of the nucleus, either at the local minimum or as a function of collective parameters q_i such as shape deformations (constrained HFB theory).

As early as 1957, Bleuler and Terreaux proposed that the single-particle wave functions of the nuclear shell model or the HF method should have no definite parity. Extension of the HF formulation to the mixed-parity Slater determinants was accomplished by Amiet and Huguenin (1963, 1966) and Müller-Schwartz (1967).

If the intrinsic parity is broken, the intrinsic state χ is not an eigenstate of the parity operator \mathcal{P} . The states with good parity p can be constructed by means of projection:

$$\Psi_p = \mathcal{N}_p(1+p\mathcal{P})\chi, \ \mathcal{N}_p^{-2} = 2(1+p\langle\chi|\mathcal{P}|\chi\rangle).$$
(20)

The energy of the projected state $E_p = \langle \Psi_p | H | \Psi_p \rangle$ is given by

$$E_{p} = E - p \frac{\langle \mathscr{P} \rangle E - \langle H \mathscr{P} \rangle}{1 + p \langle \mathscr{P} \rangle}, \qquad (21)$$

where $E = \langle \chi | H | \chi \rangle$ is the HF energy. As can be seen in Eq. (21) (Amiet and Huguenin, 1966),

$$E_{+} \leq E \leq E_{-} \quad \text{if } E < \langle H \mathscr{P} \rangle / \langle \mathscr{P} \rangle.$$
 (22)

The mean value of the parity in the intrinsic state is usually very small. The wave function of the singleparticle orbital ψ_i can be decomposed into two components with good parity

$$\psi_i = a_i^{(+)} \psi_i^{(+)} + a_i^{(-)} \psi_i^{(-)}, \ |a_i^{(+)}|^2 + |a_i^{(-)}|^2 = 1.$$
(23)

The parity content of the single-particle state is given by

$$\langle \psi_i | \mathscr{P} | \psi_i \rangle = |a_i^{(+)}|^2 - |a_i^{(-)}|^2,$$
 (24)

which yields

$$\langle \chi | \mathscr{P} | \chi \rangle = \prod_{i=1,A} (|a_i^{(+)}|^2 - |a_i^{(-)}|^2).$$
 (25)

It is instructive to analyze expressions (21), (22), and (25) using the simple two-level model proposed by Lipkin, Meshkov, and Glick (1965). The model Hamiltonian,

$$H = \epsilon K_0 - \frac{1}{2} V(K_+ K_+ + K_- K_-), \qquad (26)$$

describes a system of N fermions distributed among two levels of different parity, each having an N-fold degeneracy, and separated by an energy ϵ . The particles interact via a monopole-type residual interaction with strength V that scatters particles between upper and lower levels. In Eq. (26), K_0 and K_{\pm} are quasispin operators. The most general Slater determinant in this case is (Agassi *et al.*, 1966; Ring and Schuck, 1980) $\chi = |\phi\rangle = \exp[\tan(\phi/2)K_{\pm}]|0\rangle$. After introducing the effective coupling strength $\kappa_3 = (N-1)V/\epsilon$, we can write the HF energy and the energy of the projected states in closed form (Agassi *et al.*, 1966).

The critical value of the coupling strength in the Lipkin-Meshkov-Glick model is $\kappa_3=1$. For $\kappa_3<1$ the potential-energy curve $E(\phi)$ has only one minimum, at $\phi=0$ (vibrational limit). For $\kappa_3>1$ a stable deformation develops, $\bar{\phi}=\pm \arccos(1/\kappa_3)$. The overlap (25) in the Lipkin-Meshkov-Glick model is given by

$$\langle \phi | \mathscr{P} | \phi \rangle = \cos^N \phi, \tag{27}$$

i.e., the parity content of the HF ground state is close to zero for large values of κ_3 . The behavior of E and E_p as a function of ϕ is displayed in Fig. 5 for $\kappa_3=0.25$, 1, and 2. We see that for large values of κ_3 the minima of Eand E_{\pm} correspond to very similar values of ϕ . This is not true for smaller values of κ_3 , for which the deformation of the p=-1 state $(\bar{\phi}_-)$ is larger than that of the p=1 state $(\bar{\phi}_+)$. (For realistic calculations, see Fig. 8 below.)

1. Mean-field symmetries

Any mean-field solution, when deformed, breaks one of the basic symmetries of the nuclear Hamiltonian. However, not all symmetries of the wave functions are destroyed in the self-consistent calculations (Ring and



FIG. 5. HF energy (thick solid line) and projected energy curves (grey line, positive parity; thin line, negative parity) for the Lipkin-Mechkov-Glick model as a function of deformation ϕ for κ_3 =0.25, 1, and 2, and for N=14.

Schuck, 1980). The remaining *self-consistent symmetries* provide a simplification of the calculation as well as a convenient labeling of the individual orbitals. Among the possible remaining symmetries, the most usual are those associated with the following operators (Goodman, 1974; Bohr and Mottelson, 1975; Bohr, 1976):

- (i) time reversal (T);
- (ii) space inversion (parity) (\mathcal{P}) ;
- (iii) rotation around the symmetry axis (giving rise to the good quantum number K);
- (iv) Rotations $R_{\kappa} = \exp(-i\pi I_{\kappa})$, $\kappa = 1, 2, 3$, through the angle π around the three principal axes of the mean field (the finite symmetry group defined by these three rotations is called D_2);
- (v) reflections $S_{\kappa} = \mathscr{P} \cdot R_{\kappa}^{-1}$ through planes containing the two principal axes of the mean field.

Depending on the underlying mean-field theory and the constraints, the mean-field solution will possess none or several of these symmetries.

If the rotation is approximated by means of the cranking approach, one has to analyze the Routhian

$$H^{\omega} = H - \omega I_1, \tag{28}$$

rather than the Hamiltonian H. Due to the cranking term ωI_1 , the only self-consistent symmetries that can possibly remain in the rotating nucleus are P, R_1 , and S_1 .

To discuss the possible nuclear shapes that remain invariant with respect to different self-consistent symmetries, it is convenient to perform the multipole decomposition of the density matrix ρ :

$$\rho = \sum_{\lambda\mu} \rho_{\lambda\mu} \tag{29}$$

(Dobaczewski and Skalski, 1989). Using Eq. (29), we can express the average field [Eq. (19)] as (Ring and Schuck, 1980; Dobaczewski and Skalski, 1989)

$$\Gamma[\rho] = \sum_{\lambda\mu} \beta_{\lambda\mu} Q_{\lambda\mu}, \qquad (30)$$

where $\beta_{\lambda-\mu} = (-1)^{\mu} \beta_{\lambda\mu}^*$ and $Q_{\lambda\mu}$ transforms as a rank- λ spherical tensor. The transformation properties of spherical tensors $Q_{\lambda\mu}$ are

$$\mathscr{P}Q_{\lambda\mu}\mathscr{P}^{-1} = (-1)^{\lambda}Q_{\lambda\mu}, \qquad (31)$$

$$R_1 Q_{\lambda\mu} R_1^{-1} = (-1)^{\lambda} Q_{\lambda-\mu}, \qquad (32)$$

$$S_1 Q_{\lambda\mu} S_1^{-1} = Q_{\lambda-\mu}. \tag{33}$$

If the mean field commutes both with \mathscr{P} and R_1 , the eigenstates of H^{ω} can be characterized by the intrinsic parity π and the signature quantum number r, which is the eigenvalue of R_1 . In this case the λ -odd components vanish and $\beta_{\lambda\mu} = (-1)^{\mu} \beta_{\lambda-\mu}$, i.e., the spherical tensors $Q_{\lambda\mu}$ ($\mu \neq 0$), appear in the combinations

$$Q'_{\lambda\mu} \equiv Q_{\lambda\mu} + (-1)^{\mu} Q_{\lambda-\mu}, \ \mu > 0.$$
 (34)

Relation (34) also holds for the most general field that commutes with S_1 . However, in this case both even and odd values of λ appear in the multipole expansion given by Eq. (30). (For instance, the octupole shapes shown in Fig. 2 are S_1 invariant.)

The eigenstates of H^{ω} can be characterized by the simplex quantum number *s*, which is the eigenvalue of S_1 (Nazarewicz *et al.*, 1984a; Frauendorf and Pashkevich, 1984; Nazarewicz and Olanders, 1985a, 1985b). For the application of simplex symmetry to rotational spectra, see Sec. VI.

To take advantage of the S_1 symmetry, a new singleparticle basis must be constructed. The Goodman transformation (Goodman, 1974) from the strong-coupling basis $|k, \Omega_k\rangle$ to basis states of good simplex reads (Nazarewicz and Olanders, 1985a, 1985b)

$$|k,s=+i\rangle = \frac{1}{\sqrt{2}} [-|k,\Omega_k\rangle + (-1)^{\Omega_k - 1/2} |\overline{k,\Omega_k}\rangle], \qquad (35)$$

$$|k,s=-i\rangle = \frac{1}{\sqrt{2}} [|\overline{k,\Omega_k}\rangle + (-1)^{\Omega_k - 1/2} |k,\Omega_k\rangle], \quad (36)$$

where Ω_k is the single-particle angular momentum projection on the axis of quantization, and $|\overline{k,\Omega_k}\rangle = T|k,\Omega_k\rangle$, where $T|\pi jm\rangle = \pi(-1)^{j+m}|\pi j-m\rangle$. Symmetry properties of spherical tensors in the basis given by Eq. (35) were discussed by Leander *et al.* (1986). In particular, they demonstrated that the transition matrix element between parity partners $|k,s\rangle$ and $|k,-s\rangle$ vanishes, i.e., $\langle k,s|Q_{\lambda\mu}|k,-s\rangle=0$.

In several works, the nonaxial octupole fields have been treated in selected combinations that guarantee the presence of some self-consistent symmetries and thus facilitate calculations. Li *et al.* (1991) discussed symmetry properties of the general one-body Hamiltonian and applied them to triaxial octupole deformations (assuming one symmetry plane). Skalski (1991) assumed the presence of two mirror reflections, S_1 and S_2 ; this leaves only even- μ components in Eq. (34). In his following study, Skalski (1992) considered rather general mean fields containing axially symmetric components with λ =2, 3, 4, and triaxial octupole deformations given by Eq. (34), the latter treated separately (i.e., one triaxial deformation at a time).

As discussed by Hamamoto, Mottelson, et al. (1991), the symmetry groups of the pure-octupole fields [Eq. (34)] with $\mu=1, 2, \text{ and } 3, \text{ are respectively } C_{2v}, T_d, \text{ and }$ D_{3h} . This leads to the classification of single-particle levels in terms of corresponding irreducible representations (irreps). In the presence of axial reflectionsymmetric shapes the symmetries C_{2v} and D_{3h} are preserved, but T_d reduces to the lower symmetry D_{2d} (Skalski, 1992). Consequently, in the presence of an axial quadrupole field (i) all the single-particle levels of the same simplex interact in the Q'_{31} field (C_{2v} has one 2D spinor irrep), (ii) there are two groups of noninteracting states in the Q'_{32} field (D_{2d} has two nonequivalent 2D spinor irreps), and (iii) there are two groups of noninteracting states in the Q'_{33} field (D_{3h} has three nonequivalent 2D spinor irreps). The combination of $\mu=0$ and 2, and $\mu=1$ and 3 fields reduces the symmetry group to C_{2v} .

Examples of single-particle diagrams in the Q'_{32} field can be found in Li and Dudek (1994). Eichler and Faessler (1970) in their study of the trigonal symmetry in light nuclei considered systems invariant with respect to the trigonal C_3 symmetry, i.e., with respect to rotations of 120° about the z axis. (For the α -like nuclei the interesting symmetry point groups are $D_{\infty h}$, T_d , and D_{3h} . The group C_3 is a common lower-symmetry subgroup). The self-consistent symmetry in question, C_3 , conserves the quantum number $q=\Omega \mod 3$. A tetrahedral perturbation of the spherical harmonic-oscillator potential was considered by Elliott *et al.* (1985); they classified the single-particle levels using the irreps of T_d .

2. Shell-correction method

Mean fields, in which nucleons move as independent particles, can be obtained from a knowledge of the forces acting between nucleons using self-consistent HF theory. For a particular choice of nucleon-nucleon force and proton and neutron number, the variational principle determines whether this field is spherical or deformed.

The deformed-shell model developed in the 1950s (Rainwater, 1950; Moszkowski, 1955; Nilsson, 1955) is an approximation to the HF approach. Here, the field Γ is not determined self-consistently from Eq. (19), but is assumed to be a phenomenological average potential which contains a central part, a spin-orbit term, and a Coulomb potential for protons. All of these terms depend explicitly on a set of external deformation parameters defining the nuclear surface.

The earliest calculations employing the reflectionasymmetric deformed-shell approach (Lee and Inglis, 1957; see also Johansson, 1961) investigated the effect of spin-orbit coupling on the stability of pear-shaped deformations using wave functions of a spheroidal harmonicoscillator potential. Vogel (1968) studied the dependence of the nuclear potential energy on β_3 in the 218 $\leq A \leq 232$ region using a modified harmonic-oscillator potential with pairing, and found no cases with nonzero octupole deformation.

The deformed-shell model alone cannot be used to predict binding energies. This is because the deformedshell model energy differs from the full HF energy by the two-body interaction term. On the other hand, it is well known that binding energies are accounted for with good accuracy by the classical model of the liquid drop (Myers and Swiatecki, 1969). The two approaches are merged into the shell-correction (SC) method (also known as the macroscopic-microscopic, or Nilsson-Strutinsky method) (Strutinsky, 1967; Brack *et al.*, 1972). The main assumption of the SC method is that the total energy of a nucleus can be separated into two parts,

$$E = E_{\text{macr}} + E_{\text{shell}},\tag{37}$$

where $E_{\rm macr}$ is the macroscopic energy (depending smoothly on the number of nucleons), and $E_{\rm shell}$ is the shell-correction term, which fluctuates with particle number (reflecting the nonuniformities of the singleparticle level distribution, i.e., shell effects). The macroscopic part is usually replaced by the corresponding liquid-drop (or droplet) model value, while the shellcorrection term is calculated using the deformed independent-particle model.

Before discussing the results of the SC method, it should be emphasized that it is not a self-consistent theory. Thus it should be viewed as a practical recipe, deficient in a number of respects. Its formal justification in terms of the HF approach is given by the so-called Strutinsky energy theorem (Strutinsky, 1974; Brack and Quentin, 1981). This theorem states that the difference between the HF energy and the SC energy is of second order in the density fluctuations, provided that the deformed-shell potential gives a similar spectrum to the averaged HF potential.

Möller and Nilsson (1970) calculated the effect on the potential-energy surface in the lead and actinide region of simultaneous P_3 and P_5 degrees of freedom using the SC method with a modified harmonic-oscillator poten-

tial, and found instability with respect to octupole deformation (see also Möller, 1972, for a comprehensive description of the method as applied to an investigation into the fission process; a different approach was employed by Mustafa *et al.* (1973), who used a two-center model including asymmetric deformation). Möller *et al.* (1972) carried out similar calculations, which suggested that ²²⁴Ra has the largest equilibrium octupole deformation.

In light nuclei, the spin-orbit interaction is relatively weak and, in addition, the diffuseness of the nuclear surface is comparable to the nuclear radius. Consequently, the harmonic-oscillator model gives a fairly good approximation to the nuclear average potential. The first study of static octupole shapes in light nuclei with N=Z from ¹²C to ⁴⁴Ti, based on the SC method, was carried out by Leander and Larsson (1975). They employed the modified harmonic-oscillator potential parametrized through quadrupole (K=0,2), octupole (K=0,3), and higher-order deformations, finding minima in ε_3 for mid-shell nuclei at very large quadrupole deformation. Hellström (1977) explicitly included triaxial octupole deformations with K=2.

In the presence of the small perturbing octupole potential $V = \beta_{3\mu}Q_{3\mu}$, the shell driving force (i.e., the dependence of the shell correction on deformation) is determined by the second-order correction $\delta E_{\text{shell}}^{(2)}$, which is proportional to the square of the corresponding deformation $\beta_{3\mu}$:

$$\delta E_{\rm shell}^{(2)} = C_{3\mu} \beta_{3\mu}^2. \tag{38}$$

(The first-order term vanishes, since the expectation value of the octupole moment is zero in the parityconserving wave function. The exact expression for $\delta E_{\text{shell}}^{(2)}$ can be found in Nazarewicz *et al.*, 1995.) The shell-energy octupole-stiffness coefficient $C_{3\mu}$ determines the octupole susceptibility of the shell energy. If $C_{3\mu}$ is negative, there exists a shell force favoring stable deformations. On the other hand, if $C_{3\mu}$ is positive, the shell correction tends to restore reflection symmetry. Since for nuclei with $Z \leq 104$ the liquid-drop model favors spherical ground-state shapes, one can say that stable octupole deformations can only arise from shelleffects, i.e., from the shell driving force.

The behavior of E_{shell} obtained in the Woods-Saxon (WS) potential, as a function of β_3 and particle number (at typical ground-state deformations β_2 and β_4), is shown in Fig. 6 (for a similar plot based on the modified harmonic-oscillator model, see Leander *et al.*, 1982). We see that the shell correction favors octupole deformation ($C_{30}<0$) near particle numbers N=134 and Z=90. Other octupole-driving particle numbers expected from the SC method are 34 and 56, in nice agreement with the schematic diagram of Fig. 4. From different combinations of those particle numbers, several regions of candidates for reflection-asymmetric shapes emerge.

Möller and Nix (1981) showed that there is a substantial ground-state octupole instability in their calculations based on the folded Yukawa deformed potential and Yukawa-plus-exponential macroscopic energy, much softer to high-multipole deformations than the standard liquid drop model with sharp surfaces used previously. This was extended by Leander *et al.* (1982), who made a systematic study of the Po-U region. Nazarewicz *et al.* (1984b) employed a similar model using the WS potential to analyze octupole instability in both medium- and heavy-mass nuclei. The octupole-deformation energy curves for even-even isotopes of Rn, Ra, Th, and U obtained from the SC+WS model of Nazarewicz *et al.* (1984b) are shown in Fig. 7. In the same work, it was shown that nuclei around ¹⁴⁶Ba and ¹¹⁴Xe should also exhibit octupole instability, although the deformation is much softer than for the Ra-Th region.

Nazarewicz and co-workers applied the cranked shellcorrection method to investigate the behavior of nuclear shape as a function of spin for both the Ra-Th region (Nazarewicz *et al.*, 1984a, 1987) and the region around Z=58, N=88 (Nazarewicz and Tabor, 1992).

Chasman (1986) investigated the effect of allowing the hexacontatetrapole deformation (λ =6) parameter to vary freely, instead of being fixed as in earlier studies. He found that the binding energy increased by approximately 1 MeV for many nuclei in the mass region 220 < A < 230. Chasman and Ahmad (1986) also investigated the γ degree of freedom using the SC method with a WS potential. They found a region of γ softness for Ra and Rn isotopes with 220< A < 226 and for several nuclei with N=130.

Sobiczewski et al. (1988), Rozmej et al. (1988), and Cwiok and Nazarewicz (1989a, 1989b) demonstrated that by treating higher-order multipole deformations $(\beta_5, \beta_6, ...)$ in a self-consistent manner it is possible to further lower the octupole minima in a WS model. For instance, inclusion of higher-order deformations lowers the octupole minima in ^{144,146,148}Ba by 200-300 keV. One-quasiparticle bandheads in odd-A actinide nuclei were calculated in a deformed self-consistent WS model by Cwiok and Nazarewicz (1989a, 1991), who pointed out the importance of specialization energy (i.e., extra energy required to find a transition state at the octupole barrier with quantum numbers matching those of the reflection-asymmetric ground state) for the barrier heights. In other mass regions, deformed SC-WS calculations have also been made for ⁶⁴Ge (Ennis et al., 1991; Skalski, 1991) and for even-even nuclei near ¹¹²Ba (Skalski, 1990).

As discussed in Secs. II.A and III.B.1, the consistent treatment of triaxial octupole deformations with $\mu \neq 0$ is difficult. In an early work, Gavron *et al.* (1977) demonstrated in SC calculations that the simultaneous inclusion of triaxiality and reflection asymmetry is important around the third saddle point in actinides (see also Åberg *et al.*, 1980). Banana-type octupole deformations (μ =1) at superdeformed shapes were investigated by Chasman (1991) and Skalski (1992) using the SC method with a WS potential (see Sec. IX.B.2 for more discussion). Li and Dudek (1994) considered $\alpha_{\lambda\mu}$ deformations with λ =2, 3, 4, and 5, and predicted static α_{32} deformation (with all other deformations vanishing) in ²²²Rn.



FIG. 6. WS proton (top) and neutron (bottom) shell correction plotted versus particle number and octupole deformation β_3 . Other deformations are $\beta_2=0.15$, $\beta_4=0.08$ (Nazarewicz *et al.*, 1984b). The contours are labeled by their values in MeV.

In the context of metallic clusters, octupole instability with respect to nonaxial octupole modes $Y_{3\mu}$, $\mu \neq 0$, has been investigated by simulating the one-particle spectrum of the infinite-well potential using that of a modified oscillator potential (Hamamoto, Mottelson, *et al.* 1991; Frisk *et al.* 1994). It is found that the $\mu=3$ and $\mu=2$ octupole deformations, when treated one at a time, give rise to a strong shell structure.

3. Self-consistent methods

Early HF studies for light nuclei involving parity mixing of intrinsic wave functions were carried out by several authors. Kelson (1965) was able to reproduce approximately the negative-parity collective band in ¹⁶O by projecting out a Slater determinant containing $s_{1/2}p_{1/2}d_{3/2}$ states using the HF method with a Rosenfeld two-body interaction. The mechanism of parity mixing in single-particle orbitals in the HF framework was also investigated by several other workers, including Amiet and Huguenin (1966), Ebenhöh (1966), Röhl (1966), Parikh and Ullah (1967), and Bassichis and Svenne (1967), to improve the agreement with the observed spin-orbit splittings and magnetic moments. Müller-Schwartz (1968) pointed out that the low-lying 0^{-} state predicted by Amiet and Huguenin (1966) is unrealistic and that the ground state in ¹⁶O has no parity mixing, although excited states might have such mixing. This has been supported by Blomquist and Molinari (1968) and Burr et al. (1969), who demonstrated that the paritybreaking deformations are excluded in practically all cases for light nuclei if realistic interactions (in particular the tensor force) are included. Krappe and Wahsweiler (1967) calculated the energy surfaces of ¹⁶O,



FIG. 7. Octupole-deformation energy curves for even-even isotopes of Rn, Ra, Th, and U obtained in the SC+WS model of Nazarewicz *et al.* (1984b). Insets show experimental energies of the lowest negative-parity states in these isotopes.

²⁰Ne, and ²⁴Mg using the single-particle parity-mixed wave functions and a Hamiltonian with a simple twobody local potential of Serber-exchange type. They obtained a pronounced octupole softness. Castel and Svenne (1969) performed HF calculations with the Yale-Shakin potential for a number of sd nuclei. They calculated large values of $B(E3,3^-\rightarrow 0^+)$, 10–18 s.p.u. (single-particle units), but in each case found that the octupole deformation was zero. The most octupole-soft nuclei were predicted to be ³⁰Si and ³⁸Ar. Giraud and Sauer (1970) found parity-mixed HF solutions for ¹⁶O, ¹⁹F, ²⁰Ne, ²⁴Mg, and ²⁸Si using a Gaussian or Yukawa force with a Rosenfeld-exchange mixture. In particular, they obtained states of mixed parity and having a triaxial shape. Eichler and Faessler (1970) investigated the static octupole shapes with K=0 and 3 in ¹²C, ¹⁶O, and ²⁰Ne by means of a constrained HF method with the Yale-Shakin and Volkov potentials. They obtained small static trigonal octupole moments Q_{33} . For other paritymixed calculations for ¹⁶O, see also Do Dang et al. (1976) and Elliott et al. (1985).

Bonche *et al.* (1986) performed HF+BCS calculations, using the Skyrme III effective interaction, of the energy as a function of deformations q_2 and q_3 (defined as expectation values of the quadrupole moment and $r^3 Y_{30}$, respectively) of ²²²Ra. Significantly, they find a mini-

mum for a nonzero value of the octupole moment. Strong quadrupole-octupole coupling was also predicted by the HF+BCS calculations of Bonche (1988) for ¹⁴⁴Ba. The potential-energy curves versus q_3 , corresponding to q_2 =400, 600, and 800 fm², are displayed in Fig. 8. The HF results are compared to the parity-projected energies given by Eq. (21).

Praharaj (1986) performed parity-mixed deformed HF calculations for ^{142–148}Ba and ^{146–148}Ce, and showed that ^{142,144}Ba and ¹⁴⁸Ba are unstable to octupole deformation (see, however, the comment in Nazarewicz, 1987).

Robledo *et al.* (1987) and Egido and Robledo (1990) employed the HF+BCS method with the Gogny interaction to calculate reflection-asymmetric minima in ^{222,224}Ra, ²²²Rn, and ^{142–148}Ba. They obtained reflection-asymmetric minima in all cases. By constraining the position of the c.m. coordinate at the origin, they were able to calculate intrinsic dipole moments in these nuclei (see Sec. VII.C).

4. Particle-plus-rotor model

There have been several attempts to explain spectroscopic properties of reflection-asymmetric nuclei using concepts based on the particle-plus-rotor model. This is



FIG. 8. Potential-energy curves for ¹⁴⁴Ba obtained in HF+BCS+Skyrme calculations of Bonche (1988). The HF results (solid line) are compared with the parity-projected curves, E_+ (dashed line) and E_- (dashed-dotted line). The upper portion shows the overlaps, Eq. (25).

a highly phenomenological model that aims at reproduction and interpretation of low-lying experimental spectra. The microscopic input to this model, namely equilibrium deformations, pairing gaps, and the intrinsic single-particle Hamiltonian, is taken from self-consistent HF calculations or from the SC method.

Assuming a large deformation and the strongcoupling scheme, the particle-plus-rotor Hamiltonian can be written as (Bohr and Mottelson, 1975)

$$H = H_{\text{intr}} + H_{\text{rot}}, \tag{39}$$

where the rotor part is given by

$$H_{\rm rot} = \frac{\hbar^2}{2\mathscr{J}} (\boldsymbol{I} - \boldsymbol{j})^2 \tag{40}$$

and the intrinsic part is

$$H_{\rm intr} = H_{\rm core} + H_{\rm s.p.} \,. \tag{41}$$

The single-particle Hamiltonian can be approximated by the deformed (reflection-asymmetric) field and the BCS pairing. The space of the reflection-asymmetric particleplus-rotor model Hamiltonian is spanned by the symmetrized wave functions (Bohr and Mottelson, 1975)

$$\Psi^{\nu}_{IMKp} = \mathcal{N}(1+R_1) D^I_{MK}(1+p\mathscr{P}_c\mathscr{P}_{s.p.}) \Phi_a \chi^{\nu}_K, \quad (42)$$

where \mathscr{N} is a normalization constant, \mathscr{P}_c is the core parity operator, $\mathscr{P}_{s.p.}$ is the valence (single-particle) parity operator, $p=\pm 1$ is the total parity, Φ_a describes the core with the same orientation in space as the singleparticle potential, and χ_K^{ν} is the wave function of the valence *n*-quasiparticle configuration $(K=\Omega_1+\Omega_2$ $+\cdots+\Omega_n)$.

One of the earliest applications of this model to odd-A nuclei was that of Zaikin (1966), who calculated B(E3) rates in ¹⁹F. Leander and Sheline (1984) introduced the phenomenological core Hamiltonian that accounts for the parity splitting in the doubly even core, $E(0^-)$:

$$H_{\text{core}} = \frac{1}{2} E(0^{-})(1 - \mathcal{P}_c) = \frac{1}{2} E(0^{-})(1 - p\mathcal{P}_{\text{s.p.}}). \quad (43)$$

They applied the model to make extensive investigations of the spectroscopy of odd-*A* nuclei, and accounted for the ground-state properties of odd-*A* Ra isotopes and decoupling parameters for K=1/2 bands in Ra and Ac isotopes, using a folded Yukawa average potential to calculate single-particle levels. This model was subsequently developed by Leander and Chen (1987, 1988) using a WS deformed potential. The properties of the low-lying spectra in nuclei with A=219-229 were calculated and found to be in reasonable agreement with experiment. Finally, Sheline, Chen, and Leander (1988) applied the model to ²²³Ra.

Brink *et al.* (1987) have derived results similar to those of Leander and Sheline (1984) by taking a model in which the valence nucleons are coupled to the observed energy spectrum of the even-even core, instead of a model assuming some particular symmetry for the intrinsic system. The extension of these techniques to odd-odd nuclei has been accomplished by Afanasjev *et al.* (1991).

C. Beyond the mean field

As already discussed in Sec. III.A, correlations due to the final size are important in describing properties of transitional systems. Since even in the best cases the predicted gain in the ground-state energy due to reflectionasymmetric deformations (deformation energy) is around 1–2 MeV, i.e., rather modest, dynamical corrections beyond the mean field play an important role.

The fluctuations smooth out the transition from the reflection-symmetric (vibrational) to the reflection-asymmetric (deformed) regime. This is illustrated in Fig. 9, which shows the parity-splitting energy ΔE between the first excited state (p=-1) and the ground state



FIG. 9. The parity-splitting energy ΔE (in units of ϵ), between the first excited state (*p*=–1) and the ground state (*p*=1) of the Lipkin-Meshkov-Glick model, solved exactly as a function of κ_3 for *N*=14.

(p=1) for the exact solution of the Lipkin-Meshkov-Glick model (see Fig. 5). In contrast to the static HF theory, which exhibits a rapid phase transition at $\kappa_{3,crit}$, here the transition from vibrational to deformed regime is smooth.

Some correlations can be accommodated by means of parity projection, performed either after or (better) before variation. However, as discussed by Agassi *et al.* (1966) and Robledo (1992), parity projection does not eliminate the discrepancy around the transition point. The missing correlations have to be incorporated by other methods (see below).

1. Generator-coordinate method

A useful tool for describing the dynamical correlations is the generator-coordinate (GC) method (Ring and Schuck, 1980). The GC wave function is usually taken as a combination of many (projected) product states,

$$\Psi_i^{(\text{GC})} = \int d\mathbf{q} \sum_i f_i(\mathbf{q}) \Psi_i(\mathbf{q}), \qquad (44)$$

where the **q**'s are generator coordinates and the generating functions Ψ_i are (projected) eigenstates of the intrinsic Hamiltonian (*i*=0 denotes the vacuum, *i*=1 the first excited state, and so on). The wave function in Eq. (44) is rich enough to accommodate correlations absent in the mean-field description. In the simplest applications of this method, the generating functions can be identified with the collective coordinates. The weight functions $f_i(\mathbf{q})$ are determined by means of the variational principle from the Hill-Wheeler equation.

Marcos *et al.* (1983) studied the collective path leading from ²⁰Ne to the separated nuclei ¹⁶O + α using both the constrained HF and GC methods, and found that ²⁰Ne has softness to octupole deformation. A GC approach to molecular states was developed by, e.g., Langanke *et al.* (Langanke, 1982; Langanke *et al.*, 1984) and Baye and Descouvement (1983, 1984) who applied it to E1, E3 transitions and to reduced α -widths in light nuclei (for other works, see Sec. VIII).

Bonche et al. (1991) and Meyer et al. (1995) based their calculations on the axial HF+BCS theory with the SkM* effective interaction, and used parity projection and the GC method to investigate the octupole softness and octupole-quadrupole coupling in the superdeformed and hyperdeformed minimum of ¹⁹⁴Pb. They found an octupole excitation at about 2 MeV in both methods. A similar method, but with the SIII Skyrme parametrization, was used by Heenen et al. (1994) to study octupole excitations in light Xe and Ba nuclei around ¹¹⁴Ba. They found strong octupole correlations in this region in both GC and projected HF+BCS calculations. In the context of the parity-projected HF+BCS method and its relation to the GC method, Egido and Robledo (1991a, 1991b) investigated the importance of parity projection in the description of the negative-parity states in the light actinides. They concluded that, although it is an important effect, it is also necessary to take into account the collective correlations coming from the q_3 collective degree of freedom-and this is particularly important for the calculation of B(E3)'s. Robledo (1992) investigated the properties of the Lipkin-Meshkov-Glick model of Eq. (26). Here, the GC technique reproduces the exact solution (Ring and Schuck, 1980). To cure discrepancies in the intermediate region which appear in the projected HF theory, an alternative approach based on a twoconfiguration mixing has been proposed. As shown by Robledo (1992), this method gives very good agreement with the exact results for all deformation regimes.

Skalski *et al.* (1993a) have studied the coupling between axial quadrupole and octupole modes in $^{94-100}$ Zr using the GC+HF method with various Skyrme parametrizations.

The importance of nonaxial octupole components with $\mu=1$ and $\mu=2$ at superdeformed shapes in ^{192,194}Hg and ¹⁹⁴Pb was discussed by Skalski *et al.* (1993b). The authors concluded that the $\mu=0$ and $\mu=2$ octupole modes are to a large extent decoupled. The results of two-dimensional GC calculations for the coupled $\mu=0$ and $\mu=2$ modes are shown in Fig. 10. It is seen that the first excited state [Fig. 10(c)] can be understood in terms of the pure $\mu=0$ vibration, while the second excited state [Fig. 10(d)] shows the collective probability density characteristic of the octupole vibration with $\mu=2$.

The coupling between the axial (μ =0) quadrupole and octupole modes at normal and large deformations was addressed by two-dimensional GC+HFB+BCS calculations for ¹⁹⁴Pb (Bonche *et al.*, 1994; Meyer *et al.*, 1995). The resulting probability densities suggest that the picture of independent quadrupole and octupole excitations breaks down quickly with increasing excitation energy.

2. Time-dependent Hartree-Fock

The time-dependent Hartree-Fock (TDHF) method is a microscopic quantum-mechanical method that pro-



FIG. 10. Potential energy and probability densities of the three lowest GC-method states obtained for superdeformed ¹⁹²Hg in a two-dimensional (β_{30},β_{32}) calculation (Skalski *et al.*, 1993b): (a) energy (contour-line spacing 1 MeV); (b) ground state density; (c) first excited state density; (d) second excited state density.

vides a consistent description of all collective and singleparticle aspects of nuclear motion. The time evolution of the system is given by

$$i\dot{\rho}(t) = [h(t), \rho(t)], \tag{45}$$

where ρ is the density matrix, and *h* is the TDHF Hamiltonian, which is the sum of the kinetic energy and the time-dependent average field $\Gamma[\rho(t)]$ [Eq. (19)]. In this approach, the system itself determines the path in the multidimensional energy surface; no self-consistent symmetries, restricting the available phase space, are imposed. The wave function is a single product state, which allows for a transparent geometrical interpretation.

Strayer *et al.* (1984) used this method to study the time evolution of the $\alpha + {}^{14}C \leftrightarrow {}^{18}O$ system at energies near the Coulomb barrier, and determined the frequencies of isovector dipole and isoscalar quadrupole and octupole giant resonances in ${}^{18}O$. Umar *et al.* (1985) extended these calculations to the ${}^{12}C+{}^{12}C(0^+)$ and $\alpha + {}^{20}Ne$ systems, calculating molecular resonances in ${}^{24}Mg$.

Negele (1989) and Wolff *et al.* (1992) carried out a model calculation of fission of 32 S in three dimensions, also involving reflection-asymmetric deformations. They found a dramatic difference between the results of the static constrained HF (or adiabatic TDHF) and TDHF methods (see Sec. X.F).

3. Collective Schrödinger equation

By making the Gaussian-overlap approximation (GOA), the Hill-Wheeler equation of the GC method is reduced to the collective Schrödinger equation for the collective wave function:

$$\mathcal{H}_{\text{coll}}\Psi_i(\mathbf{q}) = E_i \Psi_i(\mathbf{q}),$$
(46)

where the collective Bohr Hamiltonian \mathcal{H}_{coll} depends on the collective mass parameters, the collective matrix, the potential energy, and the zero-point energy correction (Bohr, 1952; Ring and Schuck, 1980). If only one collective variable is considered, e.g., the octupole moment q_3 , \mathcal{H}_{coll} can be written as

$$\mathcal{H}_{\text{coll}} = -\frac{1}{\sqrt{G(q_3)}} \frac{\partial}{\partial q_3} \sqrt{G(q_3)} \frac{1}{2B(q_3)} \frac{\partial}{\partial q_3} + V(q_3)$$
$$-E_{\text{zpe}}(q_3), \qquad (47)$$

where $G(q_3)$ is the collective metric, $B(q_3)$ is the collective mass parameter, $V(q_3)$ is the collective potential, and $E_{zpe}(q_3)$ is the zero-point energy correction. The collective Schrödinger equation can also be obtained from the adiabatic TDHF method (Ring and Schuck, 1980). The resulting collective parameters are slightly different than in the GC+GOA method [for differences between the two approaches to the collective Schrödinger equation for octupole motion, see Egido and Robledo (1989, 1990)].

The collective Hamiltonian for octupole vibrations was studied by Donner and Greiner (1966). Their Hamiltonian contains a quadrupole-vibrational term, a rotational term, an octupole-vibrational term, and a quadrupole-vibrational interaction.

Zaikin (1966) and Krappe and Wille (1969) considered octupole vibrations with K=0 in the schematic collective Hamiltonian [Eq. (47)] with one collective coordinate β_3 :

$$\mathscr{H}_{\text{coll}} = -\frac{1}{2B_{30}} \frac{d^2}{d\beta_3^2} + V(\beta_3).$$
(48)

The lowest eigenstates of the Hamiltonian [Eq. (48)] with parity p are approximated [Eq. (20)] by a sum of two Gaussians centered at $\beta_3 = \pm \bar{\beta}_3$:

$$\chi(\beta_3) = \left(\frac{\lambda}{\pi}\right)^{1/4} \exp\left[-\frac{\lambda}{2}(\beta_3 - \bar{\beta}_3)^2\right].$$
(49)

The quantity λ in Eq. (49) is inversely proportional to the deformation spread of χ , and is assumed to be the variational parameter of the model. In order to obtain the variational solution, the projected ground-state energy E_+ has been minimized with respect to λ . The calculated B(E3) value,

$$B(E3) \propto |\langle \Psi_+ | \beta_3 | \Psi_- \rangle|^2 \propto \frac{\bar{\beta}_3^2}{1 - \langle \mathscr{P} \rangle^2}, \qquad (50)$$

depends mainly on the value of equilibrium deformation $\bar{\beta}_3$, and is rather insensitive to the height of the octupole barrier V_0 . On the other hand, the parity splitting shows significant dependence on V_0 (Nazarewicz and Tabor, 1992). Equation (50) demonstrates that if the nuclear potential becomes very soft with respect to the octupole deformation, the equilibrium deformation becomes a measure not of the deformation parameter but of the frequency, mass and inertia parameters describing the shape vibrations (Zaikin, 1966; Krappe and Wille, 1969; Bohr and Mottelson, 1975).

Leander and Sheline (1984) applied an intermediatecoupling scheme to the light actinides. The Hamiltonian of Eqs. (39)–(41), with H_{core} as in Eq. (48), and a particle-core coupling term approximated by $\kappa Q_{30}r^3P_3(\theta)$, was diagonalized in the reflectionsymmetric basis

$$\Psi_{IMK}^{\nu} = \mathcal{N}(1+R_1) D_{MK}^I \Phi_{n_2} \chi_K^{\nu}, \qquad (51)$$

where Φ_{n_3} is an eigenstate of $H_{\rm core}$. In particular, they obtained a very good description of the transitional nucleus ²²⁹Th. They concluded that, when applied to nuclei around ²²⁴Th, strong coupling generally works better than weak coupling.

Rohoziński and co-workers (Rohoziński, 1978; Rohoziński and Greiner, 1980; Rohoziński *et al.*, 1982; Rohoziński, 1988) extended the collective model of Donner and Greiner (1966) to take into account anharmonicities and couplings between modes in a quadrupole-octupole system. Their collective Hamiltonian can be written as

$$\mathcal{H}_{\text{coll}} = \mathcal{H}_{\text{vib}} + \mathcal{H}_{\text{rot}} + \mathcal{H}_{\text{vib-rot}}.$$
(52)

The vibrational Hamiltonian $\mathscr{H}_{\rm vib}$ describes the coupled intrinsic quadrupole-octupole vibrations, $\mathscr{H}_{\rm rot}$ is the rotational Hamiltonian, and $\mathscr{H}_{\rm vib-rot}$ is the rotationvibration Hamiltonian. After introducing the intrinsic collective variables $\alpha_{\lambda\mu}$ of Eq. (8), and the associated collective momenta and angular momenta, the Hamiltonian [Eq. (52)] acquires a simple geometric interpretation. The total collective angular momentum L is divided into the quadrupole angular momentum $L^{(2)}$ and the contribution $L^{(3)}$ coming from octupole vibrations (Donner and Greiner, 1966). The rotational Hamiltonian can be further decomposed into a rotor part, a Coriolis term (involving the coupling between L and $L^{(3)}$), and a centrifugal term. The collective Hamiltonian, Eq. (52), is diagonalized in the basis of D_{2h} -invariant wave functions of definite angular momentum L and parity.

The collective quadrupole-octupole model was applied to the limiting cases of spherical nuclei and welldeformed axially symmetric nuclei. Rohoziński and Greiner (1983) looked at the effects of Coriolis and centrifugal terms for nuclei with stable octupole deformation. They found dramatic effects on the moments of inertia and B(E3) rates. A simple collective quadrupole-octupole Hamiltonian for nuclei with static quadrupole and octupole deformations was studied by Dzyublik and Denisov (1993) and Denisov and Dzyublik (1993). They obtained simple analytic expressions for the energies of collective states in nuclei with static octupole deformations. Denisov and Dzyublik (1995) investigated the collective Hamiltonian with $\beta_2, \beta_3, \ldots, \beta_N$ deformations, considering also the case of static equilibrium deformations.

Barts *et al.* (1984) also took into account the interaction of octupole and quadrupole degrees of freedom. In their calculations the vibrational variables were separated using the Hartree method, and the wave functions determined using a variational principle. They noted that the inclusion of both degrees of freedom leads to strong breaking of the axial symmetry.

Böning *et al.* (1985) solved Eq. (46) for ²²⁶Th assuming the decoupling of odd- and even-multipole deformations (taken as collective coordinates), and taking constant mass parameters. The rotational degree of freedom was ignored in these calculations. See also Sobiczewski and Böning (1987) and Böning *et al.* (1985) for similar calculations for ²²⁴Ra. A similar approach, but with the mass parameter determined from the hopping model, was applied by Barranco *et al.* (1988) to the parity splitting in ²²²Ra.

Provoost *et al.* (1984) applied the adiabatic TDHF method to the α +¹⁶O \leftrightarrow ²⁰Ne system, and were able to reproduce the energy difference between the ground state and the lowest state of the negative-parity band.

More microscopic calculations based on the GC+GOE and adiabatic TDHF methods with the Gogny force D1 were performed in order to describe octupole correlations in the light lanthanides (Egido and Robledo, 1990, 1991a, 1991b, 1992; Martín and Robledo, 1984) and light actinides (Egido and Robledo, 1989, 1991a, 1991b; Robledo *et al.*, 1987, 1988).

Robledo *et al.* (1988) performed adiabatic TDHF calculations with q_2 and q_3 as collective variables to calculate the $(0^+) - (0^-)$ energy difference and E1,E3 transition rates for ²²²Ra. Egido and Robledo (1989) further applied adiabatic TDHF and GC+GOA techniques to a series of Ra and Th isotopes. They also applied similar methods to the ¹⁴²⁻¹⁴⁸Ba isotopes, and found that the octupole barriers are not high enough for the nuclei to be classified as octupole deformed, at least at zero spin (Egido and Robledo, 1990).

An extensive series of calculations using q_3 -constrained HF+BCS and adiabatic TDHF methods have also been reported (Egido and Robledo, 1992) for ¹⁴⁰Ba, ¹⁴²⁻¹⁵⁰Ce, ¹⁴⁴⁻¹⁵²Nd, and ¹⁴⁶⁻¹⁵⁴Sm.

4. Other methods

Microscopic many-body calculations were performed by Chasman (1979, 1980), who predicted parity doublets in several odd-A Th, Ac and Pa nuclei. These calculations employed a schematic Hamiltonian as in Eq. (13), with separable multipole-multipole forces, pairing, and projected many-body wave functions. Chasman (1989) has also investigated the role of a λ =5 particle-hole interaction, and found that the octupole correlation energy in the 1⁻ state is also independent of the nuclide, even though the corresponding changes in the correlation energy associated with the interaction of multipolarity λ =5 may be large.

In the multiphonon method, the octupole-plus-pairing Hamiltonian (13) is diagonalized in the truncated basis of multiphonon states

$$|n\rangle = \mathcal{N}_n(Q^+)^n |0\rangle. \tag{53}$$

The collective phonons Q^+ are calculated microscopically, in the Tamm-Dancoff approximation. Piepenbring (1983, 1984, 1985) and Leandri and Piepenbring (1989) applied this method to the Ra-Th isotopes, ¹⁴⁶Ba and ¹⁵²Sm. The multiphonon method gives a microscopic explanation of the low-lying $K^{\pi}=0^-$ bandhead, for which the RPA fails, and explains why the corresponding twophonon state appears at much higher energies than $2E(0^-)$. Jammari *et al.* (1983) have studied the quadrupole-octupole multiphonon method using a simple model involving two degenerate j=3/2 multiplets with different parities.

Kammuri and Kishimoto (1978) employed the Hamiltonian of Eq. (13) with quadrupole-quadrupole and octupole-octupole interactions and (monopole and quadrupole) pairing. Using the microscopic boson expansion technique, they studied properties of negativeparity bands in ¹⁰⁰Ru, ¹¹²Cd, ¹⁵⁰Sm, and ¹⁵²Gd.

Properties of negative-parity bands in ⁷²As have been studied by Petrovici *et al.* (1994) using a manydimensional variational approach based on the configuration mixing of symmetry-projected complex HFB mean fields. For the Hamiltonian, Petrovici *et al.* employed the (slightly modified) *G* matrix. So far, however, no systematic calculations of octupole states have been carried out using this method.

D. Algebraic models

Algebraic approaches to negative-parity states are usually based on the assumption that the nucleus can be well described as a system of fermion pairs with angular momentum 0, 1, 2, and 3 (s, p, d, and f pairs), often treated as phenomenological bosons. The predictive power of these models is limited; in most cases they concentrate on the reproduction of existing data.

The prescription of describing negative-parity states by the addition of an f boson to the usual s and d bosons of the interacting-boson model was first mentioned by Iachello and Arima (1974), given in more detail by Arima and Iachello (1975, 1976), and applied to the nucleus ¹⁵⁰Sm by De Voigt *et al.* (1975). The interactingboson Hamiltonian of *s*, *d*, and *f* bosons has the form

$$H_{\rm IBM} = H_{sd} + H_f + H_{\rm int}, \qquad (54)$$

where H_{sd} is the usual interacting-boson Hamiltonian describing the interacting s and d bosons,

$$H_{f} = \epsilon_{f} \sum_{\mu} f_{\mu}^{+} f_{\mu} + \frac{1}{2} \sum_{L=0,2,4,6} v_{ffff}^{L} \sum_{M} [f^{+} \times f^{+}]_{LM} [f \times f]_{LM}$$
(55)

is the general (two-body) *f*-boson Hamiltonian, and H_{int} is the general (two-body) interaction between the positive- and negative-parity bosons. For the inclusion of the *g* boson, see Dukelsky *et al.* (1983).

Because of the large number of parameters involved, in practical applications various simplifications have been made, such as assuming conservation of the *f*-boson number, or truncating terms in H_{int} . Sujkowski *et al.* (1977) and Scholten *et al.* (1978) described the excitation energies of the 1_1^- , 3_1^- , 5_1^- , and 1_2^- , 3_2^- levels in the Sm isotopes (using the *sdf* interacting-boson model) as transitional between the vibrational SU₅ limit and the rotational SU₃ limit. A systematic study of the octupole bands in rare-earth nuclei was performed by Barfield *et al.* (1986); see also Barfield *et al.* (1988).

Engel and Iachello (1985, 1987) suggested that octupole-deformed nuclei should be described by a system consisting of both p and f bosons, in addition to sand d bosons, whereas the description of octupole vibrations requires only additional f bosons. They have characterized the rotational spectra in the SU₃ and O₄ symmetry limits, and concluded that neither can be used to describe the experimental data, which are transitional between the octupole vibrational and octupoledeformed limits. Instead, they diagonalized the U₁₆ Hamiltonian with a dipole-dipole interaction added. Good fits were obtained for the energy spectra of ²¹⁸Ra and ¹⁴⁰⁻¹⁴⁸Ba (Kusnezov and Iachello, 1988; Liu et al. 1994). Figure 11 displays the spdf interactingboson model fit to the spectrum and the B(E1)/B(E2) branching ratios of ²¹⁸Ra (Engel and Iachello, 1987).

Han et al. (1985) carried out a systematic study of the negative-parity bands in even-even N=88 nuclei, using a unified set of parameters. They also found that the description of E1 transition rates and the position of 1^{-1} states could be improved by including a p boson. Otsuka (1986) investigated the structure of nuclear wave functions in deformed actinide nuclei for an intrinsic Hamiltonian containing an octupole and quadrupole (Nilsson) mean field. He found a significant number of collective dipole (1^{-}) nucleon pairs in the wave functions of lowlying states, so that p bosons should be added to the description of these nuclei. Otsuka and Sugita (1988) described the $K^{\pi}=0^+$, $K^{\pi}=0^-$ bands and transition moments in ²²⁰⁻²³²Th (and ²²⁰Ra) in terms of an spdf-boson model. In their calculations they employed the intrinsic-state formulation (Otsuka, 1986); the intrinsic wave function was assumed to be



FIG. 11. Theoretical fit to the spectrum (top) and B(E1)/B(E2) branching ratios (bottom) of ²¹⁸Ra using the *spdf* interacting-boson model (Engel and Iachello, 1987).

$$\Phi^{\alpha}(x_0s^+ + x_1p^+ + x_2d^+ + x_3f^+)^N |0\rangle, \tag{56}$$

with the amplitudes x_i determined from the variational principle (after angular momentum and parity projection). The coherent-state method was applied by Alonso *et al.* (1995) to the *spdf* SU₃ Hamiltonian with quadrupole and octupole interaction. The authors analyzed the transition to stable octupole deformations as a function of the octupole coupling constant κ_3 . Another example of the application of the coherent-state formulation based on the projected states of Eq. (56) is the 1/*N* expansion of Lac and Morrison (1995). Using the *spdf* Hamiltonian, they obtained a good fit to the low-spin ($I \le 10$) spectra of ^{226,228}Th, and to the *E*1, *E*2, and *E*3 transition rates.

Yoshinaga *et al.* (1993) applied the sdg protonneutron interacting-boson model with one f boson to a description of the negative-parity states in the Ra isotopes, assuming that there is no octupole deformation in these nuclei, at least at low spin. They were able to reproduce approximately the measured B(E1)/B(E2) branching ratios in these nuclei by adding the additional *g* boson.

Quasimolecular states have also been discussed in the framework of the algebraic vibron model (Iachello, 1981; 1984). The vibron model is a model of interacting monopole σ ($\ell^{\pi}=0^+$) and dipole π ($\ell^{\pi}=1^-$) bosons. Its group structure is given by the compact group U₄. There are two dynamical symmetry limits of the vibron model, O₄ and U₃. The former describes a system with a permanent dipole deformation, the rigid-molecule limit. The latter, the soft molecular limit, describes the rotational-vibrational spectrum around a spherical equilibrium shape. For extensions of the vibron model, see Daley and Iachello (1986), Cseh (1992). For the relation between the vibron model and the microscopic cluster model, see Cseh *et al.* (1991).

Iachello and Jackson (1982) suggested that α clustering may play an important role in the structure of heavy nuclei, and proposed a model in which the cluster states are built from s, d, and p proton and neutron pairs. This model is able to reproduce the typical low-lying negative-parity state bands seen in heavy nuclei, and small α -decay hindrance factors to members of these bands. Daley and Iachello (1983) developed this model further. The algebraic α -cluster model was used to describe the energy levels of ²²⁴Ra, and, with a different selection of parameter values, the B(E1)/B(E2) ratios in ²¹⁸Ra and ²²²Th (Daley and Gai, 1984). Extensive calculations of the values of many observables for a large range of Ra, Th, and U isotopes were presented by Daly and Barrett (1986). Daly and Nagarajan (1986) also applied the model to describe the excited collectiveparity bands in ¹⁵⁶Gd in terms of α clusters.

Alhassid *et al.* (1982) proposed that the underlying structure of an alternating-parity sequence arises from the existence of a molecular band, suggesting in particular an α +¹⁴C molecular band in ¹⁸O, and α +²¹⁴Rn in ²¹⁸Ra. They gave general expressions for sum rules for *E*1 and *E*2 transitions. Yang and Hwang (1987) obtained good agreement between predictions assuming molecular structure in a U₅ model and the observed levels in ¹⁸O and ²⁰Ne. Energy staggering in octupole bands in deformed nuclei was studied by Chou *et al.* (1992) and Casten *et al.* (1993) in the simplified *sdf* interacting-boson model. They noted that the staggered pattern of the negative-parity bands, as well as their ordering, results from underlying dynamical symmetry.

For odd-A systems, where the couplings with the odd fermion have to be taken into account, Engel *et al.* (1987) showed that the decoupling-inversion effect in ²²⁵Ra (Sec. V.D.3) is reproduced by the interacting boson-fermion model in the SU₃ limit involving a U₁₆ boson core (as described above) and 4s, 3d, 2g, and $1i_{13/2}$ single-particle orbitals (see also Alonso *et al.*, 1995). For even-even systems Chuu *et al.* (1993) coupled the *sdf* space of the interacting-boson model to a space spanned by a product of fermion pairs (occupying $f_{5/2}$ and $g_{9/2}$ orbitals) coupled to an *sd* space, to describe negative-parity energy levels in $^{64-76}$ Ge.

Ceauşescu and Raduta (1976) employed the boson expansion technique to derive microscopically the quadrupole-octupole Hamiltonian written in terms of the quadrupole and octupole boson operators. This Hamiltonian was applied by Badea *et al.* (1978) to ¹⁵⁰Sm, ¹⁵²Gd, and ²³²U using the projected coherent-state basis.

Catara *et al.* (1986) have investigated the general nature of negative-parity states in shell-model calculations truncated to generate S, P, D, and F fermion pairs, using a schematic quadrupole-quadrupole and octupole-octupole interaction and space of two degenerate orbitals with opposite parities. This collective pair approximation was extended by Yi *et al.* (1991) to describe energy levels in ¹⁸O by including particle-hole states arising from excitation of the core.

Mikhailov *et al.* (1989) applied an interacting multiboson model (*spdf* bosons) to describe the structure of ²¹⁸Ra. Dukelski *et al.* (1985) presented a description of low-lying octupole collective modes based upon a selfconsistent Hartree-Bose description involving many interacting bosons (ℓ =0, 2, and 3).

E. Cluster models

The natural model for describing cluster configurations in light nuclei is the alpha-cluster model (Dennison, 1940, 1954; Morinaga, 1956; Brink, 1957, 1966; Buck *et al.*, 1975; Rae, 1988). Wildermuth and Kanellopoulos (1958a, 1958b) considered systems composed of clusters heavier than an alpha particle, and suggested a description of ²⁰Ne in terms of an α +¹⁶O bimolecule [see also Sheline and Wildermuth (1960) and Cseh and Scheid (1992) for a discussion of various cluster configurations in light nuclei].

In the alpha-cluster model (Brink, 1966), the A-body state is constructed from N=A/4 orbitals (alpha clusters), each representing two protons and two neutrons in $1s_{1/2}$ states centered around point \mathbf{R}_j (j=1,...,N). The single-particle wave function can thus be represented by

$$\phi_i(\mathbf{r}) = (2\nu/\pi)^{3/4} \exp[-\nu(\mathbf{r} - \mathbf{R}_i)^2] \chi_i(\xi), \qquad (57)$$

where ν is the oscillator constant and $\chi(\xi)$ is the spinisospin function. The total wave function Φ is given by the parity- and angular-momentum-projected Slater determinant of the (nonorthogonal) single-particle states given by Eq. (57). In practical applications, the alpha coordinates \mathbf{R}_j are treated either as variational parameters or as collective coordinates in GC calculations.

The formation of clusters in nuclei can be treated more microscopically by antisymmetrized molecular dynamics (Horiuchi, 1991). The antisymmetrizedmolecular-dynamics wave function of the *A*-nucleon system is a parity- and angular-momentum-projected Slater determinant with the single-particle wave functions [cf. Eq. (57)]

$$\phi_j(\mathbf{r}) = (2\nu/\pi)^{3/4} \exp\left[-\nu\left(\mathbf{r} - \frac{\mathbf{Z}_j}{\sqrt{\nu}}\right)^2 + \frac{1}{2}\mathbf{Z}_j^2\right]\chi_j(\xi), \quad (58)$$

where Z_j (j=1,...,A) are the coordinate parameters of *all* nucleons, which are determined from the variational principle. A sample of antisymmetrized-molecular-dynamics calculation for ²⁰Ne is shown in Fig. 44 (see Sec. VIII).

F. Vibrational approaches

In this review we concentrate mainly on static reflection-asymmetric deformations. Consequently, vibrational approaches to octupole modes are discussed only briefly. For a comprehensive review see Rohoziński (1988) and references therein.

Historically, the observation of low-lying negativeparity excitations was explained early on as arising from octupole vibrations of the nuclear surface (Lane and Pendlebury, 1960). Early attempts were also made to reproduce energy levels and transition moments in the $K^{\pi}=0^{-}$ and $K^{\pi}=2^{-}$ bands observed in many nuclei by assuming octupole Y_{30} vibrations with additional $Y_{3\pm 2}$ asymmetric deformation (e.g., Lipas and Davidson, 1961; Davidson, 1962; Lipas, 1963; Leper, 1964). Davidson (1965) proposed a collective octupole Hamiltonian involving seven octupole degrees of freedom. For a critical discussion of those early models, see Rohoziński *et al.* (1982).

The collective model of octupole vibrations was developed by Donner and Greiner (1966), who described octupole states as arising from strong coupling of the octupole phonon to quadrupole rotational-vibrational states. They classified the resulting spectrum of energy levels, and were able to show that the anisotropy of the vibrations arises from the quadrupole-octupole interaction. The collective octupole coupling between states of opposite parity was considered by Zaikin (1966), who studied a collective Bohr Hamiltonian for octupole deformations with K=0 and ± 1 , and also considered the limit of a static octupole deformation.

As mentioned already in Sec. III.A, there have been many RPA calculations for negative-parity states in spherical nuclei. Yoshida (1962), Tamura and Udagawa (1962), Veje (1966), Raduta *et al.* (1970), and Vdovin and Soloviev (1983) employed separable octupoleoctupole interactions. Abbas and Zamick (1980) performed RPA calculations with the contact interaction, and Abbas *et al.* (1981) investigated the 3⁻ systematics in even-even nuclei with a continuum RPA+Skyrme approach [see also Bal'butsev *et al.* (1991)]. Quasiparticle RPA calculations of octupole states based on a realistic density-dependent interaction were carried out for ⁹⁶Zr by Rosso *et al.* (1993) and Fayans *et al.* (1994).

A microscopic approach based on the RPA to the nature of octupole vibrational states in deformed nuclei was developed by the Dubna group (Soloviev and Vogel, 1963; Soloviev, 1965, 1976; Ivanova *et al.*, 1976). The Hamiltonian used in their quasiparticle-phonon nuclear model is that of Eq. (13), with the single-particle Hamiltonian represented by the deformed WS potential. In this model, the microscopic phonons are built from twoquasiparticle excitations by means of the RPA, and the quasiparticle-phonon interaction is treated explicitly. Nosek *et al.* (1993a, 1993b) applied the model to parity doublets in nuclei around ¹⁵³Eu [see also Alikov *et al.* (1988) for inclusion of the Coriolis mixing]. For other calculations, see Bés (1963), Błocki and Kurcewicz (1969), Faessler *et al.* (1967), and Faessler and Plastino (1967).

Neergård and Vogel (1970a, 1970b) and Vogel (1976) generalized this approach to account for the Coriolis coupling between states of the intrinsic octupole quadruplet ($K^{\pi}=0^{-},1^{-},2^{-},3^{-}$). The Hamiltonian of a deformed nucleus was taken as

$$H = H_{\text{intr}} + H_{\text{rot}} + H_{\text{Coriolis}}, \qquad (59)$$

where

$$H_{\rm rot} = \frac{\hbar^2}{2\mathscr{J}} (\mathbf{I}^2 - I_3^2) \tag{60}$$

and

$$H_{\rm Coriolis} = -\frac{\hbar^2}{2\mathscr{J}} (I_+ J_- + I_- J_+), \tag{61}$$

where I is the total angular momentum and J is the angular momentum of the octupole phonon. The intrinsic Hamiltonian of Eq. (13) was approximated using the deformed single-particle field (modified harmonic oscillator), stretched octupole-octupole interaction, and seniority-pairing force. The octupole phonons obtained in the RPA are mixed by the Coriolis term Eq. (61), giving rise to the low-lying negative-parity excitations. The inclusion of the Coriolis coupling is crucial for the explanation of experimental B(E3) values, excitation energies of octupole states in deformed nuclei, and large moments of inertia of octupole bands. In particular, Neergård and Vogel (1970a, 1970b) demonstrated that the intrinsic J_+ matrix element in the RPA is very close to the spherical value:

$$\langle K+1|J_+|K\rangle = \sqrt{(3-K)(3+K+1)}.$$
 (62)

Vogel (1976) extended the RPA calculations to the higher-spin members of octupole bands. At high spin values, transition to the two-quasiparticle regime was predicted. After the two-quasiparticle component involving high-*j* particles becomes aligned with the rotational axis, the octupole band becomes fragmented. Leandri and Piepenbring (1993) diagonalized the Hamiltonian of Eq. (59) in the strong-coupling twoquasiparticle basis. They obtained a satisfactory description of low-spin states of negative parity in deformed nuclei.

If the rotation is treated by means of the cranking approximation, one has to consider the Routhian given by Eq. (28). Robledo *et al.* (1986) applied the cranked RPA theory based on the Hamiltonian containing the multipole-multipole interaction with λ =2, 3, and 4, with the multipole operators symmetrized with respect to signature (Sec. III.B.1), and including the monopolepairing interaction. Their calculations reproduce octupole bands in the actinides, and in particular the transition from the collective octupole regime to the quasiparticle regime discussed by Vogel (1976).

Since the RPA Hamiltonian of Eq. (13) describes particles moving in an average potential well, and interacting by schematic velocity-dependent forces, the wave functions of states with $K^{\pi}=0^{-}$ and 1^{-} contain some admixture of the c.m. motion. In order to project out this spurious component, one can add to the Hamiltonian an additional interaction guaranteeing that the translational and Galilean invariances

$$[H, \mathbf{P}_{\text{c.m.}}] = 0$$
 and $[H, \mathbf{R}_{\text{c.m.}}] = -i \frac{\hbar}{M_N} \mathbf{P}_{\text{c.m.}}$ (63)

are satisfied. In Eq. (63) the quantity $P_{c.m.}$ is the momentum vector of the c.m. In practical applications, the conditions given by Eq. (63) are satisfied on the RPA level (Pyatov and Salamov, 1977; Kvasil *et al.*, 1981, 1985; Ćwiok *et al.*, 1984). Neergård and Vogel (1970a, 1970b) and Robledo *et al.* (1986) applied this procedure and found that the c.m. correction is unimportant for the excitation energies and B(E3) rates of heavy nuclei.

In the context of the c.m. problem, of particular interest is the doubly stretched multipole-multipole separable interaction of Sakamoto and Kishimoto (1989), defined in terms of effective octupole operators

$$Q_{\lambda\mu}'' = r''^3 Y_{\lambda\mu}(\Omega''), \quad x_i'' \equiv \frac{\omega_i}{\omega_0} x_i, \quad (i = 1, 2, 3).$$
(64)

This interaction can be viewed as an improved conventional multipole-multipole force, especially when applied together with the harmonic-oscillator potential with frequencies ω_i . First, it satisfies nuclear selfconsistency rigorously, even if the system is deformed. Second, it yields the zero-energy RPA spurious modes, i.e., it automatically separates the translational and reorientation modes. Last, but not least, for this interaction the coupling between octupole and dipole modes disappears. The doubly stretched octupole interaction has been used by Mizutori et al. (1990, 1991a, 1991b), Nakatsukasa et al. (1992, 1993, 1995), and Crowell et al. (1995) to describe octupole correlations at high spins in superdeformed bands (see Sec. IX.B.2), and by Nakatsukasa et al. (1992, 1994), Nakatsukasa (1996), and Nazmitdinov and Åberg (1992) to analyze the influence of deformed-shell structures on octupole vibrations (see Sec. IX.B.1). Figure 12 (Nakatsukasa, 1996) illustrates the fragmentation of collective octupole phonons in ²³⁸U. In this figure, negative-parity Routhians obtained in the Nilsson cranked RPA method with the doubly stretched octupole interaction are shown as a function of rotational frequency. The lowest-lying rotationally aligned band built upon an octupole phonon is crossed at $\hbar \omega \approx 0.25$ MeV by a two-quasiparticle neutron band and quickly loses its collectivity. For higher-lying vibrational bands, this crossing and the following fragmentation appears at lower frequencies.

IV. EXPERIMENTAL SYSTEMATICS

The experimental evidence for reflection asymmetry in nuclei, if examined for an isolated nucleus or a par-



FIG. 12. Fragmentation of collective octupole vibrations in 238 U. Negative-parity Routhians obtained in the cranked RPA are shown as a function of rotational frequency. Large, medium, and small circles indicate RPA solutions with *E*3 transition amplitudes larger than 200 *e* fm³, between 100 *e* fm³ and 200 *e* fm³, and lower than 100 *e* fm³, respectively. Filled (open) circles indicate odd-*I* (even-*I*) band members. Experimental bands (Ward *et al.*, 1995) are denoted by stars, diamonds, and squares (Nakatsukasa, 1996).

ticular nuclear property, is not compelling. A much stronger case is evident, however, when a large body of data containing measurements of many nuclear properties in many nuclei is examined. A review of such data is presented in this and subsequent sections.

A. Low-lying 1⁻ and 3⁻ states in even-even nuclei

An obvious manifestation of reflection asymmetry in nuclei is the occurrence of low-lying negative-parity states which are collective in nature. States having such properties were first identified in Ra and Th isotopes with $N \approx 136$ by the Berkeley group (Asaro *et al.*, 1953; Stephens et al., 1954, 1955) using alpha spectroscopy. In this mass region the 1^- and 3^- states remain energetically higher than the 2^+ and 4^+ states, respectively, which rules out a simple interpretation in terms of octupole deformation. Only at higher spin do the negativeparity states become interspersed regularly with the positive-parity states (see Sec. IV.B). Figure 13 shows the variation of the energy of the 1^- , 3^- states as compared to the 2^+ , 4^+ states for selected lanthanide nuclei; Fig. 14 shows the corresponding behavior for Ra and Th isotopes. For actinide nuclei the minimum of the energy of negative-parity states is very localized in N, while there are insufficient data to determine the corresponding localization in Z. For the lanthanide region, this minimum value is attained outside the transitional region where octupole effects are strongest ($N \ge 90$). The systematic behavior of excited negative-parity states has been discussed by several authors. Neergard and Vogel (1970a) described the properties of negative-parity states in Ra and light Th isotopes in terms of the RPA (see Sec. III.F). Peker et al. (1981) also concluded that a vibrational interpretation is appropriate, and that the behavior of the negative-parity states can be explained in terms of Coriolis coupling between the $K^{\pi} = 0^{-}$, 1^- , 2^- , and 3^- bandheads. Sheline (1980) compared the behavior of these states in actinide nuclei with excited 0^+ states, and concluded that the $K^{\pi}=0^-$ bands and the excited $K^{\pi}=0^+$ bands are structurely more related to each other than to the ground-state band. However, the nature of the excited 0^+ states has not been explained satisfactorily. Zylicz (1986) excluded the interpretation of these states in terms of harmonic octupole vibrations. on the grounds that the ratio of energies of the excited 0^+ and 0^- bandheads, which should be 2 for vibrational structure, is in the range 3-4 for the light Ra nuclei.



FIG. 13. Excitation energies (keV) of the yrast 2^+ , 4^+ , 1^- , and 3^- states in the *N*=86–90 region.



FIG. 14. Excitation energies (keV) of the yrast 2^+ , 4^+ , 1^- , and 3^- states in the Z=86–92 region.

Kurcewicz *et al.* (1976, 1977) looked for lower-lying 0^+ states populated by alpha decay, without success.

A more extensive analysis of 1^- and 3^- states in actinide and lanthanide nuclei has been carried out by Cottle and Bromley (1986), who characterized the variation of excitation energy of 3^- states in terms of the onset of quadrupole deformation and the filling of the lower-energy member of a $\Delta \ell = 3$ pair. They also drew attention to the relation between the relative energies of the 3^- and the 1^- states and the ratio of energies between the 4^+ and the 2^+ states (see Fig. 15). Particularly for the lanthanides, the amount of quadrupole deformation plays a dominant role in determining the relative spacing of the 1^- and 3^- members of the $K^{\pi}=0^-$ band. This investigation has also been extended (Cottle *et al.*, 1988; Cottle, 1990a) to other regions.

Zamfir *et al.* (1989) have established a simple parametrization for the energies of 3^- states in all nuclei with $A \ge 30$ (Fig. 16):

$$E(3_1^-) = 19A^{-1/3} - 0.5N_t, \tag{65}$$

where N_t is the sum of the valence-nucleon numbers. Deviations from normal behavior characterize nuclei having the strongest octupole correlations (they are in the transitional lanthanide and actinide regions).

B. Alternating-parity rotational bands

The striking experimental feature of even-even nuclei with $Z \approx 88$, $N \approx 134$ and $Z \approx 60$, $N \approx 88$ is the interspacing of negative- and positive-parity states with the sequence $I^+, (I+1)^-, (I+2)^+, \ldots$, for states with I > 5. The first observations of such band structure in heavy nuclei were in ²¹⁸Ra (Fernández-Niello *et al.*, 1982) and ²²²Th (Ward *et al.*, 1983; Bonin *et al.*, 1983). In medium-mass nuclei, sequences of nuclear states with similar features were observed much earlier, for example, in ¹⁵²Gd (Zolnowski *et al.*, 1975), in ¹⁵⁰Sm (Sujkowski *et al.*, 1977), and in ¹⁵⁰Gd (Haenni and Sugihara, 1977), but the first connection with static octupole deformation in this mass region was made by Phillips *et al.* (1986), who studied ^{142,144,146}Ba using fission spectroscopy.



FIG. 15. The systematic behavior of $E(3_1^-) - E(1_1^-)$ in the regions (a) Z=56-70 and (b) Z=82-90. Parentheses denote tentative assignments (Cottle and Bromley, 1986).



FIG. 16. Mass-number dependence of experimental 3_1^- energies, corrected for a $-0.5\sqrt{N_t}$ dependence, where N_t is the sum of the valence-nucleon numbers, $N_t = N_p + N_n$. All non-doubly-closed shell nuclei with $A \ge 30$ and $N_t < 26$ are included. The curve is an $A^{-1/3}$ function (Zamfir *et al.*, 1989).

Tables I and II list the nuclei in which such bands have been established, together with the reactions and experimental methods employed. Figure 17 shows a typical example, ²²⁶Th, in which the sequence has been observed up to spin 20 (Schüler et al., 1986, Ackermann et al., 1993). Available data for even-even nuclei with Z=88 and 90, and N=86, 88, and 90 are shown in Figs. 18 and 19. In these figures, the data are plotted as \mathcal{J}/\hbar^2 versus $\hbar\omega$ for each parity band; here $\mathcal{J}=\hbar I_x/\omega$ is the kinematic moment of inertia, $I_x = \sqrt{(I+1/2)^2 - K^2}$ $(I, K \text{ in units of } \hbar)$ is the angular momentum perpendicular to the symmetry axis, and $\hbar \omega = dE/dI_x$ is the rotational frequency. The division between nuclei exhibiting characteristics of reflection asymmetry and reflection symmetry is apparent for N>136 for Th nuclei, for which data are most widely available. The nature of the symmetry becomes apparent in the plots of $\omega(\pi = -1)/\omega(\pi = +1)$ versus I. This ratio should equal 1 for perfectly reflection-asymmetric nuclei, but equals [4(I-3)-2)]/(4I-2) for rotation of an alignedoctupole phonon (Nazarewicz and Olanders, 1985a) (see Fig. 20).

In nuclei with $Z \approx 60$, the degeneracy of the positive and negative bands is only apparent between I=7 and I=13, and for N=86 and N=88, in the few cases where extensive data are available. In ¹⁵⁰Sm for example, the positive-parity band appears to cross a reflectionsymmetric band at $I \approx 15$ (Urban *et al.*, 1987) see Sec. VI. For N=90 nuclei the increase in quadrupole deformation pushes the positive-parity states to much lower energies than the octupole states, and the latter states are usually interpreted as being vibrational in origin (see also the discussion in Sheline and Sood, 1986).

Similar structures to that of even-even octupole nuclei are observed in transitional odd-mass and odd-odd nuclei in which the odd particles are weakly coupled to the core. Figure 21 shows the yrast sequence of 219 Ra (Cottle *et al.*, 1986) compared to its neighbors 218,220 Ra. (The term "yrast" denotes the lowest-energy state hav-

ing a given angular momentum.) The ground-state band of ²¹⁹Ra shows an alternating-parity structure, and is consistent with an interpretation in terms of an odd neutron weakly coupled to an average ²¹⁸Ra-²²⁰Ra core. Similar examples have been found in other nuclei with Z=87-90 and N<132: ²¹⁷Fr (Aïche *et al.*, 1988), ²¹⁷Ra (Roy et al., 1984) ²¹⁹Ac (Drigert and Cizewski, 1985; Khazrouni et al., 1985; Cristancho et al., 1994), ²²¹Ac (Aïche et al., 1994), ²²¹Th (Dahlinger et al., 1988), and in the odd-odd nuclei 216 Fr (Debray *et al.*, 1990), 218 Ac (Debray *et al.*, 1994), and 220 Ac (Schulz *et al.*, 1991). In the light-lanthanide region the level structure of transitional odd-A nuclei has been interpreted variously in terms of a $h_{11/2}$ proton coupled to an octupole phonon, as in the case of ¹⁵¹Eu (Vermeer et al., 1993, see also Jongman et al., 1994), and a neutron coupled to a reflection-symmetric triaxial core in the case of ¹⁵¹Sm (Khan et al., 1994; see also Basu et al., 1994). Alternating-parity structures have also been observed in ¹⁴³Ba (Zhu et al., 1995) and in ¹⁴⁹Sm (Basu et al., 1994). For nuclei farther away from the closed shell, better examples of parity doubling are seen, as discussed in Sec. V.D.1.

C. Enhanced E1 transitions

A common property of nuclei exhibiting the features of reflection asymmetry is the occurrence of relatively large E1 transition probabilities between the yrast positive- and negative-parity bands. The B(E1) values in these mass regions range from 10^{-4} to 10^{-2} s.p.u. [typical B(E1) values are less than 10^{-5} s.p.u.] Assuming the strong-coupling limit and axial shape, there is a simple relation between the E1 transition probability and the intrinsic (transition) electric dipole moment D_0 (see Sec. II.B):

$$B(E1;IK \to I'K) = \frac{3}{4\pi} D_0^2 \langle IK10 | I'K \rangle^2.$$
(66)

In the presence of Coriolis coupling and/or triaxiality, relation (66) has to be modified. For K=1/2, for instance, there appears in lowest order a signature-dependent term proportional to (Bohr and Mottelson, 1975)

$$B(E1; I \to I') = \frac{3}{4\pi} |(I \frac{1}{2} 10 |I' \frac{1}{2})D_0 + (-1)^{I+\frac{1}{2}} (I - \frac{1}{2} 11 |I' \frac{1}{2})D_1|^2, \quad (67)$$

where $D_{\mu=1}$ is the spherical component of **D** (see Sec. II.B). [A similar term appears in nuclei with low-lying octupole vibrational states, through the Coriolis coupling between K=0 and K=1 bands (see Sec. VII.E)].

In most cases, absolute values of B(E1) are not available, and D_0 has to be extracted from the known B(E1)/B(E2) branching ratios $T(E1)_{I \rightarrow I-1}/T(E2)_{I \rightarrow I-2}$, in which case the B(E2) rates are assumed to follow the rotational relationship

$$B(E2) = \frac{5}{16\pi} Q_0^2 \langle IK20 | I'K \rangle^2,$$
(68)

Nucleus	Reaction	Reference
²¹⁶ Fr (Z=87)	208 Pb(11 B,3 n)	Debrey et al. (1990)
²¹⁷ Fr	210 Pb(11 B,4 n)	Aïche et al. (1988)
217 Ra (Z=88)	208 Pb(12 C,3 <i>n</i>), 208 Pb(13 C,4 <i>n</i>)	Roy et al. (1984)
²¹⁸ Ra	208 Pb(13 C,3 <i>n</i>)	Fernández-Niello et al. (1982)
		Gono et al. (1986)
	$^{13}\mathrm{C}(^{208}\mathrm{Pb},3n)$	Gai et al. (1988)
	208 Pb(14 C,4 <i>n</i>)	Schulz et al. (1989)
	208 Pb(13 C,3 <i>n</i>)	Wieland et al. (1992b)
²¹⁹ Ra	208 Pb(14 C,3 <i>n</i>)	Cottle <i>et al.</i> (1986)
		Wieland et al. (1992a)
²²⁰ Ra	208 Pb(18 O, $\alpha 2n$)	Burrows et al. (1984)
	208 Pb(14 C,2 <i>n</i>)	Cottle <i>et al.</i> (1984)
		Celler et al. (1985)
		Shriner et al. (1985)
	208 Pb(18 O, $\alpha 2n$)	Smith <i>et al.</i> (1995)
²²¹ Ra	210 Pb(14 C,3 <i>n</i>)	Fernández-Niello et al. (1991)
²²⁴ Ra	²²⁶ Ra(⁵⁸ Ni, ⁶⁰ Ni)	Poynter et al. (1989a)
	226 Ra($\alpha, \alpha' 2n$)	Marten-Tölle et al. (1990)
²²⁶ Ra	²²⁶ Ra Coulomb excitation	Wollersheim et al. (1993)
	226 Ra(d,pn)	Ackermann et al. (1993)
²¹⁸ Ac (Z=89)	209 Bi(12 C,3n)	Debray et al. (1994)
	209 Bi(13 C,4n)	Debray et al. (1994)
²¹⁹ Ac	209 Bi(13 C,3n)	Drigert and Cizewski (1985, 1986)
		Khazrouni et al. (1985)
		Cristancho et al. (1994)
²²⁰ Ac	209 Bi(14 C,3n)	Schulz et al. (1990, 1991)
²²¹ Ac	$^{209}\text{Bi}(^{14}\text{C},2n)$	Aïche et al. (1994)
²²⁰ Th (Z=90)	208 Pb(16 O,4 <i>n</i>)	Bonin et al. (1985)
²²¹ Th	208 Pb(16 O,3 <i>n</i>)	Dahlinger et al. (1985, 1988)
²²² Th	208 Pb(18 O,4 n)	Ward <i>et al.</i> (1983)
		Bonin et al. (1985)
		Schwartz et al. (1987)
		Smith <i>et al.</i> (1995)
²²³ Th	208 Pb(18 O,3 <i>n</i>)	Dahlinger et al. (1988)
²²⁴ Th	208 Pb(18 O,2 <i>n</i>)	Schwartz et al. (1986)
	226 Ra($\alpha, 6n$)	Schüler et al. (1986), Ackermann et al. (1993)
²²⁵ Th	226 Ra(α ,5n)	Hughes et al. (1990)
²²⁶ Th	226 Ra(α ,4 n)	Schüler et al. (1986), Ackermann et al. (1993)
²²⁸ Th	226 Ra $(\alpha, 2n)$	Schüler et al. (1986), Ackermann et al. (1993)
²³⁰ U (Z=92)	230 Th(α ,4 n)	Ackermann et al. (1993)

(69)

TABLE I. Observed alternating-parity rotational bands in nuclei from the Ra-Th region, and reactions used.

where Q_0 is the transition quadrupole moment. Since many nuclei in the mass regions of interest are not good rotors, the use of the strong-coupling formulas, Eqs. (66) and (68), is questionable. Nevertheless, they provide a consistent way to extract D_0 from the data. In particular, for K=0 bands the intrinsic dipole moment is

Tables III and IV give examples of
$$B(E1)$$
 values measured for even-even nuclei in the light-lanthanide and actinide regions.

The experimental values of D_0 are shown in Fig 22, where they are compared with the theoretical values of Butler and Nazarewicz (1991) (Sec. VII.B) and of Egido and Robledo (1991b, 1992) (Sec. VII.C). The trends shown in Fig. 22 demonstrate that large fluctuations with

 $D_0 \approx \left\{ \frac{5(I-1)}{8(2I-1)} \frac{B(E1; I \to I-1)}{B(E2; I \to I-2)} \right\} Q_0.$

Nucleus	Reaction	Reference	
$\overline{^{142}\text{Ba}(Z=56)}$	²⁵² Cf fission	Phillips et al. (1986)	
	²⁵² Cf, ²⁴² Pu fission	Zhu et al. (1995)	
¹⁴³ Ba	²⁵² Cf, ²⁴² Pu fission	Zhu et al. (1995)	
¹⁴⁴ Ba	²⁵² Cf fission	Phillips et al. (1986)	
	²⁵² Cf, ²⁴² Pu fission	Zhu et al. (1995)	
¹⁴⁶ Ba	²⁵² Cf fission	Phillips et al. (1986)	
	²⁵² Cf, ²⁴² Pu fission	Zhu et al. (1995)	
¹⁴⁴ Ce (Z=58)	²⁵² Cf, ²⁴² Pu fission	Zhu et al. (1995)	
¹⁴⁶ Ce	²⁵² Cf fission	Phillips et al. (1988)	
¹⁴⁸ Ce	²⁵² Cf fission	Phillips et al. (1988)	
¹⁴⁶ Nd (Z=60)	¹⁵⁰ Nd($\alpha, \alpha' 4n$)	Urban et al. (1988)	
	136 Xe(13 C,3 <i>n</i>)	Urban et al. (1991)	
¹⁴⁸ Nd	¹⁵⁰ Nd($\alpha, \alpha' 2n$)	Urban et al. (1988)	
	²⁵² Cf fission	Durell et al. (1988)	
	¹⁴⁸ Nd Coulomb excitation	Ibbotson et al. (1991, 1993)	
¹⁵¹ Pm (Z=61)	¹⁵⁰ Nd(α ,p2n)	Vermeer et al. (1990)	
		Urban et al. (1990)	
148 Sm (Z=62)	130 Te(22 Ne,4 <i>n</i>)	Urban et al. (1991)	
¹⁴⁹ Sm	¹⁴⁸ Nd(α ,3 n)	Basu et al. (1994)	
¹⁵⁰ Sm	150 Nd(α ,4 n)	Sujkowski et al. (1977)	
		Urban et al. (1987)	
¹⁵¹ Sm	150 Nd(α ,3 n)	Basu et al. (1994)	
		Khan et al. (1994)	
¹⁴⁹ Eu (<i>Z</i> =63)	139 La(13 C,3 n)	Jongman et al. (1994)	
¹⁵⁰ Eu	136 Xe(19 F,5 <i>n</i>)		
	148 Nd(7 Li,5 <i>n</i>)	Jongman et al. (1994)	
¹⁵¹ Eu	136 Xe(19 F,4 <i>n</i>)		
	148 Nd(⁷ Li,4 <i>n</i>)	Jongman et al. (1994)	
	150 Nd(6 Li,5 <i>n</i>)	Vermeer et al. (1993)	
¹⁵³ Eu	150 Nd(⁷ Li,4 <i>n</i>)	Pearson et al. (1994)	
¹⁵⁰ Gd (Z=64)	150 Sm(α ,4 n), 152 Sm(α ,6 n)	Haenni and Sugihara (1977)	
¹⁵² Gd	¹⁵⁰ Sm(α ,2 n), ¹⁵² Sm(α ,4 n)	Zolnowski et al. (1975)	

TABLE II. Observed alternating-parity rotational bands in nuclei from the Ba-Sm region, and reactions used.

Z and N in the values of D_0 can occur due to shell effects. These fluctuations are discussed in more detail in Sec. VII.

There is less information on the systematic behavior of D_0 as a function of angular momentum. For the transitional nucleus ²¹⁸Ra, a decrease in $B(E1;I \rightarrow I-1)$ is observed in the spin range 6–8, although it is not clear whether this is associated with the large drop in B(E2) observed below spin 4 (Gai *et al.*, 1988). Gai (1988) has pointed out that the fraction of the E2 energy-weighted sum rule exhausted by the lowest 2⁺ state is unusually large for the light Ra and Th nuclei. Figure 23 displays the available data for the Ra-Th region for heavier nuclei, whose quadrupole deformation is rather stable; the rotation-induced variation in the quadrupole moment should not influence the value of D_0 . For most of these cases the value of D_0/Q_0 stays rather constant with spin, at least above I=6. One exception is ²²⁶Ra, where a dip in the value of this quantity is seen at $I\approx 4$. This feature cannot be reproduced by macroscopic-microscopic theories of D_0 (Leander *et al.*, 1986; Butler and Nazarewicz, 1991) for nuclei whose shape remains constant with spin. For ²²⁶Ra this seems to be the case, as indicated by the behavior of its quadrupole and octupole moment (see Sec. IV.D). However, fluctuations in the value of D_0 might be expected because of the cancellation effects in its macroscopic and microscopic components that are responsible for small E1 moments in ²²⁴Ra. In the lanthanides, a similar effect is seen in ¹⁴⁶Ba, which shows fluctuations with spin;



FIG. 17. Level scheme of ²²⁶Th, taken from Schüler *et al.* (1986; see also Ackermann *et al.*, 1993). The level and transition energies are in keV.

for this nucleus the overall value of D_0 is rather small, and is probably sensitive to shell effects (see Sec. VII.B). In ¹⁵⁰Sm the values of D_0 are seen to decrease markedly for transitions deexciting higher-spin members of the positive-parity ground-state band. This behavior is associated with the restoration of reflection symmetry in this nucleus above spin 10 (see Sec. VI). Similar changes in D_0 are not apparent in ¹⁴²Ba (Phillips *et al.*, 1986; Mach, Nazarewicz *et al.*, 1990; Zhu *et al.*, 1995).

The occurrence of large E1 transition strengths between low-lying states is not confined to nuclei in the "octupole" mass regions around ¹⁴⁴Ba and ²²²Th. The largest B(E1) so far observed experimentally (Millener et al., 1983) is 0.36 s.p.u., for the decay of the $1/2^{-}$ state in ¹¹Be. It is believed that this strong E1 transition is related to halo properties of ¹¹Be (see Sec. X.D). Large E1 transition strengths are also observed in sd-shell nuclei for which $Z \neq N$, for example in ¹⁸O (Gai *et al.*, 1983; see also Sec. VIII.B). In heavy nuclei, studies have revealed that large E1 transition strengths are observed between high-spin states in nuclei such as ¹⁶³Er that are well removed from the well-established octupole regions (Butler, 1990; see also Garrett, 1984; Balodis et al., 1991; Ogaza et al., 1993; Brockstedt et al., 1994; Jongman et al., 1994). Nuclear resonance fluorescence studies have shown that the summed isovector E1 strength for ground-state transitions to low-lying states remains remarkably constant for the mass region A = 150 - 174(Zilges et al., 1991; Friedrichs et al., 1992). The strongest E1 transition observed in these rare-earth nuclei is usually to the lowest 1^- state, although enhanced E1 tran-



FIG. 18. Kinematic moment of inertia \mathcal{J}/\hbar^2 (in MeV⁻¹) as a function of rotational frequency $\hbar\omega$ (in MeV) for rotational bands in even-even nuclei with N=86,88,90.

sitions have also been observed to excitations near 2.5 MeV in rare-earth nuclei (Kneissl *et al.*, 1993) and near 3.5 MeV in ^{116,124}Sn (Govaert *et al.*, 1994) (these have been interpreted as arising from two-phonon octupole- γ vibrational excitations). The experimental (*e*,*e'*) form factor for the lowest 1⁻ state in ⁴⁸Ti, ¹⁶⁴Dy, ²³²Th, and ²³⁸U has been described in terms of surface octupole vibrations (see Sec. VII.E). Large *B*(*E*1) strengths are also observed for low-lying transitions in nuclei near closed shells.

D. E3 transitions

For low-lying states in nuclei, the ground-state E3 transitions are predominantly isoscalar, and typically exhaust 4–7 % of the isoscalar energy-weighted sum rule (Kirson, 1982; Pignanelli, 1990). Consequently, low-energy B(E3) values are good measures of octupole collectivity (Rohoziński, 1988). In the limit of strong coupling, the octupole deformation β_3 is related



FIG. 19. Kinematic moment of inertia \mathcal{J}/\hbar^2 (in MeV⁻¹) as a function of rotational frequency $\hbar\omega$ (in MeV) for rotational bands in even-even nuclei with Z=88,90.

to the reduced transition probability $B(E3)\uparrow \equiv B(E3;0^+\rightarrow 3^-)$ from the ground state to the first 3^- state:

$$\beta_3 = \frac{4\pi}{ZR^3} \left[\frac{B(E3)\uparrow}{e^2} \right]^{1/2},\tag{70}$$

where the value of $B(E3)\uparrow$ can be deduced from the partial mean lifetime for E3 γ -ray emission to the ground state:

$$\tau_{E3}(s) = (0.0123) E(\text{MeV})_{3_1^{-1}}^{-7} [B(E3)\uparrow/e^2 \text{ fm}^6]^{-1}.$$
 (71)

The values of the transition probability and experimental octupole deformations deduced from $0^+_{g.s.} \rightarrow 3^-_1$ transitions have been compiled (Spear, 1989; see also Raman *et al.*, 1991) and are displayed in Fig. 24. Spear and Catford (1990) noted that the maxima in the B(E3) values at N = 34, 56, 88, and 134, and (more ambiguously) at Z=30, 40, 62, and 88 (see Fig. 24) are well correlated with the predicted regions of stable octupole deformation (Sec. III.B). The values of B(E3)are typically 30–50 s.p.u. for lanthanide nuclei in the region of octupole instability; they are plotted as a function of N in Fig. 25. It is evident that this quantity reaches a maximum for N=88–90. This might arise from the change in octupole collectivity between these neutron numbers, as suggested by the corresponding behavior in energy-level spectra (Fig. 18), or from the change in quadrupole deformation, which splits the E3 strength over the components $K^{\pi}=0^{-}$, 1⁻, 2⁻, and 3⁻ (Scholten *et al.*, 1978).

Raman *et al.* (1991) analyzed the anharmonicity of the octupole mode by inspecting the correlation between $B(E3)\uparrow$ and E_{3-} . Guided by the relation

$$B(E3)\uparrow \propto Z^2 A^{1/3} E_{3^{-1}}^{-1} \tag{72}$$

obtained in the limit of the hydrodynamical model (Bohr and Mottelson, 1975), they analyzed the data according to the expression

$$B(E3)\uparrow Z^{-2}A^{-1/3} = KE_{3^{-}}^{\eta},\tag{73}$$

where the values of K and η were obtained by means of a least-squares fit to experimental values. They obtained values of η ranging between $\eta \sim -0.7$ in spherical nuclei and $\eta \sim -0.5$ in deformed nuclei, indicating strong anharmonicities.

Although $B(E3;I \rightarrow I')$ has been measured for the transition from ground-state to the lowest 3⁻-state in many nuclei, there is less information on this quantity for higher-lying members of rotational bands in transitional and deformed nuclei. It has been realized that the yields of high-spin members of octupole bands following multiple Coulomb excitation are quite sensitive to the E3 matrix elements connecting them to the ground-state band (Butler, 1988). Figure 26 shows the experimental values for $\lambda = 2,3$ of the matrix elements

$$\langle I \| E_{\lambda} \| I' \rangle = \{ (2I+1)B(E\lambda; I \to I') \}^{1/2}$$

$$= \left\{ \frac{(2I+1)(2\lambda+1)}{16\pi} \right\}^{1/2} \langle I0\lambda 0 | I'0 \rangle \mathcal{O}_{\lambda0,c},$$
(75)

obtained for the ground-state $K^{\pi} = 0^+$ band and the lowest $K^{\pi}=0^{-}$ band in ¹⁴⁸Nd (Ibbotson *et al.*, 1993) and in ²²⁶Ra (Wollersheim et al., 1993). In Eq. (75), obtained assuming axial shape and the rotational limit, $Q_{20,c}$ = Q_0 , and $Q_{30,c}$ is the octupole moment. For ²²⁶Ra both quadrupole and octupole matrix elements can be fitted with the constant values $Q_{20,c}=750 \, \text{fm}^2$ and $Q_{30,c}$ =3100 fm³ (Wollersheim *et al.*, 1993). Application of Eqs. (6) and (7) to the measured matrix elements in ²²⁶Ra yields $\beta_3 \approx 0.10$ (Wollersheim *et al.*, 1993), in agreement with calculations by Leander et al. (1982) and Sobiczewski et al. (1988) (see Sec. III.B.2). In ¹⁴⁸Nd the corresponding values are $Q_{20,c}$ =400 fm² and $Q_{30,c}=1500 \text{ fm}^3$ (Ibbotson *et al.*, 1993), with a value of β_3 (≈ 0.12) somewhat higher than that predicted by mean-field calculations (Urban et al., 1988).

Nazarewicz and Tabor (1992) have used the collective model of Krappe and Wille (1969) for octupole deformation (Sec. III.F) to show that $B(E3;0^+ \rightarrow 3^-)$ is independent of the curvature of the nuclear potential, whereas the parity splitting is a strong function of this quantity. This implies that if the curvature varies with angular momentum, $Q_{30,c}$ should remain constant, in contrast to the parity splitting. In addition to this effect, Rohoziński and Greiner (1983) have concluded that the



FIG. 20. Plot of $\omega(\pi = -1)/\omega(\pi = +1)$ versus *I* for nuclei with *N*=86,88,90 and nuclei with *Z*=88,90. This ratio should equal unity for rotating rigidly reflection-asymmetric systems. The dashed line shows how it varies in the case of an aligned octupole phonon.

octupole matrix elements should be roughly independent of Coriolis and centrifugal effects, which gives another mechanism for the constancy of $Q_{30,c}$.

In the *sdf* version of the IBM (cf. Sec. III.D), the *E*3 operator is a one-boson operator obtained by a direct coupling of *sf* and *df* bosons to J=3:

$$Q_{3\mu} = e_3^{(sf)} \{ [s^+ \times f^*]_{3\mu} + [f^+ \times s^*]_{3\mu} \}$$

+ $e_3^{(df)} \{ [d^+ \times f^*]_{3\mu} + [f^+ \times d^*]_{3\mu} \}.$ (76)

Scholten *et al.* (1978) employed Eq. (76) to calculate B(E3) values for the Sm isotopes.

For particular closed-shell or sub-shell nuclei in which there are low-energy particle-hole E3 excitations for both protons and neutrons, large $B(E3,0^+\rightarrow3^-)$ values have been observed (Spear, 1989), for example, in ¹⁶O (14 s.p.u.), ⁴⁰Ca (31 s.p.u.), ¹³²Sn (> 7 s.p.u.; Fogelberg *et al.*, 1994), and ²⁰⁸Pb (34 s.p.u.). The largest values for observed transitions have been measured in ⁹⁶Zr (Mach, Ćwiok *et al.*, 1990; Ohm *et al.*, 1990; Hofer *et al.*, 1993; Horen *et al.*, 1993), \approx 50–60 s.p.u.(see Sec. X.A for calculations), and in ¹⁴⁸Gd (77±11 s.p.u.; Piiparinen *et al.*, 1993). Large *E*3 transition strengths have also been reported in the octupole transitional nucleus ²²⁹Th between the ground state and the octupole vibrational band built upon the ground state (Bemis *et al.*, 1988). For the *E*3 transitions in the rotational bands of ¹⁴⁸Nd and ²²⁶Ra, deduced from multiple Coulomb excitation (see above), much larger values have been measured, for example, 142±15 s.p.u. for the 9⁻→6⁺ transition in ²²⁶Ra (Wollersheim *et al.*, 1993).

V. PROPERTIES OF LOW-LYING STATES

The measured properties of nuclei in their ground state (binding energies, decay properties, and properties of the odd particle) show detailed evidence for strong octupole correlations, which is quite separate from the signature given by a rotational spectrum; a review of these properties is given in this section. For earlier reviews, the reader is referred to the seminal works of



FIG. 21. The yrast sequence of 219 Ra, taken from Cottle *et al.* (1986), compared to its neighbors 218,220 Ra.

Leander and co-authors (Leander *et al.*, 1982; Leander and Sheline, 1984; Leander and Chen, 1988).

A. Binding energies

An early theoretical indicator of intrinsic reflection asymmetry in nuclei was the necessity of including oddmass deformation parameters in calculations of groundstate masses. Möller and Nix (1981) were the first to demonstrate that the discrepancies in the calculated and experimental values of the ground-state masses of nuclei with $Z \approx 88$, $A \approx 222$ were reduced substantially by allowing the octupole-deformation parameter to vary in their calculations, which were based on the SC method with the folded Yukawa average potential and the Yukawa-plus-exponential macroscopic energy. These calculations were extended by Leander et al. (1982), who made a systematic study of nuclei with $84 \le Z \le 92$ and $130 \le N \le 140$. Leander *et al.* (1982) pointed out that the triangular closed classical orbits with $\Delta \ell = 3, 6, ...$ in a shell, which give rise to octupole deformation, may only be present with a realistic nuclear potential such as the folded Yukawa potential used by them, or a WS potential, and cannot occur for a modified harmonicoscillator potential, which can only favor quadrupole and hexadecapole distortions. The results of these calculations are shown in Fig. 27, which shows how the calculated mass-discrepancies evident for nuclei with $Z \approx 88$ and $N \approx 134$ are reduced substantially by octupole deformation. The deviations shown in Fig. 27 occur in precisely the same nuclei for which the calculated potential minimum is lowest for a reflection-asymmetric shape. Möller *et al.* (1995) have recently compiled masses calculated using both the droplet model and the folded Yukawa microscopic model, which can be used to map regions of octupole deformation throughout the periodic table.

Leander et al. (1982) found that the octupoledeformed minima are typically 1-2 MeV lower than the energy in the modified oscillator at the same deformation. This difference was further discussed by Nazarewicz et al. (1984b), who pointed out that the bindingenergy gain E_{def,asym} associated with reflectionasymmetric deformations is governed primarily by the spacing between the strongly interacting subshells with $\Delta \ell = 3$ [see Sec. III.A and Eq. (15)]. For instance, the energy separation between the spherical $2f_{7/2}$ and $1i_{13/2}$ proton shells in ²²⁴Th is 50 keV in the folded Yukawa model of Leander et al. (1982), and 600 keV in the WS model of Nazarewicz et al. (1984b). This translates into E_{def,asym}=1 MeV (folded Yukawa) and 450 keV (WS). A further increase in $E_{def,asym}$ can be achieved by allowing for higher-multipolarity deformations in calculations based on the SC method (Chasman, 1986; Sobiczewski et al., 1988; Rozmej et al., 1988; Cwiok and Nazarewicz, 1989a).

B. Alpha-decay properties

The study of alpha decay and the nature of nuclear states populated by this mechanism was important historically for the identification of low-lying negativeparity states and their subsequent interpretation in terms of octupole modes. There is little evidence that alpha decay in itself is enhanced by the presence of octupole correlations, as there is no observed correlation between alpha reduced widths and N,Z values corresponding to high octupole collectivity; see Fig. 28, taken from Toth et al. (1986). The figure shows a plot against neutron number of the reduced widths δ^2 (Rasmussen, 1959) for ground-state to ground-state alpha transitions connecting even-even nuclei with $78 \le Z \le 100$. The discontinuity at N=126 is attributed to a shell-structure effect. Theoretical approaches to alpha decay in which the four nucleons in the parent nucleus that eventually constitute the alpha particle are described by a shell model find that the inclusion of octupole deformation increases the value of δ^2 by 30% (Insolia *et al.*, 1991). Delion *et al.* (1992) used the same technique to make a systematic study of alpha decay in heavy octupole nuclei using a restricted (β_2, β_3) deformation space.

The property of alpha decay which does appear to be important in the determination of octupole collectivity is the relative decay width to different states in the same nucleus. This is usually described by the hindrance factor f, which is a measurement relative to the decay width to the ground state of even-even nuclei. The systematics of this quantity for low-lying negative-parity states in the actinides has been presented by Leander and Sheline (1984), and more recently extended by

Nucleus	I range	$ D_0 \text{ (expt.)} (e \text{ fm})$	Reference
¹⁴² Ba	I = 1	0.115(3)	Mach, Nazarewicz et al. (1990)
	I=9	0.13(2)	Phillips et al. (1986); Mowbray et al. (1989)
			Zhu et al. (1995)
¹⁴⁴ Ba	I = 7	0.071(10)	Phillips et al. (1986); Zhu et al. (1995)
	I = 8 - 11	0.14(3)	Phillips et al. (1986); Zhu et al. (1995)
¹⁴⁶ Ba	I = 1 - 3	0.06(4)	Mach, Nazarewicz et al. (1990)
	I = 5 - 7	0.009(3)	Phillips et al. (1986); Zhu et al. (1995)
¹⁴⁴ Ce	I = 7	0.17(5)	Mowbray et al. (1989)
¹⁴⁶ Ce	I = 7	0.11(2)	Phillips et al. (1988)
	I = 8 - 11	0.20(2)	Phillips et al. (1988)
¹⁴⁶ Nd	I = 1	0.14(5)	Zilges et al. (1992)
	I = 7 - 11	0.17(2)	Urban et al. (1988); Urban et al. (1991)
¹⁴⁸ Nd	I = 1	0.24(6)	Pitz et al. (1990)
	I = 1 - 4	0.13(3)	Ibbotson et al. (1993)
	I = 5 - 8	0.24(3)	Ibbotson et al. (1991, 1993)
	I = 6 - 8	0.23(3)	Urban et al. (1988)
¹⁵⁰ Nd	I = 1	0.26(5)	Pitz et al. (1990)
¹⁴⁸ Sm	I = 1	0.12(2)	Metzger et al. (1965, 1976)
	I=3	0.18(4)	Jungclaus, Börner et al. (1993)
	I < 7	0.13(1)	Urban et al. (1991)
	I > 7	0.22(2)	Urban et al. (1991)
^{150}Sm	I = 1	0.202(9)	Pitz et al. (1990)
	I = 1	0.118(5)	Jungclaus, Börner et al. (1993)
	I=3	0.185(5)	Jungclaus, Börner et al. (1993)
	I = 7 - 15	0.19(3)	Urban et al. (1987)
¹⁵² Sm	I = 1	0.24(3)	Jungclaus, Börner et al. (1993)
	I = 1	0.37(3)	Jungclaus, Belgya et al. (1993)
	I = 1	0.313(9)	Metzger (1976)
	I=3	0.34(3)	Jungclaus, Börner et al. (1993)
	I=3	0.38(7)	Jungclaus, Belgya et al. (1993)
¹⁵⁰ Gd	I=4-6	0.08(2)	Haenni and Sugihara (1977)

TABLE III. Experimental intrinsic E1 moments D_0 for even-even nuclei from the Ba-Sm region.

Poynter *et al.* (1989b); this is shown in Fig. 29. As can be seen, the values of f for decays to the 1^- level decrease as N becomes smaller, in fact becoming close to unity for Ra and Rn nuclei with $N \leq 136$. This has been interpreted (Leander and Sheline, 1984) as resulting from a change of the structure of the low-lying negative-parity states from dynamic vibration to static octupole deformation.

The systematic study of alpha transitions to excited states in odd-A nuclei was also carried out by Leander and Sheline (1984). Their studies showed that, in some cases, states having opposite parity must possess similar structures, since there are small hindrance factors to both states. This is illustrated for the case of the oddmass Ac isotopes in Fig. 30. These systematics have been extended by Sheline and Bossinga (1991; see also Sheline, 1993b), who pointed out that odd-A and odd-odd nuclei have lower hindrance factors than the corresponding even-even nuclei. The cases where alpha decay has been observed to such parity doublets (see Sec. V.D.1) with values of f < 100 are presented in Table V. Unfavored transitions typically have values of f > 100.

C. Exotic decay

The question of the possible existence of cluster structure in ground states of heavy nuclei has attracted much attention, especially because of the observed exotic decay branches via ¹⁴C, ²⁴Ne, ²⁸Mg, and others (Rose and Jones, 1984; Price, 1989). Several authors have drawn attention to the similarity in properties between alpha decay and exotic decay (Poenaru *et al.* 1984, 1985; Barwick *et al.*, 1986), although Shi and Swiatecki (1987) have pointed out that exotic decays of heavy nuclei

Nucleus	I range	$ D_0 \text{ (expt.)} (e \text{ fm})$	Reference
²¹⁸ Ra	<i>I</i> =6	0.23(5)	Gai et al. (1988)
	I = 7 - 11	0.339(17)	Gai et al. (1988)
²²⁰ Ra	I = 7 - 17	0.27(7)	Burrows et al. (1984); Cottle et al. (1984)
²²² Ra	I=3	0.38(6)	Ruchowska et al. (1992)
²²⁴ Ra	I = 3 - 5	0.028(4)	Poynter et al. (1989a); Marten-Tölle (1990)
	I = 7 - 9	< 0.11	Poynter et al. (1989a)
²²⁶ Ra	I = 1 - 5	0.06-0.10	Wollersheim et al. (1993)
	I = 7 - 12	0.12-0.21	Wollersheim et al. (1993)
	I = 7 - 11	0.16(1)	Ackermann et al. (1993)
²²⁸ Ra	I=3	0.011(1)	Ruchowska et al. (1982)
²²⁰ Th	I = 6 - 11	0.25(3)	Bonin et al. (1985)
²²² Th	I = 6 - 15	0.38(7)	Ward et al. (1983); Bonin et al. (1983)
²²⁴ Th	I = 11 - 17	0.52(2)	Ackermann et al. (1993)
²²⁶ Th	I = 9 - 19	0.30(1)	Ackermann et al. (1993)
²²⁸ Th	I = 9 - 13	0.120(3)	Ackermann et al. (1993)
²³⁰ Th	I=7-15	0.04(1)	Lauterbach et al. (1984)
²³⁰ U	<i>I</i> =11–13	0.16(5)	Ackermann et al. (1993)





TABLE IV. Experimental intrinsic E1 moments D_0 for even-even nuclei from the Ra-Th region.



FIG. 23. D_0/Q_0 plotted as a function of spin *I* for ²²⁰Ra (Smith *et al.*, 1995), ²²⁶Ra (Wollersheim *et al.*, 1993; Ackermann *et al.*, 1993), ²²²Th (Smith *et al.*, 1995), ^{224,226}Th (Ackermann *et al.*, 1993).

would occur with about the same frequency whether the parent nucleus were deformed or not.

A connection (Hussonnois et al., 1990a, 1990b, 1991; Sheline and Ragnarsson, 1991a, 1991b; Dumitrescu, 1994) has been made between the ground-state structure of reflection-asymmetric ²²³Ra and the observed (Brillard et al., 1989; Hourani et al., 1991) hindrance factors of ¹⁴C decay to excited states in ²⁰⁹Pb [although recent high-resolution studies by Hourani et al. (1995) have failed to confirm the previous observation of large hindrance factors for decay to both negative- and positiveparity states]. This tends to support the suggestion that there is a link between cluster preformation and strong octupole collectivity (Herrmann et al., 1986; Depta et al., 1986; Poenaru et al., 1994), although substantially more data are required to establish this. (For review, see Sandulescu and Greiner, 1992; Hussonnois and Ardisson, 1994.)

Delion *et al.* (1994) have performed microscopic calculations of heavy-cluster spontaneous emission using a *spherical* shell-model technique, applied previously to alpha decay (Delion *et al.*, 1992). They obtained very good agreement with experiment for ¹⁴C decay from ^{222,224,226}Ra, and made a prediction for ¹¹⁴Ba \rightarrow ¹²C+ ¹⁰²Sn decay, assuming spherical shapes. (For identification of ¹¹⁴Ba and its cluster radioactivity, see Guglielmetti *et al.*, 1995.)

Cseh *et al.* (1993) applied the vibron model coupled with the pseudo-SU₃ shell model (Sec. III.D) to describe clusterization in heavy nuclei. As an example they considered the clusterization of 224 Ra to the 210 Pb+ 14 C system.



FIG. 24. Plot of $|M(E3)|^2$, the E3 transition strength of $0_1^+ \rightarrow 3_1^-$ transitions for even-even nuclei, as a function of neutron number N (top) and proton number Z (bottom). The various symbols indicate the experimental procedures used to obtain the data (Coulomb excitation, lifetime measurements, inelastic electron scattering, deduction from β_3 values obtained from inelastic scattering of particles, and miscellaneous procedures). The peaks labeled A, B, C, D in the lower plot occur at Z values of the stable nuclei having N values corresponding to peaks A, B, C, D, respectively, in the upper figure (Spear and Catford, 1990).

D. Spectroscopic properties of the odd particle

The presence of intrinsic reflection-asymmetric deformations influences spectroscopic properties (parity splittings, ground-state spins and magnetic moments, Coriolis matrix elements, spectroscopic factors, radii, electromagnetic transitions, alpha-decay rates) of odd-*A* and odd-odd nuclei.

Leander and Sheline (1984) reviewed the properties of odd-A actinide nuclei, and outlined spectroscopic fingerprints of stable octupole deformation. For the spectroscopy of the odd-A Ra isotopes, see also the reviews by Sheline and Sood (1991) and Sheline (1993a).



FIG. 25. Values of $B(E3;0^+_1 \rightarrow 3^-_1)$ for nuclei with N=82, 84, 86, 88, and 90. The data are taken from Spear (1989) except for ¹⁴²Ce, ¹⁴⁴Nd (taken from Spear *et al.*, 1989), ¹⁴⁶Nd (Sandor *et al.*, 1993b), ¹⁴⁸Nd (Ibbotson *et al.*, 1993), and ¹⁵⁰Nd (Sandor *et al.*, 1993a).

1. Parity doublets

It was first emphasized by Chasman (1980) that a signature of intrinsic reflection asymmetry in welldeformed odd-A nuclei would be the appearance of parity doubling, i.e., for each bandhead there should be another bandhead close in energy with the same value of K and opposite parity.

In the reflection-asymmetric particle-plus-rotor of Leander and Sheline (1984), the parity splitting between members of a parity doublet in odd-A or odd-odd nuclei is [see Eq. (43)]

$$\Delta E = E(0^{-}) \langle \mathcal{P}_{s,p} \rangle; \tag{77}$$

i.e., it is always reduced compared to the value for the even-even core [see Eqs. (24) and (25); also Brink *et al.* (1987)].

All models of odd-A nuclei predict that the parity splitting of the K^{\pm} bands should be smaller than for the



FIG. 26. Plot of E2 and E3 matrix elements versus spin I deduced from Coulomb-excitation measurements for ¹⁴⁸Nd (Ibbotson *et al.*, 1993) and ²²⁶Ra (Wollersheim *et al.*, 1993). Solid lines join points calculated assuming a constant electric quadrupole or octupole moment.





FIG. 27. Top: Experimental mass minus calculated mass, with only the mass-symmetric shape coordinates ε_2 and ε_4 included in the calculations. Dashed contours indicate regions where there are no experimental data. Bottom: Experimental mass minus calculated mass, with both mass-symmetric (ε_2 and ε_4) and mass-asymmetric (ε_3) coordinates included in the calculations (Leander *et al.*, 1982).



FIG. 28. Reduced widths for *s*-wave α transitions plotted as a function of neutron number for nuclei with $78 \le Z \le 100$. Values enclosed in parentheses for Z=78 and Z=80 are based on estimated α branches (Toth *et al.*, 1986).



FIG. 29. Systematics of α hindrance factors f to the 1⁻ (full curve, round symbols), and the 3⁻ (dotted curves, square symbols), levels in the nuclei shown. N is the neutron number. Points in brackets are α hindrance factors to levels with unconfirmed I^{π} . The lines through the data points are to guide the eye only (Poynter *et al.*, 1989b).

 $K^{\pi}=0^+$ and (extrapolated) 0^- bandheads in the adjacent even-even nuclei, so that the experimental observation of parity doublets does not provide a particularly good test for these models. Also, as pointed out by Leander and Sheline (1984), closely spaced parity doublets will occur among the Nilsson orbitals in the deformedshell model for reflection-symmetric shapes, especially if pairing reduces the quasiparticle energy spacing. Nevertheless, the observation of a multiple band structure in ²²³Th (Dahlinger et al., 1988) in particular offers evidence for reflection asymmetry in an odd-mass nucleus. In this nucleus (see Fig. 31) strong E1 and M1 transitions are seen connecting the bands of different parity and signature. The energy degeneracy of these bands is highlighted in Fig. 32, where \mathcal{J}/\hbar^2 is plotted versus $\hbar\omega$ for ²²³Th and compared with adjacent even-even nuclei. Similar band structure has been observed in ²²⁵Th (Hughes et al., 1990), which is also compared with adjacent even-even nuclei in Fig. 32, in ²²¹Ra (Ackermann et al., 1989; Fernández-Niello et al., 1991) in ¹⁵¹Pm (Vermeer et al., 1990; Urban et al., 1990), and in ¹⁵³Eu (Pearson et al., 1994) (but see next subsection).

At low spin, numerous cases of parity doubling have been observed, assigned as such usually on the basis of alpha hindrance factors (see Sec. V.B), although sometimes from the nature of observed electromagnetic transitions or other spectroscopic properties. The observed features of these bands are contained within Tables V and VI, which list the locations of low-lying singleparticle states in actinide and lanthanide nuclei, respectively, together with their important experimental properties.

The K=0 parity-doublet band in the odd-odd nucleus



FIG. 30. Experimentally observed favored α decay into closely spaced parity doublets of odd-A Ac isotopes. The observed hindrance factors to the right of the levels are normalized to unity for the "natural" favored transition (Leander and Sheline, 1984). The level energies are in keV.

²²⁴Ac has been observed to have extremely small splittings between states of the same spin and opposite parity (Sheline *et al.*, 1991, 1993). In this nucleus the Newby splitting, which displaces the even- and odd-spin members in K=0 bands, in odd-odd nuclei, has been determined for each of the $K=0^+$ and $K=0^-$ bands, and shown to be consistent (see Table V) with the expectation for reflection-asymmetric systems of the same absolute value but opposite sign. Recently the *K* assignments of this band have been called into question (Ahmad *et al.*, 1994). For the nucleus ²²⁹Pa, on the edge of the region of octupole deformation, recent experiments using particle, γ , and e^- spectroscopy (Grafen *et al.*, 1991; Lösch *et al.*, 1994; Levon *et al.*, 1994) have shown that the original assignment of a $5/2^{\pm}$ parity doublet separated by 220 eV (Ahmad *et al.*, 1982) has probably no experimental basis. Sheline (1993b) has pointed out that the alpha decay of ²²⁹Pa suggests that it has octupole deformation in its ground state, and that there must exist a $5/2^{\pm}$ parity doublet in this nucleus.

2. Ground-state spins and magnetic moments

Indirect evidence for the occurrence of intrinsic reflection asymmetry comes from the spectroscopic properties of the odd particle, e.g., spin, magnetic moment, decoupling parameter, and transfer spectroscopic factor. The single-particle orbitals as a function of β_2 and β_3 are given for protons with Z>82 and neutrons with N>126 in Fig. 33, taken from Leander and Chen (1988). The energies of orbitals lying close to each other which are mixed by the octupole interaction (e.g., $g_{9/2,\Omega=5/2}$, $j_{15/2,\Omega=5/2}$ for neutrons) are strongly perturbed by octupole deformation, so that the ground-state spin and magnetic moments are expected to be different for $\beta_3=0$ and for $\beta_3 \neq 0$.

The strong-coupling expression for the rotational gyromagnetic factor g_K (K > 1/2),

$$g_K = \frac{1}{K} [Kg_\ell + (g_s - g_\ell) \langle \chi_K^\nu | s_z | \chi_K^\nu \rangle], \tag{78}$$

shows that the octupole coupling involving spin-flipped orbitals may lead to the hybridization of magnetic dipole properties. (For K=1/2 the magnetic decoupling parameter has to be considered.) Sheline and Leander (1983) applied this expression to describe the magnetic moments in the $3/2^{\pm}$ doublets of ²²⁷Ac, and Ragnarsson (1983) has applied it to the magnetic moments in ^{225,227}Ra. Sheline (1986) has pointed out that in the case of ²²³Ra, the similarity in magnetic moments of the ground state and the $3/2^-$ bandhead at 50.2 keV (see Table V) can be accounted for naturally by assuming both are part of a parity doublet. Figure 34 shows the values of $\langle \chi_K^{\nu} | s_z | \chi_K^{\nu} \rangle$ in odd-A isotopes of Ac and Pa, as calculated by Leander and Sheline (1984). The value of this quantity for 227 Ac, extracted from the measured g factor by means of Eq. (78), agrees well with the prediction of reflection-asymmetric theory.

As yet, however, the data are rather sparse for g factors measured for the many transitions observed in parity-doublet bands populated at high spin by heavyion reactions. In the cases of ¹⁵¹Pm (Vermeer et al., 1990) and ¹⁵³Eu (Pearson *et al.*, 1994), the intrinsic g factors obtained from the B(M1)/B(E2) branching ratios are systematically quite different for the transitions between the positive-parity members and between the negative-parity members, which suggests that these bands have different intrinsic structure. Indeed, calculations of magnetic moments in these nuclei using the rotor-plus-particle model are able to reproduce the experimental values without invoking static intrinsic reflection asymmetry (Afanasjev and Ragnarsson, 1995; see also Afanasjev, 1993, and Afanasjev and Mizutori, 1995). More recently, it has been observed that the magnetic properties of transitions for each parity of a band

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TABLE V. Properties of odd-mass and odd-odd nuclei in the Fr-Th region. The energies and K^{π} values of bandheads are given, together with values of the decoupling parameter *a* for K=1/2 bands, α -decay hindrance factor (h.f.) for favored transitions, and intrinsic electric dipole moment D_0 for transitions within parity-doublet bands.

Nucleus	K^{π}	Energy (keV)	а	h.f.	D_0 (e fm)	Other
²¹⁹ Fr (Liang et	t al., 1991)					
	$\frac{1}{2}$ –	81	7.03			
	$\frac{1}{2}$ +	≈330	-7.78			
	$\frac{5}{2}$ +	384.3		10		
	$\frac{2}{5}$ -	490.3		2.9		
	$\frac{2}{3}$ -	56.15		$9.3(\frac{7}{2})$		
	$\frac{2}{3}$ +	369.6		$>19(\frac{7}{2}^+)$		
²²¹ Fr (Sheline,	1988b: Liang.	Péghaire <i>et al.</i> , 1990)		(2)		
(,	$\frac{1}{2}$ –	26.0	4.3	$11(\frac{7}{2})$	0.102 ± 0.008	
	$\frac{1}{2}$ +	145.8	-2.6	$65\left(\frac{7}{2}^{+}\right)$	0.102 = 0.000	
	$\frac{2}{3}$ -	36.6	2.0	17	0.076 ± 0.003	
	$\frac{2}{3}$ +	224.6		28	0.070 - 0.000	
²²³ Fr (Kurcewi	icz et al 1992	Sheline <i>et al</i> 1995)		20		
II (Ituleewi	$\frac{3}{2}$ -	0		7.6	0.24 ± 0.04	
	$\frac{2}{3}$ +	160 45		42	0.21=0.01	
	$\frac{2}{1}$ -	55.0	-136	12		
	$\frac{1}{1}$ +	149 3	0.96			
221 R a (Ackern	$\frac{2}{100}$	R0. Liang Paris <i>et al</i> 10	0.20 000 Fernández	Niello <i>et al</i> 100	1)	
Ra (Ackelli	$\frac{5}{5}$ +		790, Permandez	-ivieno ei ui., 177.	0.36 ± 0.10	
	$\frac{2}{5}$ –	103.4			0.30 ± 0.10	
	$\frac{2}{3}$ +	221.2		1.0		
	$\frac{2}{3}$ -	321.3 485 3		1.9 ~15		
223Do (Shalino	2 10%, Shaling	403.3 Chan and Laandar 1	000. Driancon	≈ 13	1 Hodi at $al = 1002$	
Ra (Shenne	, 1960, Sheime $_{3+}$		966; Briançon	<i>ei al.</i> , 1990, Abdu	0.124 ± 0.010	-0.29 ± 0.014
	$\frac{\overline{2}}{3}$ –	0 50 10			0.124 ± 0.010	$\mu = 0.28 \pm 0.014$
	$\overline{2}_{5+}$	30.19			0.042 + 0.012	$\mu = 0.42 \pm 0.06$
	$\frac{1}{2}$	254.92			0.043 ± 0.012	
	$\overline{2}_{1+}$	369.43	1.25	= (1+)	0.070 + 0.012	
	$\overline{2}$	280.10	1.35	$5.0(\frac{1}{2})$	0.078 ± 0.012	
	1 –	250 50	2.15	$16(\frac{1}{2})$		
	$\overline{2}$	330.30	-2.15	$30(\frac{1}{2})$		
2250 (01 1)	1 1002 1	$000 \text{ n}^{-1} + 1.400 \text{ c}^{-1}$		$14(\frac{1}{2})$	1 1000)	
Ra (Shelline	et al., 1983, 1^{+}_{1+}	989; Reich <i>et al.</i> , 1980;	Heimer <i>et al.</i> , 1.90	1987; Andersen <i>et</i>	al., 1989)	
	$\frac{\overline{2}}{1}$	0	1.89		0.14 ± 0.02	
	$\frac{1}{2}$ +	55.2 1.40.0	-2.56			
	$\frac{\overline{2}}{3}$ –	149.9				
	$\frac{1}{2}$ +	225.1		1 57		
	$\overline{2}$	230.3		1.57		
227 D - (D	$\left(\frac{1}{2}\right)$	394.2	1004)	23		
Ra (Borge e	$\frac{1987}{3+}$	enne, 19890; Mach et al	., 1994)		0.000 + 0.011	
	$\frac{\overline{2}}{3}$ –	0			0.098±0.011	
	$\frac{1}{2}$	90.0	1 71			
	$\overline{2}$ 1 -	120.7	-1./1			
223 . (2	290.0	0.62			
Ac (Ahmad	<i>et al.</i> , 1989; S $_{3-}^{3-}$	heline <i>et al.</i> , 1990)		(5-)		
	$\frac{1}{2}$	4.1		$5.1(\frac{1}{2})$		
	$\frac{\overline{2}}{5}$ –	88.9		$\approx 14(\frac{5}{2})$	> 0.10	
	2 5 +	0		2.5	>0.18	
225 . (2	64.6		7.0		
Ac (Ahmad	et al., 1984, 1	987)			0.454 - 0.014	
	$\frac{5}{2}$ 3 +	0			0.171 ± 0.014	
	ž –	40.1				
	$\frac{1}{2}$	120.8		7.5		
	$\frac{5}{2}$	155.7		1.8		

Nucleus	K^{π}	Energy (keV)	а	h.f.	D_0 (e fm)	Other
²²⁷ Ac (Sheli	ne and Leande	er, 1983; Ishii <i>et al</i> ., 198	5; Martz <i>et al.</i> ,	1988)		
	$\frac{3}{2}$ -	0			0.0297 ± 0.0001	$g_K = 0.92 \pm 0.06$
	$\frac{3}{2}$ +	27.4				$g_K = 0.96 \pm 0.10$
	$(\frac{5}{2}^{-})$	273.1				
	$(\frac{5}{2}^+)$	304.6				
	$\frac{1}{2}$ -	354.6	-2.01	$13(\frac{1}{2})$		
				$5.2(\frac{7}{2})$		
	$\frac{1}{2}$ +	435.4	4.56	$>35(\frac{1}{2}^+)$		
				$83(\frac{7}{2}^+)$		
²²⁴ Ac (Sheli	ne et al., 1991,	1993; Ahmad et al. 199	94)			N 7
	(0^{-})	0				$E^{N} = 1.99 \text{ keV}$
	(0^+)	(22.0)				$E^{N} = -0.37 \text{ keV}$
	(1^{+})	66.0				
	(1^{-})	89.3				
	3-	353.9		20		
222	3+	360.2		12		
²²⁵ Th (Dahli	inger <i>et al.</i> , 198	88)				
	$\frac{5}{2}$	0			0.44 ± 0.09	
225-551 (7.7.1	2					
²²⁵ Th (Hugh	thes et al., 1990) $\frac{3}{3}$)			0.40 + 0.40	
	2 3 _	0			0.40 ± 0.10	
227 TL (T	$\overline{2}$	005)				
²² , In (Liang	g et al., 1994, 1	995)	2.07			
	$\frac{1}{2}$	0 67.2	2.97			
	$\frac{2}{3}$ +	07.2	-2.08	44 ± 5		
	$\frac{\overline{2}}{3}$ –	24.5		44 ± 3 ≈ 440		
	$\frac{2}{5}$ –	142.0		~440		
	$\frac{2}{5}$ +	547.0				
²²⁹ Pa (Levor	n et al 1004	517.0				
Ta (Level	$\frac{1}{2}$	151	-1 71		0.09 ± 0.04	
	$\frac{2}{1}$ +	10.1	1.51		0.07 - 0.01	
	$\frac{1}{2}$ -	1540	-1.75			
	$\frac{2}{1}$ +	1593	1.90			
	L					

TABLE V. (Continued).

in ¹⁴⁷Pm are similar at intermediate spin values. This has been interpreted as being consistent with this nucleus having stable octupole deformation in this spin region (Urban *et al.*, 1995).

Comparison with theory of the ground-state properties for actinide nuclei has been carried out by Ragnarsson (1983), Leander and Sheline (1984), Leander and Chen (1988), Ćwiok and Nazarewicz (1991), and Jain et al. (1990). The assignment of ground-state orbitals to these nuclei is given in Fig. 33. In general, the agreement is very good between the experimental values of $I_{g.s.}$ and $\mu_{g.s.}$ and the values calculated assuming $\beta_3 \approx 0.1$ for Rn, Fr, Ra, Ac, and Pa isotopes with $N \le 140$; but if $\beta_3 = 0$, it is difficult in some cases to find appropriate orbitals so that both $I_{g.s.}$ and $\mu_{g.s.}$ can be matched by calculations. Particularly interesting is the nucleus ²²⁵Ra, which is expected to exhibit coexistence of reflection-symmetric and reflection-asymmetric configurations (Sheline et al., 1989; Ćwiok and Nazarewicz, 1991). In this case it is difficult to explain the ground-state spin (I=1/2) (Ahmad *et al.*, 1983) without invoking octupole mixing [see, however, Piepenbring (1984)].

In the light-lanthanide nuclei, similar conclusions were reached (Leander *et al.*, 1985; Ćwiok and Nazarewicz, 1989b) in describing the ground-state spins and electromagnetic moments of ^{143,145}Cs (Coc *et al.*, 1987) and ¹⁴⁵Ba (Mueller *et al.*, 1983). The structure of parity doublets in ¹⁵¹Pm and ^{153,155}Eu was studied by Nosek *et al.* (1993a) in the quasiparticle-phonon nuclear model (Sec. III.F). They found strong octupole correlations only in the *K*=1/2 states. Spectroscopic properties of odd-*A* nuclei around ¹¹²Ba have been discussed by Heenen *et al.* (1994).

For the odd-odd nuclei ^{228,230}Pa, Herrmann *et al.* (1989) have noted that the measured ground-state magnetic moments are best fitted by including orbitals which are associated with large driving forces toward octupole shapes. Ekström *et al.* (1986) made a similar analysis of the ground-state moments of ^{223,225,227}Fr, but was unable to draw any firm conclusion on the presence of octupole deformation, as the data were also reproduced by


FIG. 31. Level scheme of ²²³Th. The spin assignment was made under the assumption of a ground state spin $I^{\pi} = 5/2^+$. The width of the arrows represents the transition intensity. Dotted transitions are unconfirmed (Dahlinger *et al.*, 1988). The level and transition energies are in keV.

particle-rotor calculations which did not include octupole deformation. For the ground-state spins of odd-odd Francium isotopes with A = 220-228, and for ^{224,226}Ac, Sheline, Chen, and Leander (1988) were able to explain the values only in terms of parity-mixed orbits.

3. Coriolis matrix elements

In the strong-coupling limit, the diagonal Coriolis matrix element for K=1/2 bands is written in terms of the decoupling parameter *a* (Zaikin, 1966; Bohr and Mottelson, 1975):

$$a = -p \langle \mathscr{P}_{s,p} \chi_{1/2}^{\nu} | j_+ | R_1 \chi_{1/2}^{\nu} \rangle, \tag{79}$$

so that opposite-parity states of a doublet having a common intrinsic structure have values of the decoupling parameter a of equal magnitude but opposite sign. That is,

$$ap = \text{const}$$
 (80)

within a doublet.

Ragnarsson (1983), and Leander and collaborators (Leander and Sheline, 1984; Leander and Chen, 1988), have made a systematic comparison of the experimental values of a with theoretical calculations using a model

which is either reflection asymmetric or symmetric. Particularly for ²²⁷Ra, reflection asymmetry appears necessary to account for the observed signs of a for the lowest $K^p = 1/2^{\pm}$ band. Figure 35 shows the comparison between experimental values and values calculated by Leander and Chen (1988) from a reflection-asymmetric particle-plus-rotor model. The theory predicts some degree of divergence from the strong-coupling limit (80), but not as much as observed, although overall the agreement between theory and experiment is quite good. For K>1/2 bands, Leander and Chen (1988) calculated moment-of-inertia and signature splitting for both parity bands, and obtained parity decoupling of the same order as that observed experimentally, although the quality of the fit is rather poor. They speculated that the observed smaller size of the Coriolis matrix elements in their calculations for deformed nuclei is brought about by the presence of octupole deformation. For ²¹⁹Fr, Liang et al. (1991) were able to reproduce the experimental values of the decoupling parameter for the K=1/2 doublet using the intermediate-coupling scheme based on the quasiparticle-phonon nuclear model (Sec. III.F), which assigns quite different quasiparticle character to the bands of different parity.

4. Spectroscopic factors

An early attempt to reproduce the measured nuclear structure factors for states populated by one-neutron transfer reactions using an octupole model was made by Løvhøiden et al. (1986), who studied the 226 Ra (t,α) 225 Ra reaction. They were able to reproduce the measured quantities for states associated with the $i_{15/2}$ intruder orbital using matrix elements calculated assuming nonzero β_3 . More extensive calculations have been carried out for ²²⁷Ac, ²²⁵Ra, and ²²⁷Ra by Leander and Chen (1988), giving reasonable agreement with experimental data for *relative* magnitudes of the nuclear structure factors, using the reflection-asymmetric particle-plus-rotor model. In contrast, Martz et al. (1988) reported that the spectroscopic strengths in ²²⁷Ac populated by the 226 Ra(α,t) and 226 Ra(3 He,d) reactions are in better agreement with models which assume reflection symmetry. Tables of spherical amplitudes C_I , useful for interpretation of one-particle transfer reactions, have been computed by Chasman (1984) for reflectionasymmetric valence-proton and -neutron single-particle states in the $A \sim 225$ mass region.

5. Radii

The introduction of laser-spectroscopic techniques has enabled many new isotope-shift measurements to be made in the actinide and lanthanide region (Aufmuth *et al.*, 1987). The isotope shift is used to derive the change $\delta \langle r^2 \rangle^{N,N'}$ in the mean-square nuclear-charge radii between nuclei having neutron numbers N and N', which is related to changes in the charge distribution between these nuclei by

$$\delta\langle r^2\rangle = \delta\langle r^2\rangle_{\rm sph} + \frac{5}{4} \pi\langle r^2\rangle_{\rm sph} (\delta\langle \beta_2^2\rangle + \delta\langle \beta_3^2\rangle + \cdots).$$
(81)



FIG. 32. Kinematic moment of inertia \mathcal{J}/\hbar^2 (in MeV⁻¹) as a function of rotational frequency $\hbar\omega$ (in MeV) for ground-state rotational bands in ²²³Th and ²²⁵Th, compared with their even-even neighbors. The groundstate spins of ^{223,225}Th are assumed to be 5/2 and 3/2 respectively. The positive (negative) parity bands are indicated by filled (open) symbols.

Figure 36, taken from Otten (1989), shows how the experimental value of $\delta \langle r^2 \rangle^{A-1,A}$ between neighboring isotopes A-1 and A varies for isotopes of Rn (Borchers et al., 1987), Fr (Coc et al., 1985, 1987), and Ra (Ahmad et al., 1983, 1985, 1988), as a function of N. This quantity shows a pronounced odd-even effect between N=132and N=138, in that $\delta \langle r^2 \rangle^{A-1,A}$ for odd A is larger than $\delta \langle r^2 \rangle^{A-1,A}$ for even A, which is opposite to what is observed all over the chart of the nuclides with very few exceptions. The Th isotopes with $N \ge 137$ do not exhibit this effect (Kälber et al., 1989). The inversion of the normal odd-even staggering in the light actinides has been associated with the occurrence of strong octupole correlations (Ahmad et al., 1984), with the suggestion that it arises from the stabilizing effect on the octupole deformation of the odd particle outside the even-even core (see also Sheline, Jain et al., 1988, 1989). It should be pointed out, however, that the experimental values for the Rn, Fr, and Ra isotopes can be well reproduced using an extended Thomas-Fermi method, even though the octupole degree of freedom is not included in the calculation (Buchinger et al., 1994).

For Ba and Cs isotopes around N=88-90, the effect is masked by the large overall increase observed with the onset of quadrupole deformation, although Sheline, Jain *et al.* (1988) have suggested that the odd-even effect is attenuated for these nuclei. For the Europium isotopes there is some evidence of an inversion of $\delta \langle r^2 \rangle^{N,N-1}$ at N=89 (Dörschel *et al.*, 1984; Ahmad *et al.*, 1985; Alkhazov *et al.*, 1990), which may be associated with octupole deformation (Afanasjev, 1993).

6. E1 transitions

Experimentally, the first evidence that the presence of octupole correlations would give rise to relatively strong E1 transitions, of order 10^{-3} to 10^{-2} s.p.u. (see Sec. II.B), came from studies of low-lying transitions in odd nuclei (for example, Ahmad *et al.*, 1984). In the strong-coupling limit, the values of the electric dipole moment D_0 for odd-mass nuclei are expected to be the same as those for the even-even core. Butler and Nazarewicz (1991) compared the experimental values for even-even and odd-A nuclei with theoretical predictions of the SC

Nucleus	K^{π}	Energy (keV)	а	D_0 (e fm)	Other	
¹⁵¹ Pm (Sood and Sheline, 1989; Urban et al., 1990; Vermeer et al., 1990)						
	$\frac{5}{2}$ +	0		0.16 ± 0.04	$\mu = 1.29 \pm 0.03$	
	$\frac{5}{2}$ -	116.8			$\mu = 2.20 \pm 0.14$	
	$\frac{3}{2}$ +	255.6				
	$\frac{3}{2}$ -	540.2				
¹⁵² Eu (Sheline and Sood, 1989)						
	4+	89.9		0.076		
	4 -	141.8				
	5+	108.1		0.034		
	5-	180.6				
¹⁵³ Eu (Shel	ine and	Sood, 1990; Pearso	n <i>et al</i> ., 199	94)		
	$\frac{5}{2}$ +	0		0.08 ± 0.02	$\mu = 1.5330 \pm 0.0008$	
	$\frac{5}{2}$ -	97.4			$\mu = 3.22 \pm 0.23$ or -0.52 ± 0.23	
¹⁵⁴ Eu (Shel	ine, 198	9a)				
	1^{+}	71.91		0.052 ± 0.008		
	1^{-}	82.82				
	3+	281.68		$0.081 \!\pm\! 0.015$		
	3-	239.29				
¹⁵⁵ Eu (Sheline and Sood, 1990)						
	$\frac{5}{2}$ +	0		0.22 ± 0.01	$\mu = 1.56 \pm 0.10$	
	$\frac{5}{2}$ -	104.3			$\mu = 2.49 \pm 0.27$	
	$\frac{1}{2}$ +	922.8	2.14			
	$\frac{1}{2}$ -	1106.7	-1.11			

TABLE VI. Similar to Table V, but for odd-mass and odd-odd nuclei in the Pm-Eu region.

method. In these calculations, the macroscopic contribution to D_0 should be similar to the even-even neighbors, although fluctuations can arise from the single-particle contribution. Leander and Chen (1988) have also carried out calculations for the B(E1) values, including effects of Coriolis mixing and odd-quasiparticle contributions.

For $K_i = K_f = 1/2$ transitions, Eq. (66) is modified by the presence of the signature-dependent term [Eq. (67)] proportional to D_1 . Reich *et al.* (1986) have determined the amplitudes D_0 and D_1 for the $1/2^{\pm}$ doublet in ²²⁵Ra.

VI. ROTATIONAL PROPERTIES OF REFLECTION-ASYMMETRIC NUCLEI

Rotational properties of octupole vibrational states were discussed in a rotor-plus-RPA model by Neergård and Vogel (1970a, 1970b) and Vogel (1976), and in the cranked RPA theory by Robledo *et al.* (1986), Mizutori *et al.* (1990, 1991a, 1991b), and Nakatsukasa *et al.* (1992, 1993, 1995) (see Sec. III.F). In this section we mainly concentrate on rotational motion of nuclei with stable reflection-asymmetric deformations.

Quasimolecular rotational bands in a reflectionasymmetric nucleus can be characterized by the "simplex" s, which is the eigenvalue of the S_1 operator [reflection through the (y,z) plane; see Sec. III.B.1]. Simplex has properties similar to those of the signature quantum number in the absence of reflection symmetry (Nazarewicz *et al.*, 1984a; Frauendorf and Pashkevich, 1984; Nazarewicz and Olanders, 1985a, 1985b).

The square of the S_1 operator is related to the total number of fermions:

$$S_1^2 = (-1)^A. (82)$$

The rotational band with simplex s is characterized by spin states I of alternating parity, (Bohr and Mottelson, 1975)

$$p = se^{-i\pi I}.$$
(83)

Thus for reflection-asymmetric systems with an even number of nucleons, one obtains

$$s = +1, I^p = 0^+, 1^-, 2^+, 3^-, \dots,$$
 (84)

$$s = -1, I^p = 0^-, 1^+, 2^-, 3^+, \dots,$$
 (85)

while for systems with odd particle number one has

$$s = +i, I^p = 1/2^+, 3/2^-, 5/2^+, 7/2^-, \dots,$$
 (86)

$$s = -i, I^p = 1/2^-, 3/2^+, 5/2^-, 7/2^+, \dots$$
 (87)

In the mirror-symmetric case the simplex becomes s=-pr (*r* is the signature quantum number).

The general structure of the cranked HFB equations and their solutions remains the same in the simplex formulation as in the usual signature-parity formulation. The simplex of the rotating vacuum is s=+1, and the simplex of an excited *n*-quasiparticle configuration becomes

$$s_{nqp} = s_1 s_2 \cdots s_n, \tag{88}$$

where s_i is the simplex of the *i*th particle.

High-spin properties of reflection-asymmetric nuclei



FIG. 33. Proton single-particle levels (top) and neutron levels (bottom) in a WS potential for β_2 =0–0.18 with β_3 =0, for β_3 =0–0.1 with β_2 =0.18, and for β_2 =0.18–0.10 with β_3 =0.1. Taken from Leander and Chen (1988).

were discussed by Nazarewicz and Olanders (1985a) in the cranking WS model. Octupole coupling between high-*j* unique-parity orbitals and normal-parity states leads to fragmentation of the aligned angular momentum over many quasiparticle states. The angular momentum content of the lowest Routhians, containing a significant component of high-*j* unique-parity states, decreases with octupole deformation. On the other hand, the average alignment of Routhians with a dominant component of normal parity increases. As a consequence, the quasiparticle-Routhian pattern becomes more uniform, and many quasiparticle Routhians have similar alignment.

The fragmentation effect is illustrated in Fig. 37, which displays the calculated quasiparticle Routhians for Z=56 and N=88 as functions of rotational frequency, without (left) and with (right) octupole deformation. As

can be seen in Fig. 37, the frequency of the first band crossing increases at the reflection-asymmetric shape as a consequence of the reduced angular momentum alignment of the lowest Routhians. In addition, the interaction between the crossing bands increases. Other diagrams of Routhians as functions of octupole deformation and rotational frequency can be found in Faber and Płoszajczak (1981), Frauendorf and Pashkevich (1984), and Åberg (1990).

At high frequencies a shape transition towards $\beta_3=0$ is expected after the alignment of the high-*j* quasiparticles. Figure 38 illustrates the influence of rotational frequency on the octupole-shell structure. The singleparticle Routhians for $N \sim 88$ are shown as functions of β_3 for $\hbar \omega = 0$, 0.3, and 0.6 MeV. At $\hbar \omega = 0.3$ MeV the octupole-shell effects are quenched, although the proton numbers Z=56 and Z=62 and neutron numbers N=88



FIG. 34. Values of $\langle s_z \rangle$ for the $3/2^{\pm}$ and $5/2^{\pm}$ parity doublets in odd-*A* isotopes of Ac and Pa. Predictions of reflectionsymmetric ($\varepsilon_3=0$) and reflection-asymmetric ($\varepsilon_3 \neq 0$) meanfield theory are indicated by dashed and solid lines, respectively (Leander and Sheline, 1984).

and N=94 are still octupole driving. However, at the higher rotational frequency $\hbar \omega = 0.6$ MeV, the octupole-shell structure is almost completely washed out.

The octupole deformation explains the gradual angular momentum alignment in the light actinides. Calculations based on the cranked reflection-asymmetric WS model reproduce the absence of sharp band crossing in nuclei around ²²²Th (Nazarewicz et al., 1984a; 1987). Figure 39 displays the alignment plot $I_x(\omega)$ for ²²²Th. Nazarewicz et al. (1987) have shown, using the particlenumber projected cranking WS model with pairing, that the octupole-deformed ground-state band ($\beta_2=0.116$, $\beta_3=0.104$) crosses a neutron-aligned band at $\hbar \omega \approx 0.20$ MeV, with a large interaction between the two bands. At the reflection-symmetric shape ($\beta_2=0.12$, $\beta_3=0$) the quasiparticle alignments are large and band interactions are small. The four-quasiparticle configurations $\nu(i_{15/2})^2 \pi(i_{13/2})^2$ become yrast at $I \approx 26$; that is, the shape transition from reflection asymmetric to reflection symmetric is expected. Experimentally, there is now evidence that the shape transition occurs at $I \approx 24$ (Smith



FIG. 35. The decoupling factor *a* times the total parity *p* for K=1/2 parity-doublet bands. Filled symbols represent measured data for the nuclei indicated on the plot; open symbols with connecting lines to the filled ones show the corresponding results of core-particle coupling calculations. The *x* axis gives the *ap* of the positive-parity band, and the *y* axis that of the negative-parity band. For strong coupling, the points would lie on the diagonal (dashed) (Leander and Chen, 1988).

et al., 1995; see also Schwartz *et al.*, 1987, 1988) while 220 Ra remains reflection asymmetric up to $I \approx 30$.

A detailed discussion of rotation-induced shape changes in reflection-asymmetric nuclei from the Ra-Th and Ba-Ce regions has been given by Nazarewicz (1987), Nazarewicz *et al.* (1987), and Nazarewicz and Tabor (1992). They employed the so-called total Routhian surface method, in which the nuclear mean field is param-



FIG. 36. Differential isotope shift $\delta \langle r^2 \rangle^{N-1,N}$ for Rn, Fr, and Ra isotopes, showing regular odd-even staggering below N=126, and inverted sense for neutron numbers N=133, 135, and 137 (Otten, 1989).



FIG. 37. Proton quasiparticle Routhians for Z=56 (top) and neutron quasiparticle Routhians for N=88 (bottom) as a function of rotational freqency. The deformation parameters used correspond to the yrast configurations of ¹⁴⁴Ba. The left panel, $\beta_3=0$, is representative of the reflection-symmetric configurations involving aligned neutron ($i_{13/2}$) and proton ($h_{11/2}$) pairs. The right panel, $\beta_3=0.1$, represents the structure of quasiparticle excitations associated with the reflection-asymmetric ground band. Levels are labeled by simplex, s=i (solid line) and s=-i (dashed line) (Nazarewicz and Tabor, 1992).

etrized by a WS single-particle potential and a BCS pair field. The energy E_{SC} of the nonrotating state, as a function of deformation $\hat{\beta} \equiv (\beta_2, \beta_3, \beta_4, ...)$, was obtained by the SC method. The total Routhian at frequency ω and deformation $\hat{\beta}$ was thus calculated as

$$E^{\omega}(\hat{\beta}) = E_{\rm SC}(\hat{\beta}) + [\langle H^{\omega}_{\hat{\beta}} \rangle - \langle H^{\omega=0}_{\hat{\beta}} \rangle].$$
(89)

The absolute minimum of the Routhian at fixed ω corresponds to the solution for a yrast state. Secondary minima correspond to other solutions, which may be yrast if they have higher angular momentum. Figure 40 shows the equilibrium deformations as a function of rotational frequency, calculated with this method for the yrast configuration in doubly even ^{220–228}Th.

Total Routhian surface calculations for nuclei predicted to have reflection-asymmetric ground states (such as ¹⁴⁴Xe, ^{144,146}Ba, ^{144,146}Ce, ^{222,224}Ra, and ^{222,224,226}Th) indicate that at medium spins the magnitude of octupole deformation increases, and the octupole minima are much better separated than in the ground state. Egido and Robledo (1990) added the rotational-energy term $I(I+1)/2 \mathscr{J}(q_3)$ to the microscopic collective Hamiltonian (valid at I=0) to investigate the behavior of ¹⁴⁶Ba at high spins. They obtained stabilization of octupole deformation at high spins.

In general, the enhancement of octupole strength with rotation (see, for example, the behavior of ²²⁰Th in Fig.

As discussed above, in the Ra and Th nuclei shape changes have been predicted to occur above I=24. In the Xe-Sm isotopes, the transition to reflection-symmetric shapes is expected to take place around I=12, which is much easier to reach experimentally; in ^{144,146}Ba, band crossings have been observed above I=12 (Zhu *et al.*, 1995).

The nuclei ^{218,220}Ra, ^{138,140}Xe, ^{140,142}Ba, ¹⁴²Ce, ¹⁴⁴Nd, and ^{146,148}Sm provide examples of strong quadrupoleoctupole coupling. Their ground-state minima are predicted to be very β_2 and β_3 soft, but they become reflection asymmetric at medium spins. This shape transition is also associated with an increase in quadrupole deformation. Experimentally, the N=86 isotones are transitional systems showing an interplay between collective and noncollective modes. In ¹⁴²Ba, ¹⁴⁶Nd, and ¹⁴⁸Sm, the structures on top of the 8^+ and 11^- yrast states have been interpreted (Urban et al., 1991; Zhu et al., 1995) in terms of noncollective multiparticle excitations. The N=84 isotones of Nd, Sm, and Gd do not exhibit rotational behavior, and their excitation spectrum can be interpreted in terms of the spherical shell model, including octupole phonons (Bargioni et al., 1995). In the Ra-Th region, spectacular examples of high-spin competition between noncollective excitations and octupole modes are the N=130 isotones ²¹⁸Ra (Schulz *et al.*, 1989) and ²¹⁹Ac (Cristancho et al., 1994). Both nuclei exhibit enhanced E1 transitions and alternating-parity sequences. But the quadrupole collectivity is weak, as shown by their irregular quasivibrational spectra. The alignment process along the yrast line of ²¹⁸Ra can be reproduced by the cranked WS calculations with pairing, assuming very small quadrupole deformation, $\beta_2=0.1$, and a large value of $\beta_3 \sim 0.09$ (Schulz *et al.*, 1989; see also Leandri and Piepenbring, 1993). The nuclei ²²⁶Ra, ²²⁸Th, ¹⁴⁶Xe, ¹⁴⁸Ba, ^{146,148}Nd,

The nuclei ²²⁶Ra, ²²⁸Th, ¹⁴⁶Xe, ¹⁴⁸Ba, ^{146,148}Nd, ¹⁵⁰Sm, and ¹⁵²Gd are calculated to have well-developed quadrupole ground-state deformations, but they are β_3 soft. At low frequencies, the negative-parity states in these nuclei can be described in terms of very collective octupole vibrations. At medium spins, however, the static theory predicts a shape transition towards $\beta_3 \neq 0$, or at least the presence of near-yrast reflection-asymmetric configurations. The experimental data on ¹⁴⁶Nd (Urban *et al.*, 1988), and ¹⁵⁰Sm (Urban *et al.*, 1987), suggest that the above scenario indeed takes place in these nuclei.

Rotational properties of nuclei around ¹¹²Ba were discussed by Heenen *et al.* (1994) in their total-Routhiansurface/WS calculations. They obtained a shape transition to $\beta_3=0$ at medium spins resulting from the alignment of $h_{11/2}$ neutrons and protons. For reflectionasymmetric cranked SC calculations for superdeformed nuclei, see Sec. IX.B.2.



FIG. 38. Single-particle WS levels for neutrons (top) and protons (bottom) plotted versus octupole deformation β_3 at fixed values of $\beta_2=0.2$ and $\beta_4=0.08$. At zero rotational frequency, $\hbar \omega = 0$, the single-particle levels are labeled by Ω . Intrinsic parity is indicated only at $\beta_3=0$ (for $\beta_3 \neq 0$ intrinsic parity is violated). At $\hbar \omega > 0$, the levels are labeled by means of simplex, s=i (solid line) and s=-i(dashed line).

Kvasil and Nazmitdinov (1985) considered the RPA using the cranking reflection-asymmetric Hamiltonian with a pairing interaction plus quadrupole-quadrupole and octupole-octupole interactions. They gave expressions for energies and E1, E2, and E3 transition moments of reflection-asymmetric nuclei. Another application of the reflection-asymmetric cranked SC method can be found in work by Faber (1981), who investigated the influence of angular momentum on the mass distribution of heavy-ion induced fission.

The angular momentum dependence of the parity splitting in the actinides was discussed by Jolos *et al.* (1993, 1994, 1995), who employed the collective Hamiltonian of Eq. (48) with the spin-dependent collective potential $V(\beta_3, I) = U(\beta_3) + A(\beta_3)I(I+1)$. The parity splitting was estimated using the WKB approximation, and the resulting phenomenological expression accurately describes the alternating-parity rotational bands in even-even actinide nuclei. Zamfir *et al.* (1994) demonstrated that there is a simple correlation between the critical angular momentum I_{oct} at which the parity splitting disappears and the energy ratio $E(3_1^-)/E(2_1^+)$.

Alonso *et al.* (1995) used the *spdf* SU₃ Hamiltonian with quadrupole and octupole interaction to describe the positive- and negative-parity yrast bands in ²²⁶Ra. They obtained a transition to reflection-asymmetric shapes at $I \sim 9$.

VII. INTRINSIC DIPOLE MOMENTS

In the presence of reflection-asymmetric deformations, a static electric dipole moment may arise in the intrinsic frame due to a shift between the center of charge and center of mass. Various theoretical treatments of this effect are discussed below.

A. Macroscopic models for the E1 moment

In the purely geometric picture, isoscalar dipole deformations are not independent degrees of freedom, but are determined from the c.m. condition, Eq. (3).

The collective electric dipole moment \mathcal{D}_{μ} appears in the second order as a product of quadrupole (λ =2) and octupole (λ =3) moments, namely,

$$\mathscr{D}_{\mu} = [\mathscr{Q}_2 \times \mathscr{O}_3]_{1\mu}. \tag{90}$$

Thus one may expect that low-energy E1 collectivity should be pronounced in nuclei with strong quadrupole and octupole correlations. If higher moments ($\lambda = 4,5,...$) are present, they also contribute to D_{μ} through the $\lambda \leftrightarrow \lambda + 1$ coupling (see below).

In the macroscopic liquid-drop model, dipole polarization results from asymmetry of the internal electric field caused by reflection-asymmetric shape deformations. Assuming an axially symmetric system, the induced dipole moment is proportional to $\beta_2\beta_3$ [see Eq.



FIG. 39. Aligned angular momentum I_x for ²²²Th, plotted as a function of rotational frequency. The calculated yrast line is plotted as a dashed line. The octupole-deformed band (thick solid line) remains yrast up to $I \sim 24$. Calculations predict, however, a backbending at higher spins, caused by crossing with a four-quasiparticle reflection-symmetric band (β_3 =0, thin solid line). The experimental values are also plotted (Nazarewicz *et al.*, 1987).

(90)]. Bohr and Mottelson (1957, 1958) and Strutinsky (1956) calculated the *E*1 moment in the liquid-drop model (Strutinsky's derivation is the technically correct one). Expressing the liquid-drop energy in terms of nucleonic densities ρ_p and ρ_n , the local volume polarization of the electric charge becomes

$$\frac{\rho_p - \rho_n}{\rho_p + \rho_n} = -\frac{1}{4C_{\text{sym}}} V_C(\mathbf{r}), \qquad (91)$$

where C_{sym} is the volume-symmetry coefficient of the liquid-drop model and V_C is the Coulomb potential. This leads to

$$D_0 = C_{LD} A Z e \beta_2 \beta_3, \tag{92}$$

where $C_{LD} \sim 0.0007$ fm (Strutinsky, 1957). (The $\beta_2\beta_3$ dependence of D_0 is often described as the "lightning rod" or charge-redistribution effect, referring to the tendency of electric charge to move toward regions of the surface with large curvature.)



FIG. 40. Total Routhian surface predictions of equilibrium deformations in doubly even $^{220-228}$ Th nuclei, at rotational frequencies ranging between 0 and 0.3 MeV/ \hbar (Nazarewicz *et al.*, 1987).

Lipas (1963) generalized Eq. (92) to triaxial shapes, and attempted to calculate values of D_0 using octupole mass parameters from experimental B(E3) values, but found poor agreement with experiment. Further improvements to the liquid-drop contribution to the electric dipole moment hinged on the development of the two-fluid liquid-drop (or droplet) model (Myers and Swiatecki, 1974), which considers not only nuclear density nonuniformities (denoted in the following by $\delta\rho \equiv \rho - \tilde{\rho}$, where $\tilde{\rho}$ is the average value of density) but also the effect of the neutron skin. In the droplet model, the macroscopic intrinsic dipole moment can be expressed as (Dorso *et al.*, 1986)

$$\boldsymbol{D} = \boldsymbol{D}_{\text{v-red}} + \boldsymbol{D}_{\text{skin}}, \tag{93}$$

where

$$\boldsymbol{D}_{\text{v-red}} = e \int \left(\frac{N}{A} \,\delta \rho_p - \frac{Z}{A} \,\delta \rho_n \right) \boldsymbol{r} d^3 \mathbf{r} \tag{94}$$

is the volume redistribution term, and

$$\boldsymbol{D}_{\text{skin}} = e \frac{NZ}{A} \frac{\Delta}{V} (\boldsymbol{R}_{\text{c.m.}} - \boldsymbol{R}_t)$$
(95)

is the neutron-skin contribution to the dipole moment. In Eq. (95), Δ is the neutron-skin layer volume, V is the nuclear volume, and \mathbf{R}_t is the location of the center of mass of the neutron skin. By further splitting the neutron-skin contribution into $\mathbf{D}_{skin}=\mathbf{D}_s+\mathbf{D}_{s-red}$, and combining the volume and surface redistribution terms into the total redistribution term, $\mathbf{D}_r=\mathbf{D}_{v-red}+\mathbf{D}_{s-red}$, one finally obtains

$$\boldsymbol{D} = \boldsymbol{D}_{\mathrm{r}} + \boldsymbol{D}_{\mathrm{s}} \,. \tag{96}$$

It turns out that the two contributions in Eq. (96) are comparable in magnitude but opposite in sign. Thus, the presence of the neutron skin leads to a reduction of the intrinsic dipole moment in the droplet model.

In the limit of small deformations, the macroscopic

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dipole moment can be evaluated using a second-order expansion (Dorso *et al.*, 1986). Considering axial shapes only, one obtains

$$D_{0} = C_{\rm r} A Z e \sum_{\lambda=2}^{\lambda_{\rm max}-1} \frac{49}{3\sqrt{35}} \frac{(\lambda^{2}-1)(8\lambda+9)}{[(2\lambda+1)(2\lambda+3)]^{3/2}} \beta_{\lambda} \beta_{\lambda+1} - C_{\rm s} \sum_{\lambda=2}^{\lambda_{\rm max}-1} \frac{\sqrt{35}}{15} \frac{(\lambda^{2}-1)(\lambda+3)}{\sqrt{(2\lambda+1)(2\lambda+3)}} \beta_{\lambda} \beta_{\lambda+1}.$$
(97)

It is seen that the neutron-skin term is very sensitive to higher multipoles. Indeed, for large λ values it behaves like λ^2 , while the expansion coefficients in the redistribution term are practically λ independent. [The validity of the estimate (97) was questioned by Denisov (1989, 1992), but this criticism has been shown to be without basis (Myers and Swiatecki, 1991).]

The magnitudes of C_r , C_s depend on the parameter values of the droplet-model mass formula. They are (Dorso *et al.*, 1986)

$$C_{\rm r} = \frac{9}{56\sqrt{35}} \frac{e^2}{\pi} \left[\frac{1}{J} + \frac{6L}{JK} I + \frac{15}{8Q} A^{-1/3} \right],$$

$$C_{\rm s} = \frac{15}{2\pi\sqrt{35}} \frac{NZ}{A} (I - \tilde{\delta}) R_0, \qquad (98)$$

where

$$I = \frac{N - Z}{A}, \quad \tilde{\delta} = \frac{I + (9e^{2}/80r_{0}Q)ZA^{-2/3}}{1 + (9J/4Q)A^{-1/3}}.$$
 (99)

In the above, J is the volume symmetry-energy coefficient, Q is the effective neutron-skin stiffness, K is the compressibility coefficient, and L is the density symmetry coefficient. The values of J, Q, K, and L are not known very accurately. They are usually assumed to be in the range 25 < J < 44 MeV, 17 < Q < 70 MeV, $K \approx 240$ MeV, and 0 < L < 100 MeV.

The relation between the exact droplet-model expressions, Eqs. (93)–(95), and the second-order formula (97) was investigated by Skalski (1994). In the absence of high-multipole deformations (e.g., if only β_2 and β_3 are present), the second-order expansion is rather accurate. If higher-multipole deformations are taken into account, the approximate expression quickly diverges from the exact result with increasing β_2 and β_3 . However, the difference can be compensated by changing the dropletmodel parameters Q and L within the range of their uncertainty.

Since the macroscopic dipole moment is proportional to β_2 , one expects D_0 to be very large for superdeformed, reflection-asymmetric systems. Indeed, for a large range of β_2 values, D_0 increases monotonically with elongation. Interestingly, for fixed values of oddmultipole deformations, the value of D_0 saturates at large elongations, $\beta_2 \sim 0.7$, and then falls with β_2 (Skalski, 1994). This effect arises from the reduction of the charge-redistribution term for very elongated shapes.

Donner and Greiner (1966), in their collective quadrupole-octupole model, suggested that the coupling between the collective amplitudes \mathscr{D}_{GDR} of the giant di-

pole resonance and the collective dipole moment of Eq. (90) gives rise to dipole transitions from collective octupole states.

Karpeshin (1992) has shown that reflectionasymmetric shapes during prompt fission give rise to anomalous E1 internal conversion.

By construction, macroscopic models for the intrinsic dipole moment do not contain shell effects. Those are discussed in the following section.

B. Shell-correction approach to E1 moments

The microscopic contribution to the E1 moment was introduced by Leander (1985b) In this approach, the E1 moment is written as the sum of a macroscopic (liquid-drop or droplet-model) term and a (renormalized) shell-correction term obtained using a singleparticle potential,

$$\boldsymbol{D} = \boldsymbol{D}_{\text{macr}} + \boldsymbol{D}_{\text{shell}}.$$
 (100)

The shell-correction contribution D_{shell} can be expressed as

$$\boldsymbol{D}_{\text{shell}} = e_p^{\text{eff}} \frac{N}{A} \langle \boldsymbol{r}_p \rangle_{\text{shell}} - e_n^{\text{eff}} \frac{Z}{A} \langle \boldsymbol{r}_n \rangle_{\text{shell}}, \qquad (101)$$

where $\langle r \rangle_{\text{shell}}$ is the shell correction to $\langle r \rangle$. In the presence of pairing correlations and rotation, it is equal to

$$\langle \boldsymbol{r} \rangle_{\text{shell}} = \sum_{i,j} \rho_{i,j} \langle i | \boldsymbol{r} | j \rangle - \sum_{i} n_{i} \langle i | \boldsymbol{r} | i \rangle, \qquad (102)$$

where $\hat{\rho}$ is the single-particle density matrix and n_i are the smoothed single-particle occupation numbers (Strutinsky, 1967; Brack *et al.*, 1972). In the absence of rotation, $\rho_{i,j} = v_i^2 \delta_{i,j}$, where v_i^2 is the BCS occupation coefficient for the single-particle state *i*. As demonstrated by Leander *et al.* (1986) and Butler and Nazarewicz (1991), the values of $\langle z \rangle_{\text{shell}}$ behave very regularly as a function of shell filling. They increase gradually until the middle of the shell, and decrease smoothly in the upper half of the shell (Fig. 41). These oscillations reflect the strength of unique-parity orbitals fragmented by the octupole interaction (Butler and Nazarewicz, 1991).

In contrast to wave functions obtained in the selfconsistent approach based on the microscopic effective interaction, the single-particle wave functions appearing in Eq. (102) do not contain correlations of the dipoledipole character. The screening effect caused by the particle-vibration coupling with the E1 giant resonance (Bohr and Mottelson, 1975) leads to renormalization of the single-particle matrix elements

$$\langle i|\mathbf{r}|j\rangle^{\text{eff}} = (1+\chi_{ij})\langle i|\mathbf{r}|j\rangle, \qquad (103)$$

where χ_{ij} is the state-dependent *E*1 polarizability coefficient [hence $e^{\text{eff}}=e(1+\chi_{ij})$]. The average value of χ_{ij} , denoted χ , can be estimated in the harmonic-oscillator model with a separable dipole-dipole interaction. Assuming that the particle-hole energies are all equal to e_{ph} , one obtains



FIG. 41. Shell correction to $\langle z \rangle$ given by Eq. (102), calculated in a WS model as a function of the numbers of particles filling the lowest orbits, for protons (filled circles, upper axis) and neutrons (open circles, lower axis). The deformation used corresponds to the calculated ground-state shape of ²²²Ra (Butler and Nazarewicz, 1991).

$$\chi(E) = \frac{e_{ph}^2 - E_{\text{GDR}}^2}{E_{\text{GDR}}^2 - E^2},$$
(104)

where E_{GDR} is the energy of the giant dipole resonance and *E* is the transition energy. The ground-state (*E*=0) value of χ is given by (Bohr and Mottelson, 1975; Leander *et al.*, 1986)

$$\chi_{\text{g.s.}} = \frac{e_{ph}^2}{E_{\text{GDR}}^2} - 1.$$
(105)

In the axial harmonic oscillator with deformation ε and frequency ω_0 , assuming the doubly stretched dipoledipole interaction (Sakamoto and Kishimoto, 1989), one obtains $e_{ph} = \hbar \omega_z = \hbar \omega_0 (1 - 2\varepsilon/3)$, $E_{\text{GDR}} \approx 1.9\hbar \omega_0 (1 - 2\varepsilon/3)$ (K=0), and $e_{ph} = \hbar \omega_\perp = \hbar \omega_0 (1 + \varepsilon/3)$, $E_{\text{GDR}} \approx 1.9\hbar \omega_0 (1 + \varepsilon/3)$ (K=1). An interesting consequence of Eq. (105) is that the deformation dependence in $\chi_{\text{g.s.}}$ cancels out. The *deformation-independent* ground-state E1 polarizability coefficient becomes $\chi_{\text{g.s.}} \approx -0.72$.

Leander *et al.* (1986) performed the analysis of E1 moments in nuclei that were predicted to have groundstate static deformations. In their calculations, the macroscopic part of the dipole moment was similar to that obtained by Strutinsky (1957). The shell correction was obtained using a deformed WS potential. In the calculations, the cranked SC method was also used to determine equilibrium shapes for different nuclei at various values of angular momentum, so that E1 moments could be evaluated as a function of Z, N, and I. The overall trend of the experimental results was reproduced in these calculations for Ra and Th nuclei. One feature observed is a cancellation of the liquid-drop and shell correction for ²²⁶Ra, which gives a local minimum in Z, N for the E1 moment. Improvements to these calculations were made by Butler and Nazarewicz (1991), who used the droplet model of Dorso *et al.* (1986) to take into account effects of the neutron skin. These calculations were able to reproduce the systematic trends seen in experimental data in both medium-mass and heavy nuclei (see Fig. 22). The cancellation of liquiddrop and shell corrections was found to occur also for ²²⁴Ra, in agreement with the experimental data (Poynter *et al.*, 1989a). The particular parametrization employed in the droplet model gave a very small macroscopic contribution to D_0 in the lanthanide region. Since the neutron and proton shell contributions can have opposite sign, zero or negative values of D_0 can arise. In particular, cancellation between proton and neutron contributions to $(D_0)_{\text{shell}}$ explains the reduced *E1* moment found in ¹⁴⁶Ba (Mach, Nazareweicz, *et al.*, 1990).

For several nuclei, such as ¹⁴⁸Ba and ²²⁶Ra, the calculations yield a negative value of D_0 . This interesting theoretical prediction can be tested experimentally by, for example, the measurement of an interference effect between the E1 and E3 $\Delta I=1 \gamma$ rays. Such a measurement would provide an excellent test for theoretical approaches aiming at reproducing enhanced E1 rates.

The macroscopic-microscopic approach to the intrinsic dipole moment was employed by Skalski (1994) to estimate the low-energy E1 collectivity in superdeformed and hyperdeformed configurations (see Sec. IX.B.2).

C. Self-consistent models

The most extensive and successful microscopic calculations of E1 moments for ground-state configurations in medium-mass and heavy even-even nuclei are those carried out by Egido, Robledo, and coworkers using the constrained HF+BCS model with the Gogny interaction. In the first study, based on a pure mean-field approach, Robledo et al. (1987) calculated intrinsic dipole moments in ^{222,224}Ra and ²²²Rn. In the following papers (Egido and Robledo, 1989, 1990, 1991a, 1991b, 1992; Robledo et al., 1988; Martín and Robledo, 1994) they solved the collective Schrödinger equation, with and without parity projection (Sec. III.C.3). They obtained excellent agreement with experimental B(E1) transition rates for nuclei from both octupole-deformed and octupole vibrational regions (see Fig. 22) and, in particular, they reproduced the very low values of D_0 in ²²⁴Ra and ¹⁴⁶Ba.

Skalski *et al.* (1993b) employed the generatorcoordinate HF+BCS model with the SkM* interaction to calculate low-energy *E*1 transitions in superdeformed Hg and Pb nuclei. For the lowest octupole μ =0 states, they predicted very large D_0 moments, ranging between 0.5 *e* fm and 0.8 *e* fm. For the μ =1 octupole modes, the calculated D_1 values were smaller and negative (~-0.25 *e* fm). In the two-mode calculations (involving coupling between the μ =0 and μ =2 octupole modes), these values were reduced by about 20%.

D. Algebraic models, boson models, and cluster models

In the spdf version of the interacting-boson model (see Sec. III.D), the E1 operator is a one-boson operator obtained by a direct coupling of df, pd, and sp bosons to J=1:

$$D_{\mu} = e_{1}^{(df)} \{ [d^{+} \times f^{*}]_{1\mu} + [f^{+} \times d^{*}]_{1\mu} \}$$

+ $e_{1}^{(pd)} \{ [p^{+} \times d^{*}]_{1\mu} + [d^{+} \times p^{*}]_{1\mu} \}$
+ $e_{1}^{(sp)} (p_{\mu}^{+} \times s + s^{+} \times p_{\mu}^{*}),$ (106)

see Han *et al.* (1985), Engel and Iachello (1987), and Otsuka and Sugita (1988). By adjustment of the effective charges $e_1^{(bb')}$ in Eq. (106), the experimental data can be fitted, although fine-tuning is required in the cases of ^{218,220}Ra (see Fig. 11; Engel and Iachello, 1987) and ¹⁴⁶Ba (Kusnezov and Iachello, 1988, Mach, Nazarewicz *et al.*, 1990; Zamfir and von Brentano, 1992; Liu *et al.*, 1994).

For vibrational states, Barfield *et al.* (1989) concluded that the *E*1 operator must include a two-body term in the *sdf* interacting-boson model in order to reproduce the transition rates measured in ¹⁴⁴Sm. A similar conclusion was reached by Zamfir *et al.* (1990), von Brentano *et al.* (1992), and Zamfir and von Brentano (1992), who specified that the two-body operator must arise from mixing with the giant dipole resonance.

The SU₃ limit of expression (106) was applied by Alonso *et al.* (1995) to *E*1 transition rates in ²²⁶Ra (Wollersheim *et al.*, 1993). The main features of the angular momentum pattern of observed D_0 values have been reproduced in the *spdf*-boson model.

In odd-A nuclei, the effective E1 operator can be constructed phenomenologically by coupling the odd quasiparticle to the interacting-boson model core. Dojnikov and Mikhailov (1994) assumed an *sd*-boson core, and performed interacting-boson-fermion model calculations for B(E1) rates in ⁷⁷Se. In their description, however, the *f*-boson contribution has been neglected.

Electric dipole rates in 150 Sm were estimated by Badea *et al.* (1978) in the coherent-state model based on the boson expansion method. For the *E*1 transition operator, they took

$$D_{\mu} = e_1 \{ [B_2^+ \times B_3^+]_{1\mu} + [B_3^+ \times B_2]_{1\mu} + \text{H.c.} \}, \quad (107)$$

where B_2 (B_3) are quadrupole (octupole) phonon operators. Kammuri and Kishimoto (1978) studied B(E1)/B(E2) ratios in ¹⁰⁰Ru, ¹¹²Cd, ¹⁵⁰Sm, and ¹⁵²Gd using the microscopic boson expansion technique. Here the E1 operator arises from the quadrupoleoctupole coupling and is given by an expression similar to Eq. (107), but with boson operators closely related to the RPA phonons. Kammuri and Kishimoto (1978) pointed out that the low-energy E1 strength in ¹⁰⁰Ru is reduced due to proton-neutron cancellation, and is enhanced in ¹⁵⁰Sm and ¹⁵²Gd. They concluded that the calculated E1 rates depend dramatically on shell structure and on the assumed shell-model space. For nonconjugate nuclei such as ¹⁸O, the interesting prediction of the cluster model and the vibron model discussed in Sec. III.D is that of large B(E1) transition strengths (Nemoto and Bandō, 1972; Buck and Pilt, 1977; Iachello, 1981), which arise if the two clusters have different charge-mass ratios. Alhassid *et al.* (1982) have developed molecular sum rules for radiative deexcitation widths in nuclei comprised of two arbitrary clusters. For the nucleus (A,Z) decomposed into two clusters (A_1,Z_1) and (A_2,Z_2) , the energy-weighted molecular sum rule is given by

$$S_{1m}(E1) = \frac{9\hbar^2 e^2}{8\pi m} \frac{(Z_1 A_2 - Z_2 A_1)^2}{A_1 A_2 A}.$$
 (108)

For very asymmetric cluster configurations (such as those involving α -particle clustering), the molecular sum rule (108) is considerably smaller than the nuclear sum rule

$$S_1(E1;A) = \frac{9\hbar^2 e^2}{8\pi m} \frac{NZ}{A}.$$
 (109)

Consequently, the molecular sum rule provides a new scale ($\sim 10^{-2}$ s.p.u.) for the B(E1) enhancement [see also the discussion in Suzuki *et al.* (1985), and Eq. (112)]. For experimental systematics see Cottle and Kemper (1991). Sum rules for soft dipole modes in halo nuclei were discussed by Sagawa *et al.* (1992) (see Sec. X.D).

E. Influence of octupole vibrations on E1 moments

Enhanced B(E1) rates can also be present in welldeformed reflection-symmetric nuclei, due to the coupling to low-lying octupole one-phonon states. This means that enhanced B(E1)'s do not always indicate octupole instability.

Kocbach and Vogel (1970) investigated the *E*1 transition rates between octupole and ground-state bands. Due to the Coriolis mixing between the *K*=0 and *K*=1 octupole phonons, Eq. (61), the *E*1 matrix elements are modified by the presence of a spin-dependent correction proportional to the ratio $\mathcal{Z}=D_1/D_0$. [For experimental values of \mathcal{Z} see, for example, McGowan and Milner (1981) (^{156,158,160}Gd, ¹⁶⁰Dy); McGowan and Milner (1993) (²³²Th); McGowan and Milner (1994) (²³⁸U); Ackermann *et al.* (1993) (^{230,232}Th).]

In the particle-vibration coupling formulation, the E1 transition operator can be written as

$$D_{\mu} = D_{\mu}^{s.p.} + e b_{\mu} r^3 Y_{3\mu}. \tag{110}$$

The second term in (110) represents the contribution coming from octupole vibrations, which is usually much larger than the single-particle term. It has been shown that after proper adjustment of b_{μ} , it is possible to reproduce the angular momentum dependence of the measured B(E1) rates.

Early calculations of E1 transition moments between single-particle states in the Nilsson model with BCS pairing grossly underestimated the experimental values, particularly for $\Delta K=0$ transitions in rare-earth nuclei (Monsonego and Piepenbring, 1964; Vergnes and Rasmussen, 1965) Improvement was achieved by the addition of octupole vibrations and particle-vibration interaction terms (Monsonego and Piepenbring, 1966; Faessler et al., 1966) and, additionally, Coriolis coupling (Bernthal and Rasmussen, 1967). Systematic calculations of this type have been carried out recently by Hagemann et al. (1993) using the method described by Hamamoto et al. (1989); they find that the amplitude of the octupole term depends not only on Z and N, but also on the pair of bands selected (see also Hamamoto, 1993). The importance of mixing into low-lying states by the giant dipole resonance has been stressed by Donner and Greiner (1966) and Iachello (1985), and the application of a model which includes octupole-octupole and dipole-dipole interaction by Alikov et al. (1988) found that the contribution from the latter interaction derived from giant-dipole-resonance parameters plays a dominant role in low-lying transitions in odd Eu and Tb nuclei, and in the core $0^+ - 1^-$ transitions. Similar conclusions have been reached by Guhr et al. (1989), who used an octupole vibration model which included a small isovector dipole component to describe e^- scattering form factors, and by Dojnikov et al. (1991), Soloviev and Sushkov (1991, 1994), Zilges et al. (1992), and Govaert et al. (1994), who paid particular attention to the large E1strengths found in low-lying states in many nuclei by nuclear resonance fluorescence studies [see Kneissl et al. (1993), Cottle (1994), and references therein]. For light nuclei, Castel et al. (1990) concluded that isospin mixing from the giant monopole resonance can be more important than from the giant dipole resonance.

VIII. MOLECULAR STATES IN LIGHT NUCLEI

The study of light nuclei from the sd region provided the first evidence for reflection asymmetry in rotating nuclei. Negative-parity bands with approximately rotational energy spacings have been observed in several nuclei. Figure 42 compares the behavior of the $K^{\pi}=0^{-1}$ and 0⁺ bands in ¹⁶O and ²⁰Ne; in ¹⁶O the negativeparity states become interleaved with the positive-parity states at the highest observed spins. Large α -decay widths are usually associated with the quasimolecular states (Horiuchi and Suzuki, 1973). Although large E3 strengths have been observed in these nuclei [typically 10-20 s.p.u. for the strongest transitions; see Häusser et al. (1971) and Krick et al. (1973)], measurements have been confined to transitions from lower-lying bands, such as the $K^{\pi}=2^{-}$ band in ²⁰Ne, or the well-known transition between the $1/2^+$ ground state and the $5/2^$ state at 1.35 MeV in ¹⁹F.

The E1 strengths are small in the self-conjugate nuclei $(<10^{-4} \text{ s.p.u.})$, but in nuclei with $Z \neq N$ large B(E1)'s have been observed (Cottle and Kemper, 1991).

The various theoretical descriptions of the negativeparity states in the light nuclei have invariably invoked a reflection-asymmetric basis to reproduce the observed experimental features. The natural model for describing



FIG. 42. The $K^{\pi}=0^{+}$ and 0^{-} rotational bands in ¹⁶O and ²⁰Ne. The bands shown have large α -decay widths or spectroscopic factors, except for the ground-state band in ²⁰Ne. The assignment of states in ¹⁶O is taken from Ajzenberg-Selove (1986), and in ²⁰Ne from Richards (1984) and Ajzenberg-Selove (1987).

the large α widths is the alpha-cluster model (Sec. III.E). Ikeda *et al.* (1968) suggested that cluster configurations would appear near the threshold energy for decay into the fragments. Figure 43 illustrates this idea, using the so-called Ikeda diagram (Horiuchi *et al.*, 1972). The alternative approach is that of the octupole-deformed mean field (Sec. III.B). For a review of the relationship between clusterlike configurations and deformed states see Rae (1988).

A. ¹⁶O

The two bands in ¹⁶O built on the 0_2^+ state at 6.05 MeV and the 1⁻ state at 9.58 MeV can be described adequately in terms of an α +¹²C bimolecular system (Roth and Wildermuth, 1960; Arima *et al.*, 1967; Horiuchi and Ikeda, 1968; Nemoto and Bandō, 1972; Suzuki, 1976; Suzuki, 1980; Fujiwara *et al.*, 1980; Baldock and Stratton, 1985; Descouvemont, 1987, 1991).

In the early HF parity-mixing calculations, Kelson (1965) discussed the spectrum of ¹⁶O in the $1p_{1/2}$, 2s, 1d shell-model space. In these calculations, the first excited 0⁺ state of O¹⁶ was described in terms of four particles in the *sd* shell with a small component in the $p_{1/2}$ state; however, Giraud and Sauer (1970) demonstrated that this parity-mixed state contained a large spurious component of center-of-mass motion. Do Dang *et al.* (1976) performed the HF calculations with parity projection before variation, and obtained the reflection-asymmetric ground state and the reflection-symmetric first excited 0⁺ state.



FIG. 43. Ikeda diagram for light nuclei. The threshold energy for each decay mode (in MeV) is indicated (Horiuchi *et al.*, 1972).

Many states in ¹⁶O can be described in terms of fouralpha-particle configurations that break intrinsic parity (tetrahedron, bent rhomb, kite); see Dennison (1940, 1954), Kameny (1956), Bertsch and Bertozzi (1971), Robson (1979), Bauhoff et al. (1984), and references therein. Bauhoff et al. (1984) investigated the structure of ¹⁶O in the alpha-cluster model with parity and angular momentum projection. They obtained tetrahedral symmetry for the ground state and several excited states, such as the 3^- state at 6.13 MeV. Elliott *et al.* (1985) performed an analysis based on the shell model perturbed by the tetrahedral potential, and concluded that the 0_1^+ , 3^- (6.13 MeV), and 4^+ (11.09 MeV) states form a tetrahedral rotational band. They were able to explain the α -transfer cross sections, and also the measured lifetime for the E3 decay of the 3^{-} state.

The energy spectrum and alpha-particle spectroscopic factors of the α +¹²C system have been discussed within the vibron model (Cseh, 1989; Cseh *et al.*, 1991).

B. ¹⁸O

The rotational band in ¹⁸O built upon the 0_2^+ (3.63 MeV) state was identified by Buck *et al.* (1977), Sakuda (1977), and Sakuda *et al.* (1978, 1979) as being an $\alpha + {}^{14}C$ dinuclear configuration.

Unlike in $n\alpha$ nuclei, where E1 transitions are isospin forbidden, the E1's within an alternating-parity band built on the 0_2^+ state in ¹⁸O are large, $\approx 10^{-2}$ s.p.u. (Gai *et al.*, 1983, 1987, 1989, 1991). The large electric transition strengths within the band, and the large α -decay widths of the 3^- and 4^+ members of this band, give strong support (Gai *et al.*, 1983) to the bimolecular interpretation.

Alhassid et al. (1982) found that for ¹⁸O having the configuration $\alpha + {}^{14}C$, the measured B(E1) transition rates within the molecular dipole band in ¹⁸O exhaust a significant fraction of the molecular sum rule, Eq. (108). Iachello (1985) has estimated that the octupole model would give much smaller values than are observed experimentally, although microscopic cluster models (Baye and Descouvemont, 1984; Assenbaum et al., 1984; Suzuki *et al.*, 1985) overestimate the B(E1) strength. The detailed calculations (Descouvemont and Baye, 1985; Funck et al., 1989) also suggest that the positiveand negative-parity cluster states in ¹⁸O have different intrinsic structure, and therefore cannot be regarded as being members of the same dipole-molecular band: the band proposed by Gai et al. (1983) consists of states belonging to different molecular bands [see also the discussion in Gai et al. (1989, 1991)]. The foundation of the molecular dipole states is not supported by recent α -cluster calculations of Reidemeister and Michel (1993), in which the parameters of the optical potential are taken from $\alpha + {}^{14}C$ elastic-scattering data. Instead they predict a negative-parity inversion-doublet band, an analogue to the cluster band in ¹⁶O. A semimicroscopic algebraic cluster model has been applied by Lévai et al. (1992). They are able to interpret most experimental states below 10 MeV, and find smaller B(E1) values than the earlier microscopic cluster calculations, while still overestimating the experimental values. The TDHF method has also been applied to the quasimolecular states in this nucleus, and predicts the existence of giant resonances in the $\alpha + {}^{14}C$ system having large quadrupole and octupole deformation (Strayer et al., 1984; Umar et al., 1985).

C. ²⁰Ne

A classic example of a reflection-asymmetric light system is the nucleus ²⁰Ne. Horiuchi and Ikeda (1968), Nemoto and Bando (1972), and Nemoto et al. (1975) have shown that the $K^{\pi}=0^{-}$ (5.78 MeV) and $K^{\pi}=0^{+}_{4}$ (8.3 MeV) bands of ²⁰Ne have a well-developed structure corresponding to a ${}^{16}O+\alpha$ dinucleus configuration, while the $K^{\pi}=0^+$ band has an intermediate character, between clusterlike and shell-like. This is consistent with the experimental behavior of the α widths (Table VII): the widths from the negative-parity states remain large, while the α width of states in the $K^{\pi} = 0_1^+$ band decrease at high spin. Tomoda and Arima (1978) have extended this description by combining a $(2s1d)^4$ shell-model space and a ${}^{16}\text{O}+\alpha$ cluster-model space. A more recent development has been the use of the GC method to calculate E1 strengths in ²⁰Ne using $\alpha + {}^{16}O$ T=0 and T=1 configurations (Descouvement and Baye, 1986). For other calculations related to the α +¹⁶O clustering in ²⁰Ne, see Hiura et al. (1969), Tanabe and Nemoto (1974), Yamamoto (1974), Matsuse et al. (1975), Fujiwara et al. (1980), Kazama et al. (1984), Kazama (1987), and Buck et al. (1995). Zhang et al. (1994) applied the

TABLE VII. Experimental behavior of the α widths in ²⁰Ne (Cseh *et al.*, 1991).

$\overline{K^{\pi}}$	I^{π}	E_x (MeV)	$ heta_{lpha}^2$
$\overline{0_{1}^{+}}$	0+	0	0.15
-	2+	1.63	0.11
	4 +	4.25	0.12
	6+	8.78	0.085
	8^{+}	11.95	0.01
0-	1-	5.79	1.03
	3-	7.16	0.87
	5-	10.26	0.90
	7-	13.69	0.84
	9-	17.43	0.48
o.+	0^+	- 97	0.70
0_4	0	≈8./	0.70
	2+	≈ 8.8	0.95
	4+	10.80	0.33

Bloch-Brink model to ²⁰Ne, and found a low-lying massasymmetric configuration. Dufour *et al.* (1994) investigated the α +¹⁶O system using a multicluster GC method. Kanada-En'yo and Horiuchi (1995a) studied the structure of the yrast line of ²⁰Ne with antisymmetrized molecular dynamics (see Sec. III.E). They found that for both positive- and negative-parity low-spin states (*I*<9) the two-cluster structure of α +¹⁶O is dominant (see Fig. 44).

Leander and Larsson (1975) obtained octupole instability for the prolate ground state of ²⁰Ne, caused by the strong octupole interaction between the [110]1/2 and [200]1/2 levels. Noto et al. (1976) applied the variational method with a parity- and angular-momentum-projected of the quadrupole-plus-octupole wave function deformed-shell model (Noto et al., 1974) and the Hamiltonian with the Volkov effective two-nucleon interaction. They obtained a static Y_{30} field in ²⁰Ne, and were able to reproduce the transition energies, the electromagnetic transition probabilities, and the charge form factors in the $0^+ \rightarrow 1^-$ (e,e') scattering. Marcos et al. (1983) performed parity-projected constrained HF calculations supplemented by GC calculations. They also found softness to octupole deformation in ²⁰Ne, and were able to reproduce the excitation energy of the 1⁻ state. A microscopic description of the clustering structure of ²⁰Ne was given by Provoost et al. (1984), using the quantized adiabatic TDHF theory with parity projection. They obtained a reflection-asymmetric ground state, and were able to reproduce the parity splitting.

The energy spectrum and the alpha-particle spectroscopic factors of the α +¹⁶O system have also been discussed within the vibron model (Cseh, 1989; Cseh *et al.*, 1991). Cseh (1993) also investigated a three-dimensional algebraic description of the α +¹⁶O cluster.

D.²⁴Mg

Microscopic calculations using two alpha clusters plus a 16 O core (Katō and Bandō, 1975, 1979) have explained

the ground-state $K^{\pi}=0^+$ band and the $K^{\pi}=0^-$ (7.55 MeV) band in terms of the parity doublet. Marsh and Rae (1986) applied the unconstrained alpha-cluster model with the Brink-Boeker *B*1 force. Several calculated energy minima have been associated with reflection-asymmetric intrinsic shapes corresponding to separation channels of ²⁰Ne+ α , ¹⁶O+2 α , ¹²C+¹²C, and ¹²C+3 α (Fig. 45). For other cluster-model calculations for ²⁴Mg, see Pilt and Wheatley (1978), Fujiwara *et al.* (1980), Katō *et al.* (1986), Descouvemont and Baye (1987, 1989), and Buck *et al.* (1990).

Noto (1980, 1981) has applied the deformed-shell model of Noto *et al.* (1974) to the rotational bands of ²⁴Mg. The calculations predict nonzero static quadrupole and octupole moments in the lowest $K^{\pi}=0^+$, 0^- , 2^+ , and 3^- bands. In particular, they reproduce the observed (Branford *et al.*, 1971) strong *E*3 transitions in the low-lying states of ²⁴Mg. The cranked SC calculations of Leander and Larsson (1975) predict a low-lying reflection-asymmetric hyperdeformed minimum in ²⁴Mg. This minimum, originating from the mixing between the [101] and [211] Nilsson states, can be associated with asymmetric ¹⁶O+ α + α or ¹⁶O+⁸Be structures seen in the ¹⁶O + ⁸Be and ²⁰Ne+ α resonances [see discussion in Rae (1988)].

E. ²⁸Si

According to the Ikeda diagram of Fig. 43, the excited states of ²⁸Si can be described in terms of various cluster configurations. Bauhoff (1982) extended the microscopic alpha-cluster model to describe the experimental band structure of the negative-parity bands in ²⁸Si. Katō *et al.* (1985) studied molecular ¹²C+¹⁶O structures (Erb and Bromley, 1981) using a semimicroscopic cluster model. Cseh (1992) applied the vibron model coupled to the SU₃ shell model to describe ²⁸Si in terms of a ²⁴Mg+ α configuration. Leander and Larsson (1975) found a hyperdeformed octupole shape in ²⁸Si having a similar configuration to that in ²⁴Mg. The same elongated structure has also been predicted to occur within the Bloch-Brink α -cluster model (Zhang *et al.*, 1993; Zhang *et al.*, 1994).

F. ³²S

Leander and Larssen (1975) found an excited minimum corresponding to $\varepsilon_3 = 0.3$, which has a configuration analogous to the hyperdeformed asymmetric minima in ²⁴Mg and ²⁸Si. The application of the alphacluster model by Zhang *et al.* (1994) has also yielded an extremely elongated minimum in this nucleus, having octupole deformation. Bauhoff *et al.* (1980), also in the alpha-cluster model, obtained intrinsic states with C_{2v} point symmetry. Experimental candidates for molecular bands having large alpha widths have been reported by Morita *et al.* (1985), using the ¹⁶O(²⁰Ne, α) reaction and (at higher energy) by Brenner (1994) using α -particle elastic scattering (see Fig. 46).



FIG. 44. Density distributions for ²⁰Ne, calculated in the parity-projected antisymmetrized molecular dynamics by Kanada-En'yo and Horiuchi (1995a). The scale of the axes is in fm.

G. ⁴⁰Ca

Alpha-cluster calculations have been carried out by Ogawa *et al.* (1977), who used the same model as Suzuki (1976); by Ohkubo and Umehara (1988), who used a method similar to that of Michel *et al.* (1986a, 1986b, 1988) (see next section); and by Reidemeister *et al.* (1990), who used an extension of the Buck-Dover-Vary model to describe cluster states. Although the negative-parity band is predicted to lie lower than in ⁴⁴Ti, and the α reduced widths are predicted to be larger (see next section), it is only recently that the parity-doublet band has been experimentally observed by Yamaya *et al.* (1993), using the ³⁶Ar(⁶Li,d) reaction.

H. ⁴⁴Ti

The *fp*-shell analogue of ²⁰Ne has received considerable attention within various types of alpha-cluster models. Michel *et al.* (1986a, 1986b, 1988), using parameters for a squared WS potential from α -particle scattering off ⁴⁰Ca, predicted the existence of a mixed-parity band at

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an excitation just above the α threshold energy. This prediction has been repeated by Ohkubo (1987, 1988) using a folding model with an effective two-body force; by Wada and Horiuchi (1987, 1988) (see also Horiuchi, 1985), using the GC method with a realistic two-body force, and squared WS potential for the imaginary component; and by Merchant et al. (1989), using the Buck-Dover-Vary model (for review, see Merchant et al., 1989). Earlier hybrid shell-cluster calculations had also predicted a negative-parity band lying at a lower excitation energy (Itonaga, 1981). The α -cluster band has been confirmed experimentally by Yamaya et al. (1990), who observed a negative-parity band at just above the α threshold energy using the (⁶Li,d) reaction (see also Artemov et al., 1995). There is also experimental evidence for a series of resonances with alternating natural parity in α -⁴⁰Ca elastic scattering at much higher excitation energy (Löhner et al., 1978; Frekers et al., 1983; Sellschop et al., 1987); these have been interpreted as a molecular band on the basis of microscopic GC calculations (Friedrich and Langanke, 1975), and by using the Brink-Boeker force within the GC method (Langanke,



FIG. 45. Energy minima in ²⁴Mg, calculated (Marsh and Rae, 1986) with the unconstrained alpha-cluster model using the Brink-Boeker B1 force.

1982). The cluster calculations of Merchant *et al.* (1989) have also reproduced these states.

IX. LARGE DEFORMATIONS

The microscopic mechanism behind reflection asymmetry at very elongated shapes (for example, at the fis-



FIG. 46. Mean values of excitation energies of resonance states in ³²S with the same spin *I*, plotted versus I(I+1). A *straight line* is fitted to the experimental points. The *broken line* is for an alpha particle orbiting the Si nucleus with a center-to-center radius $r=(1.25A^{1/3}+1.6)$ fm (Brenner, 1994).

sion barrier, and for superdeformed and hyperdeformed configurations) is twofold (Johansson, 1961; Gustafson et al., 1971; Larsson et al., 1974; Ragnarsson et al., 1978; Bengtsson et al., 1981; Nazarewicz, 1991). First, the octupole interaction Y_{30} couples the single-particle orbitals with asymptotic quantum numbers $[Nn_z\Lambda]\Omega$ and $[N+1 n_z \pm 1 \Lambda]\Omega$. The largest number of such matrix elements corresponds to states with the highest possible value of n_{\perp} , i.e., with $n_z=0$. The second mechanism is the octupole interaction between the high-N intruder orbitals and specific lower-N levels (e.g., levels belonging to the same superdeformed or hyperdeformed shell). For instance, the same pairs of orbitals, such as ([660]1/2 - [530]1/2) or ([770]1/2 - [640]1/2), which are responsible for octupole deformations in the light actinides, appear close to the Fermi level in superdeformed configurations around ¹⁴⁸Gd and ¹⁹²Hg. Another example is the pair of nearly degenerate Nilsson orbitals [624]9/2 and [512]5/2, which appear at the Fermi level of superdeformed ¹⁹⁴Hg. According to the calculations of Skalski (1992), these orbitals, strongly coupled by the Y_{32} field, are responsible for the μ =2 octupole instability in this nucleus.

A. Fission

It has long been recognized that reflection asymmetry is an important degree of freedom in the description of the fission process. Below, a short overview of this topic is given; for more detailed discussion the reader is referred to the extensive review by Bjørnholm and Lynn (1980).

1. Reflection asymmetry of the fission barrier

Early calculations which included Y_{30} deformation of the modified harmonic-oscillator potential (Johansson, 1961) were able to relate the degree of observed mass asymmetry of fission fragments as a function of Z^2/A to the degree of octupole deformation at the saddle point. Application of the SC method (Möller and Nilsson, 1970; Pashkevich, 1971; Möller, 1972) allowed improved calculations of the fission path to be made; these demonstrated that instability with respect to asymmetric distortion can occur at the second barrier peak. [According to SC calculations by Gavron et al. (1977) and Åberg et al. (1980), the reflection-asymmetric third saddle point also appears unstable with respect to triaxial deformations.] Self-consistent HF calculations including the mass-asymmetry degree of freedom in fission have also been carried out (e.g., Kolb et al., 1974); the accuracy of these calculations has been improved with better knowledge of the form of the nuclear interaction (see Berger et al., 1989, and references therein).

2. Bimodal fission

A good example illustrating the importance of the fragment shell structure in fission is the existence of the symmetric maximum in the mass distribution of fragments following the fission of nuclei approaching $^{264}_{100}$ Fm₁₆₄, and, in particular, the simultaneous appearance of two fission modes in ²⁵⁸Fm, with low and high total kinetic energy of the fragment (Hulet et al., 1986). The asymmetric two-center model of Mustafa et al. (1973) showed that the system can revert to symmetry in the path from the second saddle point to scission, for certain values of Z and N [see also Maruhn and Greiner (1974) and Lustig et al. (1980)]. Wilkins et al. (1976) associated this effect with the presence of the very strong ¹³²Sn+¹³²Sn channel. The shell-correction calculations (Möller et al., 1987; Pashkevich, 1988; Ćwiok et al., 1989), involving a rather rich space of nuclear shapes, predict the existence of several valleys on the potentialenergy surface in the vicinity of the scission point. One family of fission valleys corresponds to reflectionsymmetric configurations. However, there also exists a valley in which the asymmetric nuclear shape is close to the combination of spherical and elongated fragments.

The experimental observation of asymmetry in the fission mass distribution provides substantial clues to the nature of the mass distribution of the fissioning nucleus, but more direct evidence comes from the observation of nuclear levels at the elongated shapes. In particular, Blons and his collaborators (for a review, see Blons, 1989, and references therein) have measured fission resonances in Th and U nuclei. As discussed in Sec. IX.B.3, these have been interpreted as arising from reflection-asymmetric hyperdeformed shapes which exist in the third minimum of the fission barrier.

B. Strongly elongated configurations

1. Shell structure at strongly elongated configurations

The observation of superdeformed (SD) and hyperdeformed (HD) states constitutes an important confirmation of the shell structure of the nucleus; their unusual stability can be attributed to strong shell effects that are present in the average nuclear potential.

The SD and HD shell structures are often explained in terms of the deformed harmonic-oscillator model. At large deformations, corresponding to rational oscillatorfrequency ratios (which characterize a *rational harmonic oscillator*), strong degeneracy of the harmonic-oscillator eigenstates occurs, leading to the appearance of SD and HD magic gaps and magic numbers (Bohr and Mottelson, 1975). In the case of an axial rational harmonic oscillator with frequencies ω_z and ω_{\perp} , it is convenient to write the single-particle energies in terms of the shell frequency ω_{shell} and the shell principal quantum number N_{shell} , defined by (Bohr and Mottelson, 1975; Bengtsson *et al.*, 1981)

$$\omega_{\perp} n_{\perp} + \omega_{z} n_{z} = \omega_{\text{shell}} N_{\text{shell}}, \quad N_{\text{shell}} = n_{\perp} k_{z} + n_{z} k_{\perp},$$

$$\omega_{\text{shell}} = \frac{\tilde{\omega}}{k_{\perp} k_{z}}, \quad (111)$$

where $\omega_z k_z = \omega_\perp k_\perp = \tilde{\omega}$, and $\tilde{\omega}$ can be calculated from the volume-conservation condition. The fact that degenerate single-particle orbitals forming SD ($k_\perp = 1$, $k_z = 2$) and HD ($k_\perp = 1$, $k_z = 3$) shells have different parities has Table VIII displays the energies e_{ph} of particle-hole excitations associated with various components of the octupole tensor $Q_{3\mu}^{"}$ [Eq. (64)] (Nakatsukasa *et al.*, 1992; Nazarewicz *et al.*, 1992, 1995; Nazmitdinov and Åberg, 1992). For the SD shape, $2\omega_3-\omega_{\perp}=0$, and the particlehole energy corresponding to the $\mu=1$ mode is zero for states belonging to the same SD shell, $N_{\text{shell}}=2n_{\perp}+n_z$ [Eq. (111)]. This suggests that the triaxial modes with $\mu=1$ can play a significant role in open-shell SD nuclei (see Sec. IX.B.2). Similarly, at the SD oblate shape ($k_{\perp}=2$, $k_z=1$) the rational-harmonic-oscillator model suggests an instability with respect to $\mu=0$ and $\mu=2$ octupole modes.

In quantum-mechanical systems, strong shell effects (i.e., degeneracies) usually reflect the presence of dynamical (self-consistent) symmetries of the Hamiltonian. This offers additional quantum numbers associated with the underlying dynamical symmetry, which are the eigenvalues of the Casimir operators of the corresponding symmetry group. The connection between the SU_3 dynamical symmetry of the rational harmonic oscillator, nuclear superdeformation, and octupole modes was discussed by Nazarewicz et al. (1992) and Nazarewicz and Dobaczewski (1992). The quantum numbers of the rational harmonic oscillator allow for a transparent classification of magic numbers at large deformations. In particular, they justify the scheme of touching harmonic oscillators suggested by Bengtsson et al. (1981) and Faber and Płoszajczak (1981).

The idea of the classification scheme is illustrated in Fig. 47. At SD shapes, two kinds of closed-shell systems are expected. In the "asymmetric" case (labeled A), the magic numbers are equal to sums of two consecutive spherical magic numbers, i.e., N=1, 5, 14, 30, 55, etc. In the "symmetric" variant (B), the magic numbers are equal to doubled spherical-oscillator magic numbers, N=2, 8, 20, 40, 70, etc. The situation becomes slightly more complex at HD shapes, where the magic numbers correspond to the superposition of three sphericaloscillator magic numbers. In the "strongly asymmetric" variant (A) magic numbers N=1, 6, 18, 40, 75, etc., are obtained by combining two spherical shells $\mathcal N$ and one higher shell $\mathcal{N}+1$. In the intermediate situation (B), two spherical shells $\mathcal{N}+1$ and one lower shell \mathcal{N} are combined, i.e., N=2, 9, 24, 50, 90, etc. Finally, in the "symmetric" case (C), the magic numbers are equal to tripled spherical-oscillator magic numbers, i.e, N=3, 12, 30, 60, 105, etc.

The classification scheme shown in Fig. 47 suggests the description of deformed systems in terms of "multiclusters" of spherical magic subsystems, dictated by the decomposition of the rational-harmonic-oscillator representations into the isotropic ones (Nazarewicz *et al.*, 1992; Nazarewicz and Dobaczewski, 1992). The relation between multicluster classification and the mean-field picture was discussed by Nazarewicz and Dobaczewski (1992), who analyzed the octupole couplings in the

4	0	2	
-	v	~	

TABLE VIII. Energies of the particle-hole excitation, Δe_{ph} , associated with the octupole doubly stretched interaction $Q''_{3\mu}$.

μ	$\Delta e_{ph}/\hbar$	Optimal conditions for instability
0	$\omega_3, 2\omega_\perp - \omega_3, 2\omega_\perp + \omega_3, 3\omega_3$	superdeformed oblate shapes
1	ω_{\perp} , $2\omega_3-\omega_{\perp}$, $2\omega_3+\omega_{\perp}$, $3\omega_{\perp}$	superdeformed prolate shapes
2	$\omega_3, 2\omega_\perp - \omega_3, 2\omega_\perp + \omega_3$	superdeformed oblate shapes
3	ω_{ot} , $3\omega_{ot}$	no instability

rational-harmonic-oscillator model using the doubly stretched octupole interaction, Eq. (64).

The results of rational-harmonic-oscillator calculations for the shell-energy octupole-stiffness coefficient C_{30} [Eq. (38)] (Nazarewicz and Dobaczewski, 1992) are displayed in Fig. 48. For the spherical shape [Fig. 48(a)] the octupole-driving shell force is positive, i.e., there is no tendency to develop stable octupole deformations. The situation at the SD prolate shape is shown in Fig. 48(b). For particle numbers representing the asymmetric cases (A), C_{30} is negative. For the symmetric cases (B), there is no shell octupole-driving force toward reflection-asymmetric shapes. Finally, the HD case is illustrated in Fig. 48(c). As expected, for the systems representing the asymmetric case of Fig. 47, the shell correction decreases with octupole deformations, while no octupole-driving tendency is predicted for the symmetric case. A similar pattern has also been predicted for the $\mu \neq 0$ octupole modes (Nazarewicz, 1992; Arita and Matsuyanagi, 1993). A similar result was obtained by Nazmitdinov and Åberg (1992) and Arita and Matsuyanagi (1993) for giant octupole vibrations built on SD states, namely, a lower vibrational energy arises for the asymmetric combination of spherical magic numbers than for the symmetric combination.

Calculations based on realistic mean-field potentials confirm the prediction of the rational harmonic oscillator that the particle-number regions favoring reflectionsymmetric or reflection-asymmetric SD and HD shapes should alternate (Bengtsson et al., 1981; Åberg, 1990; Höller and Åberg, 1990; Dudek et al., 1990; Li et al., 1991). As seen in Fig. 49, based on the Nilsson-Strutinsky model, for SD shapes (ε =0.6) the tendency towards mass asymmetry is strongly favored at particle numbers around 32, 64, and 116, while for particle numbers around 44, 86, and 144, the shell correction favors reflection-symmetric shapes. For HD shapes ($\varepsilon = 0.9$), the mass asymmetry is strongly favored at particle numbers around 34, 82, and 150, while for particle numbers around 22, 48, 58, 116, and 130, the minimum shellcorrection energy is found at $\varepsilon_3=0$. Shell-correction maps as a function of particle number and triaxial octupole deformations were calculated by Li et al. (1991) for SD nuclei.

2. Superdeformations

The first experimental evidence for octupole correlations in SD configurations at high spins was found by Cullen *et al.* (1990) in 193 Hg. They explained a lowfrequency pseudocrossing in one of the observed SD bands in terms of an admixture of an octupole phonon. Further evidence for octupole vibrations in SD Hg nuclei was found by Crowell *et al.* (1994, 1995), who reported an excited SD band in ¹⁹⁰Hg, having a rather constant moment of inertia, and decaying directly into the yrast SD band. The extracted B(E1) values, of the order of 10^{-3} s.p.u., agree well with the results of RPA calculations (see below). The strongest indication of octupole correlations in the SD $A \sim 150$ mass region was found in excited bands of ¹⁵²Dy (Dagnall *et al.*, 1994), where the large-interaction band crossing in one band, and a decay branch of another band into the yrast SD band, can be understood in terms of octupole vibrations.



FIG. 47. Spectrum, in units of $\tilde{\omega}$, of the rational harmonic oscillator with $k_{\perp}=1$ and $k_z=2$ (SD prolate, left) or $k_z=3$ (HD prolate, right). Each level is labeled by the principal quantum number M and quantum numbers $(\lambda_1\lambda_2\lambda_3)$ of the rational harmonic oscillator. Except for the usual SU₃ degeneracy, $\frac{1}{2}$ (M+1)(M+2), there are no additional degeneracies present. Different positions of the Fermi level for closed-shell systems A, B, or C are indicated. The schematic diagrams in the bottom portion illustrate the number of occupied particles within each { λ } family in cases A, B, and C (Nazarewicz and Dobaczewski, 1992).



FIG. 48. Shell-correction octupole-stiffness coefficient C_{30} [in units of $7/(4\pi\omega_0^4)$] as a function of the shell quantum number, defined as $N_{\text{shell}} \equiv n_\perp k_z + n_z$. Magic particle numbers A (with spin degeneracy included) are indicated for all closed-shell configurations of the rational harmonic oscillator at (a) spherical, (b) superdeformed, and (c) hyperdeformed shapes. If C_{30} is negative (positive), then there is (is not) a shell force favoring stable octupole deformations (Nazarewicz and Dobaczewski, 1992).

Dudek *et al.* (1987, 1990) studied the mirrorasymmetric deformations in high-spin states of SD nuclei using the cranked SC method with a WS potential. They found static octupole minima in a number of nuclei in the $Z \sim 66$ and $Z \sim 80$ mass regions (cf. Höller and Åberg, 1990).

According to RPA calculations based on the cranked shell model (Mizutori et al., 1990; Mizutori et al., 1991a, 1991b; Nakatsukasa et al., 1992), low-lying octupole vibrations built on the SD shape are rather collective and, in some cases, can give rise to increased band interaction, shifts in the crossing frequency, reduced angular momentum alignment, and enhanced E1 transitions (see Sec. VI). In particular, the RPA calculations of Nakatsukasa et al. (1993), based on the doubly stretched octupole interaction of Eq. (64), demonstrate that inclusion of the coupling between quasiparticle and octupole vibrational modes is important for understanding the experimental data for SD ¹⁹³Hg. This treatment was successfully applied to SD octupole bands in ¹⁵²Dy (Nakatsukasa et al., 1995) and ¹⁹⁰Hg (Crowell et al., 1995).

Satuła *et al.* (1991) concluded, in the SC-WS model, that nuclei around ¹⁹²Hg are soft to octupole deforma-



FIG. 49. Shell energy obtained in the modified oscillator model (Höller and Åberg, 1990) versus N for SD (top) and HD (bottom) shapes at $\varepsilon_3=0$ (solid lines) and $\varepsilon_3=0.15$ (dashed lines).

tion β_3 . Bonche *et al.* (1991) and Meyer *et al.* (1995), starting from the axial HF+BCS theory with the SkM* effective interaction, used parity projection and the GC method to investigate the octupole softness in the SD minimum of ¹⁹⁴Pb. They found octupole states in the energy range of 2 MeV above the SD ground state. Delaroche *et al.* (1992) investigated the octupole softness of SD ¹⁹²Hg in the HFB+GOA model with the D1 interaction. They obtained very excited $K^{\pi}=0^{-}$ states at $E \approx 8$ MeV.

According to the rational-harmonic-oscillator model discussed in Sec. IX.B.1. triaxial octupole deformations, and especially the $\mu=1$ component, are expected to be important at SD shapes (Mottelson, 1988). Li et al. (1991) searched for triaxial octupole shapes in SD nuclei, and found stable minima in ¹⁵⁸Hf (μ =1) and ¹⁵⁶Yb (μ =2). The effects of μ =1 octupole deformations on the total-energy surface of nuclei from the A=190mass region were studied by Chasman (1991) using the cranked shell-correction method with a WS potential. In many nuclides, minima with rather small quadrupole and hexadecapole deformations and very large $\mu=1$ octupole deformations have been found to become approximately yrast at moderate spins. This result was questioned by Skalski (1992) who concluded, using a similar model, that the $\mu=1$ octupole modes are suppressed by shell effects and by rotation, and that the softest octupole modes in the SD Hg region are those with $\mu=0$ and 2. In subsequent work, Skalski *et al.* (1993b) presented GC calculations for the octupole modes having $\mu=0, 1, \text{ and } 2, \text{ using the Skyrme SkM}^*$ and SIII effective interactions. They predicted the presence of collective octupole K=0, 1, and 2 bands at energies 1.9–2.5 MeV above the SD minimum in Hg and Pb nuclei. They also calculated large B(E3) transition rates $(\approx 30 \text{ s.p.u.})$, depopulating the octupole modes.

To estimate *E*1 moments in SD nuclei, Skalski (1994) employed the SC method of Sec. VII.B, with the macroscopic part described by the droplet model and the microscopic part computed using a WS model. In the SD Gd–Dy region, the shell-correction contributions $(D_0)_{\text{shell}}$ are negative, and tend to cancel the macroscopic part. In the SD Hg-Pb region, in contrast, both contributions have equal signs, resulting in sizable dipole moments.

The GC+HF+BCS calculations by Skalski *et al.* (1993b) predicted the presence of very large E1 matrix elements [$B(E1) \sim 0.02$ s.p.u.] connecting octupole vibrational states with the yrast band in SD Hg and Pb nuclei.

3. Hyperdeformations

Hyperdeformed nuclei, i.e., nuclei with quadrupole deformations significantly larger than β_2 =0.6, are known or predicted in several mass regions. As discussed in Sec. IX.B.1, HD configurations can be stabilized in some cases by developing reflection-asymmetric deformations. Good cases of HD configurations can be found in light nuclei. For instance, the calculated (Leander and Larsson, 1975) low-lying reflection-asymmetric HD minimum in ²⁴Mg can be associated with the asymmetric ¹⁶O+ α + α (or ¹⁶O+⁸Be) structure (see Sec. VIII.D). Other examples are the HD states in ³⁶Ar (¹⁶O+¹⁶O+ α) and ⁴⁸Cr (¹⁶O+¹⁶O+¹⁶O) (Rae and Merchant, 1992). HD reflection-asymmetric structures are also discussed by Faber and Płoszajczak (1981), who performed cranked SC-WS calculations for nuclei around ²⁸Si.

According to the cranked SC method calculations by Höller and Åberg (1990), the best candidates for HD reflection-asymmetric configurations at high spins are nuclei around ¹⁴⁶Gd, ¹⁹⁴Hg, and ²⁰⁰Rn. Figure 50 (Åberg, 1993) displays the potential-energy surface for ¹⁴⁶Gd at *I*=60. The reflection-asymmetric HD band ($\varepsilon_2 \approx 0.93$, $\varepsilon_3 \approx 0.13$) in ¹⁴⁶Gd is expected to cross the SD (reflection-symmetric) band at $I \sim 80$. The single-particle properties of HD configurations around ¹⁴⁶Gd are discussed by Åberg (1993). He concluded that, because of octupole mixing, the high- \mathcal{N} classification scheme is expected to break down in HD configurations; i.e., the calculations suggest rather small differences in the moments of inertia between different HD bands.

Hyperdeformed reflection-asymmetric states in the neutron-deficient Hg and Pb isotopes with 98 < N < 110 have been calculated (Nazarewicz, 1993) to occur at excitation energies >6-7.5 MeV at I=0. These minima are rather shallow, although due to their large moments of inertia they become deeper at high angular momenta (Höller and Åberg, 1990). Energy surfaces in the $A \sim 180$ region have been investigated by Chasman and Robledo (1995), who found octupole softness in the very extended nuclear shapes in a number of nuclei.

In the actinide nuclei, the HD states are the so-called third minima around ²³²Th (Pashkevich, 1971; Möller, 1972). In these nuclei the second saddle point is split, leading to an excited reflection-asymmetric configuration with large quadrupole and octupole deformations ($\beta_2 \sim 0.90$, $\beta_3 \sim 0.35$), as predicted in a number of calculations based on the mean-field approach (Pashkevich, 1971; Möller, 1972; Möller and Nix, 1973; Howard and Möller, 1980; Bengtsson *et al.*, 1987; Berger *et al.*, 1989;



FIG. 50. Potential-energy surface in the $(\varepsilon_2, \varepsilon_3)$ plane for the $I^{\pi}=60^+$ configuration in ¹⁴⁶Gd, calculated by Åberg (1993) in the cranked modified harmonic-oscillator model. The solid contours are separated by 2 MeV, the intermediate ones by 1 MeV.

Pal, 1993; Cwiok et al., 1994; Rutz et al., 1994). Experimentally, the third minimum is indicated from a microstructure in the resonances found in the light actinides using the (n,f), (t,pf), and (d,pf) reactions (Back *et al.*, 1972; Blons et al., 1975, 1978, 1984; Bjørnholm and Lynn, 1980; Fabbro et al., 1984; Blons, 1989; Nakagome et al., 1991). The position and intensities of the resonances observed in the best-studied case, ²³¹Th (Fig. 51), are best explained as the superposition of two rotational bands of opposite parity, with K=1/2, and having very large moments of inertia ($\approx 250 \text{ MeV}^{-1}$). Evidence for intrinsic parity mixing comes from the measured asymmetry in the fission angular distributions (Baumann et al., 1989). Other evidence supporting the thirdminimum hypothesis comes from analysis of the slopes of the near-barrier photofission cross sections of actinides (Bhandari and Al-Kharam, 1989).

A systematic study of HD states in actinides using the SC method (Bengtsson *et al.*, 1987) yielded very shallow reflection-asymmetric third minima. The HD states around 232 Th also appear in self-consistent calculations based on the HFB model with the D1S interaction (Berger *et al.*, 1989), adiabatic TDHF-Skyrme-Yukawa calculations (Pal, 1993), and in relativistic mean field calculations (Rutz *et al.*, 1994).

In their systematic calculations for the even-even Rn, Ra, Th, and U isotopes, based on the SC method with a WS potential, Ćwiok *et al.* (1994) employed a manydimensional deformation space ($\beta_2 - \beta_7$), allowing for a rather general description of axially deformed reflection-asymmetric shapes. The potential-energy surfaces in the (β_2, β_3) plane for several even-even Rn, Ra, Th, and U nuclei, covering the region between the second minimum and the outer barrier, are displayed in Fig. 52. For all the nuclei shown in Fig. 52, there exist well-developed reflection-asymmetric HD minima ($\beta_2 \sim 0.9, 0.35 < \beta_3 < 0.65$). It is interesting to note that in the nuclei around ²³⁴U, the HD minimum splits into two distinct minima with very different values of β_{λ} (λ =3-7). Ćwiok *et al.* (1994) pointed out that the structure of the third minimum corresponds to a binuclear configuration involving a spherical (or nearly spherical) heavy fragment around ¹³²Sn and a well-deformed lighter fragment around ¹⁰⁰Zr. This is illustrated in Fig. 53, which shows the predicted equilibrium shapes of ²³²Th. The shape corresponding to the HD minimum looks like a superposition of the ¹³²Sn and ¹⁰⁰Zr ground-state shapes. Recently, this result has been confirmed by the self-consistent calculations of Rutz *et al.* (1994).

The clustering effect predicted in the third minima is a striking manifestation of nuclear shell structure; in the rational harmonic oscillator model (Sec. IX.B.1), the particle numbers 80 and 150 correspond to a situation which formally resembles one spherical doubly magic fragment and one well-deformed (or SD) lighter fragment. The very special role played by the ¹³²Sn structure in the fission process has been noted before in connection with mass distributions of fission fragments (Wilkins *et al.*, 1976), the analysis of cold-fission data (Asghar *et al.*, 1993; Gönnenwein, 1994), measurements of mass and kinetic-energy distributions for the photofission of ²³²Th (Piessens *et al.*, 1993), and the existence of two fission modes around ²⁵⁸Fm (see the discussion in Sec. IX.A).

The spectroscopic properties of HD minima in actinides (single-particle levels, parity doublets, decoupling parameters, g factors, moments of inertia, etc.) were discussed by Bengtsson *et al.* (1987). Skalski (1994) calculated a very large intrinsic dipole moment in the HD minimum of ²³²Th, $D_0 \approx 2.2$ *e* fm.

X. PERSPECTIVES

In this last section, a different landscape of reflection asymmetry is viewed and suggestions made as to how this topic could be further explored experimentally and theoretically.

A. Unexplored mass regions

For nuclei, manifestations of reflection asymmetry in the intrinsic system, such as interleaved positive- and negative-parity states, parity doubling in odd-mass systems, and large odd electric moments, are most pronounced in the Ra-Th region with $N \approx 134$, and, to a lesser extent, in the Ba-Sm region with $N \approx 88$. Other regions in N, Z and deformation space are expected to exhibit similar characteristics whenever there are orbitals near the Fermi surface differing by three units in ℓ and *j*. For normal deformed nuclei, two examples in the transitional region above 100Sn have been studied, ¹¹⁴Xe (Rugari et al., 1993) and ¹¹⁰Te (Paul et al., 1994). Experimental evidence for octupole correlations has also been cited for other mass regions: ⁶⁴Ge (Görres *et al.*, 1987, Ennis *et al.*, 1991); ⁷⁴Se, ⁷⁸Kr (Cottle, 1990b; Cottle *et al.*, 1990); ¹²⁸Ba (Cottle *et al.* 1989, Cottle, 1990b, 1991), and ¹²⁸Gd (Cottle, 1990a) (but see Cottle et al., 1992, 1993). So far, both SC calculations (Skalski, 1990) and self-consistent calculations (Heenen et al.,



FIG. 51. Top: (n, f) resonances in ²³¹Th. Bottom: resulting rotational band of alternating parity (Blons *et al.*, 1984).

1994) have been applied to the region near ¹¹²Ba, and lend support to a description of these transitional nuclei having strong octupole collectivity with very shallow octupole minima. The light Ba nuclei have been investigated using a microscopic treatment which requires a μ =1 octupole component (Piepenbring and Leandri, 1991).

Shell-correction calculations have also been applied to 64 Ge, predicting octupole softness at nonzero values of γ (Skalski, 1991; Ennis *et al.*, 1991). Nakatsukasa *et al.*





(1994) investigated low-energy octupole states in the even-even proton-rich Kr, Sr, and Zr nuclei using the quasiparticle RPA. They demonstrated that, unlike in the heavier nuclei, the octupole collectivity in the light Kr-Zr region is weak, and not directly correlated with the excitation energy systematics of the lowest negative-parity states.

More experimental evidence is needed, particularly for the transitional nuclei near closed shells or subshells.

These systems are rather difficult to study experimentally, being either very neutron deficient or very neutron rich. Experimental limitations also restrict at present the study of ²²⁴U, predicted to be the most octupoledeformed nucleus in its ground state (Nazarewicz *et al.*, 1984b).

In ⁹⁶Zr, the strength of the octupole coupling arises from $2p_{3/2} \rightarrow 1g_{9/2}$ proton and $2d_{5/2} \rightarrow 1h_{11/2}$ neutron particle-hole excitations. Ohm *et al.* (1990) have at-



FIG. 53. Calculated equilibrium shapes of ²³²Th at the groundstate (top), fission-isomeric (middle), and third-minimum (bottom) configuration. The values of ρ_1^{max} and ρ_2^{max} are 6.42 and 5.22 fm, respectively.

tempted to calculate both the excitation energy of the 3^{-} state at 1.897 MeV and the B(E3) strength using the quasiparticle RPA. The difficulties encountered using the quasiparticle RPA (Abbas et al., 1981) suggest that there is a large degree of anharmonicity in the octupole degree of freedom in this mass region. Octupole correlations in ⁹⁶Zr were investigated by Fayans *et al.* (1994) in the continuum quasiparticle RPA based on a densityfunctional approach. They obtained $E(3_1)=1.74$ MeV and B(E3)=39.5 s.p.u., in reasonable agreement with the recent lifetime data (Horen et al., 1993). See also Rosso et al. (1993). Mach, Cwiok et al. (1990) have performed SC calculations which suggest that, while ⁹⁶Zr contains rather complex anharmonic motion, the neighboring isotope ⁹⁸Zr should have strong octupole instability with $\hat{\beta}_3 = 0.13$ in its ground-state configuration. Skalski et al. (1993a) have made GC+HF calculations for ${}^{94-100}$ Zr. Although they were able to reproduce the energy of the lowest 3⁻ state, the calculated value of the $B(E3;0^+ \rightarrow 3^-)$ in the case of ⁹⁶Zr was gravely underestimated, irrespective of the effective force used.

B. Reflection-asymmetric shapes at large deformations

That strong octupole correlations should exist for certain SD states is supported by several theoretical calculations, and is suggested by recent experimental evidence (Sec. IX.B.2). The development of more sophisticated γ -ray arrays will enable the properties of low-lying octupole bands to be measured systematically in the SD region. Perhaps more exciting is the possible occurrence of reflection asymmetry in HD structures (Blons, 1989, Cwiok *et al.*, 1994). It remains an experimental challenge to identify bound nuclear states in HD heavy nuclei.

C. Angular momentum dependence of electromagnetic moments and reaction aspects

The determination of E3 moments between members of rotational bands has been performed for a few nuclei using Coulomb excitation techniques (see Sec. IV.D). These techniques, which demonstrate directly how octupole collectivity varies with spin, cannot be applied at present to short-lived nuclei, precluding studies for nuclei having the deepest octupole minimum.

More detailed measurements might allow the nature of the octupole minimum to be determined. Indeed, static reflection-asymmetric deformations are expected to show up in the low-energy heavy-ion induced reactions, such as Coulomb excitation or sub-barrier fusion. Catara et al. (1989) calculated fusion cross sections for ¹⁶O induced reactions on ¹⁴⁴Ba and different Ra isotopes. They concluded that the presence of static octupole deformation gives rise to a significant enhancement of the total fusion cross section. Consequently, systematic measurements of the sub-barrier fusion cross sections for the chain of isotopes (in which octupole deformation varies quickly with neutron number) could provide a signature for the presence of octupole deformation. Dasso et al. (1993) studied Coulomb-excitation patterns for systems with stable octupole deformations. According to their calculations, the probabilities for Coulomb excitation of the different members of the rotational band can discriminate between the deformed and vibrational limits of octupole deformation. In the static case, the predicted spin population does not depend on the parity of the state, whereas in the vibrational case the population of negative-parity states is strongly reduced.

D. Halos

Interesting modifications of the E1 strength distributions have already been observed in nuclei such as ¹¹Li and ¹¹Be that have very weak binding of their outermost neutrons (halo nuclei). As pointed out by Uchiyama and Morinaga (1985), in systems with loosely bound states the hindrance of low-energy E1 transitions may be removed, since the transition takes place outside the nucleus where the core polarizability vanishes (see Sec. VII.B). They introduced a cutoff model in which the effective charge vanishes inside a certain radius (the core radius plus the nuclear-force range). Hoshino et al. (1991) confirmed this assumption in large-scale shellmodel calculations for ¹¹Li and ¹¹Be (see also Hansen and Jonson, 1987; Bertsch and Foxwell, 1990; Suzuki and Tosaka, 1990; Hayes and Strottman, 1990; Hayes, 1991; Esbensen and Bertsch, 1992; Danilin et al., 1994). Sagawa *et al.* (1992) studied sum rules for soft modes in halo nuclei. The corresponding E1 sum rule is

$$S_{\text{halo}}(E1) = S(E1;A1) + S(E1;A2) + S_{1m}(E1), \quad (112)$$

where the sum rules $S_{1m}(E1)$ and S(E1;A) are given by Eqs. (108) and (109), respectively.

In one-body halo nuclei such as ¹¹Be, or in nuclei such as ⁶He or ¹¹Li that can be considered as three-body halos, reflection-asymmetric intrinsic configurations appear in calculations (see, for example, Horiuchi *et al.*, 1994; Fedorov *et al.*, 1994; Kanada-En'yo and Horiuchi, 1995b, 1995c).

E. Parity violation

Parity violation (in the laboratory frame) is caused by the parity-nonconserving component, V^{PNC} , of the weak interaction. The magnitude of this effect is of the order of $\alpha_p = G_F m_{\pi}^2 / G_S \sim 10^{-7}$, where G_F is the Fermi constant and G_S is the strong coupling constant.

The relationship between intrinsic reflection asymmetry and real parity mixing is as follows. First, in nuclei with reflection-asymmetric shapes there is a large probability of finding parity doublets very close in energy, especially in the odd-A and odd-odd systems (Leander and Sheline, 1984). This is the necessary condition for experimental studies of parity mixing (Ahmad et al., 1982; Haxton and Henley, 1983; Adelberger and Haxton, 1985). Second, it can indeed be demonstrated (Sushkov and Flambaum, 1980; Flambaum and Sushkov, 1980) that the parity-mixing amplitude can be expressed through the matrix elements of \hat{V}^{PNC} between the *intrin*sic states characterizing the parity doublet. In particular, Sushkov and Flambaum (1980) and Flambaum and Sushkov (1980) pointed out that the spatial parity violation induced by polarized thermal neutrons observed in the fission of some actinide nuclei could be related to the presence of intermediate pear-shaped states during the fission process.

The latest focus of parity-violation measurements in nuclei is on symmetry breaking in compound nuclear states (Bowman et al., 1990; Zhu et al., 1992; Frankle et al., 1992), motivated by the extremely large enhancement observed for parity violation in neutron resonances. In the statistical approach, the paritynonconserving matrix elements are treated as random variables and, in most cases, this method gives qualitative agreement with experimental data (Zhu et al., 1992). However, for one nucleus (²³²Th), strong sign correlations have been observed (Frankle et al., 1992). This phenomenon has been the subject of much study, and cannot be explained as a general property of the nucleon-nucleus weak interaction (Auerbach, 1992; Bowman et al., 1992). Recently, the nonstatistical effects in ²³²Th have been discussed (Auerbach et al., 1995; Flambaum and Zelevinsky, 1995) in terms of parity doublets resulting from reflection-asymmetric deformations, which are expected to exist in the third HD minimum of ²³²Th (see Sec. IX.B.3). In particular, Flambaum and Zelevinsky (1995) related the parity-mixing amplitude to the ratio of matrix elements involving the paritynonconserving interaction, and to the part of the interaction giving rise to nonadiabatic mixing, leading to parity splitting.

F. Reflection-asymmetric deformations in mesoscopic systems

The presence of reflection-asymmetric deformations can be attributed to strong shell effects that are present in the average potential of finite Fermi systems such as atomic nuclei or metallic clusters (Hamamoto, Mottelson, *et al.*, 1991; Frauendorf and Pashkevich, 1993; Frisk *et al.*, 1994; Koskinen *et al.*, 1995). The shapes of light nuclei and jellium clusters were studied by Koskinen *et al.* (1995) in the symmetry-unrestricted local density approximation, assuming a simplified energy functional. The shapes obtained for light N=Z nuclei show a striking similarity to those of atomic clusters. In particular, they obtained reflection-asymmetric shapes in jellium clusters of N=10, 12, 16, and 18 electrons. (For corresponding nuclear examples, see Sec. VIII.)

Figure 38 shows an example of a situation where the combined reflection-symmetric and reflectionasymmetric deformations give rise to the new nuclear shell structure. The single-particle WS levels ($\omega=0$ portions), characteristic of nuclei from the Ba-Nd region, are plotted as a function of octupole deformation β_3 . Due to octupole mixing between the $h_{11/2}$ and $d_{5/2}$ subshells, two large energy gaps open up at Z=56 and Z=62 in the proton spectrum. In the neutron system a large gap at N=88 can be seen; this is a consequence of the $i_{13/2}-f_{7/2}$ octupole interaction. A close inspection of Fig. 38 (ω =0) leads to the conclusion that, at large octupole deformations, the single-particle levels become nearly degenerate, forming quasi-j subshells. For example, at $\beta_3=0.15$, the two orbitals with $\Omega=1/2$ and 3/2are close to each other, and just below the Z=56 gap, forming a "j"=3/2 multiplet. Similar "j"=5/2 and "j"=7/2 multiplets can be seen above the Z=56 and Z=62 gaps, respectively. An analogous structure can also be observed in the neutron system.

Theoretically, the variation in the single-particle level density with shell filling (the level bunching), the existence of spherical and deformed magic numbers, and the unusual shell stability of certain shapes in mesoscopic systems, have an interpretation in terms of semiclassical periodic orbits (Balian and Bloch, 1972; Bohr and Mottelson, 1975; Strutinsky and Magner, 1976). The singleparticle level density and the shell energy can be expressed (Gutzwiller, 1967, 1971) as a sum over semiclassical periodic orbits. Consequently, the shell structure of a many-body system (and hence the presence or absence of large deformations) has its deep roots in the nonlinear dynamics of the corresponding classical Hamiltonian and the geometry of classical orbits.

Several authors (Frisk, 1990; Arita and Matsuyanagi, 1993, 1994, 1995; Arita, 1994; Heiss *et al.*, 1994, 1995a, 1995b) have analyzed the classical single-particle motion

in axially symmetric potentials with quadrupole and octupole deformations. They demonstrated that strong shell effects present at large quadrupole and octupole deformations are associated with bifurcations of short periodic orbits. Examples of classical periodic orbits for the harmonic-oscillator Hamiltonian with the doubly stretched octupole field,

$$H = h_{\rm ho} - \lambda_{30} M \omega_0^2 [r^2 Y_{30}(\Omega)]'', \qquad (113)$$

are shown in Fig. 54. Such a semiclassical analysis can shed new light on shell structures stabilizing reflectionasymmetric shapes, and on associated constants of motion.

The character of the collective dynamics (and in particular, the nature of energy dissipation) of independent classical particles moving in a deformed, axially symmetric container, oscillating with a frequency much smaller than a typical single-particle frequency, was discussed by Błocki *et al.* (1992, 1993, 1995). They demonstrated that largely ordered motion characteristic of quadrupole shapes becomes chaotic if octupole and higher-order deformations are included. Bauer *et al.* (1994) considered a harmonic-oscillator potential with a time-dependent multipole-multipole force, treated self-consistently.

The problems of strong quantum numbers, strong shell effects, and so on, appear naturally in the context of large-amplitude collective motion, which embraces such phenomena as anharmonic vibrations, shape coexistence, exotic decays, fission, fusion, and heavy-ion collisions. All those phenomena involve dynamical interaction between various stable mean fields characterized by different symmetries. The transition from one stable mean field to another goes through one of several level crossings, around which the original symmetry of the system is broken and the intrinsic quantum numbers disappear. In the context of this Review, two examples of large-amplitude collective motion can be mentioned, both related to the physics of quantum-mechanical tunneling.

The first example is the description of parity splitting in a nuclear rotational band. As discussed in Sec. IV.B and Sec. VI, very small parity splitting between the highspin members of the parity doublet has been observed in, for example, ²²⁰Ra. Since most calculations predict rather modest octupole barriers (1–2 MeV), the unusual octupole rigidity at high spins, and the variation of the parity splitting with angular momentum, are challenges for theoretical models of large-amplitude collective motion.

The second example deals with the microscopic descriptions of fission and exotic decay. There exist many descriptions of fission based on the adiabatic assumption, most employing the adiabatic TDHF theory and variations of it (such as the collective Schrödinger equation with microscopic mass tensor). By construction, adiabatic approaches cannot take into account properly the dynamics of level crossing and the associated symmetry breaking (see Sec. III.A). An interesting development is the imaginary-time mean-field theory (Levit *et al.*, 1980; Negele, 1989), which allows for a TDHF



FIG. 54. Some classical periodic orbits for the Hamiltonian (113), with the frequency ratio $\omega_{\perp} / \omega_z = \sqrt{3}$ (Arita and Matsuyanagi, 1994).

treatment of quantum tunneling. An example of this method is the description of the fission of 32 S in three dimensions (Wolff *et al.*, 1992). In particular, the crossings between positive- and negative-parity levels give rise to a dramatic difference between the results of the static constrained HF (or adiabatic TDHF) and TDHF; the corresponding collective-motion paths are strongly influenced by the breaking of reflection symmetry, and are very different in both cases.

A better understanding of the basic aspects of nuclear dynamics that govern large-amplitude collective motion will certainly be provided by the vigorous interdisciplinary interaction between nuclear physics and nonlinear dynamics.

ACKNOWLEDGMENTS

Valuable comments and suggestions from Sven Åberg and George Bertsch are gratefully acknowledged. Theoretical nuclear physics research at the University of Tennessee is supported by the U.S. Department of Energy through Contract No. DE-FG05-93ER40770. Oak Ridge National Laboratory is managed for the U.S. Department of Energy by Lockheed Martin Energy Systems under Contract No. DE-AC05-84OR21400. Nuclear physics research at the University of Liverpool is supported by grants from the U.K. Engineering and Physical Sciences Research Council (previously the Science and Engineering Research Council). P. A. Butler is grateful for support from the Joint Institute for Nuclear Research (Oak Ridge) for various visits there in which some of this work was completed; similarly W. Nazarewicz acknowledges support from the EPSRC for visits to the University of Liverpool. W. Nazarewicz thanks the Department of Energy's Institute for Nuclear Theory at

the University of Washington for its hospitality and partial support during the completion of this work.

APPENDIX: COMPILATION OF SHAPE PARAMETRIZATIONS

The standard α and β parametrization of the nuclear shape is introduced in Sec. II.A. This Appendix contains a brief compilation of other shape parametrizations of reflection-asymmetric shapes.

For the modified oscillator potential (Nilsson, 1955), the shape is defined through the equipotential surfaces of the average field:

$$\rho_t^2 = \frac{\rho_0^2}{c(\tilde{\varepsilon})} \left\{ 1 + 2\varepsilon_1 P_{10}(\theta_t) - \frac{2}{3} \varepsilon_2 P_{20}(\theta_t) + 2\sum_{\lambda} \varepsilon_{\lambda} P_{\lambda}(\theta_t) \right\}^{-1/2}.$$
(A1)

Here, the nuclear radius is denoted by ρ_t because it, together with the angles entering P_{λ} in Eq. (A1), is defined in the stretched coordinate system (Nilsson, 1955; Larsson, 1973). The dipole deformation ε_1 is determined from Eq. (3). Expressed in terms of Nilsson's deformation ε_3 , the octupole moment $Q_{30,c}$ is given by (Leander and Chen, 1988)

$$\mathcal{Q}_{30,c} = -\frac{6}{7} Z R_0^3 \varepsilon_3. \tag{A2}$$

Another way of parametrizing axial shapes has been suggested by Chasman *et al.* (1977). Here the deformations ν_{λ} are introduced via the transformation

F ()

$$r^{2} \rightarrow r^{2} \left[\exp\left(\frac{2\nu_{2}}{3}\right) \sin^{2}\theta + \exp\left(-\frac{4\nu_{2}}{3}\right) \cos^{2}\theta + \sum_{\lambda=3}^{\lambda_{\max}} \nu_{\lambda} \sqrt{\lambda + \frac{1}{2}} P_{\lambda}(\cos\theta) \right].$$
(A3)

,

When discussing asymmetric fission, straightforward expansions in spherical harmonics are not very appropriate. Parametrizations have to be introduced that are capable of describing the bifurcation of the surface into two parts. The three-quadratic surface parametrization of Nix (1969) is sometimes used at large elongations. For axial systems, the (c,h,α) cylindrical parametrization of nuclear shape is particularly useful (Nix, 1972; Brack *et al.*, 1972):

$$(x^2 + y^2)/C^2$$

$$=\begin{cases} (1-z^2/C^2)(A+Bz^2/C^2+\alpha z/C) \text{ for } B \ge 0,\\ (1-z^2/C^2)[(A+\alpha z/C)\exp(Bc^3z^2/C^2)] \text{ for } B < 0, \end{cases}$$
(A4)

where *C* is determined by the volume-conservation condition, B=2h+(c-1)/2, and $A=1/c^3-B/5$. The parameter *c* describes the elongation of the system, *h* is the neck coordinate, and α is an asymmetry parameter.

A different parametrization has been used by Pashkevich (1971). It is based on the observation that the nuclear shapes at the saddle point are well approximated by Cassinian ovals. For axial shapes, the Cassinian coordinates (R,t) are related to the Cartesian coordinates (x,y,z) by the expressions

$$R = [(x^{2} + y^{2} + z^{2})^{2} - 2s(z^{2} - x^{2} - y^{2}) + s^{2}]^{1/4},$$

$$t = \frac{\text{sgn}(z)}{\sqrt{2}} \left\{ 1 + \frac{z^{2} - x^{2} - y^{2} - s}{R^{2}} \right\}^{1/2},$$
 (A5)

where s is the squared distance from the focus of the Cassinian ovals to the origin of coordinates. The surface R(t)=const defines the Cassinian oval. In order to describe deviations from such shapes, R is written as a series in Legendre polynomials:

$$R(t) = R_0 \left[1 + \sum_m \alpha_m P_m(\cos t) \right].$$
(A6)

The reflection-asymmetric shapes correspond to odd-*m* deformations α , i.e., α_1, α_3 , and so on.

There is no unique way to relate one set of deformation parameters to another. A standard method of comparing shapes is to require equality between the collective multipole moments $\mathcal{Q}_{\lambda\mu}$ (Dudek *et al.*, 1984; Nazarewicz and Ragnarsson, 1995). For instance, in first order, the relations for $\lambda=3$ deformations are (Rohoziński, 1988)

$$\beta_3 = -2\sqrt{\frac{\pi}{7}}\varepsilon_3, \quad \nu_3 = 2\sqrt{\frac{2}{7}}\varepsilon_3$$
 (A7)

[see Eqs. (6) and (A2)].

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