

Table of Alpha-Disintegration Energies of the Heavy Elements*†

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THIS compilation is a listing of total alpha-decay energies with some supporting information pertinent to the means by which the energies were derived. The only references given will be those relevant to the energy determinations. The decay energies are the Q values for the alpha transitions and can be transformed into mass differences by including the atomic mass of He^4 .

The major uncertainty in many of the adopted Q values lies in the sphere of inadequate information on decay schemes. For example, there are relatively few cases in which measurements have been made showing that the highest-energy alpha groups are not in coincidence with gamma rays. In principle, therefore, it is possible in many cases that the alpha group used to define the decay energy does not represent the ground-state transition and that the true decay energy is greater than expressed.

There is now sufficient knowledge of the nature of alpha spectra to conclude that this uncertainty is not uniform among all of the nuclear species and probably serious for few, if any. In the first place, it is reasonably certain that for all *even-even* nuclei the most intense alpha group does indeed go to the ground state. The presence of other alpha groups of lower energy does not seriously distort the energy obtained for the main group even when the measuring instrument does not have sufficient resolution to separate the components of the complex spectrum.

For nuclei with odd nucleons, the principal alpha group will, more often than not, populate an excited state. It is in these cases that there is greatest uncertainty. However, because of the apparent close energy level spacing among nuclei of these types in the heavy element region, a state which can be heavily populated is generally found not far from the ground state. As implied here, there are selection processes which guide the population of various states, but in alpha decay the penalty for low energy is very severe and an abundant alpha group of several hundred kilo-

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volts lower energy than the total decay energy is not to be expected. There are a few exceptions to this rule, such as Cm^{243} in which the principal alpha group is known to be of 277-kev lower energy than the transition to the ground state. Further comments will be made in the following discussion of the several columns of Table I.

We have limited this compilation to alpha energies which have actually been measured and because of this arbitrary condition have not entered some which can be calculated with substantial precision. Omitted, for example, is RaE , whose alpha-decay energy can be calculated by closing a decay cycle, and the actual alpha-decay process has been detected through the growth of the decay product. There are a number of other cases in which a fairly accurate calculation of this type can be made and a large number of cases for which estimations can be made based partly on the systematic behavior of alpha energies.

It should also be noted that the so-called "long-range alpha groups" of RaC' and ThC' are not used to determine decay energies, since these represent transitions from excited states of the respective nuclei.

COLUMN 1

This column indicates the alpha emitter and its product as well as the half-life which is given solely for purposes of further identification. These are the measured half-lives and not the partial alpha-decay half-lives for those cases in which there is more than one mode of decay. Since this table is not a compilation of general decay properties, no references are given for the half-lives cited.

COLUMNS 2 AND 3

In a large fraction of the cases the "highest-energy group" of column 3 is either known to be that of the ground-state transition or is assumed to be so in the absence of information regarding a complex spectrum for the purpose of calculating the decay energy of Column 2. The Q values, unless otherwise stated under "comments," were calculated by adding to the energy of Column 2 the recoil energy, $4E/(A-4)$, where E is the alpha-particle energy and A is the mass number of the emitter.

COLUMN 4

The absence of a notation under "intensity" means that no high-resolution instrument has been used to

TABLE I.

Reaction	Adopted Q (Mev)	Highest energy group measured	Intensity (%)	Method	Energy ref.	Ident-ification	Comments
$\text{Bi}^{198} \rightarrow \text{Tl}^{194}$ 7 m	5.95	5.83		ion ch	1	B	ins evid
$\text{Bi}^{199} \rightarrow \text{Tl}^{195}$ ~25 m	5.58	5.47		ion ch range mica	1	B	ins evid
$\text{Bi}^{201} \rightarrow \text{Tl}^{197}$ 62 m	5.25	5.15		ion ch	1	B	ins evid
$\text{Bi}^{203} \rightarrow \text{Tl}^{199}$ 12 h	4.95	4.85		range emuls	2	B	ins evid
$\text{Bi}^{209} \rightarrow \text{Tl}^{205}$ ~10 ¹⁷ y	3.1	3.15 2.9		range emuls range emuls	3 4		ins evid
$\text{Bi}^{210} \rightarrow \text{Tl}^{206}$ ~10 ⁶ y	5.03	4.93		ion ch	5	A	
$\text{Bi}^{211} \rightarrow \text{Tl}^{207}$ 2,16 m	6.747	6.618 6.621	82.6	spect spect	6 7	A	
$\text{Bi}^{212} \rightarrow \text{Tl}^{208}$ 60.5 m	6.203	6.082 6.087 6.090 6.088	27.2	spect spect spect spect	8, 9 10 11 12	A	
$\text{Bi}^{213} \rightarrow \text{Tl}^{209}$ 47 m	5.97	5.86		ion ch	13	A	ins evid
$\text{Bi}^{214} \rightarrow \text{Tl}^{210}$ 19.7 m	5.610	5.505	45	spect	8	A	ins evid
$\text{Po}^{200} \rightarrow \text{Pb}^{196}$ 11 m	5.96	5.84		ion ch	14	B	e-e
$\text{Po}^{201} \rightarrow \text{Pb}^{197}$ 18 m	5.82	5.70		ion ch	14	B	ins evid
$\text{Po}^{202} \rightarrow \text{Pb}^{198}$ 52 m	5.70	5.59		ion ch	15	B	e-e
$\text{Po}^{204} \rightarrow \text{Pb}^{200}$ 3.8 h	5.48	5.37		ion ch	15	B	e-e
$\text{Po}^{205} \rightarrow \text{Pb}^{201}$ 1.5 h	5.3	5.2		ion ch	14	B	ins evid
$\text{Po}^{206} \rightarrow \text{Pb}^{202}$ 9 d	5.321	5.218	96	spect	16	B	e-e
$\text{Po}^{207} \rightarrow \text{Pb}^{203}$ 5.7 h	5.20	5.10		ion ch	14	B	ins evid
$\text{Po}^{208} \rightarrow \text{Pb}^{204}$ 2.93 y	5.208	5.108 5.109	~100	spect spect	17 16	A	
$\text{Po}^{209} \rightarrow \text{Pb}^{205}$ ~100 y	4.972	4.877	~100	spect	17	A	
$\text{Po}^{210} \rightarrow \text{Pb}^{206}$ 138.4 d	5.402	5.299 5.304 5.298	~100	spect spect spect	8, 10, 18 11 19	A	
$\text{Po}^{211} \rightarrow \text{Pb}^{207}$ 0.52 s	7.58	7.434 7.442	99	range air spect	8 20	A	It is not certain which Po^{211} group represents the ground state transition
$\text{Po}^{211} \rightarrow \text{Pb}^{207}$ 25 s		7.14		ion ch	23	D	
$\text{Po}^{212} \rightarrow \text{Pb}^{208}$ 3.04×10^{-7} s	8.949	8.776 8.780 8.786	~100	spect spect spect	8, 9 10 11	A	
$\text{Po}^{213} \rightarrow \text{Pb}^{209}$ 4.2×10^{-6} s	8.496	8.336		ion ch	24	A	ins evid
$\text{Po}^{214} \rightarrow \text{Pb}^{210}$ 1.637×10^{-4} s	7.827	7.680 7.683	~100	spect spect	9 19	A	
$\text{Po}^{215} \rightarrow \text{Pb}^{211}$ 1.83×10^{-3} s	7.52	7.365 7.383	~100	range air spect	8 25	A	
$\text{Po}^{216} \rightarrow \text{Pb}^{212}$ 0.158 s	6.903	6.774 6.775	~100	spect spect	9 26	A	
$\text{Po}^{217} \rightarrow \text{Pb}^{213}$ short	6.66	6.54		ion ch	27	C	ins evid
$\text{Po}^{218} \rightarrow \text{Pb}^{214}$ 3.05 m	6.110	5.998 5.996	~100	spect spect	9 28	A	

¹ H. M. Neumann and I. Perlman, Phys. Rev. **78**, 191 (1950).² D. C. Dunlavey and G. T. Seaborg, Phys. Rev. **85**, 757 (1952).³ H. Faraggi and A. Berthelot, Compt. rend. **232**, 2093 (1951).⁴ W. Riezler and W. Porschen, Z. Naturforsch. **7a**, 634 (1952).⁵ H. B. Levy and I. Perlman, Phys. Rev. **85**, 758 (1952); **94**, 152 (1954).⁶ M. G. Holloway and M. S. Livingston, Phys. Rev. **54**, 18 (1938).

(Summarizes work of various investigators.)

⁷ G. Vieira and L. Salgueiro, Compt. rend. **234**, 1765 (1952).

TABLE I.—Continued.

Reaction	Adopted Q (Mev)	Highest energy group measured	Intensity (%)	Method	Energy ref.	Ident-ification	Comments
$\text{At}^{203} \rightarrow \text{Bi}^{199}$ 7 m	6.22	6.10		ion ch	29	D	ins evid
$\text{At}^{205} \rightarrow \text{Bi}^{201}$ 25 m	6.02	5.90		ion ch	29	B	ins evid
$\text{At}^{207} \rightarrow \text{Bi}^{203}$ 2.0 h	5.86	5.75		ion ch	29	B	ins evid
$\text{At}^{208} \rightarrow \text{Bi}^{204}$ 1.7 h	5.76	5.65		ion ch	30	A	ins evid
$\text{At}^{209} \rightarrow \text{Bi}^{205}$ 5.5 h	5.76	5.65		ion ch	29	B	ins evid
$\text{At}^{210} \rightarrow \text{Bi}^{206}$ 8.3 h	5.626	5.519	32	spect	31	A	ins evid
$\text{At}^{211} \rightarrow \text{Bi}^{207}$ 7.5 h	5.975	5.862	~100	spect	31	A	
$\text{At}^{213} \rightarrow \text{Bi}^{209}$ short	9.4	9.2		range emuls	32	E	ins evid
$\text{At}^{214} \rightarrow \text{Bi}^{210}$ short	8.95	8.78		ion ch	33	B	ins evid
$\text{At}^{215} \rightarrow \text{Bi}^{211}$ ~10 ⁻⁴ s	8.15	8.00		ion ch	33	A	ins evid
$\text{At}^{216} \rightarrow \text{Bi}^{212}$ ~3 × 10 ⁻⁴ s	7.94	7.79		ion ch	33	A	ins evid
$\text{At}^{217} \rightarrow \text{Bi}^{213}$ 0.018 s	7.14	7.02 7.00		ion ch ion ch	24 34	A	ins evid
$\text{At}^{218} \rightarrow \text{Bi}^{214}$ ~2 s	6.75	6.63		ion ch	35	E	ins evid
$\text{At}^{219} \rightarrow \text{Bi}^{215}$ 0.9 m	6.39	6.27		ion ch	36	A	ins evid
$\text{Em}^{204} \rightarrow \text{Po}^{200}$ ~3 m	6.4	6.3		ion ch	37	D	e-e
$\text{Em}^{206} \rightarrow \text{Po}^{202}$ 6.2 m	6.37	6.24 6.25		ion ch ion ch	37 38	B	e-e
$\text{Em}^{207} \rightarrow \text{Po}^{203}$ 10 m	6.22	6.11 6.09		ion ch ion ch	37 38	B	ins evid
$\text{Em}^{208} \rightarrow \text{Po}^{204}$ 23 m	6.258	6.138		spect	39	D	e-e
$\text{Em}^{209} \rightarrow \text{Po}^{205}$ 30 m	6.14	6.02		ion ch	40	B	ins evid
$\text{Em}^{210} \rightarrow \text{Po}^{206}$ 2.7 h	6.153	6.036	~100	spect	39	B	
$\text{Em}^{211} \rightarrow \text{Po}^{207}$ 16 h	5.960	5.847	33	spect	39	A	
$\text{Em}^{212} \rightarrow \text{Po}^{208}$ 23 m	6.382	6.262	~100	spect	39	A	
$\text{Em}^{215} \rightarrow \text{Po}^{211}$ short	8.8	8.6		ion ch	41	B	ins evid
$\text{Em}^{216} \rightarrow \text{Po}^{212}$ short	8.16	8.01		ion ch	42	A	e-e
$\text{Em}^{217} \rightarrow \text{Po}^{213}$ ~10 ⁻³ s	7.89	7.74		ion ch	33	A	ins evid
$\text{Em}^{218} \rightarrow \text{Po}^{214}$ 0.019 s	7.260	7.127	~100	spect	43	A	
$\text{Em}^{219} \rightarrow \text{Po}^{215}$ 3.92 s	6.94	6.824 6.807	69	spect	44 45	A	ins evid
$\text{Em}^{220} \rightarrow \text{Po}^{216}$ 54.5 s	6.398	6.282 6.278	~100	spect spect	9 28	A	
$\text{Em}^{222} \rightarrow \text{Po}^{218}$ 3.825 d	5.587	5.486 5.482	~100	spect spect	9 28	A	
$\text{Fr}^{212} \rightarrow \text{At}^{208}$ 19.3 m	6.532	6.409	37	spect	46	A	

^a W. B. Lewis and B. V. Bowden, Proc. Roy. Soc. (London) **A145**, 235 (1934). (Summarizes the results of various investigators. The values were recalculated according to reference 6. See also reference 22.)

^b G. H. Briggs, Revs. Modern Phys. **26**, 1 (1954).

^c S. Rosenblum and G. Dupouy, Compt. rend. **194**, 1919 (1932); J. phys. et radium **4**, 262 (1933).

^d Collins, McKenzie, and Ramm, Proc. Roy. Soc. (London) **A216**, 219 (1953). (See also reference 22.)

TABLE I.—Continued.

Reaction	Adopted Q (Mev)	Highest energy group measured	Intensity (%)	Method	Energy ref.	Ident- ification	Comments
$\text{Fr}^{217} \rightarrow \text{At}^{213}$ short	8.5	8.3		range emuls	32	E	ins evid
$\text{Fr}^{218} \rightarrow \text{At}^{214}$ short	8.00	7.85		ion ch	33	B	ins evid
$\text{Fr}^{219} \rightarrow \text{At}^{215}$ 0.02 s	7.44	7.30		ion ch	33	A	ins evid
$\text{Fr}^{220} \rightarrow \text{At}^{216}$ 27.5 s	6.81	6.69		ion ch	33	A	ins evid
$\text{Fr}^{221} \rightarrow \text{At}^{217}$ 4.8 m	6.42	6.30	75	ion ch	24, 47	A	ins evid
$\text{Ra}^{213} \rightarrow \text{Em}^{209}$ 2.7 m	7.03	6.90		ion ch	48	B	ins evid
$\text{Ra}^{219} \rightarrow \text{Em}^{215}$ short	8.1	8.0		ion ch	41	B	ins evid
$\text{Ra}^{220} \rightarrow \text{Em}^{216}$ short	7.57	7.43		ion ch	42	A	e-e
$\text{Ra}^{221} \rightarrow \text{Em}^{217}$ 30 s	6.83	6.71		ion ch	33	A	ins evid
$\text{Ra}^{222} \rightarrow \text{Em}^{218}$ 38 s	6.674	6.554	~100	spect	43	A	
$\text{Ra}^{223} \rightarrow \text{Em}^{219}$ 11.2 d	5.967	5.860 5.730	weak 9	spect spect	16 16	A	ins evid
$\text{Ra}^{224} \rightarrow \text{Em}^{220}$ 3.64 d	5.784	5.681 5.679	95	spect spect	9, 26 28	A	
$\text{Ra}^{226} \rightarrow \text{Em}^{222}$ 1622 y	4.863	4.777 4.778	94.3	spect spect	28 18	A	
$\text{Ac}^{221} \rightarrow \text{Fr}^{217}$ short	7.7	7.6		range emuls	32	E	ins evid
$\text{Ac}^{222} \rightarrow \text{Fr}^{218}$ 5.5 s	7.09	6.96		ion ch	33	B	ins evid
$\text{Ac}^{223} \rightarrow \text{Fr}^{219}$ 2.2 m	6.76	6.64		ion ch	33	A	ins evid
$\text{Ac}^{224} \rightarrow \text{Fr}^{220}$ 2.9 h	6.28	6.17		ion ch	33	A	ins evid
$\text{Ac}^{225} \rightarrow \text{Fr}^{221}$ 10.0 d	5.90	5.80		ion ch	47, 24	A	ins evid
$\text{Ac}^{227} \rightarrow \text{Fr}^{223}$ 22.0 y	5.031	4.942		spect	16	A	ins evid
$\text{Th}^{223} \rightarrow \text{Ra}^{219}$ short	7.69	7.55		ion ch	41	B	ins evid
$\text{Th}^{224} \rightarrow \text{Ra}^{220}$ short	7.26	7.13		ion ch	42	A	e-e
$\text{Th}^{225} \rightarrow \text{Ra}^{221}$ 8.0 m	6.69	6.57		ion ch	33	A	ins evid
$\text{Th}^{226} \rightarrow \text{Ra}^{222}$ 30.9 m	6.450	6.336	77	spect	43	A	
$\text{Th}^{227} \rightarrow \text{Ra}^{223}$ 18.6 d	6.138	6.030	19	spect	16	A	
$\text{Th}^{228} \rightarrow \text{Ra}^{224}$ 1.90 y	5.520	5.423 5.421	71	spect spect	49 50	A	
$\text{Th}^{229} \rightarrow \text{Ra}^{225}$ 7340 y	5.11	5.02	~10	ion ch	47	A	ins evid
$\text{Th}^{230} \rightarrow \text{Ra}^{226}$ 8.0×10^4 y	4.765	4.682	76.3	spect	51	A	
$\text{Th}^{232} \rightarrow \text{Ra}^{228}$ 1.39×10^{10} y	4.064	3.994 4.006	76	ion ch range emuls	52 53	A	

¹² A. Rytz, Journal recherches centre natl. recherche sci. No. 25, December, 1953. See reference 88.

¹³ English, Cranshaw, Demers, Harvey, Hincks, Jelley, and May, Phys. Rev. **72**, 253 (1947).

¹⁴ Karraker, Ghiorso, and Templeton, Phys. Rev. **83**, 390 (1951).

¹⁵ D. G. Karraker and D. H. Templeton, Phys. Rev. **81**, 510 (1951).

¹⁶ Rosenblum, Perey, Valadares, and Guillot, mentioned in reference 21.

¹⁷ Asaro, Heiman, and Perlman (unpublished data).

¹⁸ F. Asaro and I. Perlman (unpublished data).

¹⁹ W. J. Sturm and V. Johnson, Phys. Rev. **83**, 542 (1951).

²⁰ S. Rosenblum, Compt. rend. **193**, 848 (1931). (The energy was calculated by Briggs, see reference 9.)

²¹ Hollander, Perlman, and Seaborg, Revs. Modern Phys. **25**, 469 (1953).

²² I. Perlman and F. Asaro, Ann. Revs. Nuclear Sci. **4**, (1954).

²³ F. N. Spiess, Phys. Rev. **94**, 1292 (1954).

²⁴ T. E. Cranshaw and J. A. Harvey, Can. J. Research **26A**, 243 (1948).

²⁵ Marie Curie and S. Rosenblum, Compt. rend. **194**, 1232 (1932). (This energy was calculated by Briggs, see reference 9.)

TABLE I.—Continued.

Reaction	Adopted Q (Mev)	Highest energy group measured	Intensity (%)	Method	Energy ref.	Ident-ification	Comments
$\text{Pa}^{226} \rightarrow \text{Ac}^{222}$ 1.8 m	6.93	6.81		ion ch	33	B	ins evid
$\text{Pa}^{227} \rightarrow \text{Ac}^{223}$ 38.3 m	6.58	6.46		ion ch	33	A	ins evid
$\text{Pa}^{228} \rightarrow \text{Ac}^{224}$ 22 h	6.20	6.09	75	ion ch	33	A	ins evid
$\text{Pa}^{229} \rightarrow \text{Ac}^{225}$ 1.5 d	5.79	5.69 5.66		ion ch ion ch	54 55	A	ins evid
$\text{Pa}^{231} \rightarrow \text{Ac}^{227}$ 3.43×10^4 y	5.131	5.042	11	spect	56	A	
$\text{U}^{227} \rightarrow \text{Th}^{223}$ 1.3 m	6.9	6.8		ion ch	41	B	ins evid
$\text{U}^{228} \rightarrow \text{Th}^{224}$ 9.3 m	6.79	6.67		ion ch	42	A	e-e
$\text{U}^{229} \rightarrow \text{Th}^{225}$ 58 m	6.53	6.42		ion ch	33	A	ins evid
$\text{U}^{230} \rightarrow \text{Th}^{226}$ 20.8 d	5.992	5.888	68	spect	43	A	
$\text{U}^{231} \rightarrow \text{Th}^{227}$ 4.3 d	5.55	5.45		ion ch	57	A	ins evid
$\text{U}^{232} \rightarrow \text{Th}^{228}$ 73.6 y	5.411	5.318	68	spect	18	A	
$\text{U}^{233} \rightarrow \text{Th}^{229}$ 1.62×10^5 y	4.907	4.823	83	ion ch	24	A	
$\text{U}^{234} \rightarrow \text{Th}^{230}$ 2.48×10^5 y	4.845	4.762 4.763	74	ion ch	52	A	
$\text{U}^{235} \rightarrow \text{Th}^{231}$ 7.13×10^8 y	4.66	4.58	10	ion ch	59	A	
$\text{U}^{236} \rightarrow \text{Th}^{232}$ 2.39×10^7 y	4.577	4.499	73	ion ch	60	A	
$\text{U}^{238} \rightarrow \text{Th}^{234}$ 4.49×10^9 y	4.253	4.182	77	ion ch	52	A	
$\text{Np}^{231} \rightarrow \text{Pa}^{227}$ ~50 m	6.39	6.28		ion ch	61	A	ins evid
$\text{Np}^{233} \rightarrow \text{Pa}^{229}$ 35 m	5.63	5.53		ion ch	61	A	ins evid
$\text{Np}^{235} \rightarrow \text{Pa}^{231}$ 410 d	5.15	5.06		ion ch	62	A	ins evid
$\text{Np}^{237} \rightarrow \text{Pa}^{233}$ 2.20×10^6 y	4.85	4.77		ion ch	63	A	α - γ coin
$\text{Pu}^{232} \rightarrow \text{U}^{228}$ 36 m	6.70	6.58		ion ch	64	B	e-e
$\text{Pu}^{234} \rightarrow \text{U}^{230}$ 9.0 h	6.30	6.19	86	ion ch	64	A	
$\text{Pu}^{235} \rightarrow \text{U}^{231}$ 26 m	5.95	5.85		ion ch	64	B	ins evid
$\text{Pu}^{236} \rightarrow \text{U}^{232}$ 2.7 y	5.85	5.75 5.75	80	range air ion ch	65 63	A	
$\text{Pu}^{238} \rightarrow \text{U}^{234}$ 89.6 y	5.589	5.495	72	spect	66	A	
$\text{Pu}^{239} \rightarrow \text{U}^{235}$ 24360 y	5.238	5.150 5.147	69	spect spect	67 68	A	ins evid
$\text{Pu}^{240} \rightarrow \text{U}^{236}$ 6580 y	5.249	5.162	76	spect	67	A	

²⁵ S. Rosenblum and C. Chamie, Compt. rend. 196, 1663 (1933).²⁶ Momyer, Ghiorso, and Hyde (unpublished data, 1952).²⁷ G. Bastin-Scoffier and J. Sant'ana-Dionisio, Compt. rend. 236, 1016 (1953).²⁸ Barton, Ghiorso, and Perlman, Phys. Rev. 82, 13 (1951).²⁹ Hyde, Ghiorso, and Seaborg, Phys. Rev. 77, 765 (1950).³⁰ R. W. Hoff and F. Asaro (unpublished data).³¹ J. D. Keys, Ph.D. thesis, McGill University (1951).³² Meinke, Ghiorso, and Seaborg, Phys. Rev. 81, 782 (1951).³³ Hagemann, Katzin, Studier, Ghiorso, and Seaborg, Phys. Rev. 72, 252 (1947).³⁴ B. Karlik and T. Bernert, Naturwiss. 31, 298 (1943); Z. Physik 123, 51 (1944).³⁵ E. K. Hyde and A. Ghiorso, Phys. Rev. 90, 267 (1953).³⁶ A. W. Stoner and E. K. Hyde (unpublished data, 1954).³⁷ Private communication from W. E. Burcham to G. T. Seaborg (1954).³⁸ Momyer, Asaro, and Hyde (unpublished data, 1952).³⁹ Momyer, Hyde, Ghiorso, and Glenn, Phys. Rev. 86, 805 (1952).

TABLE I.—Continued.

Reaction	Adopted Q (Mev)	Highest energy group measured	Intensity (%)	Method	Energy ref.	Ident- ification	Comments
Pu ²⁴¹ → U ²³⁷ 13.0 y	4.976	4.893	75	spect	18	A	α-γ coin
Pu ²⁴² → U ²³⁸ 1×10^6 y	4.980	4.898	80	spect	18	A	e-e
Am ²³⁷ → Np ²³³ ~1.3 h	6.11	6.01		ion ch	69	B	ins evid
Am ²³⁹ → Np ²³⁵ 12 h	5.85	5.75		ion ch	69	B	ins evid
Am ²⁴¹ → Np ²³⁷ 470 y	5.628	5.535	0.3	spect	70	A	
Am ²⁴³ → Np ²³⁹ 7600 y	5.430	5.341	0.3	spect	71	A	
Cm ²³⁸ → Pu ²³⁴ 2.5 h	6.61	6.50		ion ch	72	B	e-e
Cm ²⁴⁰ → Pu ²³⁶ 26.8 d	6.36	6.25		ion ch	73	A	e-e
Cm ²⁴¹ → Pu ²³⁷ 35 d	6.05	5.95		ion ch	74	A	ins evid
Cm ²⁴² → Pu ²³⁸ 162.5 d	6.213	6.110 6.118	73.7	spect ion ch	75 52	A	
Cm ²⁴³ → Pu ²³⁹ 35 y	6.151	5.985	6	spect	76	A	α-γ coin
Cm ²⁴⁴ → Pu ²⁴⁰ 19 y	5.895	5.798	75	spect	76	A	
Cm ²⁴⁵ → Pu ²⁴¹ 2×10^4 y	5.43	5.34		ion ch	77	B	ins evid
Bk ²⁴³ → Am ²³⁹ 4.5 h	6.83	6.72	30	ion ch	78	A	ins evid
Bk ²⁴⁵ → Am ²⁴¹ 4.95 d	6.44	6.33	23	ion ch	79	B	ins evid
Bk ²⁴⁹ → Am ²⁴⁵ ~1 y	5.5	5.4		ion ch	80, 81	A	ins evid
Cf ²⁴⁴ → Cm ²⁴⁰ 45 m	7.27	7.15		ion ch	82	B	e-e
Cf ²⁴⁶ → Cm ²⁴² 35.7 h	6.865	6.753	78	spect	83	A	
Cf ²⁴⁸ → Cm ²⁴⁴ 250 d	6.36	6.26		ion ch	84	B	e-e
Cf ²⁴⁹ → Cm ²⁴⁵ 500 y	6.10	6.00	10	ion ch	85	A	ins evid
Cf ²⁵⁰ → Cm ²⁴⁶ 11 y	6.14	6.05 6.03		ion ch ion ch	85 80	A	e-e
Cf ²⁵² → Cm ²⁴⁸ 2.1 y	6.23	6.15 6.12		ion ch ion ch	85 80	A	e-e
99 ²⁴⁷ → Bk ²⁴³ 7.3 m	7.4	7.3		ion ch	86	D	ins evid
99 ²⁵³ → Bk ²⁴⁹ 19.3 d	6.73	6.63 6.61		ion ch ion ch	85 87	A	ins evid
100 ²⁵⁴ → Cf ²⁵⁰ 3.3 h	7.31	7.22 7.17		ion ch ion ch	85 87	A	e-e
100 ²⁵⁵ → Cf ²⁵¹ 15 h	7.2	7.1		ion ch	85	D	ins evid

⁴¹ Meinke, Ghiorso, and Seaborg, Phys. Rev. **85**, 429 (1952).⁴² Orth, Ghiorso, and Seaborg (unpublished data, April, 1950).⁴³ Asaro, Slater, and Perlman (unpublished data, 1953).⁴⁴ Marie Curie and S. Rosenblum, Compt. rend. **196**, 1598 (1933). (See reference 8.)⁴⁵ Briggs (see reference 9) recalculated the energy from the original data (see reference 44) using later values for the reference energies.⁴⁶ E. K. Hyde and F. Asaro (unpublished data, 1952).⁴⁷ Hagemann, Katzin, Studier, Seaborg, and Ghiorso, Phys. Rev. **79**, 435 (1950).⁴⁸ F. F. Momoyer and E. K. Hyde (unpublished data, December, 1952).⁴⁹ Rosenblum, Valadares, and Perey, Compt. rend. **228**, 385 (1949).⁵⁰ Asaro, Stephens, and Perlman, Phys. Rev. **92**, 1495 (1953).⁵¹ Rosenblum, Valadares, and Vial, Compt. rend. **227**, 1088 (1948).⁵² H. G. Jackson and B. G. Harvey (private communication).⁵³ Philibert, Génin, and Vigneron, J. phys. et radium **15**, 16 (1954).⁵⁴ Meinke, Ghiorso, and Seaborg (unpublished data, July, 1948).⁵⁵ Hyde, Studier, and Bruehlman, reported in Argonne National Laboratory Classified Report ANL-4112 (March, 1948).⁵⁶ Rosenblum, Cotton, and Boussières, Compt. rend. **229**, 825 (1949).⁵⁷ W. W. T. Crane and I. Perlman (unpublished data, 1950).

(Footnotes concluded for Table I).

- ⁵⁸ Clark, Spencer-Palmer, and Woodward, British Declassified Report BR-522 (October, 1944).
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⁶⁰ Jaffey, Diamond, Hirsch, and Mech, Phys. Rev. **84**, 785 (1951).
⁶¹ Magnusson, Thompson, and Seaborg, Phys. Rev. **78**, 363 (1950).
⁶² James, Ghiorso, and Orth, Phys. Rev. **85**, 369 (1952).
⁶³ A. Ghiorso (unpublished data).
⁶⁴ D. A. Orth and K. Street, Jr. (unpublished data, 1951).
⁶⁵ James, Florin, Hopkins, and Ghiorso, National Nuclear Energy Series, Plutonium Project Record, Vol. 14B (McGraw-Hill Book Company, Inc., New York, 1949), p. 1604.
⁶⁶ F. Asaro and I. Perlman, Phys. Rev. **94**, 381 (1954).
⁶⁷ F. Asaro and I. Perlman, Phys. Rev. **88**, 828 (1952).
⁶⁸ Rosenblum, Valadares, and Goldschmidt, Compt. rend. **230**, 638 (1950).
⁶⁹ G. H. Higgins, Ph.D. thesis, University of California Radiation Laboratory, Unclassified Document UCRL-1796 (June, 1952).
⁷⁰ F. Asaro and I. Perlman, Phys. Rev. **93**, 1423 (1954).
⁷¹ Hummel, Asaro, and Perlman (unpublished data).
⁷² Street, Ghiorso, Orth, and Seaborg (unpublished data, October, 1948).
⁷³ G. H. Higgins and K. Street, Jr., Phys. Rev. **86**, 252 (1952).
⁷⁴ R. A. Glass, Ph.D. thesis, University of California Radiation Laboratory, Unclassified Document UCRL-2560 (April, 1954).
⁷⁵ Asaro, Reynolds, and Perlman, Phys. Rev. **87**, 277 (1952).
⁷⁶ Asaro, Thompson, and Perlman, Phys. Rev. **92**, 694 (1953).
⁷⁷ E. K. Hulet, Ph.D. thesis, University of California Radiation Laboratory, Unclassified Document UCRL-2283 (July, 1953).
⁷⁸ Thompson, Ghiorso, and Seaborg, Phys. Rev. **80**, 781 (1950).
⁷⁹ Hulet, Thompson, Ghiorso, and Street, Phys. Rev. **84**, 366 (1951).
⁸⁰ Diamond, Magnusson, Mech, Stevens, Friedman, Studier, Fields, and Huizenga, Phys. Rev. **94**, 1083 (1954).
⁸¹ A. Ghiorso *et al.* (private communication).
⁸² Thompson, Street, Ghiorso, and Seaborg, Phys. Rev. **80**, 790 (1950).
⁸³ Hummel, Chetham-Strode, Asaro, and Perlman (unpublished data).
⁸⁴ E. K. Hulet and S. G. Thompson, mentioned in reference 85.
⁸⁵ Ghiorso, Thompson, Higgins, Harvey, and Seaborg, Phys. Rev. **95**, 293 (1954).
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⁸⁷ Fields, Studier, Mech, Diamond, Friedman, Magnusson, and Huizenga, Phys. Rev. **94**, 209 (1954).
⁸⁸ In Rytz's work, the energies of the Bi^{212} alpha particles were measured relative to the main group of Po^{212} . We modified the energy of the most energetic Bi^{212} group to correspond to our adopted Q for Po^{212} .

obtain evidence on complex structure. Otherwise the entry indicates the intensity of the group believed to represent the ground-state transition. The designation "≈100" means that careful search has been made for other groups and either none has been found or that the intensities of lower-energy groups are low.

COLUMN 5

This column refers to the method of energy determination.

ion ch	ionization chamber coupled with some form of pulse-height analyzer.
range air (<i>or</i> mica)	range determination in air (<i>or</i> mica).
range emuls	range of alpha tracks in a photographic emulsion.
spect	magnetic or electrostatic spectrograph.

It is difficult to assign meaningful limits of error for most of the individual measurements so this has not been attempted. In general, those measurements made with a spectrograph may be in error by as much as 15 kev, but a large number are perhaps accurate to within 5 kev; those made by the ionization chamber method will vary in accuracy from 10 to 75 kev; and range measurements in photographic emulsions are even more difficult to assess.

COLUMN 6

References are given for the energy measurements selected.

COLUMN 7

These letter ratings give the estimated degree of certainty of the isotopic assignments according to the following code: (No designation according to the code was made for Bi^{209} because the mass number is not

uncertain, but the extremely low decay rate leaves some doubt that the few tracks observed were not spurious.)

- A Element and mass number certain;
- B Element certain and mass number probable;
- C Element probable and mass number certain or probable;
- D Element certain and mass number not well established;
- E Element probable and mass number not well established or unknown.

COLUMN 8

The comments in this column for the most part reinforce the decision on the decay energy.

ins evid	Insufficient evidence to know whether or not the alpha energy measured is that of the ground-state transition.
e-e	No direct evidence, but since the nucleus is of the even-even type it can be assumed that the measured energy is that of the ground-state transition.
α - γ coinc	This designation indicates that coincidences have been observed between alpha particles and gamma rays (or conversion electrons) which show some doubt that the highest-energy alpha group is the ground-state transition. Where the evidence is not sufficiently definite to deduce a decay energy based on anything other than the highest-energy alpha group, this is reflected by the values in Columns 2 and 3 differing only by the recoil energy. Where the evidence is sufficiently definite to deduce the decay energy, it will be found that Columns 2 and 3 differ by more than the recoil energy.