

Maps for electron clouds

Ubaldo Iriso* and Steve Peggs†

Brookhaven National Laboratory, Upton, New York 11973, USA

(Received 27 October 2004; published 10 February 2005)

The electron cloud effect has been studied by means of detailed simulation codes that typically track the particles' evolution under the influence of the corresponding electromagnetic forces and fields. In this paper we show that, for the RHIC case, the electron cloud can be treated from an abstract point of view as a bunch to bunch evolution using simple maps. Secondly, we show how this treatment yields a useful conclusion, which is otherwise difficult to obtain: for a fixed number of bunches and total beam current in RHIC, it is possible to determine the best way to distribute the bunch pattern around the ring to minimize the electron cloud formation. This application is an example of how maps become a useful tool for exploring the electron cloud evolution in parameter space.

DOI: 10.1103/PhysRevSTAB.8.024403

PACS numbers: 29.20.Dh

I. MOTIVATION

Electric fields present in many vacuum systems may accelerate electrons (produced by field emission, photo-emission, residual gas ionization, etc.) towards the wall chamber surface. If the bombarding electrons acquire enough energy, they produce secondary electrons, which in turn may be accelerated if the electric field normal to the surface is at the correct phase. These electrons may bombard another surface and again emit secondary electrons. This bouncing back and forth between surfaces is the electron *multipacting* effect. This name, derived from “resonance of multiple electron impact,” was first described by Farnsworth in 1934 [1]. If the number of emitted electrons per impinging electron, given by the secondary emission yield (SEY) of the wall surface, is greater than unity, the electron density inside the pipe increases (initially) exponentially, creating a so-called electron cloud (EC). A positively charged beam in an accelerator can produce an EC formation. This EC can eventually lead to the development of vacuum breakdown and other system failures. The proton storage ring of the INP Novosibirsk in 1965 [2] and the ISR at CERN in 1972 [3] are among the first machines suffering from electron clouds. A thorough review of EC effects in accelerators can be seen in Ref. [4].

In the 1990s, electron cloud driven instabilities became an issue in different machines [5]. Several computer simulation codes were (and still are being) developed and compared with experimental observations to study the effect. A comparison among the different codes can be seen in [6]. Typically, these codes work either by particle-in-cell methods (like CLOUDLAND) or by tracking electrons grouped into macroparticles, where each macroparticle comprises up to a maximum of around 10^5 electrons (like ECLOUD or CSEC [7]). When a macroparticle produces more electrons, its total Coulomb charge is increased. At

every time step, these detailed codes compute the necessary physical forces and fields influencing the motion of the macroparticles. If EC formation takes place, the codes track about 10^{10} electrons per meter of beam pipe (depending on the parameters). Hence, these codes use a large amount of CPU time: a complete EC simulation, depending on the input code parameters, can last from around 1 h to some days. In the cases studied here (for the parameters seen in Table I), a single simulation lasts about 1 h.

In the following, we consider that, for given beam pipe characteristics (SEY, chamber dimensions, etc.), the electron density after bunch m passes by (referred to as ρ_{m+1}) is a function only of the interaction between the bunch and the electron density before bunch m passed by (referred to as ρ_m). This is expressed by means of an iterative formalism. For instance, in a parabolic map

$$\rho_{m+1} = a\rho_m + b\rho_m^2, \quad (1)$$

where the parameters a and b are functions of beam parameters such as bunch intensity N , bunch spacing s_b , rms bunch length σ_z , and rms bunch transverse size σ_t . Ultimately, a and b are functions of the beam pipe characteristics as well: maximum SEY δ_{\max} ; electron energy at which SEY is maximum, E_{\max} ; reflectivity at zero electron energy, δ_0 ; beam pipe dimensions, etc. Therefore, the coefficients a and b summarize the EC dependence on the physical parameters. This parabolic map is sometimes called the “logistic” difference equation [8], since by introducing the dimensionless variable $X \equiv b\rho/a$, and for $a > 0$, $b < 0$, Eq. (1) can be expressed as

$$X_{m+1} = aX_m(1 - X_m), \quad (2)$$

which reproduces the logistic map formalism [8] with all its richness. For small ρ , Eq. (1) reflects the exponential growth with the bunch passage [9],

$$\rho_m \approx \rho_0 e^{(a-1)m}, \quad (3)$$

where it is clear that the electron cloud takes place for

*Electronic address: ubaldo@bnl.gov

†Electronic address: peggs@bnl.gov

TABLE I. Input parameters for electron cloud simulations testing the map hypothesis. In all cases, the simulations using CSEC and ELOUD are performed for protons bunches.

Parameter	Symbol	Unit	CSEC Value	ELOUD Value
Bunch spacing	s_b	ns	108	108
Number of bunches	M	...	60	60
rms bunch radius	σ_r	mm	2.4	2.4
Full bunch length	σ_z	ns	18	21
Protons per bunch	N	10^{10}	8–20	8–20
Revolution time	t_{rev}	μs	12.82	12.79
Beam energy	E	GeV	27.7	11.46
Beam pipe diameter	d	mm	120	120
Reflectivity at zero energy	δ_0	...	0.6	1.0
Reflectivity at infinite energy	P_∞	...	0.2	...
Rediffusion probability	P_{rd}	...	0.5	...
Reflection energy	E_{rf}	eV	60	60
Maximum SEY	δ_{max}	...	2.3	2.3
Energy for maximum SEY	E_{max}	eV	310	310
Energy for secondary electron	E_{sec}	eV	8.9	7.0
Energy width for secondary electron	σ_{sec}	eV	4.5	5.5
initial e density	ρ_{ce}	pC/m	0.2	—
Electrons generated per bunch	35 000	—
Electron generation radius	...	mm	60	—
Number of slices per bunch	60	100
Number of slices per interbunch	840	100
Initial number of macroparticles	25	—
Maximum number of macroparticles	10^5	$\approx 10^5$

$a > 1$, and otherwise the cloud collapses. Eventually, this unlimited growth is stopped by the space charge effects created by the electrons themselves.

From Eq. (1) in a parabolic mode, the saturated electron cloud density ρ_{sat} is determined as a function of the bunch intensity N simply by

$$\rho_{\text{sat}} = 0; \quad N < N_C \quad \text{or} \quad a < 1, \quad (4)$$

$$\rho_{\text{sat}} = \frac{a-1}{-b}; \quad N > N_C \quad \text{or} \quad a > 1, \quad (5)$$

where N_C marks the bunch intensity threshold for the electron cloud. Equation (1) shows a phase transition from electron cloud “off” to “on.” If a and b increase smoothly with N , the phase transition is second order. However, RHIC data show both first and second order electron cloud phase transitions [10], which is not yet well understood. The parabolic model of Eq. (1) is a mathematical tool illustrated to express our goal: to simplify the EC problem by using only a small number of mathematical parameters. In this example, these parameters are just a and b .

If the electron cloud evolution can be described using a simple map $\rho_{m+1} = f(\rho_m)$, this frees up the detailed simulation codes and enhances physical intuition through the use of simple maths. Thus, we first need to evaluate

whether or not it is possible to follow the electron density in a bunch to bunch evolution (Sec. II), and secondly we look for a suitable function to follow this evolution (Sec. III). Finally, one application of the map modeling is presented (Sec. IV).

II. BUNCH TO BUNCH EVOLUTION

A typical time evolution of the electron density is shown in Fig. 1. This evolution corresponds to a simulation where 60 bunches of 1.4×10^{11} protons each, spaced 108 ns apart, are injected into the RHIC ring. In this case the code used is CSEC [11]. The red line shows CSEC output, while the gray circles mark the average electron density between the passage of two bunches. The presence of a bunch is indicated by the black bars at the bottom of the figure, the light blue bars mark an *empty* bunch. The electron density per beam pipe meter as a function of time ρ grows exponentially until the space charge effects produce a saturation level [9]. Once the saturation level is reached the average electron density does not change significantly. In the bunch to bunch evolution, the time step is one bunch passage. Figure 1 shows that sampling the evolution on a bunch-to-bunch basis is sufficient for retaining information about the buildup and decay times, although the details of the behavior of the electron density oscillation between two bunches is lost.

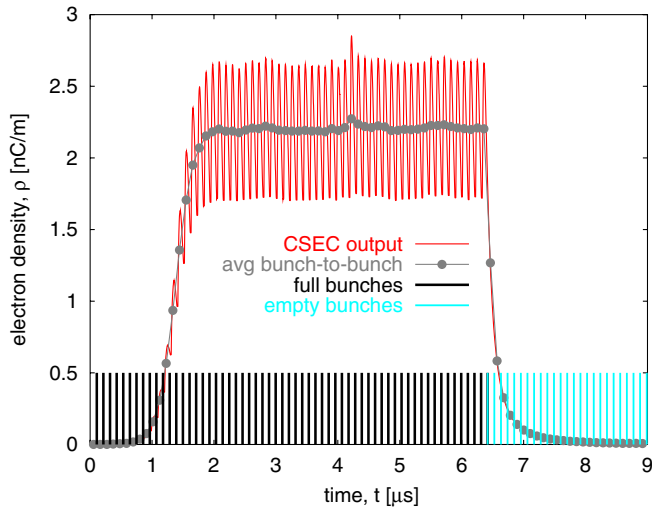


FIG. 1. (Color) Time evolution of the electron density (red line) computed with CSEC during $9 \mu\text{s}$. The RHIC revolution period is $12.82 \mu\text{s}$. This case corresponds to the injection of 60 successive bunches with a bunch spacing of 108 ns and a bunch intensity of $N = 1.4 \times 10^{11}$ protons (marked with black bars), followed by 60 empty bunches (marked with light blue bars). The gray circles mark the average electron density between two consecutive bunches.

Do existing computer simulations confirm that the electron cloud evolution can be represented by maps? For this purpose, we test this hypothesis using two codes: CSEC and ECLLOUD [12], focusing the studies on the RHIC case. Table I shows the physical parameters used for these simulations. Besides the beam characteristics, the SEY behavior as a function of the impinging electron energy is a key ingredient in the electron cloud development. All simulation codes strongly depend on the model used for the SEY behavior [13]. CSEC uses the model described in [14], where one can find detailed explanations of the parameters

named in the second part of Table I. On the other hand, ECLLOUD uses the model described in [13]. Table I compares only the most “common” surface physics parameters. Also, while ECLLOUD uses a Gaussian distribution for the emitted secondary electrons, CSEC uses a Lorentzian one (parameter σ_{sec} in Table I).

With 108 ns bunch spacing and the RHIC revolution period, one can inject up to 120 bunches (not counting the limitations given by the abort gap kickers, which decrease this number to 110). Nevertheless, for the purpose of this study we are interested in the buildup and decay evolution of the electron density, i.e., until the saturation level is reached. Therefore, we performed the simulations for bunch trains of 60 consecutive bunches (until saturation is reached) to minimize CPU time.

The bunch to bunch evolution of the electron cloud density is followed for different bunch intensities N ranging from 8×10^{10} to 2×10^{11} protons, in steps of $\Delta N = 2 \times 10^{10}$ protons using the parameters listed in Table I. Figure 2 shows how the electron density after bunch m passes by, ρ_{m+1} , behaves as a function of the previous electron density, ρ_m , for different bunch intensities N . The points in Fig. 2 show the average electron cloud density between two bunches using results from CSEC (Fig. 2, left) and ECLLOUD (Fig. 2, right). The lines correspond to cubic fits with no constant term (see below).

Figure 2 is explained as follows: starting with a small seed of electrons, electron density $\rho_0 \approx 0$ nC/m, the density grows and reaches the saturation line ($\rho_{m+1} = \rho_m$, red trace) when the space charge effects due to the electrons of the cloud itself limit further growth. In this situation, all the points (corresponding to the passage of full bunches) are in the same spot, at the 45° line. This particular line, showing $\rho_{m+1} = \rho_m$, is also called the *identity map*.

Electron cloud decay is described as the succession of bunches with a null bunch intensity, $N = 0$. Except for the

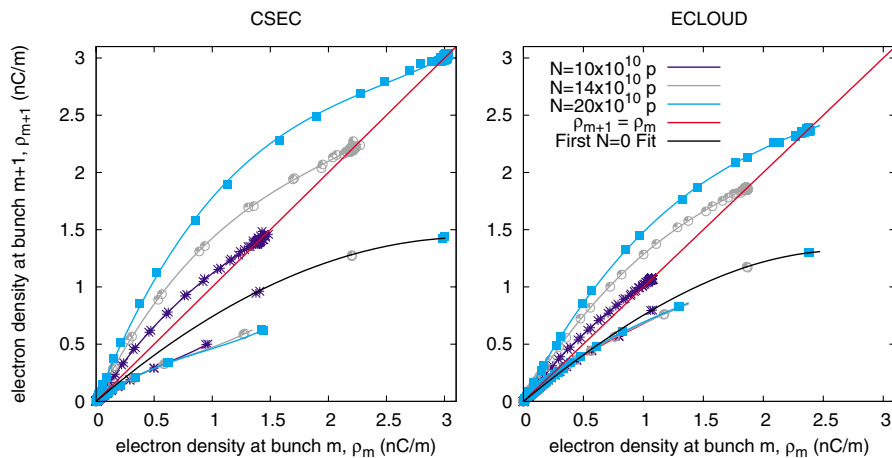


FIG. 2. (Color) Average electron density after bunch passage m , ρ_{m+1} , as a function of the electron density before bunch m passed by, ρ_m , for different bunch intensities N . The left plot shows the CSEC output, while the points on the right-hand side come from ECLLOUD simulation. In both cases, the lines correspond to cubic fits applied to the average bunch to bunch points.

point corresponding to the electron cloud density after the first empty bunch, the electron density follows a universal curve independent of the initial value of the saturated electron density. The “first empty” bunch points after the identity map (from different saturation values, ρ_{sat}) lie off the universal curve on the “first $N = 0$ ” or “first empty bunch” curve. This is arguably with the space charge effects during saturation. In other words, it takes two bunches to jump from a curve $N \neq 0$ to the decay ($N = 0$ curve).

III. MAP CANDIDATES

Different forms for the parametric maps include the above mentioned “parabolic” map [Eq. (1)], the “cubic” map (with no independent term),

$$\rho_{m+1} = a\rho_m + b\rho_m^2 + c\rho_m^3, \quad (6)$$

and an “asymptotic” map,

$$\rho_{m+1} = \frac{a\rho_m}{1 + b\rho_m}, \quad (7)$$

which is also known as the “Hassell” model in density-dependent population dynamics [15,16].

Figure 3 shows the results for the χ^2 coefficient for each fit to the data in Fig. 2, and for each bunch intensity N , for both CSEC (left plot) and ECLLOUD (right). Since the smallest χ^2 value corresponds to the cubic map, we continue the analysis using cubic maps, stating clearly that this map is valid only for electron densities within the ranges used here.

Thus, for the parameters shown in Table I, the electron density buildup for a given bunch intensity is determined by a 3-vector $\vec{A}(N) = (a, b, c)$, while decay is described by two 3-vectors, one corresponding to the first empty bunch and a second vector for the rest of them.

Figure 4 shows how the buildup coefficients (a, b, c) evolve as a function of the bunch intensity N for both

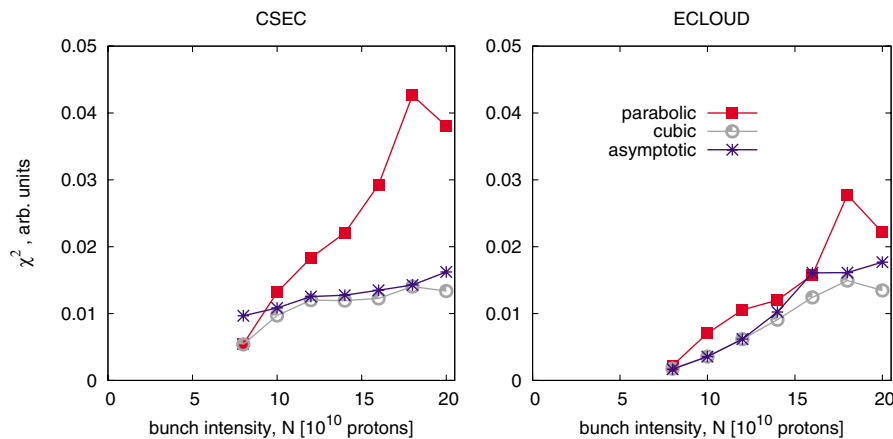


FIG. 3. (Color) Evolution of the χ^2 coefficient as a function of bunch intensity N for the different maps tested: parabolic, cubic, and asymptotic for CSEC (left plot) and ECLLOUD (right). Smallest values of χ^2 indicate a better fit quality.

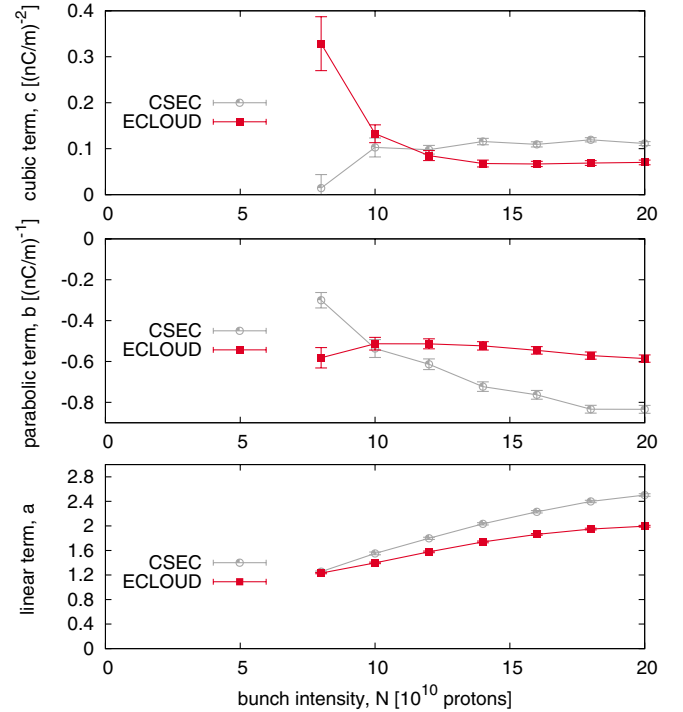


FIG. 4. (Color) Evolution of the linear a , parabolic b , and cubic c terms determining the electron cloud buildup as a function of the bunch intensity N , for both ECLLOUD (red squares) and CSEC (gray points).

CSEC (gray points) and ECLLOUD (red squares). Both codes give a similar phase transition threshold N_C around 7×10^{10} protons, when $a(N_C) = 1$. The linear coefficient a becomes larger than 1 when $N > N_C$, and increases linearly in first approximation. In all cases (different N), and using both codes, the quadratic coefficient b is negative. This gives concavity to the electron cloud density evolution in the space (ρ_m, ρ_{m+1}) and ensures a positive saturation value [see Eq. (5)]. The b coefficient decreases (increases in absolute value) for CSEC results, but using

TABLE II. Comparison of bunch patterns tested in RHIC at injection.

Parameter	Unit	Reference case	Fill no. 1	Fill no. 2	Fill no. 3
Bunch pattern	...	(6, 1, 0)	(3, 16, 4)	(3, 12, 8)	(3, 14, 6)
No. of bunches	...	56	41	69	78
Average proton/bunch N	10^{11}	1.0	1.1	1.0	0.9
Total intensity	10^{11}	56.0	44.3	68.1	70.2
Full bunch length	ns	...	16.5	17.6	34.2
Pressure rise	Yes	Yes	No
Luminosity scaling factor	...	1.00	0.88	1.23	1.13

A. Simulations for different bunch patterns

The simulation code MEC uses a cubic interpolation map to follow the bunch to bunch evolution of the electron cloud density. Simulations of different bunch patterns carried out with CSEC and reported in [17] are compared using MEC in this section. The parameters used by CSEC are reported in Table I, but in this case we fix the bunch intensity at $N = N_0 = 8 \times 10^{10}$ protons.

The use of MEC is divided into four cases, depending on the intensity of the bunches m and $m - 1$ passing by.

(i) First “full” bunch, which denotes a full bunch after an empty one, i.e., $N_m = N_0$ and $N_{m-1} = 0$, with cubic map coefficients represented by the vector $\vec{A}_{10} = (a_{10}, b_{10}, c_{10})$.

(ii) Full bunches, denoting the passage of a bunch with $N_m = N_0$ protons after another full bunch, $N_{m-1} = N_0$. The cubic map coefficients for this case are denoted by $\vec{A}_{11} = (a_{11}, b_{11}, c_{11})$.

(iii) First empty bunch, an empty bunch after a populated bunch, i.e., $N_m = 0$ and $N_{m-1} = N_0$. The corresponding cubic map coefficients are represented by $\vec{A}_{01} = (a_{01}, b_{01}, c_{01})$.

(iv) Empty bunches, succession of bunches with intensity $N_m = 0$ and $N_{m-1} = 0$. The corresponding cubic map coefficients are denoted by $\vec{A}_{00} = (a_{00}, b_{00}, c_{00})$.

The need for this subdivision requires analysis of two figures: in Fig. 2 one can see that the first $N_m = 0$ is out of the evolution of the decay curve, i.e., the curve corresponding to ghost bunches. Figure 6 justifies the case for the first $N_m = N_0$ curve. Figure 6 shows that the transition from empty to full also requires two bunches, in the same way that the transition from full bunch to empty bunch is done in two bunches.

One obtains successful results when comparing the bunch to bunch evolution using CSEC and MEC; see Fig. 7 with the different bunch patterns. Table III compares the maximum and average values for the linear electron cloud density at the last turn using CSEC and MEC. The largest difference for the maximum density is about 15% [corresponding to the case (3, 2, 0)(6, 4, 0)]; while for the average density the maximum difference is about 17%, corresponding to the case (3, 23, 17). While CSEC uses about 1 h CPU

time for each case, MEC is obviously much faster and uses only ≈ 1 ms, which represents a speed up of 7 orders of magnitude.

B. Linear approximation

Four sets of polynomial coefficients, $\vec{A}_{11}(N)$, $\vec{A}_{01}(N)$, $\vec{A}_{00}(N)$, and $\vec{A}_{10}(N)$, are required to follow the bunch to bunch evolution of the electron cloud density in MEC. Figure 6 suggests that for small electron densities the bunch to bunch evolution can be considered as linear in the (ρ_m, ρ_{m+1}) space. That is, if there is a total number of M bunches in a ring with a “bunch harmonic” number of H , the linearization of the problem gives a one turn map that is simply

$$\rho_{m+H} \approx F(N)\rho_m, \quad (8)$$

where the “one turn factor” is

$$F \equiv (a_{10}a_{01})^i a_{11}^{M-i} a_{00}^{H-M-i}, \quad (9)$$

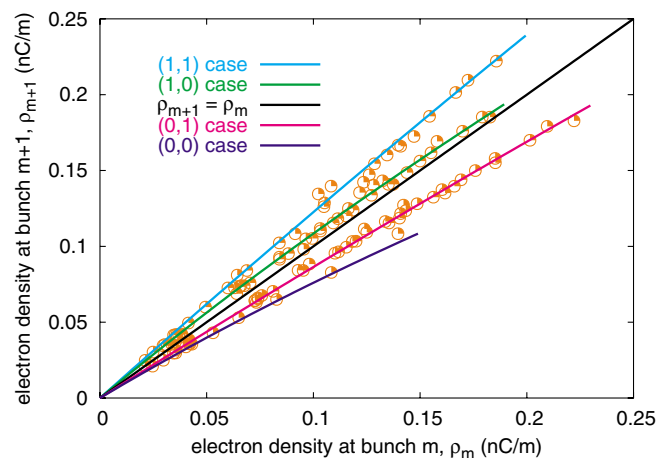


FIG. 6. (Color) Electron cloud density in (ρ_m, ρ_{m+1}) space for the bunch pattern (3, 4, 0)(6, 8, 0). The plot shows that four different behaviors are required: the case (1, 1) refers to full bunches preceded by another full bunch; the case (1, 0) refers to full bunches preceded by an empty bunch; the case (0, 1) to empty bunches preceded by a full bunch; while the case (0, 0) denotes an empty bunch preceded by another empty bunch.

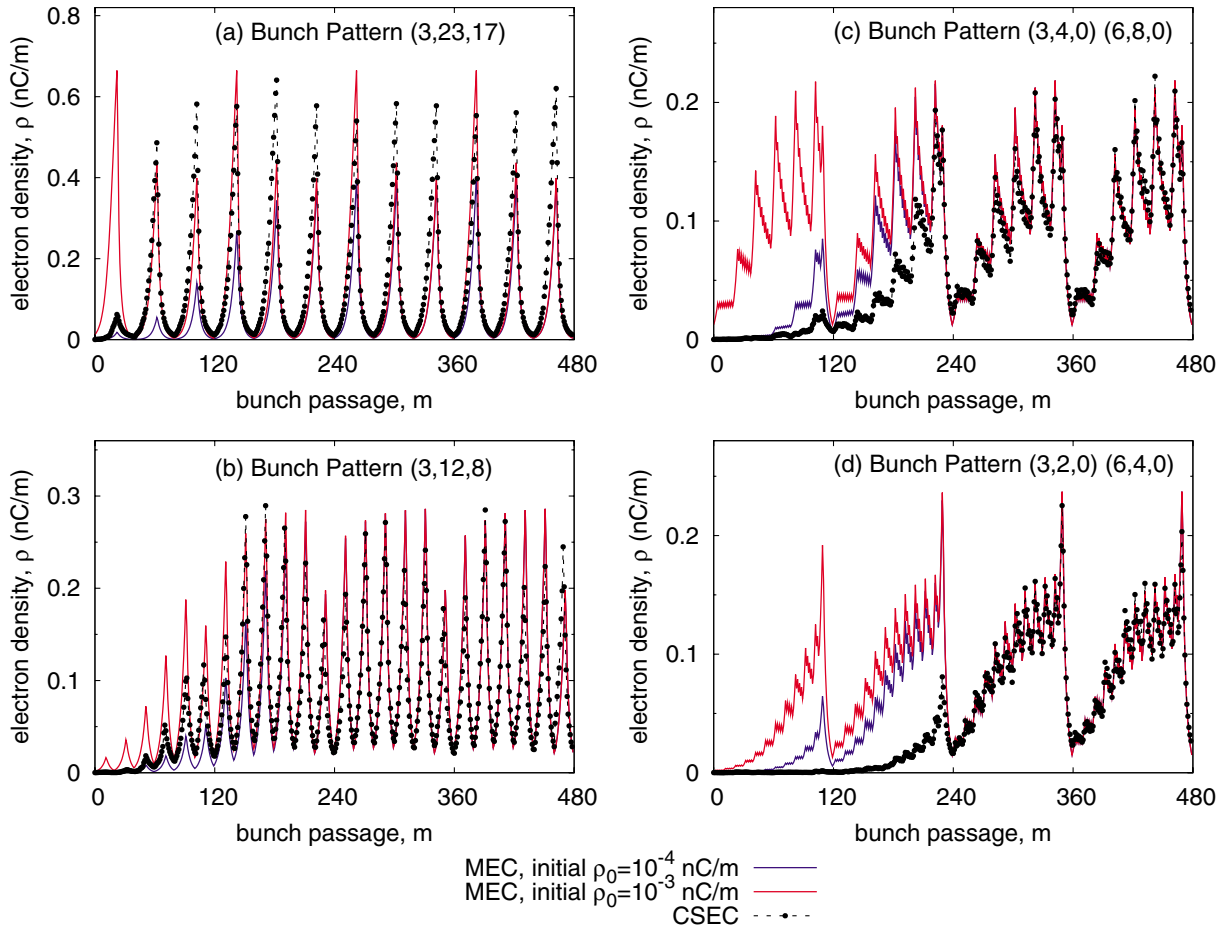


FIG. 7. (Color) Electron cloud density evolution for four of the different bunch patterns in Table III using CSEC (dashed black trace with dots) and MEC with two different initial electron densities: $\rho_0 = 10^{-4}$ nC/m (blue line) and $\rho_0 = 10^{-3}$ nC/m (red line). No matter the initial electron density, MEC results agree for the last turn (from bunch passage 360 to 480) within an error range of $\approx 15\%$ for all bunch patterns. Plots for bunch patterns (c) and (d) (on the right-hand side) have the same vertical scale while (a) and (b) differ.

and i is the number of transitions from full to empty (and empty to full) bunches. In general, the minimum possible number of transitions is $i = 1$ (if all the bunches are clumped together), and the maximum number of transitions is the smaller of M and $H - M$ (when the bunches are

spread as sparsely as possible). The special case $i = 0$ applies when there is no abort gap, $M = H$.

It is clear that if $F > 1$ then the electron cloud density increases (to some saturated value), while if $F < 1$ then the cloud disappears. When the one turn factor is rewritten as

TABLE III. Maximum ρ_{\max} and average ρ_{avg} bunch to bunch values of the linear electron density simulated with CSEC and MEC for different bunch patterns. The results agree within about 15%.

Parameter	Unit	Case no. 1	Case no. 2	Case no. 3	Case no. 4	Case no. 5
Bunch pattern	...	(3, 68, 52)	(3, 23, 17)	(3, 12, 8)	(3, 4, 0)(6, 8, 0)	(3, 2, 0)(6, 4, 0)
Number of bunches	...	68	68	68	68	68
Protons per bunch	10^{10}	8.0	8.0	8.0	8.0	8.0
ρ_{\max} using CSEC	nC/m	0.8991	0.6203	0.2849	0.2221	0.2033
ρ_{\max} using MEC	nC/m	0.9302	0.6645	0.2861	0.2184	0.2370
ρ_{avg} using CSEC	nC/m	0.3023	0.1433	0.0981	0.1006	0.0922
ρ_{avg} using MEC	nC/m	0.3216	0.1156	0.1045	0.0992	0.0924
Figure	—	—	7(a)	7(b)	7(c)	7(d)

$$F = \left(\frac{a_{10}a_{01}}{a_{11}a_{00}}\right)^i \left(\frac{a_{11}}{a_{00}}\right)^M a_{00}^H \quad (10)$$

it is clear that, for given M , H , and N , the smallest (largest) value of F occurs for the largest (smallest) allowed value of i if

$$\left(\frac{a_{10}a_{01}}{a_{11}a_{00}}\right) < 1 \quad (11)$$

and vice versa. Since Eq. (11) is valid for RHIC parameters, the most sparse distribution of a fixed number of fixed population bunches is the most stable against electron cloud growth.

Thus, from the mapping approach and using a linearized approximation we have demonstrated that the most sparse distribution of bunches in RHIC minimizes the detrimental effects of the multibunch electron cloud effects. This is not a big surprise if we consider the possibility of evenly distributed bunches, i.e., the same bunch spacing between all bunches. In this case, the most sparse distribution of bunches is equivalent to using a larger bunch spacing between them. However, Eq. (10) demonstrates that this is also valid for unevenly spaced bunches along a bunch train. This “rule of thumb” has been used in the RHIC operation when deciding the bunch pattern to inject [20,21].

V. SUMMARY

Multibunch electron cloud buildup at RHIC is modeled by simple maps. A third order polynomial map, written as $\vec{A} = (a, b, c)$, optimally reproduces the bunch to bunch evolution. For a given vacuum chamber, these coefficients are a function of the beam parameters. These parameters, (a, b, c) , as a function of the bunch intensity N are empirically determined from electron cloud simulations codes, like CSEC or ECLLOUD, as a function of the bunch intensity N .

When jumping back and forth from full to empty bunches, a memory of two bunches is found to be necessary. Therefore a complete algorithm requires four vectors: \vec{A}_{11} , \vec{A}_{10} , \vec{A}_{00} , and \vec{A}_{01} . A simulation program, MEC, uses these vectors to reproduce (within about 15%) the evolution of the electron density in a bunch to bunch approximation. The CPU time used in this case is 7 orders of magnitude smaller than that used by the conventional electron simulation codes (CSEC or ECLLOUD).

The importance of this analysis lies not only in the acceptable reproducibility of the results using MEC, but also in the ability to abstract the way to tackle electron clouds. This helps to deliver conclusions that would otherwise be difficult to obtain. For instance, using the linearized maps, actual values for the vectors analytically demonstrate that, for the straight sections of RHIC, the most sparse distribution of bunches is the most stable

against electron cloud formation, even when they cannot be evenly spaced.

In the future, it is desirable to explore how the polynomial coefficients vary as a function of the physical parameters influencing the electron cloud (SEY, chamber dimensions, bunch spacing, bunch charge, etc.) in order to obtain a better understanding of the problem. Application of maps to other machines, like B factories or the LHC, is also necessary to study the universality of map formalism.

ACKNOWLEDGMENTS

We are indebted to Wolfram Fischer for the idea of optimizing the bunch distribution around the RHIC circumference. Mike Blaskiewicz and Giovanni Rumolo substantially helped in the use of CSEC and ECLLOUD, respectively. We are very grateful for discussions with, and support from, Angelika Drees, H.C. Hseuh, Nick Luciano, Rogelio Tomás, Dejan Trbojevic, Lanfa Wang, Keith Zeno, and S. Y. Zhang.

-
- [1] P. Farnsworth, J. Franklin Inst. **218**, 411 (1934).
 - [2] G.I. Budker, G.I. Dimov, and V.G. Dudnikov, in *Proceedings of the International Symposium on Electron and Positron Storage Rings, Saclay, 1966* (Universitaires De France, Orsay, 1966), p. VIII-6-1.
 - [3] O. Gröbner, in *Proceedings of the 10th International Conference on High Energy Accelerators, Serpukhov, 1977* (USSR Academy of Science, Moscow, 1977), p. 277.
 - [4] F. Zimmermann, Phys. Rev. ST Accel. Beams **7**, 124801 (2004).
 - [5] *Proceedings of ECLLOUD'02, Mini-Workshop on Electron-Cloud Simulations for Proton and Positron Beams, Geneva, 2002*, edited by G. Rumolo and F. Zimmermann (CERN Report No. CERN-2002-001, 2002); see also <http://wwwslap.cern.ch/collective/ecloud02>
 - [6] E. Benedetto, F. Ruggiero, D. Schulte, F. Zimmermann, M. Blaskiewicz, L. Wang, G. Bellodi, G. Rumolo, K. Ohmi, S.-S. Win, M. Furman, Y. Cai, M. Pivi, V. Decyk, W. Mori, A. F. Ghalam, and T. Katsouleas, in *Proceedings of the European Particle Accelerator Conference (EPAC'04), Lucerne, 2004*, p. 2502, <http://accelconf.web.cern.ch/accelconf/e04/default.htm>
 - [7] L. Wang, G. Rumolo, and M. Blaskiewicz (private communication).
 - [8] R. M. May, Nature (London) **261**, 459 (1976).
 - [9] G. Stupakov, CERN-LHC-Project-Report-141, 1997.
 - [10] U. Iriso and S. Peggs, BNL Internal Note No. C-AD/AP/147, 2004.
 - [11] W. Fischer, M. Blaskiewicz, M. Brennan, and T. Satogata, Phys. Rev. ST Accel. Beams **5**, 124401 (2002).
 - [12] G. Rumolo and F. Zimmermann, CERN Report No. SL-Note-2002-016 AP, 2002.
 - [13] R. Cimino, I. Collins, M. Furman, M. Pivi, F. Ruggiero, G. Rumolo, and F. Zimmermann, Phys. Rev. Lett. **93**, 014801 (2004).

- [14] M. A. Furman and M. Pivi, *Phys. Rev. ST Accel. Beams* **5**, 124404 (2002).
- [15] M. P. Hassell, J. N. Lawton, and R. M. May, *J. Anim. Ecol.* **45**, 471 (1976).
- [16] J. Bascompte and R. Sole, *J. Anim. Ecol.* **63**, 256 (1994).
- [17] W. Fischer and U. Iriso, BNL Internal Note No. C-AD/AP/118, 2003.
- [18] U. Iriso, M. Blaskiewicz, A. Drees, W. Fischer, S. Peggs, and D. Trbojevic, in *Proceedings of the Particle Accelerator Conference (PAC'03), Portland, OR, 2003* (IEEE, Piscataway, NJ, 2003), p. 797.
- [19] S. Y. Zhang, M. Bai, M. Blaskiewicz, P. Cameron, A. Drees, W. Fischer, D. Gassner, J. Gullotta, P. He, H. Hseuh, H. Huang, U. Iriso, R. C. Lee, W. W. MacKay, B. Oerter, V. Ponnaiyan, V. Ptitsyn, T. Roser, T. Satogata, L. Smart, D. Trbojevic, and K. Zeno, in *Proceedings of the Particle Accelerator Conference (PAC'03), Portland, OR, 2003* (Ref. [18]), p. 54.
- [20] W. Fischer, L. Ahrens, J. Alessi, M. Bai, D. Barton, J. Beebe-Wang, M. Blaskiewicz, J. M. Brennan, D. Bruno, J. Butler, R. Calaga, P. Cameron, R. Connolly, T. D'Ottavio, J. DeLong, K. A. Drees, W. Fu, G. Ganetis, J. Glenn, T. Hayes, P. He, H.-C. Hseuh, H. Huang, P. Ingrassia, U. Iriso, R. Lee, Y. Luo, W. W. MacKay, G. Marr, A. Marusic, R. Michnoff, C. Montag, J. Morris, T. Nicoletti, B. Oerter, C. Pearson, S. Peggs, A. Pendzick, F. C. Pilat, V. Ptitsyn, T. Roser, J. Sandberg, T. Satogata, C. Schultheiss, A. Sidi-Yekhlef, L. Smart, S. Tepikian, R. Tomas, D. Trbojevic, N. Tsoupas, J. Tuozzolo, J. Van Zeijts, K. Vetter, K. Yip, A. Zaltsman, S. Y. Zhang, and W. Zhang, in *Proceedings of the European Particle Accelerator Conference (EPAC'04), Lucerne, 2004* (Ref. [6]), p. 917.
- [21] W. Fischer and U. Iriso, in *Proceedings of the European Particle Accelerator Conference (EPAC'04), Lucerne, 2004* (Ref. [6]), p. 914.