# Superstatistical approach to air pollution statistics

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Air pollution by nitrogen oxides  $(NO_x)$  is a major concern in large cities as it has severe adverse health effects. However, the statistical properties of air pollutants are not fully understood. Here, we use methods borrowed from nonequilibrium statistical mechanics to construct suitable superstatistical models for air pollution statistics. In particular, we analyze time series of nitritic oxide (NO) and nitrogen dioxide (NO<sub>2</sub>) concentrations recorded at several locations throughout Greater London. We find that the probability distributions of concentrations have heavy tails and that the dynamics is well described by  $\chi^2$  superstatistics for NO and inverse- $\chi^2$  superstatistics for NO<sub>2</sub>. Our results can be used to give precise risk estimates of high-pollution situations and pave the way to mitigation strategies.

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### I. INTRODUCTION

Complex driven nonequilibrium systems with timescale separation are often well described by superstatistical methods, i.e., by mixing several dynamics on distinct timescales and constructing effective statistical mechanics models out of this mixing [1,2]. The intensive parameter  $\beta$  that fluctuates in a superstatistical way can be an inverse temperature of the system, a fluctuating diffusion constant, or simply a local variance parameter in a given time series generated by the complex system under consideration. This formalism is in particular relevant for heterogeneous spatiotemporally varying systems and has been successfully applied to many areas of physics and beyond, most notably Lagrangian turbulence [3], defect turbulence [4], wind velocity fluctuations [5,6], share price dynamics [7], diffusion of complex biomolecules [8], frequency fluctuations in power grids [9], rainfall statistics [10,11], and many more. Here, we apply a superstatistical analysis to an important topic of high relevance, namely, the spatiotemporally dynamics of air pollution in big cities. Our main example of pollutants considered in the following are NO and NO<sub>2</sub>, but the method can be similarly applied to other substances.

NO and NO<sub>2</sub> are examples of nitrogen oxides  $(NO_x)$ , which are gaseous air pollutants that are primarily discharged through combustion [12]. In urban areas, the main cause of this is through automobile emissions [13]. These chemicals have been shown to aggravate asthma and other respiratory symptoms and increase mortality. Further, short-term exposure to NO<sub>2</sub> has been shown to correlate with ischemic and

substance within Earth's ecosystem, for example, regulating the development in plants and acting as a signaling hormone in certain processes in animals, it can also cause severe environmental damage in high concentrations [16–18]. NO may form acids, e.g., by chemically mixing with water (H<sub>2</sub>O) to form nitric acid (HNO<sub>3</sub>), or reacting further to NO<sub>2</sub> [19] it can cause acid rain [15]. Though NO<sub>x</sub> are not thought to be carcinogenic, they indicate the presence of other harmful air pollutants [13]. The data analyzed here were taken from the publicly

hemorrhagic strokes [12,14,15]. While NO is a very important

available London Air Quality Network (LAQN) website [20]. The site provides readings of pollutant levels at different locations throughout Greater London at different time intervals, with the 15-min interval data analyzed here. Inspecting the measured concentration time series for both NO and  $NO_2$ reveals that both pollutants' concentrations vary on multiple timescales (see Fig. 1).

## **II. ANALYSIS OF NO DATA**

Concentrating on the case of NO for now, we notice that the probability distribution of the measured concentration exhibits power-law tails (see Fig. 2). The distribution is well fitted by a q-exponential [21] of the form

$$p(u) = (2 - q)\lambda_q [1 + (q - 1)\lambda_q u]^{\frac{1}{1 - q}},$$
(1)

where *u* are the concentration levels  $(\mu g/m^3)$  of the pollutant, and the *q* and  $\lambda_q$  values are parameters. *q* can be regarded as an entropic index [21–23] and it is related to  $\lambda_q$  and the mean  $\mu$  by

$$\lambda_q = \frac{1}{\mu(3 - 2q)} \tag{2}$$

(see Appendixes A and B for more details).

The basic idea of the superstatistical modeling approach is that a given time series, generated by a complex driven nonequilibrium system, obeys a stochastic differential equation (SDE) where the parameters of the SDE change randomly

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FIG. 1. The concentration levels vary on several distinct timescales. (a) We display the concentration of NO, recorded at the Sir John Cass School location (London), measured in 15-min intervals. The inset reveals large variations within 1 month. (b) We plot the concentration of NO<sub>2</sub>, recorded at the North Street location (London), measured in 15-min intervals. The inset again gives the concentrations for 1 month. For both pollutants we note seasonal changes, weekly changes, and intermittent fluctuations.

as well, but on a much longer timescale [24]. In our case, the concentration time series, while being *q*-exponentially distributed as a whole, appears to be nonstationary, in the sense that it can be divided into shorter time slices, of length *T*, that each have locally an exponential density with different relaxation constant  $\lambda_e$  each (see Fig. 3). These changes in the pollutants statistics make sense due to the changing environment, e.g., because of changing weather conditions or traffic flow varying from week to week and also across seasons. Observing simple (exponential) statistics locally, while noting heavy tails in the aggregated statistics, clearly suggests a superstatistical description. One of the simplest local models is based on an exponential probability distribution of the form

$$p(u|\beta) = \beta \exp(-\beta u), \tag{3}$$



FIG. 2. Pollution statistics is not exponentially distributed, but rather *q*-exponentially distributed. We display the histogram of the NO data, together with the best exponential fit (green) and a *q*exponential fit (red). Also displayed are the value of the exponential parameter  $\lambda_e$  and the *q* and  $\lambda_q$  parameters for the *q*-exponential distribution.

where  $\beta = \lambda_e$  is a local relaxation parameter that fluctuates on the larger superstatistical timescale.

We determine this large timescale by computing the average local kurtosis  $\kappa$  of a cell of length  $\Delta t$ , similarly as done in Ref. [2] for locally Gaussian distributions. We use a local average kurtosis defined as

$$\kappa(\Delta t) = \frac{1}{t_{\max} - \Delta t} \int_0^{t_{\max} - \Delta t} dt_0 \frac{\langle (u - \bar{u})^4 \rangle_{t_0, \Delta t}}{\langle (u - \bar{u})^2 \rangle_{t_0, \Delta t}^2}, \quad (4)$$

where  $t_{\max}$  is the full length of the time series. The notation  $\langle \cdots \rangle_{t_0,\Delta t}$  indicates the expectation for the time slice of length  $\Delta t$  starting at  $t_0$ . Assuming local exponential distributions with kurtosis 9, we apply Eq. (4) to determine the long timescale *T* as the special time length  $\Delta t$  such that

$$\kappa(\Delta t) = \kappa(T) = 9. \tag{5}$$

We show an example of how *T* is calculated in Fig. 4, with a plot illustrating how the average local kurtosis depends on  $\Delta t$ . Once the time slice length *T* is calculated, we define an intensive parameter  $\beta$  for each local cell, by setting  $\beta = \lambda_e = 1/\langle u \rangle_{t_0,T}$ . The distribution of this parameter,  $f(\beta)$ , is then obtained from a histogram, as shown in Fig. 5(a). In good approximation, we find  $\beta$  to be  $\chi^2$  distributed, i.e.,

$$f(\beta) = \frac{1}{\Gamma\left(\frac{n}{2}\right)} \left(\frac{n}{2\beta_0}\right)^{\frac{n}{2}} \beta^{\frac{n}{2}-1} \exp\left(-\frac{n\beta}{2\beta_0}\right), \tag{6}$$

where *n* is the number degrees of freedom and  $\beta_0$  is the mean of  $\beta$ .

Integrating out the  $\beta$  parameter, the marginal distribution p(u) is calculated as

$$p(u) = \int_0^\infty p(u|\beta) f(\beta) d\beta, \tag{7}$$

which evaluates to

$$p(u) = (2 - q)\lambda_q [1 + (q - 1)\lambda_q u]^{\frac{1}{1 - q}},$$
(8)

with

$$-\frac{n}{2} - 1 = \frac{1}{1 - q},\tag{9}$$



FIG. 3. NO concentrations locally follow exponential distributions. (a) We plot the NO concentration for several months in 2010, with two periods of length  $\Delta t = T \approx 6$  days highlighted, whose distributions are explored in (b) and (c). The histograms of (b) and (c) are fitted with exponential distributions and we note their respective  $\lambda_e$  values. Note that the exponent in (b) is only one fourth of the exponent recovered in (c), i.e., high-pollution levels are much more likely to be observed in (b), a time period in January, than in (c), a time period in June. These different exponents quantify the observation that pollutant levels tend to be higher during the winter than during the summer (see also Fig. 1).

and

$$\frac{1}{2}(q-1)\lambda_q = \frac{\beta_0}{n}.$$
(10)

Equation (8) is the *q*-exponential distribution, which we derived from a superposition of local exponential distributions.



FIG. 4. The average kurtosis  $\kappa$  as given by Eq. (4) is plotted as a function of the time window  $\Delta t$  (blue). The crossing between the horizontal line at  $\kappa = 9$  (the kurtosis of an exponential distribution) and the  $\kappa$  vs  $\Delta t$  curve gives the value for  $\Delta t = T \approx 6$ .

Fitting the *n* parameter to find  $f(\beta)$ , as illustrated in Fig. 5(a), we determine the *q* and  $\lambda_q$  parameters from Eqs. (9) and (10), respectively. Inserting the  $\chi^2$ -distribution  $f(\beta)$ , we use (7) and (8) to compute a probability density, which both approximates the fitted *q*-exponential and the data very well, successfully serving as a consistency check of the superstatistical approach [see Fig. 5(b)].

## III. ANALYSIS OF NO2 DATA

We already noticed different statistical behavior when inspecting the trajectory of the NO<sub>2</sub> data in Fig. 1, as compared to that of NO. Following a similar procedure as for the NO data, we find that for NO<sub>2</sub> the aggregated distribution is approximated by a q-Maxwell-Boltzmann distribution, of the form

$$p(u) = \frac{1}{Z} u^2 \sigma_q^{3/2} [1 + (q-1)\sigma_q u^2]^{\frac{1}{1-q}}, \qquad (11)$$

where Z is the normalization factor given by

$$Z = \frac{\sqrt{\pi} \Gamma\left(\frac{5-3q}{2(q-1)}\right)}{4(q-1)^{3/2} \Gamma\left(\frac{1}{q-1}\right)},$$
(12)

and where q is related to the scale parameter  $\sigma_q$  and the mean  $\mu$  by

$$\sigma_q = \left(\frac{2(q-1)^{3/2}\Gamma(\frac{1}{q-1})}{\mu\sqrt{\pi}(2-q)(3-2q)\Gamma(\frac{5-3q}{2(q-1)})}\right)^2$$
(13)



FIG. 5. The superstatistical approach consistently describes the data. (a) The distribution of local exponents of the NO data, here named  $\beta$ , follows a  $\chi^2$  distribution, one of three commonly observed universality classes in general superstatistical complex systems [2]. (b) Histogram of the NO data, together with a *q*-exponential distribution with fitted *q* and  $\lambda_q$  values (red), and a *q*-exponential distribution with *q* and  $\lambda_q$  values derived from the parameters of the fitted  $\chi^2$ -distributed  $f(\beta)$  (orange).

(see Appendix C for details).

Analogously to the NO case, we explain the aggregated distribution as a superposition of simple local distributions, here chosen as ordinary Maxwell-Boltzmann distributions. Each local Maxwell-Boltzmann distribution is defined as

$$p(u) = \sqrt{\frac{16}{\pi}} u^2 \sigma_{mb}^{3/2} \exp(-\sigma_{mb} u^2), \qquad (14)$$

where  $\sigma_{mb} =: \beta$  is a scale parameter.

We determine the long timescale T by varying  $\Delta t$  so that we locally obtain the kurtosis of a Maxwell-Boltzmann distribution, which is given as

$$\kappa(T) = \frac{15\pi^2 + 16\pi - 192}{(3\pi - 8)^2} \approx 3.1082$$
(15)

(see Appendix C for the calculation). If the  $\beta$  values again follow a  $\chi^2$  distribution, the superpositioned Maxwell-Boltzmann distributions would lead to an exact *q*-Maxwell-Boltzmann distribution (see Appendix C for a detailed calculation).

The results of our superstatistical analysis for the NO<sub>2</sub> data are displayed in Fig. 6. The  $\chi^2$  distribution is only a rough fit

and the histogram of  $\beta$  is instead best fitted by an inverse- $\chi^2$  distribution

$$f(\beta) = \frac{\beta_0}{\Gamma\left(\frac{n}{2}\right)} \left(\frac{n\beta_0}{2}\right)^{\frac{n}{2}} \beta^{-\frac{n}{2}-2} \exp\left(\frac{-n\beta_0}{2\beta}\right).$$
(16)

This leads to a different superstatistics, compared to the case of NO. Integrating the conditional probability  $p(u|\beta)$ , as in Eq. (7) but with local Maxwell-Boltzmann distributions  $p(u|\beta)$  and an inverse- $\chi^2$  distribution  $f(\beta)$ , gives the marginal distribution as

$$p(u) = \sqrt{\frac{64}{\pi}} \frac{\beta_0}{\Gamma(\frac{n}{2})} \left(\frac{n\beta_0}{2}\right)^{\frac{n+1}{4}} u^{\frac{n+3}{2}} K_{\frac{1-n}{2}}(u\sqrt{2n\beta_0}), \quad (17)$$

where  $K_v(z)$  is the modified Bessel function of the second order. This inverse- $\chi^2$ -superstatistical model leads to a very good description of the data (see Fig. 6) and it approximates *q*-Maxwell-Boltzmann distributions for medium concentrations but it not exactly the same. In fact, the probability distribution function (PDF) of the observed statistics decays as  $p(u) \sim$  $\exp(-\text{const}\sqrt{u})$  for large *u* values (see also Appendix C).



FIG. 6. NO<sub>2</sub> concentrations are well described by inverse- $\chi^2$  superstatistics. (a) The extracted distribution of the scale parameter  $\beta$  for the NO<sub>2</sub> time series is very well fitted by an inverse- $\chi^2$  distribution (green). (b) Histogram of the NO<sub>2</sub> data, together with a *q*-Maxwell-Boltzmann distribution with fitted *q* and  $\sigma_q$  values (red) and a *q*-Maxwell-Boltzmann with *q* and  $\sigma_q$  values superstatistically calculated from a  $\chi^2$ -distributed *f*( $\beta$ ) (orange), and Eq. (17), which is calculated from inverse- $\chi^2$  superstatistics (green). The inverse- $\chi^2$  superstatistics fits the NO<sub>2</sub> data best.

TABLE I. Overview of the different	cases of local distributions	(exponential or Ma	xwell-Boltzmann)
and the $\beta$ distributions of the individual s	cale parameters.		

	$\chi^2$	Inv. $\chi^2$
Exponential Maxwell-Boltzmann	<i>q</i> -exp, empirical NO <i>q</i> -MB	Not observed here, but see [25] Empirical NO <sub>2</sub>

#### **IV. CONCLUSION**

To conclude, we have illustrated that superstatistical methods, originally introduced in nonequilibrium statistical mechanics and applied to fully developed turbulent flows and other complex systems, can also be used to quantitatively model the statistics of air pollution. We find excellent agreement of simple superstatistical models with the experimentally measured pollution data. A main result is that different pollutants obey different types of superstatistics (in our case  $\chi^2$  for NO and inverse  $\chi^2$  for NO<sub>2</sub>). In fact, different types of local dynamics also occur (in our case locally exponential density for NO and locally Maxwell-Boltzmann for NO<sub>2</sub>). Once the precise superstatistical model for a given pollutant has been identified, precise risk estimates of high-pollution situations can be given, by integrating the tails of the probability distribution above a given threshold. These estimates could help to design tailor-made thresholds for suitable policies to tackle the pollution problem in cities. We discuss some explicit examples in Appendix D. While the precise local statistics (exponential or Maxwell-Boltzmann) does significantly influence the likelihood to observe very low concentrations, the probability to observe high concentrations is determined by heavy tails in both cases.

The superstatistical approach presented here can easily be generalized and extended. For example, we may formulate an explicit dynamical process to match the trajectories in time (in form of a superstatistical stochastic differential equation), which could lead to synthetic dynamical models and the possibility to employ short-term predictions of the pollution concentration. Further research should be devoted to understand how and why certain local statistics, such as exponential or Maxwell-Boltzmann distributions, arise, and to potentially link these statistics to the physical and chemical properties of the pollutants.

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## APPENDIX A: SUPERSTATISTICAL CASES

These Appendixes provide calculations of some equalities used in the main text, such as relationships between mean, variance, and kurtosis for the relevant distributions. We also explain how the long timescale is determined and explicitly discuss how the knowledge of concentration probability distributions could be used to set pollution thresholds and policies. We start by providing an overview of the different cases of local distributions (exponential or Maxwell-Boltzmann) and the  $\beta$  distributions of the individual scale parameters, shown in Table I.

#### **APPENDIX B: NO**

In the case of NO concentrations, we approximate local concentrations as exponential distributions,

$$p(u) = \lambda_e \exp(-\lambda_e u), \tag{B1}$$

and the aggregated statistics as q-exponentials,

$$p(u) = (2 - q)\lambda_q [1 + (q - 1)\lambda_q u]^{\frac{1}{1 - q}},$$
 (B2)

derived from a  $\chi^2$  distribution of  $\beta = \lambda_e$ ,

*,* ,

$$f(\beta) = \frac{1}{\Gamma(\frac{n}{2})} \left(\frac{n}{2\beta_0}\right)^{\frac{n}{2}} \beta^{\frac{n}{2}-1} \exp\left(-\frac{n\beta}{2\beta_0}\right).$$
(B3)

## 1. Calculation of $\lambda_q$

Here, we show how to express the scale parameter of *q*-exponentials  $\lambda_q$  as a function of the mean of the distribution  $\mu = \langle u \rangle$  and the *q* parameter. We start by computing the mean,

$$\mu = \langle u \rangle$$
  
=  $(2 - q)\lambda_q \int_0^\infty u[1 + (q - 1)\lambda_q u]^{\frac{1}{1 - q}} du$   
=  $\frac{(2 - q)\Gamma(2)\Gamma(\frac{3 - 2q}{q - 1})}{\lambda_q (q - 1)^2 \Gamma(\frac{1}{q - 1})}$   
=  $\frac{1}{\lambda_q (3 - 2q)}$ , (B4)

which gives

$$\lambda_q = \frac{1}{\mu(3 - 2q)}.\tag{B5}$$

### 2. Calculation of mean $\mu$ of exponential distribution

We briefly recall the relationship between the exponential scale parameter  $\beta = \lambda_e$  and the mean  $\mu$  as

$$\mu = \langle u \rangle$$
  
=  $\beta \int_0^\infty u \exp(-\beta u) du$   
=  $\frac{1}{\beta}$ . (B6)

## 3. Calculation of kurtosis $\kappa$ of exponential distribution

In the main text, we determined the timescale T as the timescale where the local kurtosis of the local exponential distribution takes on the value  $\kappa = 9$ . Here, we provide the corresponding calculation,

$$\kappa = \frac{\langle (u-\mu)^4 \rangle}{\langle (u-\mu)^2 \rangle^2} = \frac{\int_0^\infty (u-\mu)^4 \beta \exp(-\beta u) du}{\left[\int_0^\infty (u-\mu)^2 \beta \exp(-\beta u) du\right]^2} = \frac{\int_0^\infty (u^4 - 4u^3\mu + 6u^2\mu^2 - 4u\mu^3 + \mu^4) \beta \exp(-\beta u) du}{\left[\int_0^\infty (u^2 - 2u\mu + \mu^2) \beta \exp(-\beta u) du\right]^2} = \frac{24 - 24\mu\beta + 12\mu^2\beta^2 - 4\mu^3\beta^3 + \mu^4\beta^4}{4 - 8\mu\beta + 8\mu^2\beta^2 - 4\mu^3\beta^3 + \mu^4\beta^4} = 9,$$
(B7)

using  $\mu = \frac{1}{\beta}$ .

## 4. Superstatistical calculation of q-exponential distribution

We show how integrating several local exponential distributions, whose exponents  $\beta$  follow a  $\chi^2$  distribution, leads to an overall *q*-exponential distribution: Each local distribution is given as

$$p(u|\beta) = \beta \exp(-\beta u). \tag{B8}$$

Integrating over all of these distributions can be expressed as

$$p(u) = \int_0^\infty p(u|\beta)f(\beta)d\beta$$
$$= \frac{1}{\Gamma\left(\frac{n}{2}\right)} \left(\frac{n}{2\beta_0}\right)^{\frac{n}{2}} \int_0^\infty \beta^{\frac{n}{2}} \exp\left[-\beta\left(\frac{n}{2\beta_0} + u\right)\right] d\beta,$$
(B9)

which we evaluate to

$$p(u) = (2 - q)\lambda_q [1 + (q - 1)\lambda_q u]^{\frac{1}{1 - q}},$$
 (B10)

if we identify

$$-\left(\frac{n+2}{2}\right) = \frac{1}{1-q},$$
 (B11)

$$\frac{1}{2}(q-1)\lambda_q = \frac{\beta_0}{n},\tag{B12}$$

where  $\beta_0$  is the mean of  $\beta$ .

We determine the long timescale T from the time series using Eq. (4) from the main text given as

$$\kappa(\Delta t) = \frac{1}{t_{\max} - \Delta t} \int_0^{t_{\max} - \Delta t} dt_0 \frac{\langle (u - \bar{u})^4 \rangle_{t_0, \Delta t}}{\langle (u - \bar{u})^2 \rangle_{t_0, \Delta t}^2}.$$
 (B13)

The long timescale *T* is defined as  $T := \Delta t$  such that  $\kappa(\Delta t) = 9$ . Here, we compute the average local kurtosis  $\kappa$  as a function of the time window  $\Delta t$  and thereby determine *T*.

## **APPENDIX C: NO<sub>2</sub>**

In the case of NO<sub>2</sub> concentrations, we approximate local concentrations as Maxwell-Boltzmann distributions,

$$p(u) = \sqrt{\frac{16}{\pi}} u^2 \sigma_{mb}^{3/2} \exp(-\sigma_{mb} u^2),$$
 (C1)

and the aggregated statistics is a *q*-Maxwell-Boltzmann distribution,

$$p(u) = \frac{1}{Z} u^2 \sigma_q^{3/2} [1 + (q-1)\sigma_q u^2]^{\frac{1}{1-q}}, \qquad (C2)$$

if  $\sigma_{mb}$  is  $\chi^2$  distributed. These *q*-Maxwell-Boltzmann distributions would strictly arise if the scale parameters  $\beta = \sigma_{mb}$  were following a  $\chi^2$  distribution. Empirically, we observe that for our data, the  $\beta$  values follow an inverse- $\chi^2$  distribution instead,

$$f(\beta) = \frac{\beta_0}{\Gamma\left(\frac{n}{2}\right)} \left(\frac{n\beta_0}{2}\right)^{\frac{n}{2}} \beta^{-\frac{n}{2}-2} \exp\left(\frac{-n\beta_0}{2\beta}\right).$$
(C3)

We obtain this inverse- $\chi^2$  distribution from a simple transformation of random variables when  $\beta^{-1}$ , rather than  $\beta$ , is  $\chi^2$  distributed.

### 1. Calculation of $\sigma_q$

Here, we show how to express the scale parameter of *q*-Maxwell-Boltzmann distributions  $\sigma_q$  as a function of the mean of the distribution  $\mu = \langle u \rangle$  and the *q* parameter. We start by computing the mean,

$$\begin{split} \mu &= \langle u \rangle \\ &= \frac{\sigma_q^{3/2}}{Z} \int_0^\infty u^3 [1 + (q-1)\sigma_q u^2]^{\frac{1}{1-q}} du \\ &= \frac{\Gamma(2)\Gamma\left(\frac{3-2q}{q-1}\right)}{2Z\sigma_q^{1/2}(q-1)^2\Gamma\left(\frac{1}{q-1}\right)} \\ &= \frac{1}{2Z(2-q)(3-2q)\sigma_q^{1/2}} \\ &= \frac{2(q-1)^{3/2}\Gamma\left(\frac{1}{q-1}\right)}{\sigma_q^{1/2}\sqrt{\pi}(2-q)(3-2q)\Gamma\left(\frac{5-3q}{q-1}\right)}, \end{split}$$
(C4)

which gives

$$\sigma_q = \left(\frac{2(q-1)^{3/2}\Gamma(\frac{1}{q-1})}{\mu\sqrt{\pi}(2-q)(3-2q)\Gamma(\frac{5-3q}{q-1})}\right)^2.$$
 (C5)

## 2. Calculation of mean $\mu$ of Maxwell-Boltzmann distribution

We briefly recall the relationship between the scale parameter  $\beta = \sigma_{mb}$  of the Maxwell-Boltzmann distribution and the mean  $\mu$  as

$$\mu = \langle u \rangle$$
  
=  $\sqrt{\frac{16}{\pi}} \beta^{3/2} \int_0^\infty u^3 \exp(-\beta u^2) du$   
=  $\sqrt{\frac{16}{\pi}} \beta^{3/2} \left(\frac{\Gamma(2)}{2\beta^2}\right) = \frac{2}{\sqrt{\pi\beta}}.$  (C6)



FIG. 7. NO<sub>2</sub> concentrations locally follow Maxwell-Boltzmann distributions. (a) We plot the NO<sub>2</sub> concentration for several months in 2010, with two periods of length  $\Delta t = T \approx 2.6$  days highlighted, whose distributions are explored in (b) and (c). The histograms of (b) and (c) are fitted with Maxwell-Boltzmann distributions and we note their respective  $\lambda_{mb}$  values, which are strongly varying in time.

## 3. Calculation of kurtosis $\kappa$ of Maxwell-Boltzmann distribution

In the main text, we determined the timescale T as the timescale where the local kurtosis of the local Maxwell-Boltzmann distribution takes on the value  $\kappa \approx 3.1082$ . Here, we provide the corresponding calculation,

$$\kappa = \frac{\langle (u-\mu)^4 \rangle}{\langle (u-\mu)^2 \rangle^2} = \frac{\sqrt{\frac{16}{\pi}} \beta^{3/2} \int_0^\infty (u-\mu)^4 u^2 \exp(-\beta u^2) du}{\left(\sqrt{\frac{16}{\pi}} \beta^{3/2} \int_0^\infty (u-\mu)^2 u^2 \exp(-\beta u^2) du\right)^2} = \frac{\frac{\mu^4 \sqrt{\pi}}{4\beta^{3/2}} - \frac{2\mu^3}{\beta^2} + \frac{9\mu^2 \sqrt{\pi}}{4\beta^{5/2}} - \frac{4\mu}{\beta^3} + \frac{15\sqrt{\pi}}{16\beta^{7/2}}}{\sqrt{\frac{16}{\pi}} \beta^{3/2} \left(\frac{\mu^2 \sqrt{\pi}}{4\beta^{3/2}} - \frac{\mu}{\beta^2} + \frac{3\sqrt{\pi}}{8\beta^{5/2}}\right)^2}$$
(C7)

$$=\frac{15\pi^2 + 16\pi - 192}{(3\pi - 8)^2} \approx 3.1082.$$
 (C8)

# 4. Superstatistical calculation of *q*-Maxwell-Boltzmann distribution

We show how integrating several local Maxwell-Boltzmann distributions, whose exponents  $\beta$  follow a  $\chi^2$  distribution, leads to an overall *q*-Maxwell-Boltzmann distribution: Each local distribution is given as

$$p(u|\beta) = \sqrt{\frac{16}{\pi}} u^2 \sigma_{mb}^{3/2} \exp(-\sigma_{mb} u^2).$$
 (C9)

Integrating over all of these distributions can be expressed as

$$p(u) = \int_0^\infty p(u|\beta) f(\beta) d\beta$$
  
=  $\sqrt{\frac{16}{\pi}} \frac{1}{\Gamma(\frac{n}{2})} \left(\frac{n}{2\beta_0}\right)^{\frac{n}{2}} u^2$   
 $\times \int_0^\infty \beta^{\frac{n+1}{2}} \exp\left[\beta\left(\frac{n}{2\beta_0} + u^2\right)\right] d\beta$ , (C10)

which can be evaluated to

$$p(u) \sim u^2 \sigma_q^{3/2} [1 + (q-1)\sigma_q u^2]^{\frac{1}{1-q}},$$
 (C11)

if we identify

$$-\left(\frac{n+3}{2}\right) = \frac{1}{1-q},$$
 (C12)

$$\frac{1}{2}(q-1)\sigma_q = \frac{\beta_0}{n},\tag{C13}$$

where  $\beta_0$  is the mean of  $\beta$ .

## 5. Calculation of inverse-χ<sup>2</sup> superstatistical distribution based on superimposed Maxwell-Boltzmann distributions

Taking the conditional density  $p(u|\beta)$  as given in Eq. (C9) and the inverse- $\chi^2$  distribution for the distribution  $f(\beta)$ , the



FIG. 8. The average kurtosis  $\kappa$  is plotted as a function of the time window  $\Delta t$ . The crossing between the horizontal line at  $\kappa = 3.1082$  (the kurtosis of a Maxwell-Boltzmann distribution) and the  $\kappa$  vs  $\Delta t$  curve gives the value  $\Delta t = T = 2.61$ .

probability density p(u) is given as

$$p(u) = \sqrt{\frac{16}{\pi}} \frac{\beta_0}{\Gamma(\frac{n}{2})} \left(\frac{n\beta_0}{2}\right)^{\frac{n}{2}} u^2$$
$$\times \int_0^\infty \beta^{-\frac{(n+1)}{2}} \exp\left(-\frac{n\beta_0}{2\beta} - \beta u^2\right) d\beta$$
$$= \sqrt{\frac{64}{\pi}} \frac{\beta_0}{\Gamma(\frac{n}{2})} \left(\frac{n\beta_0}{2}\right)^{\frac{n+1}{4}} u^{\frac{n+3}{2}} K_{\frac{1-n}{2}}(u\sqrt{2n\beta_0}). \quad (C14)$$

This PDF has still heavy tails, which decay as  $p(u) \sim \exp(-\operatorname{const}\sqrt{u})$  for large *u* values.

#### 6. Extra figures

We repeat Fig. 3 from the main text for local NO<sub>2</sub> concentrations, i.e., analyzing brief periods of the NO<sub>2</sub> trajectory. Figure 7 illustrates that in good approximation the local behavior follows Maxwell-Boltzmann distributions with fluctuating variance. Further, we also determine the long timescale *T* for the local Maxwell-Boltzmann distributions by setting  $T := \Delta t$  for the local kurtosis  $\kappa(\Delta t) \approx 3.1082$ , following Eq. (4) from the main text again. See Figs. 7 and 8.

# APPENDIX D: THRESHOLDS AND POLICIES

Policies to tackle pollution in cities often focus on thresholds and compliance with these thresholds. This then often leads to pollution just below the threshold in many regions. However, in many cases it may be more reasonable to reduce the overall exposure to pollutants [26]. Weather is responsible for many fast variations of the air pollution [26]. The focus on threshold is especially dangerous as there is strong evidence that even small concentrations of pollutants, below national standards, increase the death rate [27]. The U.K. government (Department of Health and Social Care) calls air pollution a "health emergency" and the World Health Organization (WHO) judges the situation similarly on a global scale [28]. Note that about 9500 premature deaths are attributed to air pollution in London alone every year [29]. On a global scale this number rises to about 4.9 million premature deaths, making air pollution the fifth leading cause of mortality worldwide [30].

So let us finally sketch how the knowledge of the probability density function (PDF) allows estimates of threshold crossings and total exposure. Setting thresholds on pollutant concentrations is a popular tool when setting pollution policies. Simultaneously, data coverage might not be very good in all locations. Especially if data are not available for the full time period of interest but only for a few months, estimates on how frequently thresholds are violated are difficult to obtain. As soon as we are able to determine the PDF p(u) for the concentration levels u, we can easily compute the number of days where thresholds are violated as

$$N_{u>u_{\text{threshold}}} = 365 \int_{u_{\text{threshold}}}^{\infty} p(u) du.$$
 (D1)

This integral can be easily evaluated numerically, using q-exponential distributions for NO, see Eq. (8) from the main text, and Eq. (17) from the main text for NO<sub>2</sub>, consistent with experimental observations.

The number of threshold violations explicitly depends on the parameters  $\lambda_q$  and q, which have to be determined from the data sets but can then be compared between different locations and time periods. Having these two parameters explicitly disentangles two aspects: the rate of (rare) extreme events, encoded in q, and the overall variability of the pollutant concentration, encoded in  $\lambda_q$  (or  $\lambda_{mb}$  for NO<sub>2</sub>).

Complementary to thresholds, policies could instead focus on the total pollution exposure (TPE) describing the average amount citizens are exposed to. Again, this is easily computed from the PDF as

$$TPE = 365 \int_0^\infty p(u)udu.$$
 (D2)

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