

Search for Chargino-Neutralino Production in $p\bar{p}$ Collisions at $\sqrt{s} = 1.96$ TeV

- T. Aaltonen,²³ A. Abulencia,²⁴ J. Adelman,¹³ T. Affolder,¹⁰ T. Akimoto,⁵⁵ M. G. Albrow,¹⁷ S. Amerio,⁴³ D. Amidei,³⁵ A. Anastassov,⁵² K. Anikeev,¹⁷ A. Annovi,¹⁹ J. Antos,¹⁴ M. Aoki,⁵⁵ G. Apollinari,¹⁷ T. Arisawa,⁵⁷ A. Artikov,¹⁵ W. Ashmanskas,¹⁷ A. Attal,³ A. Aurisano,⁵³ F. Azfar,⁴² P. Azzi-Bacchetta,⁴³ P. Azzurri,⁴⁶ N. Bacchetta,⁴³ W. Badgett,¹⁷ A. Barbaro-Galtieri,²⁹ V. E. Barnes,⁴⁸ B. A. Barnett,²⁵ S. Baroiant,⁷ V. Bartsch,³¹ G. Bauer,³³ P.-H. Beauchemin,³⁴ F. Bedeschi,⁴⁶ S. Behari,²⁵ G. Bellettini,⁴⁶ J. Bellinger,⁵⁹ A. Belloni,³³ D. Benjamin,¹⁶ A. Beretvas,¹⁷ J. Beringer,²⁹ T. Berry,³⁰ A. Bhatti,⁵⁰ M. Binkley,¹⁷ D. Bisello,⁴³ I. Bizjak,³¹ R. E. Blair,² C. Blocker,⁶ B. Blumenfeld,²⁵ A. Bocci,¹⁶ A. Bodek,⁴⁹ V. Boisvert,⁴⁹ G. Bolla,⁴⁸ A. Bolshov,³³ D. Bortoletto,⁴⁸ J. Boudreau,⁴⁷ A. Boveia,¹⁰ B. Brau,¹⁰ L. Brigliadori,⁵ C. Bromberg,³⁶ E. Brubaker,¹³ J. Budagov,¹⁵ H. S. Budd,⁴⁹ S. Budd,²⁴ K. Burkett,¹⁷ G. Busetto,⁴³ P. Bussey,²¹ A. Buzatu,³⁴ K. L. Byrum,² S. Cabrera,¹⁶^a M. Campanelli,²⁰ M. Campbell,³⁵ F. Canelli,¹⁷ A. Canepa,⁴⁵ S. Carillo,¹⁸^b D. Carlsmith,⁵⁹ R. Carosi,⁴⁶ S. Carron,³⁴ B. Casal,¹¹ M. Casarsa,⁵⁴ A. Castro,⁵ P. Catastini,⁴⁶ D. Cauz,⁵⁴ M. Cavalli-Sforza,³ A. Cerri,²⁹ L. Cerrito,^{31,m} S. H. Chang,²⁸ Y. C. Chen,¹ M. Chertok,⁷ G. Chiarelli,⁴⁶ G. Chlachidze,¹⁷ F. Chlebana,¹⁷ I. Cho,²⁸ K. Cho,²⁸ D. Chokheli,¹⁵ J. P. Chou,²² G. Choudalakis,³³ S. H. Chuang,⁵² K. Chung,¹² W. H. Chung,⁵⁹ Y. S. Chung,⁴⁹ M. Cilijak,⁴⁶ C. I. Ciobanu,²⁴ M. A. Ciocci,⁴⁶ A. Clark,²⁰ D. Clark,⁶ M. Coca,¹⁶ G. Compostella,⁴³ M. E. Convery,⁵⁰ J. Conway,⁷ B. Cooper,³¹ K. Copic,³⁵ M. Cordelli,¹⁹ G. Cortiana,⁴³ F. Crescioli,⁴⁶ C. Cuena Almenar,⁷^a J. Cuevas,¹¹^b R. Culbertson,¹⁷ J. C. Cully,³⁵ S. DaRonco,⁴³ M. Datta,¹⁷ S. D'Auria,²¹ T. Davies,²¹ D. Dagenhart,¹⁷ P. de Barbaro,⁴⁹ S. De Cecco,⁵¹ A. Deisher,²⁹ G. De Lentdecker,⁴⁹^c G. De Lorenzo,³ M. Dell'Orso,⁴⁶ F. Delli Paoli,⁴³ L. Demortier,⁵⁰ J. Deng,¹⁶ M. Deninno,⁵ D. De Pedis,⁵¹ P. F. Derwent,¹⁷ G. P. Di Giovanni,⁴⁴ C. Dionisi,⁵¹ B. Di Ruzza,⁵⁴ J. R. Dittmann,⁴ M. D'Onofrio,³ C. Dörr,²⁶ S. Donati,⁴⁶ P. Dong,⁸ J. Donini,⁴³ T. Dorigo,⁴³ S. Dube,⁵² J. Efron,³⁹ R. Erbacher,⁷ D. Errede,²⁴ S. Errede,²⁴ R. Eusebi,¹⁷ H. C. Fang,²⁹ S. Farrington,³⁰ I. Fedorko,⁴⁶ W. T. Fedorko,¹³ R. G. Feild,⁶⁰ M. Feindt,²⁶ J. P. Fernandez,³² R. Field,¹⁸ G. Flanagan,⁴⁸ R. Forrest,⁷ S. Forrester,⁷ M. Franklin,²² J. C. Freeman,²⁹ I. Furic,¹³ M. Gallinaro,⁵⁰ J. Galyardt,¹² J. E. Garcia,⁴⁶ F. Garberson,¹⁰ A. F. Garfinkel,⁴⁸ C. Gay,⁶⁰ H. Gerberich,²⁴ D. Gerdes,³⁵ S. Giagu,⁵¹ P. Giannetti,⁴⁶ K. Gibson,⁴⁷ J. L. Gimmell,⁴⁹ C. Ginsburg,¹⁷ N. Giokaris,^{15,a} M. Giordani,⁵⁴ P. Giromini,¹⁹ M. Giunta,⁴⁶ G. Giurgiu,²⁵ V. Glagolev,¹⁵ D. Glenzinski,¹⁷ M. Gold,³⁷ N. Goldschmidt,¹⁸ J. Goldstein,^{42,b} A. Golossanov,¹⁷ G. Gomez,¹¹ G. Gomez-Ceballos,³³ M. Goncharov,⁵³ O. González,³² I. Gorelov,³⁷ A. T. Goshaw,¹⁶ K. Goulianos,⁵⁰ A. Gresele,⁴³ S. Grinstein,²² C. Grossi-Pilcher,¹³ R. C. Group,¹⁷ U. Grundler,²⁴ J. Guimaraes da Costa,²² Z. Gunay-Unalan,³⁶ C. Haber,²⁹ K. Hahn,³³ S. R. Hahn,¹⁷ E. Halkiadakis,⁵² A. Hamilton,²⁰ B.-Y. Han,⁴⁹ J. Y. Han,⁴⁹ R. Handler,⁵⁹ F. Happacher,¹⁹ K. Hara,⁵⁵ D. Hare,⁵² M. Hare,⁵⁶ S. Harper,⁴² R. F. Harr,⁵⁸ R. M. Harris,¹⁷ M. Hartz,⁴⁷ K. Hatakeyama,⁵⁰ J. Hauser,⁸ C. Hays,⁴² M. Heck,²⁶ A. Heijboer,⁴⁵ B. Heinemann,²⁹ J. Heinrich,⁴⁵ C. Henderson,³³ M. Herndon,⁵⁹ J. Heuser,²⁶ D. Hidas,¹⁶ C. S. Hill,^{10,b} D. Hirschbuehl,²⁶ A. Hocker,¹⁷ A. Holloway,²² S. Hou,¹ M. Houlden,³⁰ S.-C. Hsu,⁹ B. T. Huffman,⁴² R. E. Hughes,³⁹ U. Husemann,⁶⁰ J. Huston,³⁶ J. Incandela,¹⁰ G. Introzzi,⁴⁶ M. Iori,⁵¹ A. Ivanov,⁷ B. Iyutin,³³ E. James,¹⁷ D. Jang,⁵² B. Jayatilaka,¹⁶ D. Jeans,⁵¹ E. J. Jeon,²⁸ S. Jindariani,¹⁸ W. Johnson,⁷ M. Jones,⁴⁸ K. K. Joo,²⁸ S. Y. Jun,¹² J. E. Jung,²⁸ T. R. Junk,²⁴ T. Kamon,⁵³ P. E. Karchin,⁵⁸ Y. Kato,⁴¹ Y. Kemp,²⁶ R. Kephart,¹⁷ U. Kerzel,²⁶ V. Khotilovich,⁵³ B. Kilminster,³⁹ D. H. Kim,²⁸ H. S. Kim,²⁸ J. E. Kim,²⁸ M. J. Kim,¹⁷ S. B. Kim,²⁸ S. H. Kim,⁵⁵ Y. K. Kim,¹³ N. Kimura,⁵⁵ L. Kirsch,⁶ S. Klimenko,¹⁸ M. Klute,³³ B. Knuteson,³³ B. R. Ko,¹⁶ K. Kondo,⁵⁷ D. J. Kong,²⁸ J. Konigsberg,¹⁸ A. Korytov,¹⁸ A. V. Kotwal,¹⁶ A. C. Kraan,⁴⁵ J. Kraus,²⁴ M. Kreps,²⁶ J. Kroll,⁴⁵ N. Krumnack,⁴ M. Kruse,¹⁶ V. Krutelyov,¹⁰ T. Kubo,⁵⁵ S. E. Kuhlmann,² T. Kuhr,²⁶ N. P. Kulkarni,⁵⁸ Y. Kusakabe,⁵⁷ S. Kwang,¹³ A. T. Laasanen,⁴⁸ S. Lai,³⁴ S. Lami,⁴⁶ S. Lammel,¹⁷ M. Lancaster,³¹ R. L. Lander,⁷ K. Lannon,³⁹ A. Lath,⁵² G. Latino,⁴⁶ I. Lazzizzera,⁴³ T. LeCompte,² J. Lee,⁴⁹ J. Lee,²⁸ Y. J. Lee,²⁸ S. W. Lee,^{53,o} R. Lefèvre,²⁰ N. Leonardo,³³ S. Leone,⁴⁶ S. Levy,¹³ J. D. Lewis,¹⁷ C. Lin,⁶⁰ C. S. Lin,¹⁷ M. Lindgren,¹⁷ E. Lippeles,⁹ A. Lister,⁷ D. O. Litvintsev,¹⁷ T. Liu,¹⁷ N. S. Lockyer,⁴⁵ A. Loginov,⁶⁰ M. Loreti,⁴³ R.-S. Lu,¹ D. Lucchesi,⁴³ P. Lujan,²⁹ P. Lukens,¹⁷ G. Lungu,¹⁸ L. Lyons,⁴² J. Lys,²⁹ R. Lysak,¹⁴ E. Lytken,⁴⁸ P. Mack,²⁶ D. MacQueen,³⁴ R. Madrak,¹⁷ K. Maeshima,¹⁷ K. Makhoul,³³ T. Maki,²³ P. Maksimovic,²⁵ S. Malde,⁴² S. Malik,³¹ G. Manca,³⁰ A. Manousakis,^{15,a} F. Margaroli,⁵ R. Marginean,¹⁷ C. Marino,²⁶ C. P. Marino,²⁴ A. Martin,⁶⁰ M. Martin,²⁵ V. Martin,^{21,g} M. Martínez,³ R. Martínez-Ballarín,³² T. Maruyama,⁵⁵ P. Mastrandrea,⁵¹ T. Masubuchi,⁵⁵ H. Matsunaga,⁵⁵ M. E. Mattson,⁵⁸ R. Mazini,³⁴ P. Mazzanti,⁵ K. S. McFarland,⁴⁹ P. McIntyre,⁵³ R. McNulty,^{30,f} A. Mehta,³⁰ P. Mehtala,²³ S. Menzemer,^{11,h} A. Menzione,⁴⁶ P. Merkel,⁴⁸ C. Mesropian,⁵⁰ A. Messina,³⁶ T. Miao,¹⁷ N. Miladinovic,⁶ J. Miles,³³ R. Miller,³⁶ C. Mills,¹⁰ M. Milnik,²⁶ A. Mitra,¹ G. Mitselmakher,¹⁸ A. Miyamoto,²⁷ S. Moed,²⁰ N. Moggi,⁵ B. Mohr,⁸ C. S. Moon,²⁸ R. Moore,¹⁷ M. Morello,⁴⁶ P. Movilla Fernandez,²⁹ J. Mühlstädt,²⁹ A. Mukherjee,¹⁷ Th. Muller,²⁶

- R. Mumford,²⁵ P. Murat,¹⁷ M. Mussini,⁵ J. Nachtman,¹⁷ A. Nagano,⁵⁵ J. Naganoma,⁵⁷ K. Nakamura,⁵⁵ I. Nakano,⁴⁰
A. Napier,⁵⁶ V. Necula,¹⁶ C. Neu,⁴⁵ M. S. Neubauer,⁹ J. Nielsen,^{29,bl} L. Nodulman,² O. Norniella,³ E. Nurse,³¹ S. H. Oh,¹⁶
Y. D. Oh,²⁸ I. Oksuzian,¹⁸ T. Okusawa,⁴¹ R. Oldeman,³⁰ R. Orava,²³ K. Osterberg,²³ C. Pagliarone,⁴⁶ E. Palencia,¹¹
V. Papadimitriou,¹⁷ A. Papaikonomou,²⁶ A. A. Paramonov,¹³ B. Parks,³⁹ S. Pashapour,³⁴ J. Patrick,¹⁷ G. Paulette,⁵⁴
M. Paulini,¹² C. Paus,³³ D. E. Pellett,⁷ A. Penzo,⁵⁴ T. J. Phillips,¹⁶ G. Piacentino,⁴⁶ J. Piedra,⁴⁴ L. Pinera,¹⁸ K. Pitts,²⁴
C. Plager,⁸ L. Pondrom,⁵⁹ X. Portell,³ O. Poukhov,¹⁵ N. Pounder,⁴² F. Prakoshyn,¹⁵ A. Pronko,¹⁷ J. Proudfoot,²
F. Ptohos,^{19,bl} G. Punzi,⁴⁶ J. Pursley,²⁵ J. Rademacker,^{42,bl} A. Rahaman,⁴⁷ V. Ramakrishnan,⁵⁹ N. Ranjan,⁴⁸ I. Redondo,³²
B. Reisert,¹⁷ V. Rekovic,³⁷ P. Renton,⁴² M. Rescigno,⁵¹ S. Richter,²⁶ F. Rimondi,⁵ L. Ristori,⁴⁶ A. Robson,²¹ T. Rodrigo,¹¹
E. Rogers,²⁴ S. Rolli,⁵⁶ R. Roser,¹⁷ M. Rossi,⁵⁴ R. Rossin,¹⁰ P. Roy,³⁴ A. Ruiz,¹¹ J. Russ,¹² V. Rusu,¹³ H. Saarikko,²³
A. Safonov,⁵³ W. K. Sakumoto,⁴⁹ G. Salamanna,⁵¹ O. Saltó,³ L. Santi,⁵⁴ S. Sarkar,⁵¹ L. Sartori,⁴⁶ K. Sato,¹⁷ P. Savard,³⁴
A. Savoy-Navarro,⁴⁴ T. Scheidle,²⁶ P. Schlabach,¹⁷ E. E. Schmidt,¹⁷ M. P. Schmidt,⁶⁰ M. Schmitt,³⁸ T. Schwarz,⁷
L. Scodellaro,¹¹ A. L. Scott,¹⁰ A. Scribano,⁴⁶ F. Scuri,⁴⁶ A. Sedov,⁴⁸ S. Seidel,³⁷ Y. Seiya,⁴¹ A. Semenov,¹⁵
L. Sexton-Kennedy,¹⁷ A. Sfyrla,²⁰ S. Z. Shalhout,⁵⁸ M. D. Shapiro,²⁹ T. Shears,³⁰ P. F. Shepard,⁴⁷ D. Sherman,²²
M. Shimojima,^{55,bl} M. Shochet,¹³ Y. Shon,⁵⁹ I. Shreyber,²⁰ A. Sidoti,⁴⁶ P. Sinervo,³⁴ A. Sisakyan,¹⁵ A. J. Slaughter,¹⁷
J. Slaunwhite,³⁹ K. Sliwa,⁵⁶ J. R. Smith,⁷ F. D. Snider,¹⁷ R. Snihur,³⁴ M. Soderberg,³⁵ A. Soha,⁷ S. Somalwar,⁵² V. Sorin,³⁶
J. Spalding,¹⁷ F. Spinella,⁴⁶ T. Spreitzer,³⁴ P. Squillacioti,⁴⁶ M. Stanitzki,⁶⁰ A. Staveris-Polykalas,⁴⁶ R. St. Denis,²¹
B. Stelzer,⁸ O. Stelzer-Chilton,⁴² D. Stentz,³⁸ J. Strologas,³⁷ D. Stuart,¹⁰ J. S. Suh,²⁸ A. Sukhanov,¹⁸ H. Sun,⁵⁶ I. Suslov,¹⁵
T. Suzuki,⁵⁵ A. Taffard,^{24,bl} R. Takashima,⁴⁰ Y. Takeuchi,⁵⁵ R. Tanaka,⁴⁰ M. Tecchio,³⁵ P. K. Teng,¹ K. Terashi,⁵⁰
J. Thom,^{17,bl} A. S. Thompson,²¹ E. Thomson,⁴⁵ P. Tipton,⁶⁰ V. Tiwari,¹² S. Tkaczyk,¹⁷ D. Toback,⁵³ S. Tokar,¹⁴
K. Tollefson,³⁶ T. Tomura,⁵⁵ D. Tonelli,⁴⁶ S. Torre,¹⁹ D. Torretta,¹⁷ S. Tourneur,⁴⁴ W. Trischuk,³⁴ S. Tsuno,⁴⁰ Y. Tu,⁴⁵
N. Turini,⁴⁶ F. Ukegawa,⁵⁵ S. Uozumi,⁵⁵ S. Vallecorsa,²⁰ N. van Remortel,²³ A. Varganov,³⁵ E. Vataga,³⁷ F. Vazquez,^{18,bl}
G. Velev,¹⁷ C. Vellidis,^{46,bl} G. Veramendi,²⁴ V. Veszpremi,⁴⁸ M. Vidal,³² R. Vidal,¹⁷ I. Vila,¹¹ R. Vilar,¹¹ T. Vine,³¹
M. Vogel,³⁷ I. Vollrath,³⁴ I. Volobouev,^{29,bl} G. Volpi,⁴⁶ F. Würthwein,⁹ P. Wagner,⁵³ R. G. Wagner,² R. L. Wagner,¹⁷
J. Wagner,²⁶ W. Wagner,²⁶ R. Wallny,⁸ S. M. Wang,¹ A. Warburton,³⁴ D. Waters,³¹ M. Weinberger,⁵³ W. C. Wester III,¹⁷
B. Whitehouse,⁵⁶ D. Whiteson,^{45,bl} A. B. Wicklund,² E. Wicklund,¹⁷ G. Williams,³⁴ H. H. Williams,⁴⁵ P. Wilson,¹⁷
B. L. Winer,³⁹ P. Wittich,^{17,bl} S. Wolbers,¹⁷ C. Wolfe,¹³ T. Wright,³⁵ X. Wu,²⁰ S. M. Wynne,³⁰ A. Yagil,⁹ K. Yamamoto,⁴¹
J. Yamaoka,⁵² T. Yamashita,⁴⁰ C. Yang,⁶⁰ U. K. Yang,^{13,bl} Y. C. Yang,²⁸ W. M. Yao,²⁹ G. P. Yeh,¹⁷ J. Yoh,¹⁷ K. Yorita,¹³
T. Yoshida,⁴¹ G. B. Yu,⁴⁹ I. Yu,²⁸ S. S. Yu,¹⁷ J. C. Yun,¹⁷ L. Zanello,⁵¹ A. Zanetti,⁵⁴ I. Zaw,²² X. Zhang,²⁴
J. Zhou,⁵² and S. Zucchelli⁵

(CDF Collaboration)

¹Institute of Physics, Academia Sinica, Taipei, Taiwan 11529, Republic of China²Argonne National Laboratory, Argonne, Illinois 60439, USA³Institut de Fisica d'Altes Energies, Universitat Autònoma de Barcelona, E-08193, Bellaterra (Barcelona), Spain⁴Baylor University, Waco, Texas 76798, USA⁵Istituto Nazionale di Fisica Nucleare, University of Bologna, I-40127 Bologna, Italy⁶Brandeis University, Waltham, Massachusetts 02254, USA⁷University of California, Davis, Davis, California 95616, USA⁸University of California, Los Angeles, Los Angeles, California 90024, USA⁹University of California, San Diego, La Jolla, California 92093, USA¹⁰University of California, Santa Barbara, Santa Barbara, California 93106, USA¹¹Instituto de Fisica de Cantabria, CSIC-University of Cantabria, 39005 Santander, Spain¹²Carnegie Mellon University, Pittsburgh, Pennsylvania 15213, USA¹³Enrico Fermi Institute, University of Chicago, Chicago, Illinois 60637, USA¹⁴Comenius University, 842 48 Bratislava, Slovakia; Institute of Experimental Physics, 040 01 Kosice, Slovakia¹⁵Joint Institute for Nuclear Research, RU-141980 Dubna, Russia¹⁶Duke University, Durham, North Carolina 27708, USA¹⁷Fermi National Accelerator Laboratory, Batavia, Illinois 60510, USA¹⁸University of Florida, Gainesville, Florida 32611, USA¹⁹Laboratori Nazionali di Frascati, Istituto Nazionale di Fisica Nucleare, I-00044 Frascati, Italy²⁰University of Geneva, CH-1211 Geneva 4, Switzerland²¹Glasgow University, Glasgow G12 8QQ, United Kingdom²²Harvard University, Cambridge, Massachusetts 02138, USA

²³Division of High Energy Physics, Department of Physics, University of Helsinki and Helsinki Institute of Physics,
FIN-00014, Helsinki, Finland

²⁴University of Illinois, Urbana, Illinois 61801, USA

²⁵The Johns Hopkins University, Baltimore, Maryland 21218, USA

²⁶Institut für Experimentelle Kernphysik, Universität Karlsruhe, 76128 Karlsruhe, Germany

²⁷High Energy Accelerator Research Organization (KEK), Tsukuba, Ibaraki 305, Japan

²⁸Center for High Energy Physics: Kyungpook National University, Taegu 702-701, Korea;
Seoul National University, Seoul 151-742, Korea;
SungKyunKwan University, Suwon 440-746, Korea

²⁹Ernest Orlando Lawrence Berkeley National Laboratory, Berkeley, California 94720, USA

³⁰University of Liverpool, Liverpool L69 7ZE, United Kingdom

³¹University College London, London WC1E 6BT, United Kingdom

³²Centro de Investigaciones Energeticas Medioambientales y Tecnologicas, E-28040 Madrid, Spain

³³Massachusetts Institute of Technology, Cambridge, Massachusetts 02139, USA

³⁴Institute of Particle Physics: McGill University, Montréal, Canada H3A 2T8;
and University of Toronto, Toronto, Canada M5S 1A7

³⁵University of Michigan, Ann Arbor, Michigan 48109, USA

³⁶Michigan State University, East Lansing, Michigan 48824, USA

³⁷University of New Mexico, Albuquerque, New Mexico 87131, USA

³⁸Northwestern University, Evanston, Illinois 60208, USA

³⁹The Ohio State University, Columbus, Ohio 43210, USA

⁴⁰Okayama University, Okayama 700-8530, Japan

⁴¹Osaka City University, Osaka 588, Japan

⁴²University of Oxford, Oxford OX1 3RH, United Kingdom

⁴³Istituto Nazionale di Fisica Nucleare, University of Padova, Sezione di Padova-Trento, I-35131 Padova, Italy

⁴⁴LPNHE, Université Pierre et Marie Curie/IN2P3-CNRS, UMR7585, Paris, F-75252 France

⁴⁵University of Pennsylvania, Philadelphia, Pennsylvania 19104, USA

⁴⁶Istituto Nazionale di Fisica Nucleare Pisa, Universities of Pisa, Siena and Scuola Normale Superiore, I-56127 Pisa, Italy

⁴⁷University of Pittsburgh, Pittsburgh, Pennsylvania 15260, USA

⁴⁸Purdue University, West Lafayette, Indiana 47907, USA

⁴⁹University of Rochester, Rochester, New York 14627, USA

⁵⁰The Rockefeller University, New York, New York 10021, USA

⁵¹Istituto Nazionale di Fisica Nucleare, Sezione di Roma 1, University of Rome "La Sapienza," I-00185 Roma, Italy

⁵²Rutgers University, Piscataway, New Jersey 08855, USA

⁵³Texas A&M University, College Station, Texas 77843, USA

⁵⁴Istituto Nazionale di Fisica Nucleare, University of Trieste/Udine, Italy

⁵⁵University of Tsukuba, Tsukuba, Ibaraki 305, Japan

⁵⁶Tufts University, Medford, Massachusetts 02155, USA

⁵⁷Waseda University, Tokyo 169, Japan

⁵⁸Wayne State University, Detroit, Michigan 48201, USA

⁵⁹University of Wisconsin, Madison, Wisconsin 53706, USA

⁶⁰Yale University, New Haven, Connecticut 06520, USA

(Received 17 July 2007; published 9 November 2007)

We present a search for associated production of the chargino and neutralino supersymmetric particles using up to 1.1 fb^{-1} of integrated luminosity collected by the CDF II experiment at the Tevatron $p\bar{p}$ collider at $\sqrt{s} = 1.96 \text{ TeV}$. We analyze events with a large transverse momentum imbalance and either three charged leptons or two charged leptons of the same electric charge. The numbers of observed events are consistent with standard model expectations. Upper limits on the production cross section are derived in different theoretical models.

DOI: 10.1103/PhysRevLett.99.191806

PACS numbers: 14.80.Ly, 12.60.Jv, 13.85.Rm

Supersymmetry (SUSY) [1] is one of the most appealing theories for physics beyond the standard model (SM). SUSY predicts the existence of a superpartner (sparticle) for each SM particle, sharing the same quantum numbers but differing by half a unit of spin. SUSY addresses several problems of the SM: it can solve the "hierarchy problem" [2], it can provide a good candidate for the cold dark matter in the Universe [3], and it makes possible a unification of

the fundamental forces at high energies [4]. A natural solution to the hierarchy problem and the prospect of gauge coupling unification suggest that sparticle masses are near the electroweak scale and thus may be observable at the Tevatron.

We present a search for the associated production of the lightest chargino $\tilde{\chi}_1^\pm$ and the second-to-lightest neutralino $\tilde{\chi}_2^0$, the mass eigenstates of the superpartners of the elec-

tritweak gauge and Higgs boson. Chargino-neutralino production, $p\bar{p} \rightarrow \tilde{\chi}_1^\pm \tilde{\chi}_2^0 + X$, is one of the most interesting SUSY processes at the Tevatron in models in which the gauginos are light. It can be detected through the observation of isolated charged leptons [5] and a large imbalance in the transverse energy (\cancel{E}_T) from the decays $\tilde{\chi}_1^\pm \rightarrow \ell^\pm \nu \tilde{\chi}_1^0$ and $\tilde{\chi}_2^0 \rightarrow \ell^+ \ell^- \tilde{\chi}_1^0$, where $\ell = e, \mu, \tau$, and $\tilde{\chi}_1^0$ is the lightest SUSY particle, assumed to be stable and to escape detection. This signature has the experimental advantage that at hadron colliders leptons are relatively rare compared to the copiously produced jets, they are well identifiable, and the SM backgrounds are rather small, as they arise primarily from electroweak processes. We use three benchmark models based on the minimal supersymmetric standard model (MSSM) to interpret the data. The models differ mostly in the leptonic branching ratios of the $\tilde{\chi}_1^\pm$ and $\tilde{\chi}_2^0$ and in the kinematic properties of the leptons.

Previous searches at LEP excluded $\tilde{\chi}_1^\pm$ masses below 103.5 GeV/ c^2 [6] at the 95% confidence level (C.L.). This constraint is very robust and does not change much within minimal supergravity-inspired SUSY models. The D0 Collaboration recently constrained the $\tilde{\chi}_1^\pm$ mass to be larger than 117 GeV/ c^2 at 95% C.L. in a specific scenario [7].

This Letter summarizes and interprets the results from seven individual search channels with different lepton flavors and kinematic properties in the final state. The analyses use $p\bar{p}$ collision data taken by the CDF II detector at the Tevatron accelerator with a center-of-mass energy of $\sqrt{s} = 1.96$ TeV. The data, collected between March 2002 and February 2006, correspond to an integrated luminosity between 0.7 and 1.1 fb $^{-1}$, depending on the decay signature.

The CDF II detector [8] is cylindrically symmetric around the beam pipe in which the protons and antiprotons collide [5]. The transverse momentum of charged particles, p_T , is measured by a tracking system composed of an eight-layer silicon strip detector and a 96-layer drift chamber; both are located inside a solenoid providing a magnetic field of 1.4 T aligned along the beam axis. Electromagnetic and hadronic calorimeters surrounding the solenoid measure the energies of particles up to $|\eta| < 3.6$. Wire chambers and scintillators are installed around the hadronic calorimeter to detect muons with $|\eta| < 1.4$. Gas Cherenkov counters [9] measure the average number of $p\bar{p}$ inelastic collisions per bunch crossing and thereby determine the beam luminosity.

We now outline the seven individual analyses that are then combined to achieve maximum sensitivity. The like-sign (LS) analyses ($e^\pm e^\pm, e^\pm \mu^\pm, \mu^\pm \mu^\pm$) [10] require two leptons of the same electric charge and do not require the detection of a third lepton. In the trilepton analyses, the third lepton candidate can either be a fully reconstructed electron or muon ($e\ell\ell, \mu\ell\ell$ [11], $\mu\mu\ell$ [12], $\ell = e, \mu$) or a “track” from a charged particle (eet). The eet analysis gains acceptance in detector areas where the electron or

muon detection is less efficient and adds sensitivity to hadronic decays of τ leptons. There is up to 30% overlap among the individual analyses which is accounted for in the combination, as explained later in this Letter.

The dominant SM background sources are diboson production with three or more prompt leptons (WZ, ZZ) and Drell-Yan (DY) events in which the third lepton results either from the conversion of a bremsstrahlung photon ($Z + \gamma$, with $\gamma \rightarrow e^+ e^-$) or from a misidentified hadron. With Z we denote the production of a Z or a virtual photon γ^* . For the LS analyses $W\gamma$ production with a photon conversion is also a significant background. Smaller background contributions arise from $t\bar{t}$ and $b\bar{b}$ production with semileptonic b - and c -hadron decays. All these background sources are modeled using Monte Carlo (MC) event generators. Backgrounds from $t\bar{t}$, ZZ , and DY production are generated using PYTHIA [13]. The $W\gamma$ and WZ backgrounds are generated using for the hard-scattering process a program by Baur and Berger [14] and MADEVENT [15], respectively. In all cases PYTHIA is used for the parton showering and hadronization and the parton distribution functions are parametrized using CTEQ5 [16]. All MC events are subsequently passed through the CDF II detector simulation based on the GEANT3 [17] framework, and reconstructed and analyzed in the same way as the data. The $b\bar{b}$ background was determined using a combination of MC and data estimates [12] and is negligible in most of the analyses. An additional background source is hadrons that are misidentified as leptons (“mis-id”). We determine the misidentification probability in jet data samples as a function of the lepton p_T or E_T using jets and tracks and apply it to the two leptons (one lepton) data sample for the trileptons (LS) analyses [10,11].

In the LS, $e\ell\ell$, and $\mu\ell\ell$ analyses, events are triggered on one well-identified central electron with $E_T > 18$ GeV or muon with $p_T > 18$ GeV/ c . In the eet ($\mu\mu\ell$) analysis events are triggered on two central electrons (muons) with $E_T > 4$ GeV ($p_T > 4$ GeV/ c). We select additional electrons in the central and the forward calorimeters. They are required to have a shower shape consistent with that expected for an electron and a track matched to the calorimeter cluster. Muons are required to deposit an amount of energy in the calorimeter consistent with the expectation for a minimum ionizing particle; additionally, trigger muons must have associated hits in the muon detectors. Dedicated algorithms reject photon conversions and cosmic rays [11]. We require all leptons to be isolated from other particles in the event. For the electrons and muons in the $e\ell\ell$, $\mu\ell\ell$, eet , and $\mu\mu\ell$ analyses the isolation requirement is based only on the calorimeter energy deposits. For the track of the eet analysis it is based only on charged tracks, and for electrons and muons in the LS analyses it is based on both.

We exclude events in which two leptons form an invariant mass $m_{\ell\ell} < 15$ GeV/ c^2 or $76 < m_{\ell\ell} < 106$ GeV/ c^2 in

order to remove DY and diboson events. The lower mass threshold additionally removes $b\bar{b}$ background. For the $e\ell\ell$ (LS) analysis the lower mass value is raised to $20(25)\text{ GeV}/c^2$. For the LS analyses the mass interval near the Z resonance is changed to $66 < m_{\ell\ell} < 116\text{ GeV}/c^2$. Backgrounds from DY production are further reduced by requiring $\cancel{E}_T > 15\text{ GeV}$ (for the eet analysis $\cancel{E}_T > 20\text{ GeV}$). The $t\bar{t}$ background is reduced by vetoing events with large hadronic jet activity. Requirements on the angles between the leptons and \cancel{E}_T are placed in order to reduce the cosmic-ray background. A detailed description of the selection requirements is given elsewhere [10–12].

To illustrate the model sensitivity of the search we use three example SUSY models. The first is the scenario of minimal supergravity (mSUGRA) [18], a grand unified theory including gravity which has five independent parameters, fully determining all the masses and couplings of the SUSY particles. Since the present analysis is most sensitive to the common gaugino mass $m_{1/2}$, we fix the other four parameters: the common scalar mass is set to $m_0 = 60\text{ GeV}/c^2$, the higgsino mixing parameter (μ) is taken to be positive, the trilinear coupling (A_0) is set to 0, and the ratio of the vacuum expectation values of the two Higgs fields ($\tan\beta$) is set to 3. In mSUGRA the lightest slepton is a SUSY partner of the τ leading to a larger branching ratio (\mathcal{B}) into τ leptons: about 90% of the events contain at least one τ lepton. The second model we call “MSSM (W/Z model)”: all the parameters are taken to be the same as in the above model, but the \mathcal{B} of $\tilde{\chi}_1^\pm(\tilde{\chi}_2^0)$ into leptons is fixed to be the same as the \mathcal{B} of the $W(Z)$ gauge boson into leptons. In the third model, the “MSSM (no $\tilde{\ell}$ mixing)”, the slepton chirality eigenstates are the mass eigenstates, resulting in nearly equal branching ratios of $\tilde{\chi}_1^\pm\tilde{\chi}_2^0$ to all three lepton flavors. In this model about 30% of the events contain no τ leptons. For all three scenarios the relationship between the masses of the gauginos is $m_{\tilde{\chi}_1^\pm} \approx m_{\tilde{\chi}_2^0} \approx 2m_{\tilde{\chi}_1^0}$. The slepton mass value is approximately $0.31 \times m_{\tilde{\chi}_1^\pm} + 67\text{ GeV}/c^2$. For the first and second models, the most stringent constraint on the $\tilde{\chi}_1^\pm$ mass to date is

the LEP limit of $103.5\text{ GeV}/c^2$ while for the third model the most stringent limit is $117\text{ GeV}/c^2$ [7].

For the signal simulation we use SOFTSUSY [19] and ISAJET [20] to compute the sparticle mass spectrum; PYTHIA is used to generate the hard-scattering events, the parton radiation, and the hadronization. CTEQ5 is used for the parton distribution functions. The CDF II detector’s response to these events is then simulated. The signal acceptance for the process $p\bar{p} \rightarrow \tilde{\chi}_1^\pm\tilde{\chi}_2^0 + X \rightarrow \ell\ell\ell\nu\tilde{\chi}_1^0\tilde{\chi}_1^0 + X$, ($\ell = e, \mu, \tau$), in the MSSM (no $\tilde{\ell}$ mixing) scenario varies from 2.7% at $m_{\tilde{\chi}_1^\pm} \approx 105\text{ GeV}/c^2$, where only the LS analyses are sensitive, to 6.2% at $m_{\tilde{\chi}_1^\pm} \approx 160\text{ GeV}/c^2$, where the acceptance of the trilepton analyses is maximal. The acceptance in the MSSM (W/Z model) rises from 3.0% at low $m_{\tilde{\chi}_1^\pm}$ to 4.0% at $m_{\tilde{\chi}_1^\pm} \approx 150\text{ GeV}/c^2$. The acceptance for mSUGRA is only 1.0%, independent of $m_{\tilde{\chi}_1^\pm}$, due to the enhanced decay of $\tilde{\chi}_1^\pm\tilde{\chi}_2^0$ into τ leptons.

There are systematic uncertainties on both the signal acceptance and the background prediction that have been evaluated by many comparisons of data and simulation [10–12]. The signal acceptance has an uncertainty due to the lepton selection (1.5%–13%) and trigger (<0.5%) efficiencies, the modeling of QCD radiation (2%–12%), the parton distribution functions (1%), the integrated luminosity (6%), and the jet energy measurement (1%–5%). The background is also affected by the lepton- and jet-related uncertainties and the luminosity uncertainty. Additionally, we consider uncertainties on the lepton misidentification rate (50% for the trilepton and 10%–20% for the LS analyses), the conversion background (3%–16%), and the theoretical cross sections (7% for diboson, 10% for $t\bar{t}$, and 5% for DY production). Finally, the statistical uncertainties on the Monte Carlo samples are taken into account.

In Table I, the number of observed events is compared to the background contributions for each analysis. The number of data events is consistent with the background expectation in all analyses and no evidence of non-SM

TABLE I. The numbers of expected and observed events for each analysis before events which are shared by more than one analysis are assigned to a single analysis. For the eet analysis the background from misidentified jets is included in the $Z + \gamma$ background estimate of this analysis.

	$e^\pm e^\pm$	$e^\pm \mu^\pm$	$\mu^\pm \mu^\pm$	$e\ell\ell$	$\mu\ell\ell$	eet	$\mu\mu\ell$
$Z + \gamma$	0.49	0.62	—	0.18	0.30	0.54	0.06
$W\gamma$	1.54	1.63	—	—	—	—	—
$t\bar{t} + b\bar{b}$	0.01	0.03	0.01	0.04	0.03	0.22	0.06
WW, WZ, ZZ	0.32	0.82	0.53	0.39	0.66	0.21	0.09
Mis-id	0.60	0.90	0.38	0.14	0.27	—	0.20
Total	2.96	4.00	0.92	0.75	1.26	0.97	0.41
Uncertainty	± 0.48	± 0.57	± 0.12	± 0.36	± 0.27	± 0.28	± 0.11
Observed	4	8	1	0	1	3	1

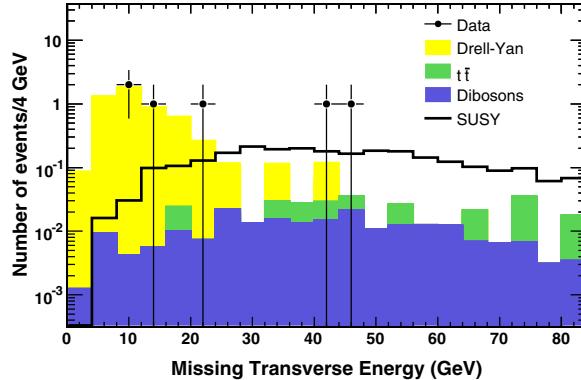


FIG. 1 (color online). \cancel{E}_T distribution for the eet analysis after all the selection criteria are applied except for $\cancel{E}_T > 20$ GeV. The observed data (points) are compared to the sum of the SM contributions (stacked, filled histograms). The open histogram shows the expected SUSY contribution for the mSUGRA scenario described in the text with $m_{\tilde{\chi}_1^\pm} = 113$ GeV/ c^2 and $\sigma \times \mathcal{B} = 0.16$ pb.

physics is observed. There is a slight excess of the data in the $e^\pm \mu^\pm$ and eet analyses. Figure 1 shows the \cancel{E}_T distribution for the eet analysis. The observed data agree well with SM predictions at low values of \cancel{E}_T where the background is dominant.

The data can be used to constrain the cross section times branching ratio, $\sigma \times \mathcal{B}$, and the allowed mass range of charginos for the SUSY scenarios discussed earlier. To remove the overlaps among the analyses, events which are selected by more than one analysis are assigned to the channel with the highest *a priori* signal-to-background ratio to obtain the best sensitivity. The frequentist-based “CL_s” method [21,22] is then used on the resulting non-overlapping analyses. The correlations between the systematic uncertainties of the individual analyses have been evaluated for both the signal and the background and are accounted for in the limit calculation. A mass limit for the $\tilde{\chi}_1^\pm$ is derived from the cross section limit by comparing the observed limit to the next-to-leading order calculation for the cross section [23]. An uncertainty of 10% on the signal theoretical cross section is included in the limit calculation [24].

Figure 2 shows the 95% C.L. upper limit on the $\sigma \times \mathcal{B}$ as a function of the $\tilde{\chi}_1^\pm$ mass in the three theoretical scenarios mSUGRA, MSSM (W/Z model), and MSSM (no $\tilde{\ell}$ mixing). In each scenario the observed and the expected limits are compared with the theoretical $\sigma \times \mathcal{B}$ predictions. The expected limit is defined to be the median limit one would obtain in a sample of independent experiments in which no signal is present. It is used to estimate the *a priori* sensitivity of the search since it does not depend on the observed data. The observed limit is typically a factor of 2 larger than the expected, as the number of events observed is higher than the predicted background in the eet and $e^\pm \mu^\pm$ analyses. In the mSUGRA scenario

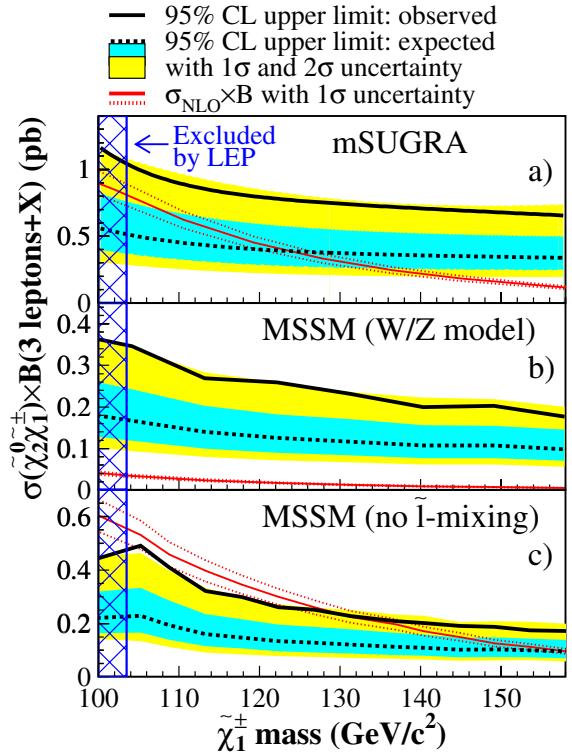


FIG. 2 (color online). $\sigma \times \mathcal{B}$ for $p\bar{p} \rightarrow \tilde{\chi}_1^\pm \tilde{\chi}_2^0 + X \rightarrow \ell\ell\ell + X$ as function of the $\tilde{\chi}_1^\pm$ mass. Shown are the observed and expected experimental upper limits and the theoretical cross section for the mSUGRA (a), the MSSM (W/Z model) (b), and the MSSM (no $\tilde{\ell}$ mixing) (c) scenario described in the text. Also shown are the 1σ and 2σ ranges for the expected limit.

[Fig. 2(a)] $\sigma \times \mathcal{B} > 1$ pb is excluded for all $m_{\tilde{\chi}_1^\pm}$. The limit improves with increasing mass down to 0.8 pb at $m_{\tilde{\chi}_1^\pm} = 150$ GeV/ c^2 . In the MSSM W/Z model and no $\tilde{\ell}$ mixing scenarios [Fig. 2(b) and 2(c)] the limits on $\sigma \times \mathcal{B}$ range between 0.2 and 0.4 pb. In the MSSM (no $\tilde{\ell}$ mixing) scenario [Fig. 2(c)] we set a 95% C.L. lower limit on the $\tilde{\chi}_1^\pm$ mass of 129 GeV/ c^2 , which is the most stringent limit to date. In the mSUGRA and MSSM (no $\tilde{\ell}$ mixing) scenarios the expected $\tilde{\chi}_1^\pm$ mass limits are 122 GeV/ c^2 and 157 GeV/ c^2 , respectively.

In conclusion, we present a search for the associated production of charginos and neutralinos using events with isolated leptons and large \cancel{E}_T . The observed data are consistent with standard model predictions. Limits are derived on the cross section times the leptonic branching ratio for this process and on the mass of the lightest chargino under different theoretical assumptions. In the MSSM (no $\tilde{\ell}$ mixing) scenario, chargino masses below 129 GeV/ c^2 are excluded at the 95% C.L., extending the previous limits by 12 GeV/ c^2 .

We thank the Fermilab staff and the technical staffs of the participating institutions for their vital contributions. We thank T. Plehn and B. Allanach for their help on the theoretical interpretation. This work was supported by the

US Department of Energy and National Science Foundation; the Italian Istituto Nazionale di Fisica Nucleare; the Ministry of Education, Culture, Sports, Science, and Technology of Japan; the Natural Sciences and Engineering Research Council of Canada; the National Science Council of the Republic of China; the Swiss National Science Foundation; the A.P. Sloan Foundation; the Bundesministerium für Bildung und Forschung, Germany; the Korean Science and Engineering Foundation and the Korean Research Foundation; the Particle Physics and Astronomy Research Council and the Royal Society, UK; the Institut National de Physique Nucléaire et Physique des Particules/CNRS; the Russian Foundation for Basic Research; the Comisión Interministerial de Ciencia y Tecnología, Spain; the European Community's Human Potential Programme under Contract No. HPRN-CT-2002-00292; and the Academy of Finland.

^aVisiting scientist from University of Athens, 15784 Athens, Greece.

^bVisiting scientist from University of Bristol, Bristol BS8 1TL, United Kingdom.

^cVisiting scientist from University Libre de Bruxelles, B-1050 Brussels, Belgium.

^dVisiting scientist from Cornell University, Ithaca, NY 14853, USA.

^eVisiting scientist from University of Cyprus, Nicosia CY-1678, Cyprus.

^fVisiting scientist from University College Dublin, Dublin 4, Ireland.

^gVisiting scientist from University of Edinburgh, Edinburgh EH9 3JZ, United Kingdom.

^hVisiting scientist from University of Heidelberg, D-69120 Heidelberg, Germany.

ⁱVisiting scientist from Universidad Iberoamericana, Mexico D.F., Mexico.

^jVisiting scientist from University of Manchester, Manchester M13 9PL, United Kingdom.

^kVisiting scientist from Nagasaki Institute of Applied Science, Nagasaki, Japan.

^lVisiting scientist from University de Oviedo, E-33007 Oviedo, Spain.

^mVisiting scientist from University of London, Queen Mary College, London, E1 4NS, United Kingdom.

ⁿVisiting scientist from University of California, Santa Cruz, Santa Cruz, CA 95064, USA.

^oVisiting scientist from Texas Tech University, Lubbock, TX 79409, USA.

^pVisiting scientist from University of California, Irvine, Irvine, CA 92697, USA.

^qVisiting scientist from IFIC (CSIC-Universitat de Valencia), 46071 Valencia, Spain.

- [1] J. Wess and B. Zumino, Nucl. Phys. **B70**, 39 (1974).
- [2] E. Witten, Nucl. Phys. **B188**, 513 (1981); N. Sakai, Z. Phys. C **11**, 153 (1981); S. Dimopoulos, Nucl. Phys. **B193**, 150 (1981).
- [3] J. Ellis, J.S. Hagelin, D.V. Nanopoulos, K. Olive, and M. Srednicki, Nucl. Phys. **B238**, 453 (1984); H. Goldberg, Phys. Rev. Lett. **50**, 1419 (1983).
- [4] S. Dimopoulos, S. Raby, and F. Wilczek, Phys. Rev. D **24**, 1681 (1981).
- [5] We use a cylindrical coordinate system where the z axis is along the proton beam direction and θ is the polar angle. Pseudorapidity is $\eta = -\ln \tan(\theta/2)$, while transverse momentum is $p_T = |p| \sin \theta$ and transverse energy is $E_T = E \sin \theta$. Missing transverse energy, \not{E}_T , is defined as the magnitude of $-\sum_i E_T^i \hat{n}_i$, where \hat{n}_i is the unit vector in the azimuthal plane that points from the beam line to the i th calorimeter tower. The quantity E_T^{iso} (p_T^{iso}) is defined as the scalar sum of the transverse energies (momenta) within a cone of $\Delta R = \sqrt{\Delta \eta^2 + \Delta \phi^2} < 0.4$ around the selected electron (muon or track). For electrons (muons, tracks) we require the E_T^{iso} to be less than 2 GeV or 10% of the electron E_T (muon or track p_T).
- [6] LEP SUSY Working Group Report No. LEPSUSYWG/01-03.1, 2001.
- [7] V. Abazov *et al.* (D0 Collaboration), Phys. Rev. Lett. **95**, 151805 (2005).
- [8] D. Acosta *et al.* (CDF Collaboration), Phys. Rev. D **71**, 032001 (2005), and references therein.
- [9] D. Acosta *et al.*, Nucl. Instrum. Methods Phys. Res., Sect. A **461**, 540 (2001).
- [10] A. Abulencia *et al.* (CDF Collaboration), Phys. Rev. Lett. **98**, 221803 (2007).
- [11] A. Aaltonen *et al.*, Phys. Rev. D. (to be published).
- [12] A. Aaltonen *et al.*, Phys. Rev. D. (to be published).
- [13] T. Sjöstrand, L. Lönnblad, and S. Mrenna, LU-TP-01-21, arXiv:hep-ph/0108264 (we use version 6.3).
- [14] U. Baur and E.L. Berger, Phys. Rev. D **41**, 1476 (1990).
- [15] F. Maltoni and T. Stelzer, J. High Energy Phys. 02 (2003) 027.
- [16] H.L. Lai *et al.*, Eur. Phys. J. C **12**, 375 (2000).
- [17] R. Brun and F. Carminati, CERN Program Library Long Writeup Report No. W5013, 1993.
- [18] H.P. Nilles, Phys. Rep. **110**, 1 (1984).
- [19] B.C. Allanach, Comput. Phys. Commun. **143**, 305 (2002); arXiv:hep-ph/0104145 (we use version 2.0.7).
- [20] F. Paige *et al.*, arXiv:hep-ph/0312045 (we use version 7.71).
- [21] T. Junk, Nucl. Instrum. Methods Phys. Res., Sect. A **434**, 435 (1999).
- [22] A.L. Read, J. Phys. G **28**, 2693 (2002).
- [23] W. Beenakker, R. Hopker, and M. Spira, arXiv:hep-ph/9611232.
- [24] T. Plehn and M. Spira (private communication).