## **Time-Dependent Nuclear Measurements of Mix in Inertial Confinement Fusion**

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The first time-dependent nuclear measurements of turbulent mix in inertial confinement fusion have been obtained. Implosions of spherical deuterated-plastic shells filled with pure  ${}^{3}$ He gas require atomic-scale mixing of the shell and gas for the D- ${}^{3}$ He nuclear reaction to proceed. The time necessary for Rayleigh-Taylor (RT) growth to induce mix delays peak nuclear production time, compared to equivalent capsules filled with a D<sub>2</sub>- ${}^{3}$ He mixture, by 75  $\pm$  30 ps, equal to half the nuclear burn duration. These observations indicate the likelihood of atomic mix at the tips of core-penetrating RT spikes.

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Ignition and high gain in inertial confinement fusion (ICF) [1,2] are critically dependent on mitigation of the Rayleigh-Taylor (RT) instability. ICF capsules typically consist of a spherical shell filled with a gaseous fuel, and are imploded using lasers (direct-drive) or x rays (indirectdrive) to rapidly deposit energy and ablate the capsule surface. The RT instability, which is the growth of nonuniformities at a density interface when a low-density material accelerates a high-density material, occurs during two distinct intervals in ICF implosions. During the acceleration phase, the low-density ablating plasma accelerates the solid shell inwards, and perturbations seeded by energy deposition nonuniformities or initial capsule surface roughness feeds through to the inner fuel-shell surface. During the deceleration phase, shortly before the time of maximum capsule compression, growth of the RT instability at the fuel-shell interface quickly saturates, resulting in small-scale, turbulent eddies that leads to atomic-scale mixing of the fuel and shell [3]. RT growth and the resulting mixing processes disrupt the formation of the hot-spot in the fuel, lowering its temperature and reducing its volume, which may prevent the capsule from igniting. Understanding the nature and timing of RT growth and mix under different conditions is an important step toward mitigating their adverse effects.

Substantial and sustained efforts to understand the RT instability and mix have been ongoing for many decades [4]. This Letter presents the first time-dependent nuclear burn measurements of the mix region in ICF implosions. Although it has been previously demonstrated that there is no mix in the burn region at shock-bang time [5,6], it was unknown how long after shock collapse it takes for atomic mixing to occur. Other relevant work on the mix region in ICF implosions includes time-integrated nuclear yield measurements in both direct-drive [6–10] and indirect-drive [11] configurations, as well as time-dependent

x-ray measurements of capsules doped with tracer elements [12]. Time-dependent nuclear measurements obtained from implosions of CD-shelled capsules filled with nearly pure tritium have recently been reported [13].

This Letter reports results from direct-drive experiments conducted with the OMEGA laser system [14], with 60 fully smoothed [15] UV ( $\lambda = 351$  nm) beams in a 1 ns flattop pulse and a total energy of 23 kJ. The on-target illumination uniformity was typically  $\leq 2\%$  rms. The spherical plastic target capsules had diameters between 860 and 880  $\mu$ m, a total shell thickness of 20  $\mu$ m, and a 0.1  $\mu$ m rms outer surface roughness. "CH" capsules had shells made of plastic (CH), and a gaseous fill of deuterium and helium-3 (D<sub>2</sub> and <sup>3</sup>He, equimolar by atom). "CD" capsules had gaseous fills of pure <sup>3</sup>He, and a shell made mostly of CH, except for a 1  $\mu$ m layer of deuterated plastic (CD) on the inner surface (Fig. 1). The fill pressures of the D<sub>2</sub>-<sup>3</sup>He and the pure <sup>3</sup>He mixtures in CH and CD capsules were chosen to give equal initial fill mass densities  $\rho_0$ , at values of 0.5 or 2.5 mg/cm<sup>3</sup>. Because fully ionized D and <sup>3</sup>He have the same value of (1 + Z)/A, mixtures with the same mass density have the same total particle density when fully ionized, and can be considered hydrodynamically equivalent [16].

Implosions of CH and CD capsules were observed using simultaneous measurements of products from two distinct primary nuclear reactions to study the nature and timing of mix. The D<sup>3</sup>He reaction, D +  $^3$ He  $\rightarrow$   $^4$ He + p, and the DD-n reaction, D + D  $\rightarrow$   $^3$ He + n, have dramatically different composition and temperature sensitivities [16], which is used herein to constrain possible mix scenarios. The D<sup>3</sup>He reaction depends much more strongly on temperature due to the doubly-charged  $^3$ He reactant, and when the reactant species are initially separated, such as in CD capsules, they must be mixed before nuclear production will occur [17].

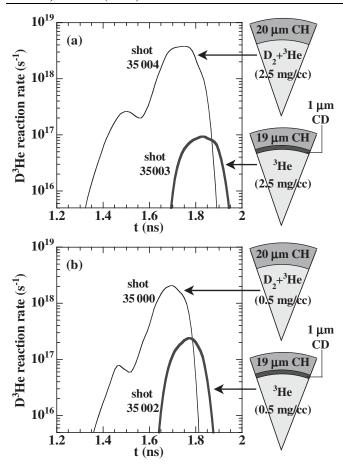


FIG. 1. Measurements of the D<sup>3</sup>He nuclear reaction history from implosions of spherical plastic (CH) shells filled with an equimolar D<sub>2</sub>-<sup>3</sup>He mixture, and of equivalent CD-layer capsules filled with pure <sup>3</sup>He. The gaseous fuel was filled to initial densities of (a) 2.5 mg/cm<sup>3</sup> and (b) 0.5 mg/cm<sup>3</sup>. The CH capsule histories show distinct times of D<sup>3</sup>He nuclear production corresponding to the shock (at  $\sim 1.5$  ns) and compression ( $\sim 1.75$  ns) burns. CD capsule implosions require mixing of the fuel and shell on the atomic scale for D<sup>3</sup>He production, and the histories show that no such mix has occurred at shock-bang time. The time necessary for hydroinstabilities to induce fuel-shell mix results in a typical 75  $\pm$  30 ps delay in the peak D<sup>3</sup>He reaction rate in CD capsules compared to equivalent CH capsules. In addition, nuclear production in CD implosions continues even after the compression burn ends in CH capsules, staying well above the typical noise level of  $3 \times 10^{15}/s$  for an additional 50 ps.

Possible scenarios of atomic mix are constrained using spectral measurements of nascent 14.7 MeV D<sup>3</sup>He protons. D<sup>3</sup>He protons experience energy loss from their birth energy as they pass through the compressed shell plasma on their way out of the capsule. The proton-emission-path-averaged capsule areal density  $\rho L$  is inferred using the mean-energy downshift of measured proton spectra [18]. For implosions with the same mean radial areal density  $\rho R$ , the value of  $\rho L$  depends on the spatial distribution of the proton source and shell mass [18]. A larger correction factor is needed as the mean source radius approaches the

mean shell radius, as protons traverse longer paths through the shell. For example, a quasi-one-dimensional scenario of atomic mix that consists of a spherical mixing layer just inside of a compressed spherical shell will require a much larger correction factor than a three-dimensional scenario involving turbulent mix induced at the tips of RT spikes driven into the hot core.

The dynamics of RT growth are of essential importance for any mix scenario. These dynamics are studied using temporal measurements of the emission of D<sup>3</sup>He protons, obtained using the proton temporal diagnostic (PTD) [19]. The PTD primarily consists of a 1-mm thick BC-422 scintillator, an optical transport system, and a fast streak camera. Optical fiducial pulses simultaneously recorded by the streak camera give an absolute timing accuracy of ~25 ps. The time history of the proton arrival at the scintillator is obtained by deconvolution of the detector response from the streak camera image. D<sup>3</sup>He proton spectral measurements [18] are then used to infer the D<sup>3</sup>He reaction rate history from the proton current at the scintillator. Additional details on PTD instrumentation and data processing are described by Frenje *et al.* [19].

Temporal measurements of 2.45-MeV neutrons from the DD-*n* nuclear reaction were obtained using the neutron temporal diagnostic (NTD) [20], which works on the same principle as the PTD, but is optimized for neutron detection. Although the DD reaction rate in CD capsule implosions is below the noise floor of the NTD, integrated DD yields were readily obtained using time-of-flight neutron detectors [21].

Implosions of CH capsules with D<sup>3</sup>He fuel characteristically emit D<sup>3</sup>He protons at two distinct times (Fig. 1). The shock burn is induced by the collapse of an ingoing spherical shock, and occurs before the imploding shell starts to decelerate. About 250 ps later, during the deceleration phase, the compression burn occurs as the imploding capsule compresses and reheats the fuel. In contrast to the two stages of proton emission observed in CH capsule implosions with D<sup>3</sup>He fills, CD capsules emit protons only during the later phase (Fig. 1), confirming the hypothesis that there is no mix at the time of shock collapse [22] first presented by Petrasso *et al.* [5].

Measurements of time-integrated nuclear yields demonstrate that capsules with lower  $\rho_0$  have an increased susceptibility to mix [6,9]. Yields increased with lower  $\rho_0$  for CD capsule implosions, even though low  $\rho_0$  is less favorable for nuclear production in the capsule core, as seen through the decrease in yields for CH capsules (see Fig. 1 and Table I). The increase in yields for lower  $\rho_0$  cannot be attributed to a difference in the temperature profile because both DD-n and D<sup>3</sup>He yields increased by about the same factor of 1.8, despite markedly different composition and temperature dependence. Additional mix of <sup>3</sup>He with the CD shell in low  $\rho_0$  implosions must be invoked to explain the yield trends [23].

The time necessary for RT growth to induce turbulent, atomic-scale mixing of the fuel and shell results in a delay

TABLE I. The number of shots in different ensembles of implosions of D<sup>3</sup>He-filled CH capsules and <sup>3</sup>He-filled CD capsules with two values of initial fill density  $\rho_0$  is shown, along with ensemble averages and standard errors of the mean for several experimental observables: bang time and burn duration for DD-n and D<sup>3</sup>He nuclear reaction histories, time-integrated DD-n and D<sup>3</sup>He yields ( $Y_n$  and  $Y_p$ ), and areal density  $\rho L$ . Standard errors are quoted in the same units as the averages, except for the yields, which are expressed as a percent. Only the compression component is included for  $Y_p$  and  $\rho L$  in CH capsules.

Туре	$ ho_0 \ (\text{mg/cm}^3)$	N shots	DD bang (ps)	DD burn (ps)	D <sup>3</sup> He bang (ps)	D <sup>3</sup> He burn (ps)	$Y_n \ (\times 10^8)$	Error (%)	$Y_p \ (\times 10^7)$	Error (%)	$\rho L \over (\text{mg/cm}^2)$
СН	2.5	8	$1749 \pm 24$	$157 \pm 10$	$1734 \pm 19$	$155 \pm 11$	129	6	61	10	54 ± 2
CH	0.5	8	$1697 \pm 22$	$148 \pm 11$	$1704 \pm 15$	$123 \pm 12$	29	9	30	16	$61 \pm 2$
CD	2.5	7	-	-	$1817 \pm 31$	$154 \pm 15$	5.1	9	1.7	11	$64 \pm 4$
CD	0.5	5	-	-	$1772 \pm 15$	$153 \pm 13$	9.4	7	3.0	13	66 ± 4

in the bang time (defined as the time of peak D³He reaction rate) of CD capsules compared to equivalent CH capsules of 83  $\pm$  37 ps and 69  $\pm$  21 ps for high and low  $\rho_0$  (Fig. 2 and Table I), respectively; this is equal to about half the typical 150 ps burn duration (defined as the full temporal width above half peak reaction rate). The delay is calculated as the difference between the ensemble averages of CD and CH capsule bang times, and the error is calculated as the quadrature sum of the standard errors of the mean for each ensemble average.

Measurements of DD-n bang time in CH capsules closely match the observed D<sup>3</sup>He bang time (Fig. 2). However, the DD reaction rate in CD capsules was too low for robust timing measurements.

The observed bang-time delay is not an artifact of limitations of the diagnostics or experimental setup. The timing jitter of the PTD is the same for CH and CD implosions, and is less than 20 ps, while bang-time errors of only 10 ps are introduced in the deconvolution process by proton energy spectrum uncertainties. A small systematic difference in shell thickness between CH and CD capsules was corrected using a 13 ps adjustment to the bang-time delay

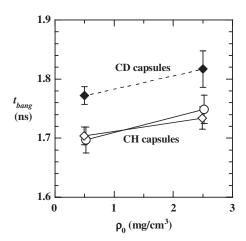


FIG. 2. Mean and standard error of  $D^3He$  (diamonds) and DD-n (circles) compression-bang times from CH (open markers) and CD (solid markers) capsule implosions as a function of initial fill density. In CD capsules,  $D^3He$  bang time consistently occurs  $\sim$ 75 ps later than in CH capsules.

[24], and it has been demonstrated that bang time does not depend on potential differences in implosion dynamics between capsules with pure  ${}^{3}$ He fuel and capsules with  $D_{2}$ - ${}^{3}$ He mixtures [16].

The observed delay of the peak reaction rate for CD capsules is likely due to the difference in how mix affects nuclear production in CH and CD capsules. Whereas mix tends to quench nuclear production in CH capsules through dilution and cooling of the hot fuel, in CD capsules, mix enhances nuclear production by the addition and heating of the D reactant from the shell. Nuclear production in CD capsules does not occur until later in the deceleration phase, when the growth of the RT instability has had time to induce turbulent mixing. Enhancement of reactant densities in CD capsules by continued mix in the later

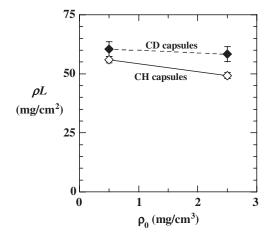


FIG. 3. Mean and standard error of proton-emission-pathaveraged areal densities  $(\rho L)$  for CH (open markers) and CD (solid markers) implosions as a function of initial fill density.  $D^3He$  proton spectral measurements are used to infer this compression-burn averaged  $\rho L$ , where the shock component of CH implosion spectra has been excluded. For CH capsules, the radial areal density  $(\rho R)$  can be obtained from  $\rho L$  using a small correction  $(\rho R \sim 0.93 \rho L)$  which depends on the shell aspect ratio. The relation between  $\rho R$  and  $\rho L$  for CD capsules sensitively depends on the source profile as the mean source radius approaches the mean shell radius—that  $\rho L$  is not much higher than in CH capsules suggests that the source profile is still centrally peaked.

stages of compression, combined with the larger total mass of fuel in such capsules, is enough to prolong nuclear production even after production would have been quenched in a CH implosion (Fig. 1).

Furthermore, systematically later nuclear production in CD capsule implosions leads to higher expected  $\rho L$ . The mean radial areal density  $\rho R$  increases throughout the deceleration phase as the shell continues to compress, so protons will selectively sample higher  $\rho R$  (and  $\rho L$ ) if they are emitted later in time. This effect is in addition to the potentially higher  $\rho L$  for CD capsules from geometric effects due to a noncentralized proton source profile, described above.

As seen in Fig. 3 and Table I,  $\rho L$  is 9% and 18% higher for implosions of CD capsules than for equivalent CH capsules with low and high  $\rho_0$ , respectively [25]. These values are not much higher, suggesting that one or both of the effects described above might not be as significant as expected. On this basis we conjecture that the source of protons in CD capsules may be dominated by atomic mixing at the tips of RT spikes from the shell that drive into the hot core, which would result in a more central proton emission profile, and a smaller increase in  $\rho L$ .

In summary, temporal measurements of D<sup>3</sup>He protons emitted from ICF implosions of CD-shelled, <sup>3</sup>He-filled capsules offer new and valuable insights into the dynamics of turbulent mixing induced by saturation of the Rayleigh-Taylor instability. The first such measurements have demonstrated that bang time is substantially delayed as RT growth saturates to produce mix. The  $83 \pm 37$  ps bangtime delay of CD implosions compared to D<sup>3</sup>He-filled, CH implosions for high initial fill densities  $(\rho_0)$  is equal to half the burn duration. Reducing  $\rho_0$  by a factor of 5 increases the susceptibility of the implosion to mix, and does not significantly affect the bang-time delay, observed to be  $69 \pm 21$  ps. Continued mixing of the fill gas and shell prolongs nuclear production in CD capsules even after it is quenched in equivalent CH capsules. Finally, the relatively small increase in areal density  $\rho L$  of CD compared to CH capsules, despite the later bang time, suggests that nuclear production is dominated by mixing induced at the tips of RT spikes driven into the core.

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- [22] Or that such mix, if present, has been insufficiently heated to give nuclear production.
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- [24] The 13 ps reduction in the delay corrects for a  $1/3~\mu m$  systematic difference in the total thickness of the CH and CD shells, where the timing of each burn history was adjusted by  $(40~ps) \times (20-\Delta)$ , with  $\Delta$  as the capsule thickness in  $\mu m$ . The 40 ps/ $\mu m$  correction factor was obtained by a linear fit of CH capsule bang times over a range of thicknesses from 15 to 27  $\mu m$ .
- [25] The slightly higher (<2%) initial shell mass in CD capsules due to the high density of the 1  $\mu$ m thick CD layer and the systematic thickness difference has a minimal impact on  $\rho L$  (<1%).