

Superradiant Rayleigh Scattering and Collective Atomic Recoil Lasing in a Ring Cavity

S. Slama, S. Bux, G. Krenz, C. Zimmermann, and Ph. W. Courteille

Physikalisches Institut, Eberhard-Karls-Universität Tübingen, Auf der Morgenstelle 14, D-72076 Tübingen, Germany

(Received 25 October 2006; published 1 February 2007)

Collective interaction of light with an atomic gas can give rise to superradiant instabilities. We experimentally study the sudden buildup of a reverse light field in a laser-driven high-finesse ring cavity filled with ultracold thermal or Bose-Einstein condensed atoms. While superradiant Rayleigh scattering from atomic clouds is normally observed only at very low temperatures (i.e., well below 1 μK), the presence of the ring cavity enhances cooperativity and allows for superradiance with thermal clouds as hot as several 10 μK . A characterization of the superradiance at various temperatures and cooperativity parameters allows us to link it to the collective atomic recoil laser.

DOI: [10.1103/PhysRevLett.98.053603](https://doi.org/10.1103/PhysRevLett.98.053603)

PACS numbers: 42.50.Gy, 03.75.-b, 34.50.-s, 42.60.Lh

Cold atoms are most interesting candidates for studying collective instabilities. On one hand, they can easily be prepared and controlled in large numbers accessible to thermodynamic descriptions. On the other hand, individual atoms follow the rules of quantum mechanics and quantum statistics at ultralow temperatures, which predestinates them for studies of quantum synchronization effects [1].

Collective instabilities in clouds of cold and ultracold atoms driven by light have recently been observed in various situations [2–4], the most prominent of which are superradiant Rayleigh scattering (SRyS) [2] and collective atomic recoil lasing (CARL) [3]. The signature of CARL is the sudden buildup of a probe field oriented reversely to a strong pump interacting with an atomic gas together with a bunching of the atoms [5]. The underlying runaway amplification mechanism is particularly strong, if the reverse probe field is recycled by a ring cavity. In previous Letters, we have discussed the relationship between CARL, recoil-induced resonances [3] and Kuramoto-type self-synchronization [6]. But CARL also has a close analogy with SRyS, since they both share the same gain mechanism [7].

SRyS has first been observed by shining a short laser pulse into a Bose-Einstein condensate (BEC) [2]. The scattering atoms collectively form motional side modes that are coupled out of the condensate, while at the same time a burst of light is emitted into the long axis of the condensate. The phenomenon has initially been explained as a four-wave mixing process between optical and matter-wave modes with bosonic stimulation through the macroscopically populated final atomic momentum state. Although it was realized that bosonic quantum degeneracy was not essential for SRyS, the bosonic enhancement picture, which seemed corroborated by the fact that SRyS was (at first) not observable for thermal atomic clouds, instigated a controversy about the role of quantum statistics. Theoretical work [8] and the very recent experimental observation of superradiant Raman scattering with thermal atoms [9] showed, however, that the gain mechanism is independent of quantum statistics. What counts is

not the quantum nature of the particles, but their *cooperative* behavior.

The difficulty with thermal atomic ensembles is that the coherence time ruling the observability of SRyS is Doppler-limited [9]. It is therefore surprising that collective gain can be seen for CARL with thermal atoms as hot as a few 100 μK [3]. In the present Letter we demonstrate that SRyS is also possible at high temperature, and we show how the presence of an optical high-finesse ring cavity preserves coherence and cancels out the deleterious effects of thermal motion. In fact, the superradiant gain process itself causes a rapid diffusion in momentum space leading to decoherence, unless the cavity restricts the density-of-states available to the scattered light mode. The same mechanism works for thermal diffusion. The spectrum of the scattered light, which is Doppler broadened by thermal atomic motion, is filtered out by the cavity. The experimental signatures we present thus demonstrate the common root of CARL and SRyS.

Our experiment represents the first study of a BEC stored in a macroscopic optical cavity. We produce a cloud of ultracold ^{87}Rb in a magnetic trap and then move the cloud into the mode volume of a high-finesse optical ring cavity. The scheme of our experiment is shown in Fig. 1(a). We load the atoms from a two-dimensional magneto-optical trap (MOT) into a standard MOT located in an ultrahigh vacuum chamber. From here the atoms are transferred into a magnetic quadrupole potential operated with the same coils as the MOT. The potential is then compressed and the atoms are transferred via a second into a third quadrupole trap. With two pairs of thin wires separated by 1 mm and running parallel to the symmetry axis of the third quadrupole trap a Ioffe-Pritchard type potential is created [10]. The atoms can easily be shifted up and down along the Ioffe wires.

In the Ioffe-Pritchard trap the Rb cloud is cooled by forced evaporation: A microwave frequency resonantly tuned to the ground state hyperfine structure couples the trapped Zeeman state $|2, 2\rangle$ and the untrapped $|1, 1\rangle$. After 15 s of down ramping the microwave, we reach the threshold to quantum degeneracy at $T_c = 800$ nK with about

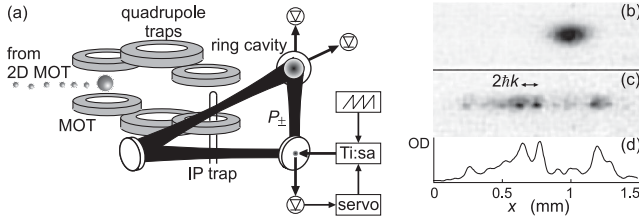


FIG. 1. (a) Schematic view of the experimental setup. A two-dimensional MOT (2D MOT) feeds a MOT in the main chamber. From here the cloud is transferred adiabatically in several intermediate steps into a Ioffe-Pritchard (IP) type magnetic trap overlapping with the ring cavity mode volume. A Ti:sapphire laser resonantly pumps the cavity mode P_+ . Both cavity modes P_{\pm} are observed via the light fields leaking out through one of the cavity mirrors. The atomic cloud can be visualized by absorption imaging. Typical images of a condensate cloud at $T = 0.5T_c$ having and not having interacted with the cavity are shown in (c) and (b), respectively. The images are recorded after 10 ms of free expansion. Curve (d) shows the vertically integrated optical density (OD) of image (c).

$N = 5 \times 10^5$ atoms. Typical trap frequencies at the end of the evaporation ramp are $\omega_x \approx \omega_y \approx 2\pi \times 200$ Hz and $\omega_z \approx 2\pi \times 50$ Hz, obtained at a bias field of 2 G. Cooling down further yields almost pure condensates of 2×10^5 Rb atoms. The long axis of the cigar-shaped condensate is parallel to the ring cavity axis.

The ring cavity is very similar to the one used in Refs. [3,6,11]. It consists of one plane and two curved mirrors. It is 8.5 cm long corresponding to a free spectral range of $\delta_{\text{fsr}} = 3.5$ GHz and has a beam waist of $w_0 = 107 \mu\text{m}$. One mode is continuously pumped by a titanium-sapphire laser. The laser can be stabilized to this mode using the Pound-Drever-Hall method. The ring cavity can be operated in two ways depending on the polarization of the incoupled light. For p -polarized light a finesse of 87 000 is determined from a measured intensity decay time of $\tau \approx 3.8 \mu\text{s}$. For s -polarized light the finesse is 6400.

We measure the intracavity light powers in the pump mode, P_+ , and the reverse mode, P_- , via the fields leaking through one of the cavity mirrors [see Fig. 1(a)]. To prevent atom losses due to light scattering, we switch off the pump laser while the atoms are driven into the resonator. As soon as the magnetically trapped atoms are located within the mode of the ring cavity, we switch on the Ti:sapphire laser tuned between 0.5 and 2 nm to the red of the D_1 line at 794.9 nm and ramp the laser frequency across one of the cavity resonances. When the laser crosses a cavity resonance, the fast branch of the Pound-Drever-Hall servo (controlling an acousto-optic modulator) quickly pulls the laser frequency to the center of the cavity resonance and tightly locks its phase. After about $50 \mu\text{s}$ the pump light is turned off. The buildup time for the cavity mode α_+ is limited by the bandwidth of the locking servo to about $\tau_{\text{bw}} = 20 \mu\text{s}$, which is a bit longer than the cavity decay time.

Another observable is the atomic density distribution measured by time-of-flight absorption imaging. A typical image of an expanded condensate having interacted with the cavity fields is shown in Fig. 1(c). The intensity of the various Bragg diffraction orders tell the population of the momentum states.

Figures 2(a) and 2(b) show a typical recorded time evolution of the powers P_{\pm} . They illustrate the central phenomenon studied in this Letter. The pumped mode fills up with light on the time scale τ_{bw} . In contrast, the power buildup in the reverse mode is delayed and suddenly increases exponentially, until it reaches a level of 0.1 to 1% of the pumped mode power. This superradiant burst of light is followed by several revivals and decays. The delay, the height of the first peak, and the subsequent evolution depend on certain experimentally tunable parameters. Those are the atom number N , the pump power P_+ , the temperature T of the atom cloud, and the coupling strength $U_0 = g^2/\Delta_a$, where $g = \sqrt{3}\Gamma\delta_{\text{fsr}}/kw_0 \approx 2\pi \times 84$ kHz is the one-photon Rabi frequency in our cavity, Δ_a the laser detuning from the atomic resonance, whose decay width is $\Gamma = 2\pi \times 6$ MHz, and $k = 2\pi/\lambda$. The rise time of the superradiant burst is about $1 \mu\text{s}$, which is considerably faster than the cavity mode buildup time. The $2 \mu\text{K}$ cold atoms are not Bose-condensed. Our experiments do not show a qualitative difference in the reverse mode behavior below T_c .

The coupled dynamics is governed by the interdependence of only a few characteristic quantities [5,12,13]. Those are the cavity decay rate $\kappa = (2\tau)^{-1} \approx 2\pi \times 20$ kHz, the recoil shift $\omega_r = 2\hbar k^2/m \approx 2\pi \times 14$ kHz, the Doppler width of the atomic velocity distribution $\sigma_v = 2k\sqrt{k_B T/m}$, and the (small signal) collective gain G . The gain describes the exponential increase of the number of

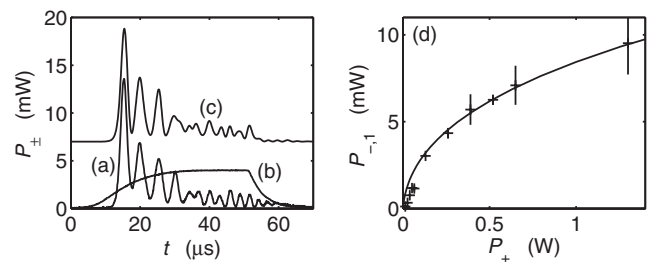


FIG. 2. (a) Measured time evolution of the reverse power P_- . The pump laser power is $P_+ = 4$ W. The cavity is operated at high finesse. The atom number is $N = 1.5 \times 10^6$ and the laser wavelength is $\lambda = 797.3$ nm. Curve (b) marks the time evolution of the recorded pump laser power scaled down by 1000. Curve (c) shows (offset by 7 mW) a numerical simulation of the reverse power using the above parameters (see text). To account for the finite switch-on time of the pump laser power, its experimentally recorded time evolution is plugged into the simulations, where we assume that the pump laser frequency is fixed and resonant to a cavity mode. (d) Measured and calculated (solid line) height $P_{-,1}$ of the first peak as a function of pump power P_+ . Here $N = 2.4 \times 10^6$ and $\lambda = 796.1$ nm.

recoiling atoms and of the scattered light intensity due to cooperation. For conventional SRyS, i.e., in the absence of a cavity, the relevant lifetime of the optical fields is limited by the size L of the condensate along the axis into which the light is scattered, $\tau_{\text{sr}} = (2\kappa_{\text{sr}})^{-1} = L/c$. The gain is given by $G = N\Omega^2 g_{\text{sr}}^2 / 2\kappa_{\text{sr}} \Delta_a^2$, where Ω is the Rabi frequency generated by the incident laser beam and $g_{\text{sr}} = \sqrt{3\Gamma\kappa_{\text{sr}}}/kw_{\text{sr}}$ is the coupling strength of the effective cavity formed by the condensate to the atomic D_1 transition. w_{sr} is the radial width of the condensate. Typically, the gain $G \approx 10^5 \text{ s}^{-1}$ is much smaller than κ_{sr} [2], i.e., scattered photons leave the interaction region before the next photons are scattered. For collective gain to take place, the coherence time must be longer than the mean time delay between subsequent scattering events, which is only possible if the coherence is stored as a Raman coherence between momentum states of the individual atoms before and after scattering.

In contrast, if the decay of the scattered light mode is slowed down by a cavity, the cavity plays a major role in storing the coherence. The gain is then estimated by the same formula as above, but replacing g_{sr} by g and κ_{sr} by κ . Furthermore, since the pump and the reverse beam now share the same mode volume, we may set $\Omega_i = \sqrt{n}g$. We obtain a gain, $G = nN\Omega_0^2/2\kappa$, which can experimentally be tuned between $G \approx 10^3\text{--}10^9 \text{ s}^{-1}$.

The various regimes in which CARL and SRyS may occur are characterized by the size of the collective gain bandwidth $\Delta\omega_G$ (which is the spectral regime in which the gain is larger than half its maximum value) [7] as compared to the decay width κ (or κ_{sr}) and the recoil frequency ω_r . In particular, one may distinguish two situations called the *superradiant* or *bad-cavity limit*, where $\Delta\omega_G \ll \kappa$ and the *good-cavity limit* for which $\Delta\omega_G \gg \kappa$ [5]. Because of the fact that the field decay rate in our ring cavity is 6 to 7 orders of magnitude smaller than in the superradiance experiments, both regimes are accessible to our experiment by operating the cavity at either high or low finesse and by appropriately tuning the gain via the parameters P_+ , N , and λ . At a given cavity decay width the superradiant regime is attained for weak collective coupling, i.e., small atom numbers and low pump powers, while the good-cavity limit is realized for large atom numbers and high powers [12].

Figure 2(d) shows the dependence of the height of the first peak as a function of the pump laser power in the good-cavity limit, i.e., for the high-finesse case. The theoretical curves in this and all other figures are calculated based on the CARL model already used in Refs. [3,5,12,14], which treats all degrees of freedom classically. It also adiabatically eliminates the electronically excited states, which is a good approximation far from resonance [15]. The simulations are done with 100 atoms, each atom representing $N/100$ atoms. Finally, we tested that the classical model coincides with a generalized description treating the atomic motion as quantized to a good approximation [13].

Clear features that allow us to distinguish between the good-cavity and the superradiant regime are the dependences of the peak height on the pump power and on atom number: $P_{-,1} \propto N^2 P_+$ for the superradiant case and $P_{-,1} \propto N^{4/3} P_+^{1/3}$ for good cavities [5]. The different N dependences are demonstrated in Figs. 3(a) and 3(b). Interestingly, the superradiant limit exhibits a thresholdlike behavior. It is due to backscattering at the mirror surfaces. This phenomenon, which also accounts for the light being in the reverse mode below the threshold, is discussed below.

Differently from earlier SRyS experiments [2,9] we observe not only a single light burst, but a train of self-similar pulses reminiscent to nonlinear ringing, which demonstrates how the cavity preserves the coherence against diffusion in momentum space induced by the collective coupling. The cavity is, however, also able to neutralize diffusion due to thermal motion. In fact, we observe SRyS for temperatures of the atomic cloud as high as $40 \mu\text{K}$ (which is 2 orders of magnitude larger than previously reported [9]). We studied the regime of high temperatures, $\sigma_v \gg \kappa$, ω_r , earlier [3], but we are now able to study the CARL dependence on temperature down to the level of quantum degeneracy.

We experimentally tune the temperature via forced evaporation ensuring a fixed final atom number of $N = 10^6$ for all measurements. Figure 4(a) shows how the temperature affects the evolution of the reverse power. At low temperatures, strong superradiant bursts are observed followed by a large number of revivals. As the temperature rises, the revivals gradually fade out, until at $T \approx 20 \mu\text{K}$ only a single burst survives. The amount of power scattered into the reverse mode rapidly decreases at higher temperatures, as seen in Fig. 4(b).

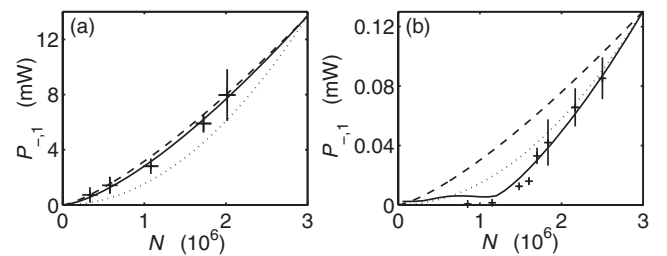


FIG. 3. (a) Measured height of the first peak as a function of atom number N . The cavity is operated at high-finesse (good-cavity limit), the pump laser is set to $P_+ = 0.5 \text{ W}$ and $\lambda = 796.1 \text{ nm}$. The solid line shows a simulation. To emphasize the N dependence we also plot curves going like N^2 (dotted line) and $N^{4/3}$ (dashed line). Those dependences are expected in the superradiant and in the good-cavity limit, respectively. (b) Same as (a) but the cavity is now operated at low finesse (superradiant limit), the pump laser is set to $P_+ = 70 \text{ mW}$. To compensate for the loss in cooperativity the laser is tuned closer to resonance at $\lambda = 795.3 \text{ nm}$. The amount of light in the reverse mode due to mirror backscattering in steady state without atoms is $P_{-,s} = 1.8 \times 10^{-4} P_+$.

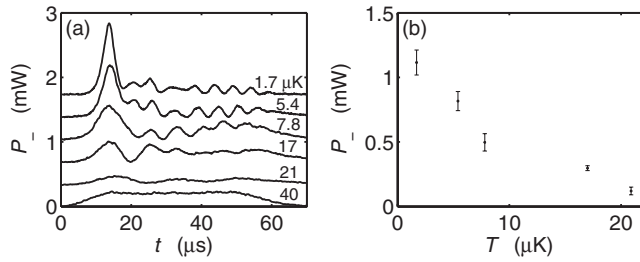


FIG. 4. (a) Measured time-evolutions of the reverse mode at various temperatures. The cavity is operated at high finesse. For clarity the curves are vertically shifted from one another by 0.35 mW. The atom number is the same for all data points, $N = 10^6$. The pump power is $P_+ = 1$ W. The reverse power observed at high temperatures is there even without atoms and is caused by mirror backscattering. (b) Measured SRyS peak height as a function of temperature.

At even higher temperature beyond 40 μK we observe that the reverse power P_- approaches a stationary value within the time τ_{bw} without discernible oscillations. In fact, in this case the evolution of the reverse power is the same as for the empty cavity. It is due to residual backscattering from impurities at the mirror surfaces. The large amount of mirror backscattering, which adopts values between 0% and 1% [16], suggests that it plays a major role in the dynamics. This is true indeed for small atom numbers and in the very beginning of the light pulse. But as soon as the reverse mode starts to build up, the collective dynamics takes over completely, and the rise time of the reverse mode intensity is much shorter than the cavity buildup time. The height of the superradiant peaks turns out to be quite insensitive to the amount of mirror backscattering, and at longer times the average value of the light intensity is smaller than in the absence of atoms, i.e., the collective atomic motion suppresses mirror backscattering.

In conclusion, our observation of superradiance in the bad-cavity regime of the CARL experimentally demonstrates the intrinsic link between both phenomena. Together with the earlier observation of CARL with atomic clouds as hot as several 100 μK [3] this proves that the gain process underlying both, SRyS and CARL, is not based on quantum statistics, but on cooperativity [8]. These results clarify the intricate relationship between CARL and superradiance in a regime, where the coupling between radiative and matter-wave modes is completely coherent. This is also the regime for which the prospect of a robust quantum entanglement of the modes has been pointed out [17].

Although the experiment reported here is the first one to study the interaction of Bose-Einstein condensates with macroscopic cavities, the quantum degeneracy plays no role in the observed effects. Future studies, however, may focus on the role of quantum statistics and interatomic interaction [18].

This research has been supported by the Deutsche Forschungsgemeinschaft (DFG) under Contract No. Co 229/3-1.

- [1] O. V. Zhirov and D. L. Shepelyansky, *Eur. Phys. J. D* **38**, 375 (2006).
- [2] S. Inouye *et al.*, *Science* **285**, 571 (1999); M. Kozuma *et al.*, *Science* **286**, 2309 (1999); L. Fallani *et al.*, *Phys. Rev. A* **71**, 033612 (2005).
- [3] D. Kruse, Ch. von Cube, C. Zimmermann, and Ph. W. Courteille, *Phys. Rev. Lett.* **91**, 183601 (2003).
- [4] B. Nagorny, Th. Elsässer, and A. Hemmerich, *Phys. Rev. Lett.* **91**, 153003 (2003); A. T. Black, H. W. Chan, and V. Vuletic, *Phys. Rev. Lett.* **91**, 203001 (2003); G. Labeyrie, F. Michaud, and R. Kaiser, *Phys. Rev. Lett.* **96**, 023003 (2006).
- [5] R. Bonifacio and L. De Salvo, *Nucl. Instrum. Methods Phys. Res., Sect. A* **341**, 360 (1994); R. Bonifacio, L. De Salvo, L. M. Narducci, and E. J. D'Angelo, *Phys. Rev. A* **50**, 1716 (1994).
- [6] C. von Cube *et al.*, *Phys. Rev. Lett.* **93**, 083601 (2004).
- [7] N. Piovella, R. Bonifacio, B. W. J. McNeil, and G. R. M. Robb, *Opt. Commun.* **187**, 165 (2001).
- [8] M. G. Moore and P. Meystre, *Phys. Rev. Lett.* **86**, 4199 (2001); W. Ketterle and S. Inouye, *Phys. Rev. Lett.* **86**, 4203 (2001).
- [9] Y. Yoshikawa, Y. Torii, and T. Kuga, *Phys. Rev. Lett.* **94**, 083602 (2005).
- [10] C. Silber *et al.*, *Phys. Rev. Lett.* **95**, 170408 (2005).
- [11] D. Kruse *et al.*, *Phys. Rev. A* **67**, 051802(R) (2003).
- [12] R. Bonifacio, G. R. M. Robb, and B. W. J. McNeil, *Phys. Rev. A* **56**, 912 (1997).
- [13] N. Piovella, M. Gatelli, and R. Bonifacio, *Opt. Commun.* **194**, 167 (2001).
- [14] M. Gangl and H. Ritsch, *Phys. Rev. A* **61**, 043405 (2000); M. Perrin, Zongxiong Ye, and L. M. Narducci, *Phys. Rev. A* **66**, 043809 (2002).
- [15] For the same reason radiation pressure is small, even though not completely negligible. We observe related effects for measuring times longer than 100 μs and account for it in the simulations.
- [16] We experimentally observe a strong dependence of the amount of mirror backscattering on the eigenfrequency of the cavity to which the pump laser is locked. This phenomenon, which is due to interference of the light backscattered from all three cavity mirrors, will be discussed elsewhere. From the experimental point of view, the most interesting feature is that backscattering can be avoided by a proper choice of the resonant cavity mode. The amount of mirror backscattering is measured independently and plugged into the simulations.
- [17] M. G. Moore and P. Meystre, *Phys. Rev. A* **59**, R1754 (1999); N. Piovella, M. Cola, and R. Bonifacio, *Phys. Rev. A* **67**, 013817 (2003).
- [18] P. Horak, S. M. Barnett, and H. Ritsch, *Phys. Rev. A* **61**, 033609 (2000).