## Measurement of the Negative Muon Anomalous Magnetic Moment to 0.7 ppm

G.W. Bennett,<sup>2</sup> B. Bousquet,<sup>9</sup> H. N. Brown,<sup>2</sup> G. Bunce,<sup>2</sup> R. M. Carey,<sup>1</sup> P. Cushman,<sup>9</sup> G. T. Danby,<sup>2</sup> P. T. Debevec,<sup>7</sup> M. Deile,<sup>11</sup> H. Deng,<sup>11</sup> S. K. Dhawan,<sup>11</sup> V. P. Druzhinin,<sup>3</sup> L. Duong,<sup>9</sup> F. J. M. Farley,<sup>11</sup> G.V. Fedotovich,<sup>3</sup> F. E. Gray,<sup>7</sup> D. Grigoriev,<sup>3</sup> M. Grosse-Perdekamp,<sup>11</sup> A. Grossmann,<sup>6</sup> M. F. Hare,<sup>1</sup> D.W. Hertzog,<sup>7</sup> X. Huang,<sup>1</sup> V.W. Hughes,<sup>11,\*</sup> M. Iwasaki,<sup>10</sup> K. Jungmann,<sup>5</sup> D. Kawall,<sup>11</sup> B. I. Khazin,<sup>3</sup> F. Krienen,<sup>1</sup> I. Kronkvist,<sup>9</sup> A. Lam,<sup>1</sup> R. Larsen,<sup>2</sup> Y.Y. Lee,<sup>2</sup> I. Logashenko,<sup>1,3</sup> R. McNabb,<sup>9</sup> W. Meng,<sup>2</sup> J. P. Miller,<sup>1</sup> W. M. Morse,<sup>2</sup> D. Nikas,<sup>2</sup> C. J. G. Onderwater,<sup>7</sup> Y. Orlov,<sup>4</sup> C. S. Özben,<sup>2,7</sup> J. M. Paley,<sup>1</sup> Q. Peng,<sup>1</sup> C. C. Polly,<sup>7</sup> J. Pretz,<sup>11</sup> R. Prigl,<sup>2</sup> G. zu Putlitz,<sup>6</sup> T. Qian,<sup>9</sup> S. I. Redin,<sup>3,11</sup> O. Rind,<sup>1</sup> B. L. Roberts,<sup>1</sup> N. Ryskulov,<sup>3</sup> Y. K. Semertzidis,<sup>2</sup> P. Shagin,<sup>9</sup> Yu. M. Shatunov,<sup>3</sup> E. P. Sichtermann,<sup>11</sup> E. Solodov,<sup>3</sup> M. Sossong,<sup>7</sup> L. R. Sulak,<sup>1</sup> A. Trofimov,<sup>1</sup> P. von Walter,<sup>6</sup> and A. Yamamoto<sup>8</sup>

(Muon (g - 2) Collaboration)

<sup>1</sup>Department of Physics, Boston University, Boston, Massachusetts 02215, USA

<sup>2</sup>Brookhaven National Laboratory, Upton, New York 11973, USA

<sup>3</sup>Budker Institute of Nuclear Physics, Novosibirsk, Russia

<sup>4</sup>Newman Laboratory, Cornell University, Ithaca, New York 14853, USA

<sup>5</sup>Kernfysisch Versneller Instituut, Rijksuniversiteit Groningen, NL 9747 AA Groningen, The Netherlands

<sup>6</sup>Physikalisches Institut der Universität Heidelberg, 69120 Heidelberg, Germany

<sup>7</sup>Department of Physics, University of Illinois at Urbana-Champaign, Illinois 61801, USA

<sup>8</sup>KEK, High Energy Accelerator Research Organization, Tsukuba, Ibaraki 305-0801, Japan

<sup>9</sup>Department of Physics, University of Minnesota, Minneapolis, Minnesota 55455, USA

<sup>10</sup>Tokyo Institute of Technology, Tokyo, Japan

<sup>11</sup>Department of Physics, Yale University, New Haven, Connecticut 06520, USA (Received 10 January 2004; published 23 April 2004)

The anomalous magnetic moment of the negative muon has been measured to a precision of 0.7 ppm (ppm) at the Brookhaven Alternating Gradient Synchrotron. This result is based on data collected in 2001, and is over an order of magnitude more precise than the previous measurement for the negative muon. The result  $a_{\mu^-} = 11\,659\,214(8)(3) \times 10^{-10}$  (0.7 ppm), where the first uncertainty is statistical and the second is systematic, is consistent with previous measurements of the anomaly for the positive and the negative muon. The average of the measurements of the muon anomaly is  $a_{\mu}(\exp) = 11\,659\,208(6) \times 10^{-10}$  (0.5 ppm).

DOI: 10.1103/PhysRevLett.92.161802

PACS numbers: 13.40.Em, 12.15.Lk, 14.60.Ef

The anomalous magnetic moments of the muon and the electron have played an important role in the development of the standard model. Compared to the electron, the muon anomaly has a relative sensitivity to heavier mass scales which typically is proportional to  $(m_{\mu}/m_e)^2$ . At the present level of accuracy, the muon anomaly gives an experimental sensitivity to virtual W and Z gauge bosons as well as a potential sensitivity to other, as yet unobserved, particles in the few hundred GeV/ $c^2$  mass range [1].

We report our result for the negative muon anomalous magnetic moment  $a_{\mu^-} = (g-2)/2$  from data collected in early 2001. The measurement is based on muon spin precession in a magnetic storage ring with electrostatic focusing. The same experimental technique was used as in our most recent measurements of  $a_{\mu^+}$  [2,3], and a similar precision of 0.7 ppm was achieved. Detailed descriptions of the apparatus may be found elsewhere [4–8].

For polarized muons moving in a uniform magnetic field  $\vec{B}$  perpendicular to the muon spin and to the plane of

161802-1 0031-9007/04/92(16)/161802(5)\$22.50

the orbit and in an electric quadrupole field  $\vec{E}$ , which is used for vertical focusing [8], the angular frequency difference,  $\omega_a$  between the spin precession frequency and the cyclotron frequency, is given by

$$\vec{\omega}_{a} = \frac{e}{mc} \bigg[ a_{\mu} \vec{B} - \bigg( a_{\mu} - \frac{1}{\gamma^{2} - 1} \bigg) \vec{\beta} \times \vec{E} \bigg].$$
(1)

The dependence of  $\omega_a$  on the electric field is eliminated by storing muons with the "magic"  $\gamma = 29.3$  [9], which corresponds to a muon momentum p = 3.09 GeV/c. Hence, measurement of  $\omega_a$  and of *B*, in terms of the free proton NMR frequency  $\omega_p$  and the ratio of muon to proton magnetic moments  $\lambda$ , determines  $a_{\mu}$ . At the magic  $\gamma$ , the muon lifetime is approximately 64.4 µs and the (g - 2) precession period is 4.37 µs. With a field of 1.45 T in our storage ring [4], the central orbit radius is 7.11 m.

The difference frequency  $\omega_a$  was determined by counting the number N(t) of decay electrons above an energy threshold. The time spectrum of decay electrons is then given by

$$N(t) = N_0 e^{-t/(\gamma \tau)} [1 + A \sin(\omega_a t + \phi_a)].$$
(2)

The normalization  $N_0$ , asymmetry A, and phase  $\phi_a$  vary with the chosen energy threshold.

The measurement of the magnetic field frequency  $\omega_p$  is based on proton NMR in water. A trolley with 17 NMR probes was moved typically every three days throughout the entire muon storage region. About 150 fixed NMR probes distributed around the ring in the top and bottom walls of the vacuum chamber were used to interpolate the field between trolley measurements. The system was calibrated with respect to a standard probe with a spherical H<sub>2</sub>O sample [5]. The homogeneity of the field in 2001 (Fig. 1) was similar to that achieved for the opposite polarity field in 2000 [3].

The field  $\langle B \rangle$  weighted with the analyzed event sample was obtained from two largely independent analyses, whose results were found to agree to within 0.05 ppm. Its final value is expressed in terms of the free proton resonance frequency and is given by  $\omega_p/(2\pi) =$ 61791400(11) Hz (0.2 ppm). Table I lists the uncertainties. The improved 2001 uncertainties resulted from refinements in the calibration measurements, and from an upgraded system to determine the azimuthal trolley position in the storage ring.

The 2001  $\omega_a$  data collection was similar to that in 2000. However, the hardware energy threshold of the detectors was kept lower and equal for all counters at 0.9 GeV compared to 1.0–1.4 GeV in 2000. This was made possible by reducing the intensity of the injected beam, which in turn reduced the light flash in the detectors [2,3]. These factors allowed all the detectors to be turned on and be stable by 32  $\mu$ s after beam injection, as opposed to 50  $\mu$ s in 2000. As a result of the reduced rates, the fraction of overlapping signals (pileup) after 32  $\mu$ s in



FIG. 1. A two-dimensional multipole expansion of the 2001 field averaged over azimuth from one out of 20 trolley measurements. Half ppm contours with respect to a central azimuthal average field  $B_0 = 1.451269$  T are shown. The multipole amplitudes relative to  $B_0$  are given at the beam aperture, which had a radius of 4.5 cm and is indicated by the circle.

2001 was comparable to the pileup fraction after 50  $\mu$ s in 2000. In 2000, the field focusing index *n*, which is proportional to the electric field gradient, was n = 0.137, corresponding to a horizontal coherent betatron oscillation frequency (CBO) of 466 kHz [3]. This frequency was close to twice the (g - 2) frequency of 229 kHz, which resulted in a sizable uncertainty in the fitted  $\omega_a$  value [3]. In 2001, we used two different *n* values, n = 0.122 and n = 0.142, which resulted in CBO frequencies, 419 and 491 kHz, that are further from twice the (g - 2) frequency (see Fig. 2). Consequently, the uncertainty caused by CBO is smaller. Furthermore, it also reduced the correlation between the CBO and detector gain effects in the fits to the time spectrum.

Two independent implementations of the algorithm to reconstruct the electron times and energies from the calorimeter signals were used. The frequency  $\omega_a$  was determined by fitting the time distribution of decay electrons. Five independent analyses were performed in order to probe the systematic uncertainties and, of course, to protect against mistakes. All five results agreed within the expected statistical deviations due to different data selection and weightings. These analyses are described below.

Two of the analyses used slightly different parametrizations [2,3] that included CBO modulations and fitted the combined electron spectrum in the energy range 1.8– 3.4 GeV. In the third analysis, the counts were weighted with the experimentally determined energy-dependent modulation asymmetry, which optimized the statistical power of the data. This method permitted the analyzed energy range to be extended. We used an energy range of 1.5 to 3.4 GeV, which together with the asymmetry weighting resulted in a 10% improvement of the statistical uncertainty. As in the first two analyses, the resulting spectrum of *weighted* counts was fitted to a function that parametrized all known and statistically significant perturbations.

The remaining analyses fit the ratio [2,3] formed by randomly assigning the data to four statistically independent subsets  $n_1$  to  $n_4$ . The subsets were rejoined in  $u(t) = n_1(t) + n_2(t)$  and  $v(t) = n_3(t - \tau_a/2) + n_4(t + \tau_a/2)$ , where  $\tau_a$  is an estimate of the (g - 2) period, and then

TABLE I. Systematic uncertainties for the  $\omega_p$  analysis.

Source of uncertainty	Size (ppm)
Absolute calibration of standard probe	0.05
Calibration of trolley probe	0.09
Trolley measurements of $B_0$	0.05
Interpolation with fixed probes	0.07
Uncertainty from muon distribution	0.03
Others <sup>a</sup>	0.10
Total systematic error on $\omega_p$	0.17

<sup>a</sup>Higher multipoles, trolley temperature and voltage response, eddy currents from the kickers, and time-varying stray fields.



FIG. 2. The Fourier spectrum of the residuals of a fit to the five free parameters given in Eq. (2) for the high (top) and low (bottom) *n*-value data. The corresponding CBO frequencies, located at 491 kHz (top) and 419 kHz (bottom) as well as their (g - 2) sidebands are clearly visible. Dashed lines indicate the (g - 2) frequency.

combined to form the time spectrum r(t) = [u(t) - v(t)]/[u(t) + v(t)]. The (g - 2) rate modulation of v is 180° degrees out of phase compared to that of u, and to sufficient precision r(t) can be described by  $A \sin(\omega_a t + \phi_a)$ . The ratio r(t) is largely insensitive to changes of observed counts on time scales larger than  $\tau_a = 2\pi/\omega_a \sim 4 \ \mu$ s.

In one of the ratio analyses, the sensitivity to CBO was reduced by combining the data from both *n* values and all detectors prior to fitting. The data were fitted from 32  $\mu$ s after injection when all detectors were on. In the second ratio analysis, the data were fitted separately for each calorimeter and *n* value. The fits began between 24 and 32  $\mu$ s, and required the parametrization of the CBO effects in the fit function.

Changes in the radial and vertical muon distributions with time were quantified, and were found to have a negligible effect on  $\omega_a$ . A small reduction in the pulsed electrostatic quadrupole voltages [8] during the measurement period could change the vertical muon distribution. Analysis of the data from scintillator counter hodoscopes placed in front of the calorimeters combined with a beam tracking calculation and a GEANT based simulation set a systematic error limit of 0.03 ppm. The muon radial distribution is determined by the magnetic field and the momentum distribution [2,10]. The magnetic field does not change with time after injection, except due to the field from eddy currents induced by the fast kicker [7]. This was measured, and found to have a negligible effect on the muon radial distribution. Muons of lower momenta decay earlier in the laboratory frame than muons of higher momenta. The momentum distribution of the stored beam thus changes during the 600  $\mu$ s measurement period. The effect on  $\omega_a$  due to this change was studied in simulation, and was found to be 0.03 ppm.

The results for  $\omega_a$  for the two *n* values are consistent (see Fig. 3) and were combined for each of the analyses. The values for  $\omega_a$  from the five analyses are in agreement to within variations expected from the differences in the analyzed event samples and the treatment of the data. The analysis techniques are expected to have somewhat different sensitivities to different systematic effects. Detailed comparisons of the results, using all analyzed data as well as only the data in overlap, showed no evidence for unaccounted systematic differences. The five resulting values for  $\omega_a$  were combined in a simple arithmetic mean to obtain a single value for  $\omega_a$ .

The resulting frequency value is  $\omega_a/(2\pi) = 229\,073.59(15)(5)$  Hz (0.7 ppm), which includes a correction of +0.77(6) ppm for contributions to Eq. (1) caused by vertical oscillations (0.30 ppm) and for the effect of the horizontal electric fields on muons with  $\gamma \neq 29.3$  (0.47 ppm). The stated uncertainties account for strong correlations among the individual results, both statistical and systematic. Table II lists the systematic uncertainties in the combined result with these correlations taken into account.

After the  $\omega_p$  and  $\omega_a$  analyses were finalized separately and independently,  $a_\mu$  was evaluated. The result is

$$a_{\mu^{-}} = \frac{R}{\lambda - R} = 11659214(8)(3) \times 10^{-10} (0.7 \text{ ppm}),$$
 (3)



FIG. 3. Comparison of the  $\omega_a$  values from the five analyses for the low-*n* (filled) and high-*n* (open) data sets. Analysis 4 used only the combined low and high-*n* data (square). The divisions on the vertical axis are separated by 1 ppm, and the indicated uncertainties are statistical. The systematic uncertainties are considerably smaller.

TABLE II. Systematic uncertainties for the combined  $\omega_a$  analysis.

Source of errors	Size (ppm)
Coherent betatron oscillations	0.07
Pileup	0.08
Gain changes	0.12
Lost muons	0.09
Others <sup>a</sup>	0.11
Total systematic error on $\omega_a$	0.21

<sup>a</sup>Alternating gradient synchrotron background, timing shifts, E field and vertical oscillations, beam debunching/randomization, binning, and fitting procedure.

where  $R_{\mu^-} \equiv \omega_a / \omega_p = 0.0037072083(26)$  and  $\lambda = \mu_{\mu} / \mu_p = 3.18334539(10)$  [11]. This new result is in good agreement with the average of  $R_{\mu^+} = 0.0037072048(25)$  [2] as predicted by the CPT theorem. The difference  $\Delta R = R_{\mu^-} - R_{\mu^+} = (3.5 \pm 3.4) \times 10^{-9}$ . The new average is  $R_{\mu} = 0.0037072063(20)$  and

$$a_{\mu}(\exp) = 11\,659\,208(6) \times 10^{-10}$$
 (0.5 ppm), (4)

in which the total uncertainty consists of  $5 \times 10^{-10}$  (0.4 ppm) statistical uncertainty and  $4 \times 10^{-10}$  (0.3 ppm) systematic uncertainty. The correlation of systematic uncertainties between the data sets has been taken into account. The combined result for the positive muon [3],  $a_{\mu^+}(\exp) = 11\,659\,203(8) \times 10^{-10}$  (0.7 ppm) has a statistical uncertainty of  $6 \times 10^{-10}$  (0.6 ppm) and a systematic uncertainty of  $5 \times 10^{-10}$  (0.4 ppm). It is shown in Fig. 4 together with the new result for the negative muon and their average.

The standard model (SM) prediction for  $a_{\mu}$  consists of QED, hadronic, and weak contributions. The uncertainty on the standard model value is dominated by the uncertainty on the lowest-order hadronic vacuum polarization. This contribution can be determined directly from the annihilation of  $e^+e^-$  to hadrons through a dispersion integral [12]. The indirect determination using data from hadronic  $\tau$  decays, the conserved vector current hypothesis, plus the appropriate isospin corrections, could in principle improve the precision of  $a_{\mu}$  (had). However, discrepancies between the  $\tau$  and the  $e^+e^-$  results exist [13,14]. The two data sets do not give consistent results for the pion form factor. Using  $e^+e^-$  annihilation data, the corresponding theoretical value is  $a_{\mu}(SM) =$  11659181(8) × 10<sup>-10</sup> (0.7 ppm). The value deduced from  $\tau$  decay is larger by 15 × 10<sup>-10</sup> and has a stated uncertainty of  $7 \times 10^{-10}$  (0.7 ppm). The difference between the experimental determination of  $a_{\mu}$  and the standard model theory using the  $e^+e^-$  or  $\tau$  data for the calculation of the hadronic vacuum polarization is  $2.7\sigma$ and 1.4 $\sigma$ , respectively.

This is the final analysis of the anomalous magnetic moment from experiment E821 at the Brookhaven



FIG. 4 (color online). Measurements of  $a_{\mu}$  by E821 with the SM predictions (see text for discussion). Uncertainties indicated on the measurements are total uncertainties.

Alternating Gradient Synchrotron. We aim to substantially improve our result in a new measurement and look forward to continued efforts to improve the theoretical evaluation.

We thank T. Kirk, D. I. Lowenstein, P. Pile, and the staff of the BNL Alternating Gradient Synchrotron for the strong support they have given this experiment. This work was supported in part by the U.S. Department of Energy, the U.S. National Science Foundation, the U.S. National Computational Science Alliance, the German Bundesminister für Bildung und Forschung, the Russian Ministry of Science, and the U.S.-Japan Agreement in High Energy Physics.

\*Deceased.

- V.W. Hughes and T. Kinoshita, Comments Nucl. Part. Phys. 14, 341 (1985).
- [2] Muon (g 2) Collaboration, H. N. Brown *et al.*, Phys. Rev. D **62**, 091101 (2000); Muon (g 2) Collaboration, H. N. Brown *et al.*, Phys. Rev. Lett. **86**, 2227 (2001).
- [3] Muon (g 2) Collaboration, G.W. Bennett *et al.*, Phys. Rev. Lett. 89, 101804 (2002).
- [4] A. Yamamoto *et al.*, Nucl. Instrum. Methods Phys. Res., Sect. A **491**, 23 (2002); G. T. Danby *et al.*, Nucl. Instrum. Methods Phys. Res., Sect. A **457**, 151 (2001); S. I. Redin *et al.*, Nucl. Instrum. Methods Phys. Res., Sect. A **473**, 260 (2001); R. Prigl *et al.*, Nucl. Instrum. Methods Phys. Res., Sect. A **374**, 118 (1996).
- [5] X. Fei, V.W. Hughes, and R. Prigl, Nucl. Instrum. Methods Phys. Res., Sect. A **394**, 349 (1997).
- [6] J. Ouyang *et al.*, Nucl. Instrum. Methods Phys. Res., Sect. A **374**, 215 (1996); S. A. Sedykh *et al.*, Nucl. Instrum. Methods Phys. Res., Sect. A **455**, 346 (2000).
- [7] E. Efstathiadis *et al.*, Nucl. Instrum. Methods Phys. Res., Sect. A **496**, 8 (2003).
- [8] Y. K. Semertzidis *et al.*, Nucl. Instrum. Methods Phys. Res., Sect. A **503**, 458 (2003).
- [9] J. Bailey et al., Nucl. Phys. B150, 1 (1979).

- [10] Y. Orlov, C. S. Özben, and Y. K. Semertzidis, Nucl. Instrum. Methods Phys. Res., Sect. A 482, 767 (2002).
- [11] W. Liu *et al.*, Phys. Rev. Lett. **82**, 711 (1999); Particle Data Group, K. Hagiwara *et al.*, Phys. Rev. D **66**, 010001 (2002).
- [12] B. E. Lautrup, A. Peterman, and E. de Rafael, Phys. Rep. 3C, 193 (1972).
- [13] M. Davier, S. Eidelman, A. Höcker, and Z. Zhang, Eur. Phys. J. C **31**, 503 (2003).
- [14] S. Ghozzi and F. Jegerlehner, Phys. Lett. B 583, 222 (2004).