## Spin Coupling between Cold Atoms and the Thermal Fluctuations of a Metal Surface

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We describe an experiment in which Bose-Einstein condensates and cold atom clouds are held by a microscopic magnetic trap near a room-temperature metal wire 500  $\mu$ m in diameter. The lifetime for atoms to remain in the microtrap is measured over a range of distances down to 27  $\mu$ m from the surface of the metal. We observe the loss of atoms from the microtrap due to spin flips. These are induced by radio-frequency thermal fluctuations of the magnetic field near the surface, as predicted but not previously observed.

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The ability to control cold atom clouds in microscopic magnetic traps [1–3] and waveguides [4–6] has created the new field of miniaturized atom optics [7,8]. With the use of microstructured surfaces (atom chips) it becomes possible to control cold atoms on the  $\mu$ m length scale and to anticipate the construction of integrated atom interferometers [9–11]. Ultimately there is the possibility of controlling the quantum coherences within arrays of individual atoms for use in quantum information processing [12,13]. For these kinds of applications it is important to avoid fluctuating or inhomogeneous perturbations, which tend to destroy the quantum coherences.

Clouds within 100  $\mu$ m of a current-carrying wire and cooled below a few  $\mu K$  have recently revealed three surface-related decoherence effects. First, the clouds break into fragments along the length of the wire as a result of a corrugated trapping potential [14,15]. The corrugations are caused by a small spatially alternating magnetic field *parallel* to the wire [16], which is presumably due to a small transverse component of the current. The second effect is heating of the cloud [10,14] due to audio frequency technical noise in the currents that form the microtrap, which cause it to shake. Finally, trapped atoms are lost [10,14] through spin flips induced by radiofrequency technical noise in the wire currents. Some of these effects have recently been elucidated by Leanhardt et al. through a comparison of magnetic and optical traps near a surface [17].

In addition to these essentially technical decoherence effects, there is a more fundamental limitation associated with the thermal fluctuations of the magnetic field. When the cold atoms are far from any surfaces in their room-temperature environment, they interact with the blackbody radiation field. This has very little power at the resonance frequencies of the atoms, making the thermalization times exceedingly long and the decoherence effects correspondingly small. However, atoms trapped some tens of  $\mu$ m above a metal interact with the thermally fluctuating near field of the surface, whose spectrum is very different from the blackbody spectrum. Recent calculations [18] have shown that spin flips in-

duced by this near field can cause atoms to be ejected from a magnetic trap in less than a second. In this Letter we present measurements of this fundamental effect, which, in the presence of technical losses, has not previously been accessible [14,17].

A diagram of the atom chip used to form our microtrap is shown in Fig. 1. The main wire is a 500  $\mu$ m diameter guide wire along the z direction. It consists of a central core of Cu with 185  $\mu$ m radius, a 55  $\mu$ m thick Al layer, and a 10  $\mu$ m thick ceramic sheath. This wire is glued by high-vacuum epoxy (Bylapox 7285) into the 200  $\mu$ m-deep channel formed by a glass substrate and two glass cover slips. Below the guide wire there are four transverse wires, 800  $\mu$ m in diameter. The cover slips are coated with 60 nm of gold so that they reflect 780 nm light. In order to make a Bose-Einstein condensate (BEC) in the microtrap we first collect <sup>87</sup>Rb atoms using a magneto-optical trap (MOT) whose beams are reflected from the gold surface. This MOT collects  $1 \times 10^8$  atoms at a height of 4 mm above the surface and cools them to 50  $\mu$ K. The MOT is pulled down to a height of 1.3 mm by passing a current of 3.2 A through the guide wire and adding a uniform magnetic field  $B_x$  of 6 G along the x direction. This compresses the cloud into a cylindrical shape and increases the phase space density of the atoms to  $2 \times 10^{-6}$  [19].

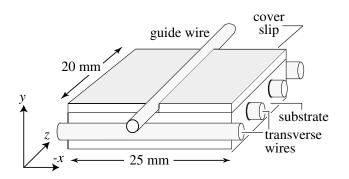


FIG. 1. Construction of the atom chip (not to scale). Atoms are trapped near the surface of the guide wire.

The light and the anti-Helmholtz coils of the MOT are then switched off and the atoms are optically pumped into the  $|F, m\rangle = |2, 2\rangle$  state. We collect  $2 \times 10^7$  of these atoms in the magnetic guide formed by the guide wire (8 A along z), and the transverse bias field  $B_x$  (10 G along x). Axial confinement is provided by the inner transverse wires (15 A each along -x), and the outer transverse wires (15 A each along x). The field at the center of this trap is partly canceled by an axial bias field  $B_z$  (6 G along z). Next, the trap is adiabatically compressed over 0.5 s by increasing  $B_x$  and  $B_z$  to 29 and 11 G, respectively, and reducing the guide current to 6.9 A. This brings the trap to a distance of 225  $\mu$ m from the wire and raises the radial and axial trap frequencies to 840 and 26 Hz. The elastic collision rate is now  $\sim$ 54 s<sup>-1</sup>, which is high enough for forced rf evaporative cooling to be efficient. We sweep the rf frequency logarithmically over 12.5 s from 13 MHz to a final frequency near 600 kHz. This cools the cloud down to well below the 380 nK critical temperature for Bose-Einstein condensation and produces up to  $5 \times 10^4$  atoms in the condensate [19].

We find that the number of atoms in the microtrap decays exceedingly slowly with time. We cannot leave the trap on long enough to measure this lifetime precisely because the vacuum feedthroughs that carry the trap currents overheat, causing a sudden increase of pressure after  $\sim 20$  s. However, it is well in excess of 100 s. For many of our measurements we stop the evaporation after 6 s at a temperature in the range of 1–5  $\mu$ K. These thermal clouds also have lifetimes well over 100 s.

To bring the atoms closer to the surface, we smoothly reduce the current flowing in the guide wire during the last 1 s of evaporation. The arrival of the atoms at the desired height coincides with the end of the evaporation ramp. The cloud is then viewed on a CCD camera using absorption imaging with the probe beam propagating along the x direction. At distances above 100  $\mu$ m from the surface the axial profile of the cloud (i.e., along the zdirection) is Gaussian, as shown in Fig. 2. However, as the cloud approaches the surface of the wire, it breaks into fragments, similar to those observed by other groups close to Cu wire [14,15]. At a height of 27  $\mu$ m, the potential wells responsible for the fragmentation shown in Fig. 2 are  $\sim 4 \ \mu K$  deep and are due to a magnetic field parallel to the wire [16] that alternates between  $\sim \pm$ 30 mG. Relative to the expected field of the wire this anomalous part is  $\sim 10^{-3}$ . Compared with Refs. [14,15], the current density in our wire is much lower, making it unlikely that the fragmentation is a result of current instability at high density, as has been suggested [16]. It remains an open question whether this phenomenon is due to some fundamental physics, such as arrangement of the spins in the conduction electrons [14,16], or simply to imperfections in the geometry of the wire.

Figure 3 shows the total number of atoms in the (fragmented) microtrap versus time, measured at a height of

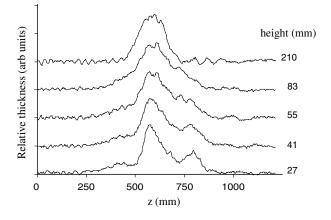


FIG. 2. Optical thickness of a 5  $\mu$ K atom cloud measured by absorption imaging. As the cloud approaches the wire it begins to fragment along the z direction. This effect has previously been seen above Cu wires.

29(4)  $\mu$ m above the Al surface of the wire. A small number of atoms is used in order to avoid the unnecessary complication of 3-body loss from the trap. This curve yields a lifetime of 5.1(0.3) s—very much shorter than the lifetime far from the wire. The increased loss rate is due to spin-flip transitions  $|2, 2\rangle \rightarrow |2, 1\rangle \rightarrow |2, 0\rangle$  driven by thermal fluctuations of the magnetic field close to the wire. The resonant frequency  $f_0$  for this spin-flip transition is  $\frac{1}{2}\mu_B B_0/h$ , where  $\mu_B$  is the Bohr magneton, *h* is Planck's constant, and  $B_0$  is the magnetic field at the center of the trap, where the cold atoms reside. This field is produced by the combination of the transverse wires and the axial bias field  $B_z$ , and it is controlled by adjusting  $B_z$ .

Since the evaporative cooling uses these same spin-flip transitions, we are able to determine the transition frequency  $f_0$  by making several evaporations with different values of the final rf frequency. The final temperature, as measured by the mean square length of the thermal cloud,

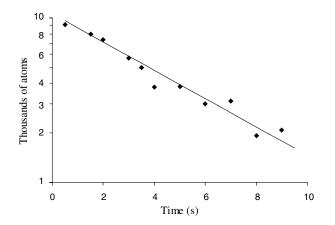


FIG. 3. Number of atoms remaining in the microtrap versus time at a distance of 29  $\mu$ m from the conducting surface of the wire. The cloud is initially at a temperature of 1  $\mu$ K.

is linearly related to this final frequency. Extrapolation to zero cloud length yields the spin-flip frequency  $f_0$  of atoms at the center of the trap. The measurements shown in Fig. 3 were taken with  $f_0 = 1.8$  MHz, corresponding to a field of  $B_0 = 2.6$  G.

Figure 4 shows the lifetimes measured by similar decay curves at a variety of distances from the wire. There are two series of measurements using two different values of the spin-flip frequency:  $f_0 = 1.8(1)$  MHz (open circles), and  $f_0 = 560(10)$  kHz (filled squares). The contribution of the transverse wires to  $B_0$  varies slightly with height, but we have measured this and have compensated for it by adjusting  $B_{z}$  so as to maintain constant  $f_{0}$ . In each series, the lifetime for atoms to remain in the trap exhibits a strong dependence on the distance from the surface, decreasing by an order of magnitude as the distance is reduced from 80 to 30  $\mu$ m. We note that the radial trap frequency is by contrast almost constant over this height range (735–865 Hz for the larger  $f_0$  and 1270–1560 Hz for the smaller), because it is inversely proportional to the distance r from the center of the wire. At a given height, atoms with a lower spin-flip frequency have a shorter lifetime.

The solid (dashed) curves in Fig. 4 show calculated thermal spin-flip lifetimes  $\tau$  for rubidium atoms in the  $|2, 2\rangle$  state above a thick plane slab of Cu (Al) with  $f_0 = 560$  kHz and 1.8 MHz. In calculating these curves, we have found that the analytical expression given in Eq. 23 of Henkel *et al.* [18] is in error by a factor of 2–3 over this range. We therefore integrated their Eq. 22 numerically to

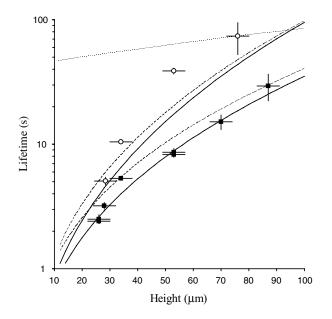


FIG. 4. Lifetime of trapped atoms versus distance from the Al surface of the wire. Filled squares (open circles): measurements with spin-flip frequency  $f_0 = 560$  kHz ( $f_0 = 1.8$  MHz). Solid (dashed) lines: calculated lifetimes above a thick slab of copper (aluminum) for these two spin-flip frequencies. Dotted line: expected scaling for technical noise.

obtain the results shown [20]. The main point we report here is that our experiment shows lifetimes limited only at the level of the thermal spin flips: the lifetime has not been unexpectedly cut short by some other mechanism. This is in itself a significant experimental result for those interested in developing atom chips for quantum coherent manipulations of atoms. Beyond that, our results provide a preliminary test of the validity of the thermal spin-flip theory. To the best of our knowledge, the theory for this new effect above a metal surface has been worked out only for the case of a plane slab, as in Ref. [18], and since our experiment involves atoms above a cylindrical wire, we cannot expect the points in Fig. 4 to lie exactly on the lines. However, the atom-surface distance is less than half the radius of the wire, so it is reasonable to expect some correspondence between this theory and our experiment, and indeed that is seen in Fig. 4. At the lower frequency,  $f_0 = 560$  kHz, the skin depth of the aluminum is 110  $\mu$ m, whereas the thickness of the aluminum outer layer is 55  $\mu$ m. We therefore expect that fluctuations in the copper core will contribute to the decay rate and that the lifetime will be somewhere between the two lowest theory curves. At the higher frequency, the skin depth is only 61  $\mu$ m, and the lifetime should be closer to the aluminum value shown in the topmost curve. The agreement we see between theory and experiment seems to us remarkably good given the differences in geometry of the two. A more precise test of the theory will have to await a proper cylindrical calculation.

This result has been achieved after taking considerable care to control technical noise in the currents being carried by the wires. In the bandwidth 20 Hz-300 kHz, our current regulation circuits have rms noise below 20  $\mu$ A, corresponding to a part per 10<sup>6</sup> of the dc current. At present, however, the reference voltages that control  $B_x$ and the guide wire current are provided by waveform generators whose outputs do not have this level of stability. They increase the total noise current at audio frequencies by a factor of  $\sim 100$ . We observe a heating rate in the range 0.1–0.5  $\mu$ K/s that is thought to be the consequence of this. To look for rf technical noise close to the spin-flip frequency we initially used the antenna that drives the evaporative cooling as a receiver to search for stray fields. A spectrum analyzer revealed some very weak rf signals that were removed after paying attention to ground loops and shielding. The atoms themselves then proved to be much more sensitive spectrum analyzers. With the atoms at 225  $\mu$ m above the wire, a scan of  $f_0$  revealed a resonant dip in the lifetime at approximately 1 MHz. This was found to be from a parasitic oscillation in the optical shutter drivers that appeared while the shutters were closed. Once it was removed, no other rf radiation could be detected. If there were any residual rf current in the guide wire, the resulting field would vary as 1/r, giving a lifetime proportional to  $r^2$ , as verified by [17]. This power law is indicated by the dotted line in Fig. 4, arbitrarily

placed to pass through one of our data points. Clearly the height dependence we observe is much stronger than that. Spin flips induced by rf pickup cause a loss rate proportional to the rf power at the spin-flip frequency  $f_0$ . This leads to a constant ratio of lifetimes, not the height-dependent ratio that we observe.

For the short-range points presented in Fig. 4 we checked that the lifetime did not change when the size of the cloud was increased by raising the temperature to 5  $\mu$ K. We are therefore confident that the lifetimes presented here are not influenced at all by atoms hitting the surface, despite its close proximity. Presumably the BEC provides reliable spin-flip data at shorter distances than the 27  $\mu$ m presented here because it has a smaller radius than the thermal cloud, typically 1–2  $\mu$ m. However, we do not present any shorter range data here because we have no way other than changing the cloud size to distinguish surface loss from spin-flip loss.

Atom chips aim to control quantum superposition and entanglement in neutral atoms. From this point of view, it is very undesirable that thermal fluctuations relax the atomic spins over a few seconds at a distance of 30  $\mu$ m from the surface. This problem can be reduced if the metallic parts of the chip surface are restricted to being thin films. Many atom chips are made by covering a substrate with a 3–5  $\mu$ m thick metal layer, which is then patterned lithographically. Since the fields radiated by thermal current fluctuations must propagate through the conductor to reach the atoms, we can consider that the current fluctuations of interest in a thick slab are localized within a skin depth of the surface, i.e., in a layer  $\sim 100 \ \mu m$  thick. Hence, metal layers that are 3–5  $\mu m$ thick generate 20-30 times less noise power at these frequencies and produces a correspondingly lower spinflip loss rate. Recently we used a different apparatus to hold rubidium atoms in a magnetic trap 30  $\mu$ m away from a 200 nm-thick gold film on a videotape substrate, where this argument predicts a spin-flip lifetime due to the gold of many hundreds of seconds. We did not know what to expect from the videotape. We see no effect of the surface [21], indicating a thermal spin-flip lifetime of at least 50 s at a height of 30  $\mu$ m. Leanhardt *et al.* [17] have also observed long lifetimes at 70  $\mu$ m from a 5  $\mu$ m Cu film.

Even so, the coupling between the atoms and the substrate poses an important technical difficulty for atom chips if the atoms are to approach the surface much more closely than ~10  $\mu$ m. One method could be to cool the chip to suppress the thermal fluctuations, and to use superconducting wire to avoid dissipating heat. An alternative is to avoid current-carrying wires altogether on the surface of the chip by using permanent magnets to produce the required microscopic structures. Videotape provides a simple way of making structures down to ~5  $\mu$ m in size [7]. In our laboratory we loaded atoms into microtraps formed above sinusoidally magnetized videotape, where they can readily be cooled by evaporation and manipulated by externally applied magnetic fields [21]. In order to reach an even smaller length scale, we are now exploring the use of magneto-optical films for making magnetic traps. These are metallic, but can be as thin as 30 nm and still produce traps several mK deep a few  $\mu$ m away from the surface.

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