

## Measurement of the Interaction Energy near a Feshbach Resonance in a ${}^6\text{Li}$ Fermi Gas

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We investigate the strongly interacting regime in an optically trapped  ${}^6\text{Li}$  Fermi mixture near a Feshbach resonance. The resonance is found at 800(40) G in good agreement with theory. Anisotropic expansion of the gas is interpreted by collisional hydrodynamics. We observe an unexpected and large shift (80 G) between the resonance peak and both the maximum of atom loss and the change of sign of the interaction energy.

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The achievement of Bose-Einstein condensation (BEC) in dilute atomic gases [1] has naturally triggered research on cooling of Fermi gases to quantum degeneracy. Several groups have now reached Fermi degeneracy in  ${}^{40}\text{K}$  and  ${}^6\text{Li}$  with temperatures down to about 0.1 to 0.2 of the Fermi temperature  $T_F$  [2–7]. One of the major goals of this research is to observe the transition to a superfluid phase [8], the analogue of the superconducting phase transition in metals [9]. Very low temperatures and strong attractive interactions in a two-component Fermi gas are favorable conditions to reach this superfluid state. The interactions in atomic gases at low temperature are usually described by a single parameter, the  $s$ -wave scattering length  $a$ . This quantity can be tuned near a Feshbach resonance where the sign and magnitude of  $a$  can be adjusted by means of an external magnetic field  $B$  [10,11]. Therefore, when entering the regime of strong interactions, dilute Fermi gases offer unique opportunities to test new theoretical approaches. For instance, near resonance, the critical temperature for superfluidity has been predicted to be as high as 0.25–0.5  $T_F$  [12], a temperature range experimentally accessible.

Both  ${}^{40}\text{K}$  and  ${}^6\text{Li}$  possess Feshbach resonances at convenient magnetic fields [13,14]. In Fig. 1 is plotted the theoretical scattering length for the mixture of the two lowest spin states of  ${}^6\text{Li}$ ,  $|F, m_F\rangle = |1/2, 1/2\rangle$ ,  $|1/2, -1/2\rangle$ , calculated from updated potentials consistent with the recently measured zero crossing of  $a$  at  $B \approx 530$  G [15,16]. Beside the broad resonance near 855 G on which we concentrate in this work, there is a very sharp resonance at 545 G, and increased atom losses have been measured peaking around 545 and 680 G [5]. Recently, strong interactions in this Fermi mixture have been demonstrated at a fixed magnetic field of 930 G through anisotropic expansion of the gas [7]. In this Letter we study the atomic interactions through the entire range of the Feshbach resonance, from 600 G up to 1.3 kG. We measure three physical quantities (i) the interaction energy with positive and negative values, (ii) the anisotropy

of the atomic cloud during the expansion, and (iii) the atom loss. We find the resonance at 800(40) G, where the trapped gas is most strongly interacting, with a ratio between the interaction energy and the kinetic energy reaching  $-0.3$ . Surprisingly, we observe a large shift ( $\sim 80$  G) between this position, and both the location of maximum loss in the gas and the change of sign of the interaction energy. Our results are in partial agreement with the physical picture of a Feshbach resonance in the region of strong interactions but also reveal unexpected effects. Finally, they point toward the best experimental conditions to search for the superfluid transition in  ${}^6\text{Li}$ .

Our experimental approach to measure the interaction energy of a two-component Fermi gas is based on the analysis of time of flight (TOF) images of atoms released from an anisotropic trap. The energy of the trapped gas, prepared at any value of the  $B$  field, is the sum of potential, kinetic, and interaction energies,  $E_{\text{pot}} + E_{\text{kin}} + E_{\text{int}}$ . Switching off abruptly the trapping potential ( $E_{\text{pot}} \rightarrow 0$ ), the gas is released and expands for a variable time before an absorption image is recorded. At long expansion time, the spatial distribution of the cloud reflects the velocity distribution. This procedure is done routinely for BEC studies [1] and has been recently investigated theoretically for Fermi gases in [17]. The novelty of our approach

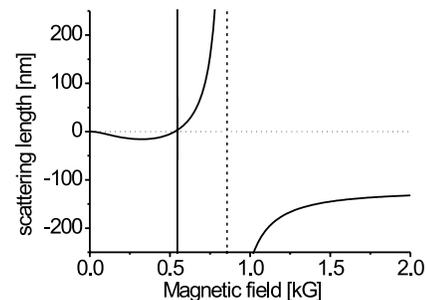


FIG. 1. Predicted scattering length versus magnetic field in  ${}^6\text{Li}$   $|F, m_F\rangle = |1/2, 1/2\rangle, |1/2, -1/2\rangle$  mixture.

is in what follows. Because of the low inductance of the coils used to produce the magnetic field  $B$ , this field can be switched off rapidly ( $\leq 20 \mu\text{s}$ ) at any desired time during the expansion of the atomic cloud. For  $B = 0$  the atoms have negligible interactions since  $a \approx 0$  (Fig. 1). As a consequence, the expansion of the gas can be recorded without atomic interactions during the time of flight period ( $B = 0$ ), or with interactions during the TOF period ( $B \neq 0$ ). Since expansions with  $B = 0$  are ballistic, they reflect the kinetic energy of the initially trapped gas,  $E_{\text{kin}}$ . On the other hand, for TOF with  $B \neq 0$ , the interaction energy is converted into kinetic energy during the early stage of the expansion ( $\approx 150 \mu\text{s}$ ). TOF images at long time reflect the released energy,  $E_{\text{rel}} = E_{\text{kin}} + E_{\text{int}}$  [18]. Comparisons between TOF images with  $B = 0$  and  $B \neq 0$  allow us to simply deduce the ratio  $E_{\text{int}}/E_{\text{kin}}$ . We operate at temperatures between  $0.5 T_F$  and  $T_F$ , where Fermi degeneracy does not play an important role. Even in this nearly classical regime, the gas is found to be strongly interacting. The observed expansions with  $B \neq 0$  are anisotropic because of collisional hydrodynamics, as predicted in [19,20], and observed in [7,21,22].

In our experiment, a gas of  $3 \times 10^5$   $^6\text{Li}$  atoms is prepared in state  $|1/2, 1/2\rangle$  in a Nd:YAG crossed beam optical dipole trap at a temperature of  $10 \mu\text{K}$  [4,23]. The horizontal beam (respectively, vertical) propagates along  $x$  ( $y$ ), has a maximum power of  $2.5 \text{ W}$  ( $2.5 \text{ W}$ ), and a waist of  $\sim 25 \mu\text{m}$  ( $\sim 30 \mu\text{m}$ ). Using a radio frequency field, we drive the Zeeman transition between  $|1/2, 1/2\rangle$  and  $|1/2, -1/2\rangle$  to prepare a balanced mixture of the two states at any chosen value of  $B$  between  $2 \text{ G}$  and  $1.3 \text{ kG}$ . Using a Stern-Gerlach method we check that the populations in both states are equal to within 10%.

We observe that atom losses occur with very different rates below and above resonance. Near  $720 \text{ G}$  the lifetime of the gas is on the order of  $10 \text{ ms}$ , whereas near  $900 \text{ G}$ , the lifetime is in excess of  $10 \text{ s}$ . The latter is surprisingly large in comparison with similar situations for bosons near a Feshbach resonance [24]. Therefore, we performed two sets of experiments—one with the spin mixture prepared above resonance ( $1060 \text{ G}$ ) and the other below resonance ( $5 \text{ G}$ ). In the first set, evaporative cooling is performed by lowering the power of the vertical beam. It produces  $2N = 7 \times 10^4$  atoms at  $T \approx 0.6 T_F$  with  $k_B T_F = \hbar \bar{\omega} (6N)^{1/3}$  and  $\bar{\omega} = (\omega_x \omega_y \omega_z)^{1/3}$ . The trap is nearly cigar shaped with frequencies  $\omega_x/2\pi = 1.1 \text{ kHz}$ ,  $\omega_y/2\pi = 3.0 \text{ kHz}$ , and  $\omega_z/2\pi = 3.2 \text{ kHz}$ . The magnetic field is adiabatically ramped in  $50 \text{ ms}$  to a final value where TOF expansion images are taken. The effect of strong interactions is illustrated in Fig. 2, where both typical images and a plot of the expanded cloud Gaussian sizes  $r_x$  and  $r_y$  are displayed. In Fig. 2(a), expansions with  $B = 0$  reveal the initial momentum distribution of the cloud which is isotropic as expected. We deduce the total kinetic energy  $E_{\text{kin}}$  of the trapped gas mixture from Gaussian fits to this distribution. For our moderate quantum degener-

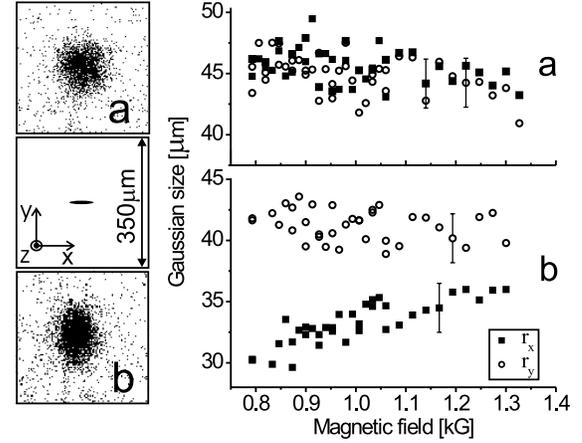


FIG. 2. Left: Geometry of the trapped atomic cloud (center) and expansion images without magnetic field (a) ( $B = 0$ ), and with magnetic field (b) ( $B \neq 0$ ). Right: Corresponding Gaussian sizes of the expanded clouds along  $x$  (squares) and  $y$  (open circles) versus magnetic field. (a) Time of flight images after  $650 \mu\text{s}$  expansion with  $B = 0$ . (b) Images after  $400 \mu\text{s}$  expansion with  $B \neq 0$  and  $250 \mu\text{s}$  with  $B = 0$ . Images shown on the left are for  $B = 900 \text{ G}$ .

acy, the temperature can be estimated from  $k_B T = 2E_{\text{kin}}/3$ . We find  $T \approx 3.5 \mu\text{K} \approx 0.6 T_F$ , constant for fields between  $0.8$  and  $1.3 \text{ kG}$ .

In Fig. 2(b), expansions with  $B \neq 0$  are anisotropic. Little expansion is seen along the weak axis of the trap. The ellipticity of the cloud is inverted because of hydrodynamic behavior during the expansion [7]. Collisions then redistribute the gas energy in the direction of maximum density gradient. The anisotropy  $r_y/r_x$  ranges from  $1.1$  at large  $B$  field to  $1.4$  near  $0.8 \text{ kG}$ , whereas the hydrodynamic scaling equations [19] predict an anisotropy of  $1.53$  in the fully hydrodynamic regime. The usual criterion for hydrodynamicity is found from the ratio  $R$  of the mean free path  $\lambda_0 = (n_0 \sigma)^{-1}$ , where  $n_0$  is the peak density, over the radial size  $r_{\text{rad}}$  of the cloud. For a classical gas neglecting mean-field interactions we find

$$R = \frac{\lambda_0}{r_{\text{rad}}} = \frac{(2\pi)^{3/2}}{N\sigma} \left( \frac{k_B T}{m\bar{\omega}^2} \right) \frac{\omega_{\text{rad}}}{\bar{\omega}}.$$

$R \ll 1$  ( $R \gg 1$ ) corresponds to the hydrodynamic (collisionless) regime. With the predicted value of  $a = -185 \text{ nm}$  at  $1060 \text{ kG}$ , and  $\sigma = 4\pi a^2$ ,  $R = 0.03$ . Therefore, in the early stages of the expansion, the gas is hydrodynamic as in [7,21,22]. Furthermore, with these large values of  $a$  and our typical temperatures, the scattering cross section is energy dependent. At  $B = 1060 \text{ G}$ ,  $|ka| = 0.95$ , where  $k = \sqrt{mk_B T/2\hbar^2}$  is the typical relative momentum of two colliding atoms.  $\sigma$  becomes unitarity limited,  $\sigma = 4\pi a^2/(1 + k^2 a^2) \approx 4\pi/k^2$  for  $|ka| \gg 1$  [25]. Consequently, as the magnetic field is decreased below  $1.3 \text{ kG}$ , the gas gradually enters deeper in the unitarity regime.

This effect has an important consequence on the gas behavior during hydrodynamic expansion: since the relative momentum  $k$  of the colliding atoms decreases as  $k \propto n^{1/3}$  (where  $n$  is the density) [26],  $\sigma$  increases. For a cigar-shaped cloud the expansion is mostly 2D, and  $R$  decreases as  $n^{1/6}$ . The gas becomes more hydrodynamic as the expansion proceeds until this model breaks down when relative momenta  $k$  become too small to remain in the unitarity limit [27]. Then  $R$  increases as  $n^{-1/2}$ . Consequently, the larger  $|a|$ , the longer the expansion remains hydrodynamic and the stronger the anisotropy is. When the  $B$  field is decreased in Fig. 2, the anisotropy increase indicates that the scattering length becomes more and more negative. Therefore, this puts an upper bound of  $\approx 800$  G for the position of the Feshbach  $s$ -wave resonance.

In the second set of experiments, we focused on the lower magnetic field values between 550 and 820 G, the region where losses have been observed. The mixture is prepared at low magnetic field, and the field is ramped up to 320 G, where the scattering length versus  $B$  has a local minimum of  $a = -8$  nm. Evaporative cooling is performed there, leading to  $T = 2.4$   $\mu$ K and  $\omega_x/2\pi = 0.78$  kHz,  $\omega_y/\pi = 2.1$  kHz, and  $\omega_z/2\pi = 2.25$  kHz. Finally, the  $B$  field is ramped to 555 G in 50 ms where  $a \approx 0$ , and then in 10 ms to different values between 600 and 850 G where TOF expansions are recorded. The Gaussian widths and number of detected atoms versus  $B$  are plotted in Fig. 3. Comparing with the data of Fig. 2, we observe several new features. First, the number of detected atoms has a pronounced minimum near 725 G, a position compatible with previously published results [5]. These losses are associated with a strong heating of the cloud [Fig. 3(a)].

Second, in Fig. 3(b) for  $B \leq 675$  G, the expansion is isotropic, consistent with a collisionless regime. Here  $a$  is predicted to be  $0 \leq a \leq 55$  nm. Since  $ka \leq 0.35$ , the scattering cross section is independent of energy and the gas is not hydrodynamic ( $R \geq 1$ ). Above 700 G, a pronounced asymmetry in TOF images appears abruptly. The gas enters the hydrodynamic regime (and unitarity limit) with the same inversion of ellipticity as in Fig. 2, a signature of strong interactions. The maximum anisotropy, which we measure at  $B = 800(40)$  G, locates the peak of the Feshbach resonance, in agreement with the predicted position of 855(30) G. A unique feature of the resonance is the very large shift ( $\approx 80$  G) between the resonance peak and the maximum of loss and heating. This can be qualitatively explained by the creation for  $a \geq 0$  of weakly bound molecules by three-body recombination. According to [28], in this process the binding energy of the molecule  $\hbar^2/ma^2$  for large  $a$  is dissipated into kinetic energy of the atom + molecule system. For large  $a$ , the binding energy is small and the heating associated with three-body recombination is negligible, and the molecules remain optically trapped. For small or

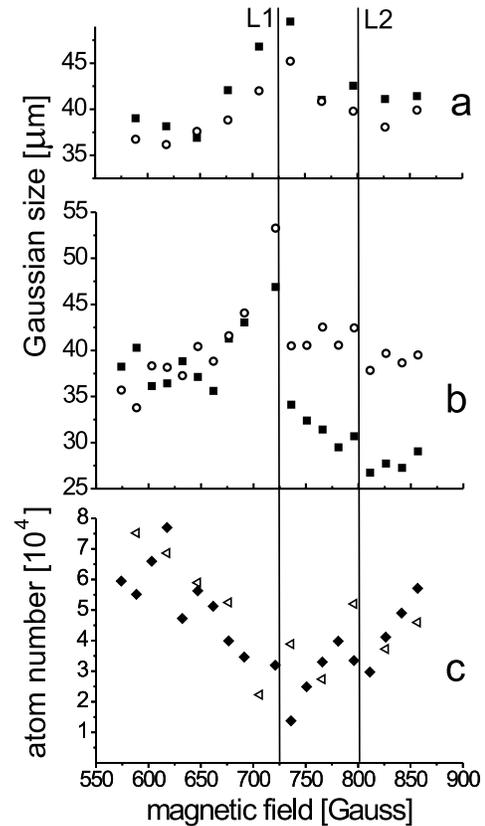


FIG. 3. (a),(b) Gaussian sizes along  $x$  and  $y$  versus magnetic field. The detection conditions and symbols are the same as in Fig. 2. (a) Expansion with  $B = 0$ , (b) expansion with  $B \neq 0$ , (c) number of detected atoms versus magnetic field. Open triangles and black diamonds correspond to graphs (a) and (b), respectively. Line L1 corresponds to the maximum of loss, and L2 to the resonance position.

intermediate values of  $a$ , the collision products are highly energetic, resulting in strong heating and loss. At the maximum of loss (720 G),  $a = 102$  nm and  $\hbar^2/k_B m a^2 = 7$   $\mu$ K is on the order of the trap depth in the weakest direction  $x$ .

As explained above, the interaction energy can be calculated directly from the difference of release energies, obtained from the Gaussian sizes of Figs. 2 and 3. In Fig. 4 is plotted  $E_{\text{int}}/E_{\text{kin}}$  versus magnetic field. As  $B$  is increased above 550 G,  $E_{\text{int}}$  first goes up, due to increasing  $a$ , in accordance with mean-field theory. Near 720 G,  $E_{\text{int}}$  changes sign abruptly. The ratio  $E_{\text{int}}/E_{\text{kin}}$  then exhibits a plateau near  $-0.12$  up to 850 G. Note that for the data at low field (Fig. 3) the number of atoms and temperature changes with  $B$ . For the data at higher fields in Fig. 4, we find  $E_{\text{int}}/E_{\text{kin}} = -0.3$  at resonance, a value comparable to the  $T = 0$  calculation of [29]. The interaction energy becomes less and less negative above 800 G. The discrepancy between the two sets of data in the overlap region can be due to different temperature, confinement, and atom number. The solid/dashed curve in Fig. 4 shows

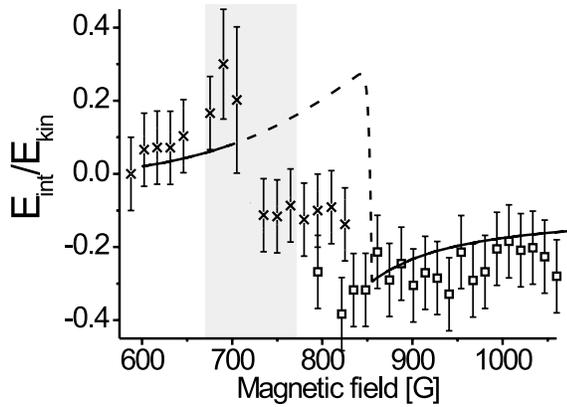


FIG. 4. Ratio of the interaction to kinetic energy versus magnetic field. Open squares: samples prepared above resonance as in Fig. 2. Crosses: average of three sets of data, recorded in conditions of Fig. 3. Solid/dashed curve: mean field calculation for  $7 \times 10^4$  atoms and kinetic energy  $E_{\text{kin}}/k_B = 5.25 \mu\text{K}$ . The gray area indicates the region of losses.

the result of a mean-field calculation based on a full energy dependent phase shift for the atom-atom scattering. The interaction and kinetic energies are obtained from a self-consistent distribution function that contains the interparticle interaction and trapping potential. These calculations agree with experiment for  $B \geq 850$  G where  $a < 0$  and weakly bound molecular states do not exist, and for  $B \leq 720$  G ( $a > 0$ ) where such molecules escape from the trap after being formed. For  $720 < B < 850$  G ( $a > 0$ ) this approach also gives  $E_{\text{int}} > 0$  (dashed curve in Fig. 4). The observed negative ratio  $E_{\text{int}}/E_{\text{kin}}$  in this region is obtained including the presence of weakly bound molecules in the gas, in analogy with a negative value of the second virial coefficient for a Boltzmann gas [30]. Details of our calculations will be published elsewhere.

In summary, we have studied a Fermi gas mixture in the strongly interacting regime near a Feshbach resonance. New features have been observed which may be the signature of richer physics, for instance molecule formation [31]. Anisotropic expansions are observed both for repulsive and attractive mean-field interactions, in a moderately degenerate Fermi gas. This is interpreted in terms of collisional hydrodynamics without invoking Fermi superfluidity.

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- [1] Proceedings of the CXL International School of Physics "Enrico Fermi," edited by M. Inguscio, S. Stringari, and C. Wieman (Italian Physical Society, Bologna, 1999).
- [2] B. DeMarco and D. Jin, *Science* **285**, 1703 (1999).
- [3] A. G. Truscott *et al.*, *Science* **291**, 2570 (2001).
- [4] F. Schreck *et al.*, *Phys. Rev. Lett.* **87**, 080403 (2001).
- [5] K. Dieckmann *et al.*, *Phys. Rev. Lett.* **89**, 203201 (2002).
- [6] G. Modugno *et al.*, *Science* **297**, 2240 (2002).
- [7] K. O'Hara *et al.*, *Science* **298**, 2179 (2002); M. E. Gehm *et al.*, *Phys. Rev. A* **68**, 011401 (2003).
- [8] A. G. Leggett, *J. Phys. (Paris)* **C7**, 19 (1980).
- [9] J. Bardeen, L. Cooper and J. Schrieffer, *Phys. Rev.* **108**, 1175 (1957).
- [10] H. Feshbach, *Ann. Phys. (N.Y.)* **5**, 357 (1958); **19**, 287 (1962).
- [11] E. Tiesinga, B. J. Verhaar, and H. T. C. Stoof, *Phys. Rev. A* **47**, 4114 (1993).
- [12] M. Holland, S. J. J. M. F. Kokkelmans, M. L. Chiofalo, and R. Walser, *Phys. Rev. Lett.*, **87**, 120406 (2001).
- [13] M. Houbiers, H. T. C. Stoof, W. I. McAlexander, and R. G. Hulet, *Phys. Rev. A* **57**, R1497 (1998).
- [14] T. Loftus *et al.*, *Phys. Rev. Lett.* **88**, 173201 (2002).
- [15] K. O'Hara *et al.*, *Phys. Rev. A*, **66**, 041401(R) (2002).
- [16] S. Jochim *et al.*, *Phys. Rev. Lett.*, **89**, 273202 (2002).
- [17] C. Menotti, P. Pedri, and S. Stringari, *Phys. Rev. Lett.* **89**, 250402 (2002).
- [18] We derive  $E_{\text{rel}}$  as if the whole expansion had been ballistic. Simulations show that this approximation induces a maximum error of 3%.
- [19] Y. Kagan, E. L. Surkov, and G. V. Shlyapnikov, *Phys. Rev. A*, **55**, R18 (1997).
- [20] E. Arimondo, E. Cerboneschi, and H. Wu, in Ref. [1], p. 573.
- [21] C. A. Regal and D. Jin, *Phys. Rev. Lett.* **90**, 053201 (2003).
- [22] I. Shvarchuck *et al.*, *Phys. Rev. Lett.* **89**, 270404 (2002).
- [23] L. Khaykovich *et al.*, *Science* **296**, 1290–1293 (2002).
- [24] S. Inouye *et al.*, *Nature (London)* **392**, 151 (1998).
- [25] Modification of this formula in the strongly degenerate regime; see R. Combescot, cond-mat/0302209.
- [26] In an adiabatic transformation of the gas, the phase space density is constant during the time of flight.
- [27] G. V. Shlyapnikov, in *Proceedings of the 18th International Conference on Atomic Physics*, edited by H. R. Sadeghpour, E. J. Heller, and D. E. Pritchard (World Scientific, Singapore, 2002).
- [28] D. S. Petrov, *Phys. Rev. A* **67**, 010703 (2003); L. Pricoupenko, cond-mat/0006263.
- [29] H. Heiselberg, *Phys. Rev. A* **63**, 043606 (2001).
- [30] L. D. Landau and E. M. Lifshitz, *Statistical Physics*, Part 1 (Pergamon Press, Oxford, 1980); L. P. Pitaevskii and S. Stringari (private communication).
- [31] E. A. Donley, N. R. Claussen, S. T. Thompson, and W. E. Wieman, *Nature (London)* **417**, 529 (2002).