Magnetic Propulsion of Intense Lithium Streams in a Tokamak Magnetic Field

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This paper describes the effect and gives the theory of magnetic propulsion which allows driving free surface plasma facing liquid lithium streams in tokamaks. In the approximation of a thin flowing layer the MHD equations are reduced to one integrodifferential equation which takes into account the propulsion effect, viscosity, and the drag force due to magnetic pumping and other interactions with the magnetic field. A stability criterion is obtained for stabilization of the "sausage" instability of the streams by centrifugal force.

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For fusion, liquid lithium has unique physical properties [low melting temperature of $T_{\rm melt} \simeq 180\,^{\circ}$C, low density $\rho \simeq 0.53\ {\rm g/cm^3}$, low viscosity $\nu \simeq 0.0004$ Pa sec in combination with high thermal capacity $c_p \simeq 4200\ {\rm J/(kg\,K)}$ and thermal conductivity $\kappa \simeq 47\ {\rm W/(m\,K)}$ [1]] which would make it potentially the best coolant for blankets of thermonuclear devices. Nevertheless, because of the rather high electric conductivity, $\sigma \simeq 3.3 \times 10^6 (\Omega\,{\rm m})^{-1}$ [2] liquid lithium has been essentially dropped from the design of magnetic fusion reactors. In the strong magnetic field of tokamaks driving a lithium flow in metal pipes or channels consumes energy and creates high pressure in the wall, while insulated wall pipes still require a substantial technology development for the fusion reactor environment. Existing proposals for using so-called "liquid walls" for power extraction from the tokamak reactor plasma [3] rely on passive lithium flow which is sensitive to a variety of MHD effects. A recently discovered (December, 1998 [4]) effect of magnetic propulsion suggests another way of using lithium, i.e., with the actively driven free surface streaming along the inner walls of the vacuum chamber.

This paper outlines the theoretical basis of magnetic propulsion. Figure 1(a) shows two lithium layers of about 1 cm thick, which are held to the surface by the $\mathbf{I}_{\rm pol} \times \mathbf{B}_{\rm tor}$ electromagnetic force, where $\mathbf{I}_{\rm pol}$ represents an externally driven electric current through the layers and $\mathbf{B}_{\rm tor}$ is the toroidal magnetic field of the tokamak. The pressure p in the fluid, created by the electromagnetic force, is inhomogeneous in both normal and longitudinal directions. In the normal direction it is analogous to the atmospheric (highly magnified) pressure of the Earth. It vanishes at the free surface of the liquid and reaches maximum (about one or several atmospheres) at the surface of the wall. Its inhomogeneity along the wall surface propels the fluid (as a wind) from the high field side of the chamber, where the $\mathbf{I}_{\rm pol} \times \mathbf{B}_{\rm tor}$ force is largest, to the low field side.

Neglecting viscosity and possible electromagnetic drag forces, the fluid velocity $\mathbf{V}$ in each flowing layer obeys the Bernoulli law

$$p + \rho g z + \rho \frac{\mathbf{V}^2}{2} = {\rm const}, \qquad p = |\mathbf{I}_{\rm pol} \times \mathbf{B}_{\rm tor}|, \quad (1)$$

where z is the vertical coordinate and $g = 9.8\ {\rm m/sec^2}$ is the acceleration of gravity. Note that the tokamak magnetic field $\mathbf{B}_{\rm tor} = 5$ T, reasonable for reactors, corresponds to a magnetic pressure of 100 atm. Using only a small fraction of it (e.g., 1%–2%) it is possible to create a pressure drop along the flow larger than 1 atm, which corresponds to $\simeq 20$ m/ sec of the liquid lithium velocity.

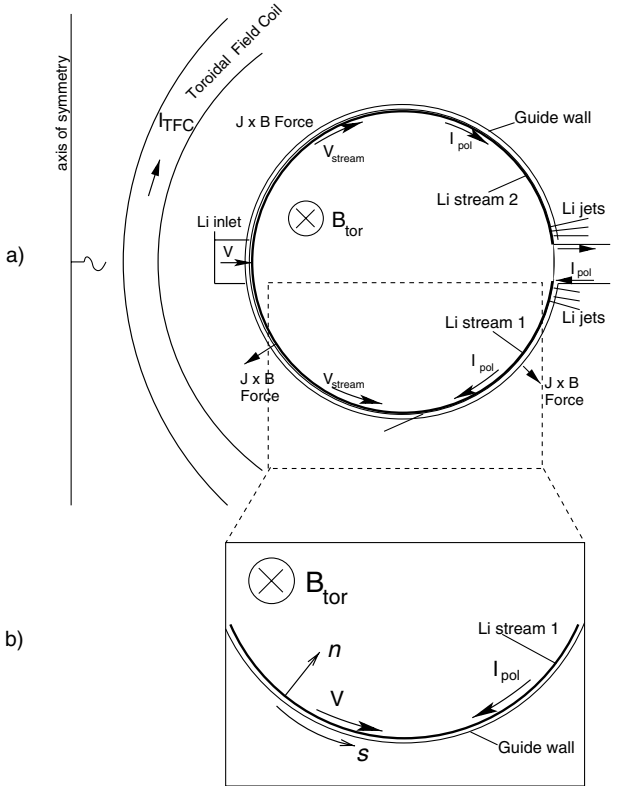

FIG. 1. (a) Cross section of a tokamak with the toroidal field coils (TFC), a vacuum chamber with two lithium streams at the top and bottom halves of the vacuum chamber. (b) Local coordinate system n, s.

Such a velocity is sufficient to provide reactor relevant power extraction from the plasma of the order of 3.5 MW/m^2. For comparison, in the International Thermonuclear Experimental Reactor project this number is about 20 times smaller [5].

This paper describes the propulsion itself. Two closely related problems of how to supply the liquid to the inlet (e.g., in the form of fast thin jets and their conversion into the flow) or to remove it from the chamber (e.g., by converting the flow into a jet shower) are not considered here.

The dynamics of the liquid metal is determined by a standard set of MHD equations

$$\rho\frac{\partial \mathbf{V}}{\partial t} + \rho(\mathbf{V}\cdot\nabla)\mathbf{V} = -\nabla p - \rho g\nabla z + (\mathbf{j}\times\mathbf{B}) + \nu\Delta\mathbf{V}, \tag{2}$$

$$\frac{\partial \mathbf{A}}{\partial t} - \nabla\varphi_E + (\mathbf{V}\times\mathbf{B}) = \frac{\mathbf{j}}{\sigma}, \tag{3}$$

where $\mathbf{A}$ is the vector potential of the magnetic field, φ_E is the electric potential, $\mathbf{j}$ is the current density in the metal, and ν is the viscosity.

The toroidal geometry of tokamaks justifies the consideration of axisymmetric flow with no side walls limiting the streams. Note, that even with a few separating walls present, the magnetic propulsion is not sensitive to the resulting Hartmann drag force [6].

We use the orthogonal curvilinear coordinates n, s, ϕ

$$r = r_w - z_w' n, \qquad z = z_w + r_w' n, \tag{4}$$

suitable for describing the thin liquid flow along a guide plate as shown in Fig. 1(b). Here, r, ϕ, and z are cylindrical coordinates, the functions $r_w = r_w(s)$, $z_w = z_w(s)$ describe the geometry of the wall surface in terms of the length s along the surface, $r_w'^2 + z_w'^2 = 1$. The normal coordinate n is 0 at the guide surface and $n = h$ at the free surface of the flow [here, $h = h(s)$ is the thickness of the metal layer]. It is also convenient to introduce the normalized coordinate $\bar{n}$ (conformal to the flow surface) and the poloidal curvature κ_w of the guide wall surface

$$\bar{n} \equiv \frac{n}{h}, \qquad \kappa_w \equiv |r_w'' z_w' - z_w'' r_w'|. \tag{5}$$

In thin layer approximation we adopt the relations

$$\kappa_w h \ll 1, \qquad \frac{h}{R} \ll 1, \qquad h\frac{\partial}{\partial s} \ll 1. \tag{6}$$

We also can neglect the normal component of velocity

$$V \equiv \mathbf{V}\nabla s, \qquad \mathbf{V}\nabla n \ll V. \tag{7}$$

The height h of the stream is determined by the continuity equation

$$\dot{M} = 2\pi\rho\int_0^h rV dn = 2\pi\rho h\int_0^1 rV d\bar{n}, \tag{8}$$

where $\dot{M}$ is the mass flow in the stream.

The electric current can be represented through its stream functions as

$$\mu_0\mathbf{j} = (\nabla F\times\nabla\phi), \qquad j_n = \frac{F_s'}{\mu_0 r}, \qquad j_s = -\frac{F_{\bar{n}}'}{\mu_0 h r}. \tag{9}$$

The stream function of the electric current F includes three components

$$F = F_w + F_{\mathrm{ext}}\bar{n} + i(\bar{n}, s), \qquad F_w = rB_{\mathrm{tor}} = \mathrm{const}. \tag{10}$$

The major term $F_w = 0.2 I_{\mathrm{TFC}}$ is related to the total current I_{TFC} in the toroidal field coils, $F_{\mathrm{ext}}\bar{n}$ represents the poloidal current (uniformly distributed across the stream), $F_{\mathrm{ext}} = 0.2 I_{\mathrm{pol}}$, driven through the stream by an external source, and i is the stream function of the eddy currents. Assuming wall thickness smaller than h and neglecting the current sharing between the wall and the stream, i satisfies the boundary conditions $i(0, s) = i(1, s) = 0$.

In thin layer approximation the MHD equations can be written in terms of their projections as

$$\rho\frac{\partial V_n}{\partial t} + \rho\kappa_w V^2 = -\frac{p_{\bar{n}}'}{h} - \rho g\frac{z_{\bar{n}}'}{h} - \frac{FF_{\bar{n}}'}{\mu_0 h r^2}, \tag{11}$$

$$\rho\frac{\partial V}{\partial t} + \rho V V_s' = -p_s' - \rho g z_w' - \frac{FF_s'}{\mu_0 r^2} + \frac{\nu}{h^2}V_{\bar{n}\bar{n}}'', \tag{12}$$

$$-\frac{1}{h}\varphi_{E,\bar{n}}' + \frac{VF}{r} = 0, \qquad -\varphi_{E,s}' = -\frac{F_{\bar{n}}'}{\mu_0\sigma h r}. \tag{13}$$

All s derivatives here are taken at $\bar{n} = \mathrm{const}$. Equations (11)–(13) together with Eq. (8) represent a full set of MHD equations. We kept the time derivative of the normal component of velocity in the first equation in order to cover also possible instabilities of the streams. Assuming that the flow is directed along the magnetic surfaces of the equilibrium configuration, we neglected here the interaction of the flow with the poloidal magnetic field. It can be easily included as an additive effect.

The normal component of Ohm's law in the leading approximation gives the distribution of the scalar potential

$$\varphi_E = \varphi_0(s) - \frac{UF_w}{r^2}, \qquad U \equiv -\int_0^{\bar{n}} hrV d\bar{n}, \tag{14}$$

while the longitudinal component

$$\varphi_0' - F_w\left(\frac{U}{r^2}\right)_s' = \frac{F_{\mathrm{ext}} + i_{\bar{n}}'}{\mu_0\sigma h r} \tag{15}$$

determines the stream function $i(\bar{n}, s)$ and eddy currents

$$\frac{i}{\mu_0\sigma h r} = -F_w\left(\frac{1}{r^2}\int_0^{\bar{n}}\tilde{U}d\bar{n}\right)_s', \quad \tilde{U} \equiv U - \int_0^1 U d\bar{n}, \tag{16}$$

$$j^n = -\frac{\sigma F_w}{r}\left[hr\left(\frac{1}{r^2}\int_0^{\bar{n}}\tilde{U}d\bar{n}\right)_s'\right]_s', \quad j^s = \sigma F_w\left(\frac{\tilde{U}}{r^2}\right)_s'. \tag{17}$$

The normal component of the equation of motion in leading order (neglecting the centrifugal force) gives the pressure distribution inside the fluid

$$p = \frac{F^2(1) - F^2}{2\mu_0 r^2} = F_w \frac{F_{\rm ext}(1-\bar{n}) - i}{\mu_0 r^2}. \qquad (18)$$

Its substitution into the longitudinal equation of motion leads to a magnetic propulsion equation (MPE)

$$\rho \frac{\partial V}{\partial t} + \rho V V_s' = -\rho g z_w' - \frac{I_{\rm pol} F_w^2 (1-\bar{n})}{I_{\rm TFC}\mu_0}\left(\frac{1}{r^2}\right)_s' - \sigma h r F_w^2 \left(\frac{1}{r^2}\int_0^{\bar{n}} \tilde{U} d\bar{n}\right)_s' \left(\frac{1}{r^2}\right)_s' + \frac{\nu}{h^2} V_{\bar{n}\bar{n}}'' - (I_\phi B_s)_s'(1-\bar{n}) - \sigma V B_n^2, \qquad (19)$$

which contains only one unknown variable $V(\bar{n}, s)$.

The first term in the right-hand side (RHS) of the MPE is gravity; the second is the magnetic propulsion term, which accelerates the fluid toward the low magnetic field side of the guide wall. The third term in the right-hand side of the MPE represents the diamagnetic drag force or "magnetic pumping," the fourth is viscosity. The last two terms, written without derivation, describe interaction with the poloidal magnetic field. The first is the propulsion effect due to interaction of a toroidal current (if it is present) in the stream with the tangential component of the poloidal field B_s. There, I_ϕ is the linear current density. The second is the classical drag force due to interaction of the flow with the normal (axisymmetric) B_n component (which may appear, e.g., due to plasma displacement).

The diamagnetic drag force with respect to the propulsion term can be assessed assuming a uniform velocity profile across the stream. For values averaged over the cross section, their ratio is determined by the propulsion parameter R_D as

$$\frac{\text{drag force}}{\text{propulsion}} \simeq \frac{R_D r_w'}{3}, \qquad R_D \equiv \Re_2 \frac{I_{\rm TFC}}{I_{\rm pol}}, \qquad (20)$$

$$\Re_2 \equiv \frac{\mu_0 \sigma h^2 V}{r}. \qquad (21)$$

The same propulsion parameter is responsible for possible peeling of the flow surface due to the appearance of "negative" pressure in the flow, if at some point $R_D r_w' > 1$.

For liquid lithium $\mu_0\sigma \simeq 4$ and for typical rector parameters, e.g., $r = 6$ m, $V \simeq 20$ m/sec, $h \simeq 1$ cm, the diamagnetic drag force has only a small effect on flow. Even for $I_{\rm pol} = 0.01 I_{\rm TFC}$, the propulsion parameter $R_D \ll 1$. It is important to notice that the Reynolds number $\Re_2$, introduced here, is different from the "electromagnetic" Reynolds number $\Re_1 \equiv \mu_0\sigma h V$, which plays a role in plasma stabilization by the lithium streams [7–9]. This fact makes lithium dynamics consistent with the possible plasma stabilization by intense lithium streams.

The interaction with the normal B_n component of the magnetic field is characterized by the conventional Reynolds number $\Re_0$

$$\Delta\rho \frac{V^2}{2} = \Re_0 \frac{1}{L}\int_0^L \frac{B_n^2}{2\mu_0} ds, \qquad \Re_0 \equiv \mu_0 \sigma L V, \qquad (22)$$

where L is the length of the flow. Because $\Re_0 \gg 1$, the flow is very sensitive to the quality of alignment of the flow with magnetic surfaces. Nevertheless, the magnetic propulsion allows a certain level of misalignment of the flow with the magnetic surfaces and of the normal magnetic field.

We leave out of the scope of the paper consideration of obstacles to the flow, such as walls around ports in the vacuum chamber. Their drag effect can be assessed based on Hartmann layer theory [6] (assuming electrically insulated walls)

$$\Delta\rho \frac{V^2}{2} \simeq \frac{4\Re_0}{Ha} \frac{B_{\rm tor}^2}{2\mu_0}, \qquad Ha \equiv Bw\sqrt{\frac{\sigma}{\nu}}, \qquad (23)$$

where w is the distance between walls in the flow. Even for a high speed of lithium streams this pressure drop can be made small compared to the driving pressure.

The possible electromagnetic instabilities appear in the next approximation when the centrifugal force and the dependence of the magnetic field line tension on the major radius are taken into account in the equation for pressure

$$\left(p - \frac{F^2(1) - F^2}{2\mu_0 r^2}\right)_{\bar{n}}' = -\rho g r_w' h - \frac{2p}{r} z_w' h - \rho \kappa_w h V^2 \qquad (24)$$

due to explicit dependence of the RHS on the height h of the flow. Thus, the first term drives the gravitational instability (at the top of the vacuum chamber) [10], while the second one drives the "sausage" instability at the high field side, where $z_w' < 0$ [11]. The third term is the stabilizing effect of the centrifugal force.

The centrifugal force robustly stabilizes the sausage instability of the stream if

$$\int_{-1}^{\bar{n}} \left(\rho\kappa_w V^2 - 2p\frac{z_w'}{r}\right) d\bar{n} > 0, \qquad \rho\frac{\langle V^2\rangle}{2} > \frac{p_w}{2\kappa_w r} z_w', \qquad (25)$$

where p_w is the pressure of the stream on the wall surface. Because $|z_w'| < 1$ and the poloidal curvature is determined by the minor radius a of the plasma, $k_w \simeq 1/a$, this stability condition can be fulfilled for sufficiently high inlet stream velocity.

Much smaller flow speed is required for stabilization of the gravitational interchange instability

$$\langle V^2\rangle > \frac{g|r_w'|}{k_w} \qquad (26)$$

for the stream moving along the upper half of the vacuum chamber.

An example of a stationary solution to the MPE (19) for liquid lithium is shown in Fig. 2. The velocity, averaged over the cross section, and shown in Fig. 2(a), rises

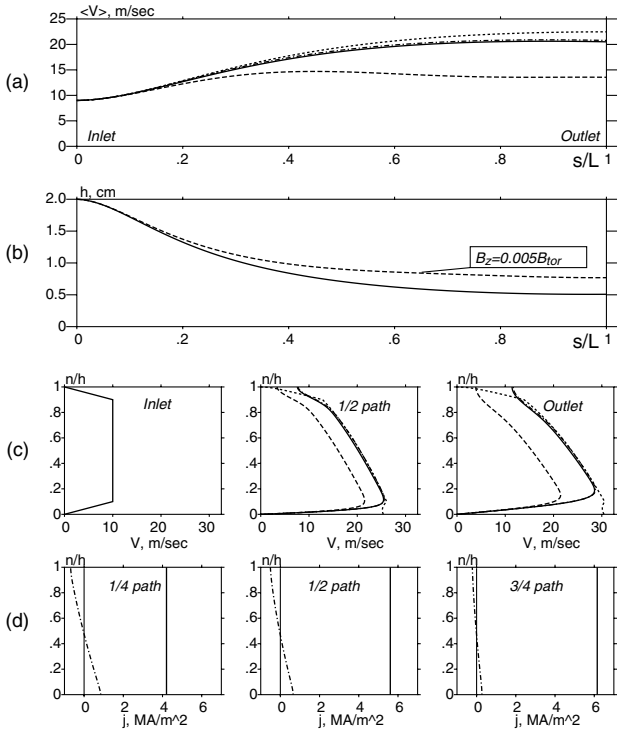

FIG. 2. Numerical solution of the MPE for $R = 6\,\mathrm{m}$, $a = 1.6\,\mathrm{m}$, $B_{\mathrm{tor}} = 5\,\mathrm{T}$ (at $r = R$), $I_{\mathrm{ext}} = 0.01 I_{\mathrm{TFC}} = 1.5\,\mathrm{MA}$, $h_{\mathrm{inlet}} = 2\,\mathrm{cm}$. (a) Averaged velocity and (b) height h as a function of longitudinal coordinate, (c) flow velocity profile in three cross sections along the stream. Solid lines represent solutions of the MPE (19) for $B_n = 0$, dashed lines for $B_n = 0.005 B_{\mathrm{tor}} \cos(s/a)$, dotted line corresponds to the "ideal" solution with neither viscosity nor magnetic pumping, and dot-dashed lines the solution with no magnetic pumping. (d) Profiles of externally driven current density (solid lines) and of eddy current (dashed lines).

from an inlet value to the characteristic level of 20 m/sec at the middle of the path. The dotted line solution was obtained neglecting viscosity and the diamagnetic drag force in MPE. Neglecting diamagnetic drag force (dot-dashed line) shows the very small effect of diamagnetic drag (in contrast to the case of Ref. [12], where this effect was exaggerated). The normal component of magnetic field does affect the propulsion, as shown by a dashed line solution. Note that in this case the imposed vertical magnetic field would correspond 2–3 $h(0)$ in the plasma displacement (depending on the plasma current).

The flow becomes significantly thinner on the way to the outlet point [Fig. 2(b)], Fig. 2(c) shows a comparison of the velocity profile for the cases without the effect of diamagnetic drag force, without both viscosity and drag force and with a normal magnetic field. Figure 2(d) shows a comparison of the driven current density (solid lines) with the eddy currents (dot-dashed lines), which are significantly smaller at all cross sections of the flow. The Ohmic power losses (dominant ones) in the streams are inversely proportional to their height and in the examples are 23 MW (no vertical field) and 18 MW (vertical field present). They are in the acceptable range for power reactors.

The effect of viscosity on magnetic propulsion in the laminar approximation is very weak and for typical velocities 10–20 m/sec of the streams is localized in a submillimeter boundary layer near the surface of the wall. Note that, while magnetic propulsion drives the inner layer of the flow, viscosity propels the surface layer.

Summarizing, the magnetic propulsion of intense lithium streams represents a means for driving controlled lithium streams in a tokamak magnetic field. With the development of the technology of injection/ejection of the streams it can be used for efficient power extraction from the plasma. The magnetic propulsion equation (19) describes the 2-dimensional MHD of the liquid lithium flow and shows robustness of the propulsion effect to viscous drag and diamagnetic drag force, characterized by dimensionless parameters $R_D \ll 1$ and $\Re_2 \ll 1$. While the flow is sensitive to the normal components of the magnetic fields, in the case of magnetic propulsion misalignment by a few thicknesses of the flow is still tolerable. We also have shown that the streams can be stabilized by centrifugal force against sausage and gravitational instabilities. The thin layer approach can be extended to cases which would include disturbances of axisymmetry (like penetrations or nonsymmetric normal fields).

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