## **Benchmarks for the Forward Observables at RHIC, the Tevatron-Run II, and the LHC**

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We present predictions on the total cross sections and on the  $\rho$  parameter for present and future  $pp$ and  $\bar{p}p$  colliders, and on total cross sections for  $\gamma p \rightarrow$  hadrons at cosmic-ray energies and for  $\gamma \gamma \rightarrow$ hadrons up to  $\sqrt{s} = 1$  TeV. These predictions are based on an extensive study of possible analytic hadrons up to  $\sqrt{s} = 1$  TeV. These predictions are based on an extensive study of possible analytic parametrizations invoking the largest hadronic dataset available at  $t = 0$ . The uncertainties on total cross sections reach 1*:*9% at the Relativistic Heavy Ion Collider, 3*:*1% at the Tevatron, and 4*:*8% at the Large Hadron Collider, whereas those on the  $\rho$  parameter are, respectively, 5.4%, 5.2%, and 5.4%.

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In recent works [1,2], we have performed an exhaustive  
study of the analytic parametrizations of soft forward  
data. For this purpose, we gathered the largest available  
set of data at 
$$
t = 0
$$
, which includes all measured total  
cross sections and ratios of the real part to the imaginary  
part of the elastic amplitude ( $\rho$  parameter) for the scat-  
tering of  $pp$ ,  $\overline{p}p$ ,  $\pi^{\pm}p$ ,  $K^{\pm}p$ , and total cross sections for  
 $\gamma p$ ,  $\gamma \gamma$ , and  $\Sigma^{-}p$  [3,4].

Several experiments are under way [6], or being planned, to measure the hadronic amplitudes at  $t = 0$ . Some authors [7–9] also presented what they feel are reference values for the total  $\gamma p$  and  $\gamma \gamma \rightarrow$  hadrons cross sections. Thus it is timely and appropriate to present independently our predictions for the forward observables at the Relativistic Heavy Ion Collider (RHIC), the Tevatron-run II, and the Large Hadron Collider (LHC) as well as for  $\gamma p$  total cross section at cosmic-ray energies and for  $\gamma\gamma$  total cross sections up to 1 TeV.

We can summarize [10] the general form of the parametrizations by quoting the total cross sections for the scattering of *m* on *n*, from which the  $\rho$  parameter is obtained via analyticity. The ingredients are the contribution  $y^{mn}(s)$  of the highest meson trajectories ( $\rho$ ,  $\omega$ ,  $a$ , and *f*) and the rising  $C = +1$  term  $h^{mn}(s)$  from the pomeron contribution to the total cross section:

$$
\sigma_{\text{tot}}^{mn} = (y^{mn}(s) + h^{mn}(s))/s.
$$
 (1)

The first term is parametrized via Regge theory, and we allow the lower trajectories to be partially degenerate, i.e., our experience shows that it is enough to introduce one intercept for the  $C = +1$  trajectories, and another one for the  $C = -1$  [11]. A further lifting of the degeneracy is certainly possible, but does not seem to modify significantly the results [12]. Hence we use

$$
y^{mn}(s) = Y_{+}^{mn}(s/s_1)^{\alpha_{+}} \pm Y_{-}^{mn}(s/s_1)^{\alpha_{-}}, \qquad (2)
$$

with  $s_1 = 1$  GeV<sup>2</sup>. The contribution of these trajectories is represented by RR in the model abbreviations.

As for the part rising with energy, we consider here two main options: it can rise as a lns, or as a  $\ln^2 s$ , with in each case the possibility to add a constant term. We shall not consider the simple-pole parametrization [13], because, after refitting, it is excluded at the the 98% C.L. if one lowers the energy cutoff to 5 GeV [14]. In the following, we shall refer explicitly only to our preferred parametrization of *hmn*, noted PL2:

$$
h^{mn}(s) = s(B^{mn}\ln^{2}(s/s_{0}) + P^{mn}),
$$
 (3)

where  $s_0$  is a universal scale parameter (to be determined by the fits) identical for all collisions.

We consider several possible constraints on the parameters of Eqs. (2) and (3): degeneracy of the reggeon trajectories ( $\alpha_+ = \alpha_-$ ); universality of rising terms (*B<sup>mn</sup>*) independent of the hadrons) [15–17]; factorization for the residues in the case of the  $\gamma\gamma$  and  $\gamma p$  cross sections  $(h_{\gamma\gamma} = \delta h_{\gamma p} = \delta^2 h_{pp})$ ; quark counting rules [18] (predicting the  $\Sigma p$  cross section from  $pp$ ,  $Kp$ , and  $\pi p$ ); and finally the Johnson-Treiman-Freund [19] relation for the cross section differences.

Out of the 256 possible variants, we keep only models which have an overall  $\chi^2$ /dof  $\leq 1.0$  and a non-negative pomeron contribution at all energies. This leaves us 24 point possible models for fits to  $\sigma_{tot}$  for  $\sqrt{s} \ge 10$  GeV, and 5 for  $\sqrt{s} > 4$  GeV/(222 Table XI from [11). If we include the s possible models for fits to  $\sigma_{tot}$  for  $\sqrt{s} \ge 10$  GeV, and 3 for  $\sqrt{s} \ge 4$  GeV(see Table XI from [1]). If we include the  $\rho$ data, we are left with 20 (respectively 4) variants for  $\sqrt{s} \ge 10$  (respectively 5) GeV (see Table XIV from [1]).



FIG. 1. Predictions for total cross sections. The black error band shows the statistical errors to the best fit, the closest curves near it give the sum of statistical and systematic errors to the best fit due to the ambiguity in Tevatron data, and the highest and lowest curves show the total errors bands from all models considered in this Letter (note that the upper curve showing the systematic error is indistinguishable from the highest curve in this case).

Here we shall neither give the list of models nor spell out ranking criteria based on new indicators that quantify certain qualities of the fits, but simply mention that the triple-pole parametrization  $RRP_{\text{nf}}L2_{\text{u}}$  [15,16] was determined to be the highest-ranking model leading to the most satisfactory description of the data. This parametrization has a universal (u)  $B \ln^2(s/s_0)$  term, a nonfactorizing (nf) constant term and nondegenerate lower trajectories.

We start by giving the predictions of this model, adwe start by giving the predictions of this model, adjusted for  $(\sqrt{s} \ge 5 \text{ GeV})$ , with updated data points from ZEUS [5]. The quality of the fit is shown in Table I, column 2. For  $pp$  and  $\bar{p}p$ , the central value of the fit gives

$$
\sigma_{\text{tot}}^{\bar{p}p,pp} = 42.6 \, s^{-0.46} \pm 33.4 \, s^{-0.545} + 35.5
$$

$$
+ 0.307 \ln^2 \left(\frac{s}{29.1}\right),\tag{4}
$$

with all coefficients in mb and *s* in GeV2.

The following predictions include statistical errors calculated from the full error matrix  $E_{ij}$ . We define

$$
\Delta Q = \sum_{ij} E_{ij} \frac{\partial^2 Q}{\partial x_i \partial x_j},\tag{5}
$$

with  $Q = \sigma_{\text{tot}}$  or  $\rho$  and  $x_i$  the parameters of the model. These errors are shown in Figs. 1 and 2 by a filled band, and in Tables II and III.

In these figures and tables, we also give our estimate of the systematic uncertainty coming from the discrepancy between different FNAL measurements of  $\sigma_{\text{tot}}$ : we fit  $RRP_{\text{nf}}L2_{\text{u}}$  either to the high data (CDF) or to the low



FIG. 2. Predictions for the  $\rho$  parameter. The curves and band are as in Fig. 1.

ones (E710/E811), and get two error bands. The distances from the central value of the combined fit to the upper (respectively lower) border of these bands give us the positive (respectively negative) systematic errors. We

TABLE I. Summary of the quality of the fits at different stages of the Review of Particle Physics (RPP) database (DB): DB02, the 2002 RPP DB; DB02Z, the 2002 RPP DB with new ZEUS data; DB02Z-CDF, with the CDF point removed; DB02Z-E710/E811, with E710/E811 points removed. The first line gives the overall  $\chi^2$ /dof for the global fits; the other lines give the  $\chi^2$ /nop for data subsamples; the last line gives in each case the parameter controlling the asymptotic form of cross sections.

	DB <sub>02</sub>	DB02Z	DB02Z	DB02Z		
Sample			$-CDF$	$-E710/E811$		
Total	0.968	0.966	0.964	0.951		
Total cross sections						
$\overline{p}p$	1.15	1.15	1.12	1.05		
p p	0.84	0.84	0.84	0.84		
$\pi^- p$	0.96	0.96	0.96	0.96		
$\pi^+ p$	0.71	0.71	0.71	0.71		
$K^- p$	0.62	0.62	0.62	0.61		
$K^+ p$	0.71	0.71	0.71	0.71		
$\Sigma^- p$	0.38	0.38	0.38	0.38		
$\gamma p$	0.61	0.58	0.58	0.58		
$\gamma\gamma$	0.65	0.64	0.64	0.63		
Elastic forward Re/Im						
$\overline{p} p$	0.52	0.52	0.52	0.53		
p p	1.83	1.83	1.83	1.80		
$\pi^- p$	1.10	1.10	1.09	1.14		
$\pi^+ p$	1.50	1.50	1.52	1.46		
$K^- p$	1.00	0.99	1.01	0.96		
$K^+ p$	1.07	1.07	1.10	0.98		
Values of the parameter B						
	0.307(10)	0.307(10)	0.301(10)	0.327(10)		

TABLE II. Predictions for  $\sigma_{\text{tot}}$  and  $\rho$ , for  $\bar{p}p$  (at  $\sqrt{s}$ 1960 GeV), and for *pp* (all other energies). The central values and statistical errors correspond to the preferred model  $RRP<sub>nf</sub> L2<sub>u</sub>$ , and the systematic errors come from the consideration of two choices between CDF and E710/E811  $\overline{p}p$  data in the simultaneous global fits.

$\sqrt{s}$ (GeV)	$\sigma$ (mb)	
100	$46.37 \pm 0.06$ $-0.09$	$+0.0040$ $0.1058 \pm 0.0012$ $-0.0021$
200	$51.76 \pm 0.12 \begin{array}{c} +0.39 \\ 0.2 \end{array}$	$+0.0051$ $0.1275 \pm 0.0015$ 0.0026
300	$+0.57$ $55.50 \pm 0.17$ $-0.30$	$+0.0055$ $0.1352 \pm 0.0016$ $-0.0028$
400	$+0.71$ $58.41\pm0.21$ $-0.36$	$+0.0056$ $0.1391 \pm 0.0017$ $-0.0030$
500	$60.82 \pm 0.25$ $-0.45$	$+0.0057$ $0.1413 \pm 0.0017$ $-0.0030$
600	$62.87 \pm 0.28$ $-0.48$	$+0.0058$ $0.1416 \pm 0.0018$ –0.0031
1960	$+1.85$ $78.27 \pm 0.55$ –0.96	$+0.0057$ $0.1450 \pm 0.0018$ -0.0030
10000	$+3.6$ $105.1 \pm 1.1$ $-1.9$	$+0.0047$ $0.1382 \pm 0.0016$ $-0.0027$
12000	$108.5 \pm 1.2 \frac{+3.8}{-2.0}$	$+0.0046$ $0.1371 \pm 0.0015$ -0.0026
14000	$111.5 \pm 1.2 + 4.1$	$0.1361 \pm 0.0015$ 0.0025

estimate the total errors as the sum of the systematic and of the statistical uncertainties [20].

One can see that the total errors on total cross sections are of the order of 1*:*9% at RHIC, of the order of 3*:*1% at the Tevatron, and as large as 4*:*8% at the LHC and dominated by the systematic errors. The errors on the  $\rho$  parameter are much larger, reaching 5*:*4% at RHIC, 5*:*2% at the Tevatron, and 5*:*4% at the LHC. This comes from the fact that experimental errors are bigger, hence less constraining, and also from the incompatibility of some lowenergy determinations of  $\rho$  [1]. This means that the systematic error is always bigger than the statistical one.

Concerning the contradictory data, we are forced to use them in our fits until the discrepancy is resolved by further experiments. In the case of the Tevatron data, the discrepancy results in a big shift (of more than  $1\sigma$ ) in the central value of the coefficient *B* of the  $\ln^2 s$  term, which controls the asymptotic behavior, and hence that asymptotic predictions are appreciably weakened by the present situation. The opportunities to measure  $\sigma_{\text{tot}}$  and  $\rho$ will be scarce in the future, thus any new measurement at RHIC, the Tevatron-run II and the LHC should not be missed. Unfortunately, the recent publication of E811 [21] does not clear the problem as their value for  $\rho$  is fully compatible with our preferred model, whereas their number for  $\sigma_{\text{tot}}$  (which is highly correlated with  $\rho$ ) has hardly changed.

TABLE III. Predictions for  $\sigma_{\text{tot}}$  for  $\gamma p \rightarrow$  hadrons and for  $\gamma\gamma \rightarrow$  hadrons for cosmic ray energies. The central values, the statistical errors, and the systematic errors are as in Table II.

$p_{\text{lab}}^{\gamma}$ (GeV)	$\sigma_{\gamma p}$ (mb)	$\sqrt{s}$ (GeV)	$\sigma_{\gamma\gamma}$ ( $\mu$ b)
$0.5 \times 10^{6}$	$0.24 \pm 0.01 \pm 0.01$	200	$0.55 \pm 0.03 \pm 0.03$
	$1.0 \times 10^6$ 0.26 $\pm$ 0.01 $\pm$ 0.01	300	$0.61 \pm 0.04 \pm 0.04$
	$1.0 \times 10^7$ 0.33 $\pm$ 0.02 $\pm$ 0.02	400	$0.66 \pm 0.04 \pm 0.04$
	$1.0 \times 10^8$ 0.42 ± 0.02 $^{+0.03}_{-0.02}$	500	$0.70 \pm 0.05 \pm 0.05$
	$1.0 \times 10^{9}$ $0.52 \pm 0.03 \frac{+0.04}{-0.03}$	1000	$0.84 \pm 0.07 \pm 0.07$

It is interesting to note that the choice of one FNAL result or the other leads to a variation of the overall fit quality, as shown in Table I (last two columns) [22]: the variant with CDF data has slightly better overall  $\chi^2$ /dof and better  $\chi^2$ /nop distribution over subsamples. We can consider this as an indication that the global picture emerging from fits to all data on forward observables supports the CDF data and disfavors the E710/E811 data supports the CDF data and dista<br>at  $\sqrt{s}$  = 1.8 TeV (see also [23]).

Finally, we also present in Figs. 1 and 2 our estimate of the region where new physics would be discovered. For each of the 20 parametrizations which satisfy our criteria each of the 20 parametrizations which satisfy our criteria<br>for applicability [1] for  $\sqrt{s} \ge 10$  GeV, and which obey the Froissart-Martin bound [24], we construct error bands according to Eq. (5). This gives us 20  $1\sigma$ -error bands. Their union represents the ''allowed region'' where analytic models built according to (1) can reproduce the data. A measurement outside this region would imply that new physics ingredients are needed.

To conclude, we believe that we have given here the best possible estimates for present and future  $pp$  and  $\bar{p}p$ facilities. Although one might be tempted to use only data in an energy range close to the one measured, one must realize that analytic parametrizations are constrained both by lower-energy data, and by their asymptotic regime. Because the pomeron mixes (physically and numerically) with the *f* trajectory, fits to all data help to disentangle the two contributions.

A sharpening of our error bars would enable one to decide if the unitarization plays an essential role and what form it takes. This in turn can have an impact on the determination of the survival of probability gaps in hard scattering, and on the usefulness of pomeron exchange as a detection tool.

Any significant deviation from the predictions based on model  $RRP_{nf}L2_u$  will lead to a reevaluation of the hierarchy of models and presumably change the preferred parametrization to another one. A deviation from the 'allowed region'' would be an indication that strong interactions demand a generalization of the analytic models discussed so far, e.g., by adding odderon terms, or new pomeron terms, as suggested by pQCD.

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*Note added.—*After this study was completed, Menon [25] pointed out to us that one of the Akeno points (at  $p_{lab} = 1.85 \times 10^8$  GeV) should be changed to 100 mb instead of the 93 mb we used. We have checked that this does not significantly modify our results—the main effect being to reduce the lower systematic errors [10].

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