Effect of ³He on Submonolayer Superfluidity

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We have studied the superfluid response of ${}^{3}\text{He}{}^{4}\text{He}$ mixture films adsorbed onto porous gold for a wide range of ${}^{3}\text{He}$ and ${}^{4}\text{He}$ coverages, focusing on submonolayer superfluidity. At T = 0, ${}^{3}\text{He}$ appears to float on top of ${}^{4}\text{He}$ and can be viewed as a second substrate that induces its own inert layer. Depending on the ${}^{4}\text{He}$ content, the zero-temperature superfluid mass and the superfluid onset temperature either saturate or vanish with the addition of ${}^{3}\text{He}$. The T = 0 superfluid-insulator phase boundary, which can be described by a simple function, is found in the ${}^{3}\text{He}{}^{4}\text{He}$ coverage plane.

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The superfluid transition in pure ⁴He films adsorbed onto amorphous substrates is a subject of considerable theoretical [1] and experimental [2–6] interest. There have also been a number of experiments exploring the nature of superfluidity and ordering of ³He-⁴He mixture films on such substrates [7–19]. In most of these studies, ³He ranging from a small fraction of a monolayer to multilayers was added to pure ⁴He films with superfluid coverage on the order of and exceeding 1 monolayer (the monolayer coverage is 12.9 μ mol/m² for ⁴He and 10.6 μ mol/m² for ³He). It was found that superfluidity is suppressed [7,8] with the addition of ³He and that in the T = 0 limit ³He tends to reside near the free surface of the film [10–14]. For multilayer ³He this configuration has been termed the "superfluid sandwich" [11].

An interesting question is how does ³He affect superfluidity of a ⁴He film when the total superfluid coverage is only a small fraction of a monolayer. In this Letter we report a systematic study on the suppression of superfluidity by up to 8.5 monolayers of ³He in ⁴He films with superfluid coverage ranging from 4% of a layer to 1 monolayer. For films with superfluid coverage of more than 0.40 monolayers of ⁴He we found that the superfluid fraction decays exponentially with the addition of ³He and superfluidity persists in the low temperature limit no matter how much ³He is added to the film. The measurements allow us to deduce quantitatively the T = 0 phase diagram.

The experimental cell is the same as the one used in Ref. [5]. The substrate is porous gold of 70% porosity consisting of multiply interconnected gold strands of 0.06 μ m in diameter. After each new ³He and/or ⁴He dose, the adsorbed film is annealed at a temperature at which there is considerable vapor pressure (>0.1 torr), then it is cooled to the superfluid onset temperature T_c . This procedure yields consistent results [19]; by repeating this procedure several times without adding any new helium to the film, T_c reproduces to within 0.5 mK.

The 12 different experimental runs were performed by keeping the ⁴He coverage n_4 constant and incrementally adding ³He to the mixture. Figure 1 shows a data set for $n_4 = 29.35 \ \mu \text{mol/m}^2$ with n_{4s} , the superfluid coverage in the absence of ³He, of 0.32 layers. The superfluid cover-

age, n_{4s} , is n_4 less the minimum coverage necessary for superfluidity $n_0 = 25.25 \ \mu \text{mol/m}^2$, often called the inert layer coverage. The areal superfluid density ρ_s , measured by torsional oscillator technique, has been corrected for the tortuosity [20] of the porous gold sample. The curve with the highest T_c corresponds to the pure ⁴He film. One can determine T_c by either extrapolating the superfluid density to zero or by locating the dissipation peak that accompanies the onset of superfluidity. For films with $T_c > 30$ mK, the values obtained by the two different methods are always within 6% of each other. For films with $T_c < 15$ mK we cannot reliably determine T_c for all films from the dissipation signal.

Figure 2 shows the dependence of T_c on the ³He coverage n_3 . For films with $n_{4s} \ge 0.42$ monolayers, T_c decreases rapidly then saturates with the addition of ³He. Such a behavior has been observed in earlier measurements [8–10]. Films showing this behavior are marked by filled symbols in Fig. 2. Films with lower starting ⁴He coverages, however, show a vanishing T_c at a finite n_3 ,



FIG. 1. Superfluid density and dissipation versus temperature for several mixture films having $n_4 = 29.35 \ \mu \text{mol/m}^2$. The ³He coverages for curves a-g are 0, 1.91, 3.83, 7.34, 9.15, 12.55, and 14.91 $\mu \text{mol/m}^2$. For clarity, the magnitude of dissipation for curves d-g has been doubled.

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FIG. 2. Superfluid onset temperature versus n_3 for eight mixture films of different n_4 . The starting superfluid coverages n_{4s} at $n_3 = 0$ are 0.09, 0.16, 0.26, 0.37, 0.42, 0.55, 0.63, and 0.85 monolayers. Lines are guides to the eye and are anchored to data points of higher ³He coverages. The inset shows the same data renormalized by the critical curve $\Theta(n_3)$.

as shown by curves with open symbols. For example, for the curve with $n_{4s} = 0.09$ and 0.26 layers, superfluidity is completely suppressed when n_3 exceeds 0.17 and 0.94 layers, respectively. A critical curve $\Theta(n_3)$ with n_{4s} that is between 0.37 and 0.42 monolayers, which separates the two different classes of behavior, should exist. This curve will smoothly decrease with n_3 and extrapolate to $T_c = 0$ in the limit of a very large n_3 . Since the curves with $n_{4s} = 0.37$ and 0.42 monolayers collapse quite well when vertically shifted, this critical curve $\Theta(n_3)$ can be approximated by subtracting from the $n_{4s} = 0.42$ layer curve 18 mK, the asymptotic T_c value at large n_3 . The critical curve has a total ⁴He coverage $n_4^{crit} = 30.4 \pm 0.2 \ \mu \text{mol/m}^2$ and a starting superfluid onset temperature $T_c = 430 \pm 20$ mK.

The qualitatively different behaviors at large n_3 are accentuated if we divide all curves of Fig. 2 by the critical curve $\Theta(n_3)$. The result, shown in the inset in Fig. 2, bears an interesting resemblance with that for amorphous superconducting Bi films [21]. This comparison illustrates that the ground state at T = 0 of mixtures with large ³He content and that of Bi films changes from insulating to superfluid with just a very small change in the parameter favoring superfluidity (i.e., ⁴He and Bi surface coverages).

Replotting our data in the n_3 , n_4 , T parameter space, the complete phase diagram can be constructed. The onset of superfluidity at different n_3 and n_4 appears as a sheet in Fig. 3. This onset sheet separates the superfluid and insulating phases, the superfluid being stable below the sheet. The curve in the $n_3 = 0$ plane is the phase boundary for pure ⁴He films and it has been explored earlier [6]. For T_c between 0.2 and 0.6 K, this phase boundary is linear. Below 0.2 K, however, there is a significant deviation toward a smaller n_4 from this linear behavior [2,6]. While the intercept based on data with $T_c \leq 0.12$ K of this curve



FIG. 3. The phase diagram in the n_3 , n_4 , T space of ³He-⁴He mixture films adsorbed onto porous gold. The onset sheet separates the superfluid (below) and the insulating (above) region. Note the disparity of scales; the range of n_3 is almost 10 times that of n_4 . The T = 0 phase boundary and the critical curve $\Theta(n_3)$ are shown as thick lines.

with the n_4 axis is $n_0 = 25.25 \pm 0.1 \ \mu \text{mol/m}^2$, the linear region with $T_c > 0.2$ K extrapolates to zero at an apparent inert coverage n_0^* that is larger than n_0 by a few percent of a monolayer.

The top panel in Fig. 4 shows the T = 0 superfluid density, $\rho_{s0} = \rho_s(T = 0)$, as a function of n_3 . ρ_{s0} is deduced by fitting $\rho_s(T)$ measured in the temperature range of $T < 0.6T_c$ to the form $\rho_s(T) = \rho_{s0} - aT^2$ [3]. A quantitative analysis of the result shows that the functional dependence of ρ_{s0} on n_3 is identical for <u>all</u> data sets that have $n_{4s} \ge 0.42$ monolayers. The fits to these data sets (filled symbols) are given by

$$\rho_{s0}(n_3, n_4) = \rho_{s0}(n_3 = 0, n_4) - A[1 - \exp(-n_3/B)],$$
(1)

with $A = 5.1 \pm 0.2 \ \mu \text{mol/m}^2$, $B = 9.5 \pm 0.5 \ \mu \text{mol/m}^2$. Because of the limited range of n_3 , it is not possible to get meaningful fits to vanishing data sets, i.e., for $n_{4s} \le 0.37$ monolayers, shown with open symbols. Nevertheless Eq. (1), with identical parameters A and B, appears to give a good description of these data sets. It is therefore reasonable to conclude that Eq. (1) provides a "universal" description of the depletion of superfluidity with ³He in the T = 0 limit. We note that while curves of Fig. 2 resemble those of Fig. 4, there is no simple exponential dependence of T_c on n_3 that fits all data. This is the case because the solubility of ³He in ⁴He and ⁴He in ³He is highly temperature dependent [10,11].

The inset in the top panel in Fig. 4 shows that the ρ_{s0} curves at different n_3 , obtained by interpolating curves of Fig. 4, increase linearly with n_4 . Previous studies [2–4] have established that for pure ⁴He films this dependence is linear when n_{4s} exceeds a tenth of a layer. ρ_{s0} was also found to be linear with n_4 for mixture films with $n_3 = 12$ layers [11]. The slope of this linear function is



FIG. 4. The top panel shows the evolution with n_3 of ρ_{s0} for mixture films of different ⁴He content. ⁴He coverages are the same as in Fig. 2. Solid lines have the same functional form given by Eq. (1). The inset shows ρ_{s0} as a function of n_4 for ³He coverages, from left to right, of 0, 0.28, 0.75, 1.7, and 8.5 layers. The lower panel depicts the T = 0 phase boundary. The extrapolated data (stars) are well approximated by an exponential function (solid line), particularly for n_3 more than a monolayer.

the same as that for pure ⁴He films [11]. The inset in Fig. 4 shows that the same linear dependence of ρ_{s0} versus n_4 with the same slope is found for all, including submonolayer, ³He coverages. This is a direct consequence of Eq. (1). Therefore, for submonolayer superfluid films, as in thicker films [10–12], there is no evidence of ³He dissolving into ⁴He at T = 0.

The parameter A is the height of the exponential decays in Fig. 4 and its physical interpretation can be obtained by letting $n_3 \rightarrow \infty$. We find

$$A = \rho_{s0}(n_3 = 0) - \rho_{s0}(n_3 \to \infty), \qquad (2)$$

which means that *A* is the superfluid mass that will turn normal upon addition of a large ³He dose to a pure ⁴He film thicker than n_4^{crit} . The depletion of superfluidity, i.e., the increase of the localized ⁴He induced by the ³He, is governed by the exponential function $\exp(-n_3/B)$. *B* is nearly a monolayer of ³He.

The intersection of the superfluid onset sheet with the T = 0 plane separates the superfluid and insulating phases in the n_3 - n_4 plane. Stars in the lower panel in Fig. 4 show this phase boundary as obtained by interpolating the T_c versus n_3 curves in Fig. 2 and then extrapolating the results to T = 0 for all n_3 . For large ³He doses this phase boundary approaches n_4^{crit} asymptotically. To quantitatively obtain the T = 0 phase boundary one should solve either $T_c(n_3, n_4) = 0$ or $\rho_{s0}(n_3, n_4) = 0$. Using Eq. (1), the solution for the latter equation can be obtained only if we know the explicit dependence of $\rho_{s0}(n_3 = 0, n_4)$ on n_4 . Our measurement cannot resolve the deviation of $\rho_{s0}(n_3 = 0, n_4)$ from linearity described above. We adopt the simplest (linear) approximation $\rho_{s0}(n_3 = 0, n_4) \approx n_4 - n_0^*$. With this approximation, Eq. (1) yields

$$n_4 \simeq n_0^* + A[1 - \exp(-n_3/B)].$$
 (3)

The approximation used for ρ_{s0} overestimates the inert coverage. We can correct for this by replacing n_0^* with n_0 in Eq. (3), which is equivalent to a 0.25 μ mol/m² shift of the n_4 scale. Such a shift is within the uncertainty of our measurements. The resulting curve is shown in the lower panel in Fig. 4 as a solid line. The important conclusion is that the T = 0 phase boundary mirrors the exponential decay of ρ_{s0} from Eq. (1). While the shifted Eq. (3) describes well the extrapolated data for large n_3 , at low n_3 there is a significant discrepancy. The phase boundary shows a step near $n_3 = 6 \ \mu$ mol/m². It is not clear if this feature is related to those seen in the damping of third sound [15], magnetization [16], NMR relaxation times [17], and heat capacity [18] of mixture films at higher ⁴He coverages.

We found that the amount of localized ⁴He increases from 25.25 μ mol/m² with no ³He in the film and saturates exponentially at 30.4 μ mol/m² when a very large amount of ³He, up to 8.5 layers, is added to the film. The reason for the saturation is that ³He atoms farther (i.e., beyond the first monolayer) from the ³He-⁴He interface are expected to have a diminishing effect in localizing the superfluid. The difference of 5.15 μ mol/m² is identical within error to the value of A obtained from Eq. (1). This number is within 13% of the previously measured value [11].

What is the microscopic configuration of a ³He-⁴He mixture film in the T = 0 limit? Complete phase separation is found theoretically for an ideal two dimensional mixture film [22]. On a real substrate, however, where vertical displacement is allowed, the different zero point energies of ³He and ⁴He will separate the two isotopes in the van der Waals field perpendicular to the surface. Such a mechanism is indeed responsible for the superfluid sandwich model and it appears to be valid, according to the results summarized in Fig. 4, irrespective to the surface coverages of ³He and ⁴He.

In a recent study [6] of superfluid ⁴He films absorbed onto various substrates, no correlation was found between the inert coverage and the strength of the long range van der Waals tail of the ⁴He-substrate interaction potential. The inert coverages of 6.1, 8.7, 10.3, 19.4, 22.5, and $25.3 \ \mu \text{mol/m}^2$ on, respectively, H₂, HD, D₂, Ne, Ar, and Au substrates, however, were found to scale with the well depths of the potentials. This suggests that the short range ⁴He-substrate forces play the crucial role in determining n_0 . We think the ³He layer that resides at the free surface of the superfluid film can be viewed as a substrate which induces another inert layer that saturates at 5.1 μ mol/m². If we approximate the well depth of the interaction of one ⁴He atom with 8.5 layers of ³He to be 8.7 K, the value calculated for one ⁴He atom and a semi-infinite space of ⁴He [23], then the inert coverage of 5.1 μ mol/m² on ³He follows the trend found for the aforementioned substrates. This suggests that the mechanism of inducing the non-superfluid or inert ⁴He close to the ³He overlayer at T = 0 is similar to that on an amorphous solid substrate. We note that an inert layer that is less than the monolayer coverage, like that on H₂ and ³He, is clearly a phenomenological concept. It is a measure of the localization effect of the substrate on a superfluid film.

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- J. A. Hertz, L. Fleishman, and P. W. Anderson, Phys. Rev. Lett. 43, 942 (1979); M. P. A. Fisher *et al.*, Phys. Rev. B 40, 546 (1989); D. K. K. Lee and J. M. F. Gunn, Phys. Rev. B 46, 301 (1992).
- [2] P. A. Crowell, F. W. Van Keuls, and J. D. Reppy, Phys. Rev. B 55, 12 620 (1997), and references therein.
- [3] D. J. Bishop *et al.*, Phys. Rev. B **24**, 5047 (1981); H. Cho and G. A. Williams, J. Low Temp. Phys. **101**, 463 (1995).
- [4] B.C. Crooker *et al.*, Phys. Rev. Lett. **51**, 666 (1983);
 G. Agnolet, D. F. McQueeney, and J. D. Reppy, Phys. Rev. B **39**, 8934 (1989).
- [5] G. A. Csáthy et al., Phys. Rev. Lett. 80, 4482 (1998).
- [6] G. A. Csáthy and M. H. W. Chan, J. Low Temp. Phys. 121, 451 (2001); (to be published).

- [7] E. N. Smith *et al.*, J. Phys. (Paris), Colloq. **39**, C6-342 (1978); E. Webster, G. Webster, and M. Chester, Phys. Rev. Lett. **42**, 243 (1979); J. P. Laheurte, J. C. Noiray, J. P. Romagnan, and H. Dandache, Phys. Rev. B **22**, 4307 (1980); G. Agnolet, D. F. McQueeney, and J. D. Reppy, in *Proceedings of the 17th International Conference on Low Temperature Physics*, edited by U. Eckern, A. Schmid, W. Weber, and H. Wuhl (North-Holland, Amsterdam, 1984), p. 965.
- [8] D. F. McQueeney, Ph.D. thesis, Cornell University, 1988.
- [9] X. W. Wang and F. M. Gasparini, Phys. Rev. B 34, 4916 (1986).
- [10] J. P. Romagnan et al., Phys. Rev. B 37, 5639 (1988).
- [11] D. McQueeney, G. Agnolet, and J. D. Reppy, Phys. Rev. Lett. 52, 1325 (1984).
- [12] F. M. Ellis et al., Phys. Rev. Lett. 46, 1461 (1981).
- [13] M. J. DiPirro and F. M. Gasparini, Phys. Rev. Lett. 44, 269 (1980).
- [14] B. K. Bhattacharyya and F. M. Gasparini, Phys. Rev. B 31, 2719 (1985).
- [15] F. M. Ellis and R. B. Hallock, Phys. Rev. B 29, 497 (1984).
- [16] R. H. Higley, D. T. Sprague, and R. B. Hallock, Phys. Rev. Lett. 63, 2570 (1989).
- [17] N. Alikacem, D. T. Sprague, and R. B. Hallock, Phys. Rev. Lett. 67, 2501 (1991).
- [18] P.-C. Ho and R. B. Hallock, J. Low Temp. Phys. 121, 501 (2001).
- [19] G.A. Csáthy and M.H.W. Chan, Physica (Amsterdam) 284B-288B, 333 (2000).
- [20] P. A. Crowell et al., Phys. Rev. B 51, 12721 (1995).
- [21] Y. Liu *et al.*, Phys. Rev. Lett. **67**, 2068 (1991); N. Marković *et al.*, Phys. Rev. B **60**, 4320 (1999).
- [22] R. A. Guyer and M. D. Miller, Phys. Rev. B 22, 142 (1980).
- [23] M. Wagner and D. M. Ceperley, J. Low Temp. Phys. 94, 185 (1994).