Structure of Odd-N Superheavy Elements

S. Ćwiok,^{1,2} W. Nazarewicz,^{3,4,5} and P. H. Heenen^{2,6}

¹Institute of Physics, Warsaw University of Technology, ul. Koszykowa 75, PL-00662 Warsaw, Poland

³Department of Physics, University of Tennessee, Knoxville, Tennessee 37996

⁴Physics Division, Oak Ridge National Laboratory, Oak Ridge, Tennessee 37831

⁵Institute of Theoretical Physics, University of Warsaw, ul. Hoża 69, PL-00-681 Warsaw, Poland

⁶Service de Physique Nucléaire Théorique, U.L.B - C.P. 229, B-1050 Brussels, Belgium

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Structure of the odd-*N* superheavy elements with $Z \leq 118$ and $N \leq 175$ is investigated using the self-consistent Skyrme-Hartree-Fock-Bogoliubov method with pairing. Microscopic analysis of one-quasiparticle neutron states, alpha-decay energies, and deformations is performed. Good agreement was obtained with the recently reported α -decay chains of 289 114 and 293 118.

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One of the fundamental and persistent questions of nuclear science concerns the maximum charge and weight that the nucleus may attain. Amazingly, even after thirtyor-so years of our quest for the superheavy elements, the borders of the upper-right end of the nuclear chart are unknown. Theoretically, the mere existence of the heaviest elements with $Z \gtrsim 104$ is entirely due to quantal shell effects. Indeed, for these nuclei the classical nuclear droplet, governed by surface tension and Coulomb repulsion, fissions immediately due to the huge electric charge. At the end of the sixties, a number of theoretical predictions were made that pointed towards the existence of long-lived superheavy (SHE) nuclei [1-5]. However, it is only during recent years that significant progress in the production of the heaviest nuclei has been achieved [6,7]. Notably, three new elements, Z = 110, 111, and 112, were synthesized by means of both cold and hot fusion reactions [8-10]. These heaviest isotopes decay predominantly by groups of α particles (α chains) as expected theoretically [11–13].

All the heaviest elements found recently are believed to be well deformed. Indeed, the measured α -decay energies, along with complementary syntheses of new neutron-rich isotopes of elements Z = 106 and Z = 108, have furnished confirmation of the special stability of the deformed shell at N = 162 predicted by theory [14,15]. Still heavier and more neutron-rich elements are expected to be spherical and even more strongly stabilized by shell effects.

There is no consensus among theorists with regard to the center of the shell stability in the region of spherical SHE. For the neutrons, most calculations predict a magic gap at N = 184. However, because of different treatments of the large Coulomb potential and spin-orbit interaction, various models yield different predictions for the position of the magic proton gap (cf. discussion in Ref. [16]). While Z = 114 was predicted to be the next magic proton number in the macroscopic-microscopic models [11,15], the recent self-consistent models [16,17] favor spherical shell closures at Z = 126, 124, and 120. In this context, the synthesis of the N = 175 isotones reported by the Dubna and Livermore (DL) Collaboration (Z = 114) [18] and the Berkeley and Oregon (BO) team (Z = 118) [19] is both exciting and important. In both cases, the evidence is based on the observed α -decay chains (one in the DL experiment and three in the BO experiment). Because of the large neutron number of the compound nucleus, these α chains do not terminate at some known systems. Consequently, until the charge-mass identification is made, one can use theoretical arguments to support/disprove the experimental assignment. This is the subject of this Letter.

Recently, we have performed a large-scale selfconsistent study of properties of even-even, A-odd, and odd-odd heaviest nuclei. Detailed results of this work will be published elsewhere [20]. Here, we present the small subset of our results relevant to the DL and BO experiments. In our work, we have used the Hartree-Fock-Bogoliubov (HFB) method with a Skyrme interaction in the particle-hole (p-h) channel and a delta force in the pairing channel. The details of calculations closely follow Ref. [16]. The HFB equations have been solved in the coordinate space according to the method of Ref. [21]. The actual mesh is very dense; it consists of 17³ points with a mesh step of 0.8 fm. This has allowed a very accurate determination of binding energies. In the p-hchannel, the Skyrme effective interaction SLy4 has been used. This parametrization [22] has been, in particular, adjusted to reproduce long isotopic sequences; hence one can expect it to have good isospin properties. In order to minimize unphysical fluctuations of the pairing field in regions of low level density, an approximate particle number projection was implemented by means of the Lipkin-Nogami method [23,24]. The pairing strengths of the interaction have then been adjusted to reproduce the average proton and neutron pairing gaps in even-even nuclei around ²⁵⁴Fm. The properties of one-quasiparticle states were calculated by means of the self-consistent blocking. In addition, the odd-T terms that appear in the intrinsic Hamiltonian due to nonzero spin [25] were

²Joint Institute for Heavy Ion Research, Oak Ridge, Tennessee 37831

properly considered. To the best of our knowledge, this work is the first self-consistent study of one-quasiparticle states in SHE that properly treats the polarization due to the odd (unpaired) particle.

Any theoretical model aiming at making predictions in an unknown territory should first be tested in known regions of the chart of the nuclides. Since the main decay mode for the known heaviest elements is α decay, one of the crucial quantities determining lifetimes of the heaviest and superheavy elements is the α -decay energy. The Q_{α} values for the heaviest elements calculated with the HFB-SLy4 model are displayed in Fig. 1. They are compared with known α -decay energies. It is seen that the agreement with experimental data is very good. The only deviation is for the N = 152 isotones of Cm, Cf, and Fm, where calculations overestimate the experimental values of Q_{α} by ~500 keV. (This is due to the N = 152 deformed shell closure which is slightly underestimated by theory.) The good agreement with experimental data that we obtain convinces us that the HFB-SLy4 model can be used to predict properties of unknown SHE.

In order to gain some understanding about the singleparticle shell structure around ²⁸⁴112, we show in Fig. 2 the predicted proton and neutron levels along the α chain of even-even nuclei starting at ²⁶⁰Fm and ending at ²⁹⁶118. The ground-state (g.s.) mass quadrupole deformation β_2 (defined according to Ref. [16]) decreases along this chain from the value of $\beta_2 \approx 0.26$ at the deformed subshell closure at N = 162 to $\beta_2 \approx 0$ when approaching the magic gap N = 184 [16]. Around N = 174, quadrupole deformation takes intermediate values, $\beta_2 \approx 0.13$.



FIG. 1. Q_{α} values for even-even nuclei with $96 \le Z \le 118$ obtained in the HFB-SLy4 model. The experimental Q_{α} values (filled circles) are taken from Refs. [7,26].

As far as the protons are concerned, no significant shell effects are seen between the well-known deformed gap at Z = 108 and that at Z = 118, which appears at higher neutron numbers and smaller deformations. In particular, in the region of particle numbers considered, *no* shell effect at Z = 114 is present. On the other hand, there appears a very interesting shell structure in the neutron system. For the nuclei with Z > 110 the 175th neutron occupies the isolated Nilsson orbital $[707]\frac{15}{2}^-$ originating from the $1j_{15/2}$ shell. This level separates the large single-particle gaps at N = 174 and 176. For nuclei with 164 < N < 174, there are several single-neutron levels around the Fermi level, namely the positive-parity orbitals $[611]\frac{1}{2}^+$, $[606]\frac{11}{2}^+$, $[604]\frac{9}{2}^+$, $[611]\frac{3}{2}^+$, and $[613]\frac{5}{2}^+$, and the



FIG. 2. Proton (top) and neutron (bottom) single-particle levels calculated in the HFB-SLy4 model for the α chain of even-even nuclei: (Z, N) = (100, 160), (102, 162),..., (118, 178). Note that in the self-consistent method each of these nuclei has different ground-state deformations.

negative-parity level $[716]\frac{13}{2}^{-}$. As discussed below, the level bunching below N = 174 and the presence of the high- Ω orbital at N = 175 has significant consequences for the α -decay properties of ²⁸⁹114 and ²⁹³118.

The calculated one-quasiparticle structures and the g.s.to-g.s. values of Q_{α} in the ²⁸⁹114 and ²⁹³118 α -decay chains are shown in Tables I and II, respectively. Using this data, other Q_{α} values for all the low-lying states shown in Tables I and II can easily be obtained. For instance, for the α transition between the excited [611] $\frac{1}{2}^+$ level in ²⁸⁹114 and the [611] $\frac{1}{2}^+$ g.s. of ²⁸⁵112 $Q_{\alpha} =$ 9.64 MeV + 0.52 MeV = 10.16 MeV.

According to our calculations, the g.s-to-g.s. α decays of ²⁸⁹114 and ²⁹³118 are structurally forbidden since the g.s. properties of parent and daughter nuclei differ dramatically. The allowed transitions to the excited $[707]\frac{15}{2}^{-1}$ level in ²⁸⁵112 and ²⁸⁹116 correspond to $Q_{\alpha} = 9.0$ MeV and $Q_{\alpha} = 10.9$ MeV, respectively, and are considerably lower than the experimental values (DL: 9.9 MeV; BO: 12.6 MeV). Consequently, the most probable candidates for the first α transitions are the 10.2 MeV (²⁸⁹114) and 11.8 MeV (²⁹³118) lines associated with the allowed $[611]\frac{1^+}{2^+} \rightarrow [611]\frac{1^+}{2^+}$ decays. By the same token, the $[611]\frac{1^+}{2^+}$ g.s. of N = 173 isotones is expected to decay to the excited $[611]^{\frac{1}{2}^+}$ level in the N = 171 daughters. For ²⁸⁵112, the corresponding Q_{α} energy, 8.76 MeV, is very close to the experimental energy of the second α transition, 8.84 MeV, reported by the DL group. For the ²⁸⁹116 decay we obtain $Q_{\alpha} = 10.14$ MeV, i.e., we underestimate the energy of the second α particle (11.8 MeV) from the BO experiment. This is the worst deviation from experiment obtained in our calculations.

The $[611]_{2}^{1^{+}}$ level is also expected to be the ground state of the N = 169 isotones, and our prediction for the allowed $[611]_{2}^{1^{+}} \rightarrow [611]_{2}^{1^{+}}$ decays is 9.4 MeV for ²⁸¹110 and 10.6 MeV for ²⁸⁵114. Again, we obtain excellent agreement with the DL data (9.0 MeV) and underestimate the energy of the third α transition in the BO experiment (11.5 MeV).

For the α transitions in the ²⁸¹112 $\rightarrow \cdots \rightarrow$ ²⁶⁵104 chain, allowing the positive-parity Nilsson orbitals with very similar quantum numbers, we obtain the following values of Q_{α} : 10.7 MeV (BO: 10.8 MeV), 11.1 MeV (BO: 10.3 MeV), 9.9 MeV (BO: 9.9 MeV), 8.0–8.5 MeV (BO: 8.9 MeV). An alternative route is possible that involves α transitions between $j_{15/2}$ orbitals. Here, for the last two Q_{α} values we obtain 9.8 and 8.7 MeV. In both cases we obtain good agreement with BO data. Table I also shows the predicted structure of ²⁹³116. The g.s. of this transitional nucleus ($\beta_2 \approx 0.09$) is a [604] $\frac{7}{2}^+$ Nilsson level. The Q_{α} value for the allowed α decay to the excited [604] $\frac{7}{2}^+$ state in ¹¹⁴289 is 9.7 MeV.

According to Ref. [18], the maximum cross section for producing evaporation residues in the DL experiment corresponds to the 3n-evaporation channel but the 2n (²⁹⁰114) and 4n (²⁸⁸114) channels cannot be excluded.

The Q_{α} values calculated for the g.s. \rightarrow g.s. α -decay chains of ²⁹⁰114 and ²⁸⁸114 are 9.7, 8.5, 8.9 MeV and 9.4, 9.4, 9.8 MeV, respectively. It is seen, therefore, that the experimental data are also consistent with the α -decay chain of ²⁹⁰114. On the other hand, for the α evaporation channel in the BO experiment [19] (the ²⁹⁰116 α chain) our calculations yield 10.2, 9.6, 10.8, 10.6, 9.9, 9.0, 6.9 MeV; i.e., the agreement is considerably worse compared to the ²⁹³118 scenario.

In summary, we present the first self-consistent analysis of one-quasineutron states in SHE. The calculated equilibrium deformations of one-quasiparticle states in N = 175 isotones are rather small ($\beta_2 \approx 0.11$), and they increase along the α -decay chain. This trend reflects the influence of the N = 184 magic neutron gap for the heaviest systems and the deformed N = 162 gap for the

TABLE I. Predicted structure of the lowest one-quasi-particle excitations in the ²⁹³116 α -decay chain. Each excitation is characterized by the Nilsson quantum numbers of the odd neutron, excitation energy, and quadrupole deformation. The number in parentheses indicates the Q_{α} value for the ground-state to ground-state transition.

Nucleus	Orbital	Energy (MeV)	$oldsymbol{eta}_2$
²⁹³ 116	$[604]\frac{7}{2}^+$	0	0.09
(10.47)	$[602]\frac{5}{2}^+$	0.31	0.09
	$[611]\frac{1}{2}^+$	0.52	0.08
	$[707] \frac{15}{2}^{-}$	0.93	0.09
²⁸⁹ 114	$[707] \frac{15}{2}^{-}$	0	0.12
(9.64)	$[611]\frac{1}{2}^+$	0.52	0.11
	$[604]\frac{7}{2}^+$	0.79	0.13
	$[602]\frac{5}{2}^+$	1.17	0.12
²⁸⁵ 112	$[611]\frac{1}{2}^+$	0	0.14
(8.88)	$[611]\frac{3}{2}^+$	0.60	0.14
	$[707] \frac{15}{2}^{-}$	0.62	0.13
	$[606]\frac{11}{2}^+$	0.65	0.15
	$[604]\frac{9}{2}^+$	0.72	0.15
²⁸¹ 110	$[604]\frac{9}{2}^+$	0	0.19
(9.32)	$[606]\frac{11}{2}^+$	0.07	0.19
	$[611]\frac{1}{2}^+$	0.12	0.18
	$[611]\frac{3}{2}^+$	0.59	0.17
	$[613]\frac{5}{2}^+$	0.65	0.17
	$[716] \frac{13}{2}^{-}$	0.94	0.17
277 108	$[611]\frac{1}{2}^+$	0	0.21
	$[604]\frac{9}{2}^+$	0.04	0.20
	$[613]\frac{5}{2}^+$	0.31	0.21
	$[716] \frac{13}{2}^{-}$	0.36	0.21
	$[611]\frac{3}{2}^+$	0.38	0.20

TABLE II. Same as in Table I, except for the $^{293}118~\alpha$ -decay chain.

Nucleus	Orbital	Energy (MeV)	$oldsymbol{eta}_2$
²⁹³ 118	$[707] \frac{15}{2}^{-}$	0	0.11
(11.59)	$[611]\frac{1}{2}^+$	0.16	0.10
	$[602]\frac{5}{2}^+$	0.84	0.12
	$[604]\frac{7}{2}^+$	0.98	0.10
²⁸⁹ 116	$[611]\frac{1}{2}^+$	0	0.13
(10.18)	$[606] \frac{11}{2}^+$	0.62	0.14
	$[611]\frac{3}{2}^+$	0.66	0.13
	$[604]\frac{9}{2}^+$	0.69	0.14
	$[707] \frac{15}{2}^{-}$	0.72	0.13
²⁸⁵ 114	$[606] \frac{11}{2}^+$	0	0.16
(10.60)	$[611]\frac{1}{2}^+$	0.04	0.16
	$[611]\frac{3}{2}^+$	0.15	0.16
	$[604]\frac{9}{2}^+$	0.16	0.16
	$[613]\frac{5}{2}^+$	0.28	0.15
	$[716] \frac{13}{2}^{-}$	0.73	0.16
²⁸¹ 112	$[611]\frac{1}{2}^+$	0	0.19
(10.85)	$[604]\frac{9}{2}^+$	0.07	0.19
	$[611] \frac{3}{2}^+$	0.37	0.19
	$[613]\frac{5}{2}^+$	0.41	0.19
	$[606] \frac{11}{2}^+$	0.42	0.19
	$[716] \frac{13}{2}^{-}$	0.51	0.19
²⁷⁷ 110	$[716] \frac{13}{2}^{-}$	0	0.22
(11.07)	$[613]\frac{5}{2}^+$	0.02	0.21
	$[611]\frac{3}{2}^+$	0.07	0.21
	$[611]\frac{1}{2}^+$	0.15	0.22
	$[604]\frac{9}{2}^+$	0.27	0.21
	$[606] \frac{11}{2}^+$	0.83	0.21
²⁷³ 108	$[611]\frac{3}{2}^+$	0	0.23
(9.93)	$[613]\frac{5}{2}^+$	0.04	0.23
	$[716] \frac{13}{2}^{-}$	0.05	0.23
	$[604]\frac{9}{2}^+$	0.54	0.22
	$[611]\frac{1}{2}^+$	0.55	0.23
²⁶⁹ 106	$[611]\frac{3}{2}^+$	0	0.24
(8.45)	$[613]\frac{5}{2}^+$	0.15	0.24
	$[716] \frac{13}{2}^{-}$	0.22	0.22
	$[604]\frac{9}{2}^+$	0.60	0.24
²⁶⁵ 104	$[725]\frac{11}{2}^{-}$	0	0.25
	$[604]\frac{9}{2}^+$	0.09	0.25
	$[613]\frac{7}{2}^+$	0.42	0.25
	$[622]\frac{3}{2}^+$	0.79	0.25

lightest nuclei in the chain. As far as Q_{α} values, our calculations are consistent with the recent experimental findings of Refs. [18] and [19]. In particular, the DL data

can be interpreted in terms of the α chains of ²⁸⁹114 or ²⁹⁰114. We also obtain good agreement with the BO data assuming the ²⁹³118 α chain. The ground state of the N = 175 isotones is calculated to be a high- Ω *isomeric* state [707] $\frac{15}{2}^{-}$. Because of structural arguments, this state is probably bypassed. (The same is true for the high- Ω ground state of N = 171.) This result may explain the observed lower cross section for the production of ²⁹³118 (~2 pb) as compared to calculations by Smolańczuk [13] (~670 pb).

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