Measurement of the Energy Spectrum of ν_e from Muon Decay and Implications for the Lorentz Structure of the Weak Interaction

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The KARMEN experiment uses the reaction ${}^{12}C(\nu_e, e^{-}){}^{12}N_{g.s.}$ to measure the energy distribution of ν_e emitted in muon decay at rest $\mu^+ \rightarrow e^+ + \nu_e + \overline{\nu}_{\mu}$. The ν_e analog ω_L of the famous Michel parameter ρ has been derived from a maximum-likelihood analysis of events near the kinematic end point, $E_{\text{max}} = 52.8$ MeV. The result, $\omega_L = (2.7^{+3.8}_{-3.3} \pm 3.1) \times 10^{-2}$, is in good agreement with the standard model prediction $\omega_L = 0$. We deduce a 90% confidence upper limit of $\omega_L \leq 0.113$, which corresponds to a limit of $|g_{RL}^S + 2g_{RL}^T| \leq 0.78$ on the interference term between scalar and tensor coupling constants. [S0031-9007(98)06631-9]

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Experimental results from nuclear β decay and muon decay form the basis of the V - A hypothesis, which is an essential feature of the standard model (SM) of electroweak interactions. The rate of muon decay, the purely leptonic process $\mu^+ \rightarrow e^+ + \nu_e + \overline{\nu}_{\mu}$, has been used to determine the universal Fermi coupling constant G_F . Precise measurements of the shape of the e^+ energy spectrum, the decay asymmetry between the μ^+ spin and e^+ momentum, and the polarization vector of the e^+ have led to bounds on the scalar, vector, and tensor coupling constants, which form the Lorentz structure of the charged weak interaction. These results combined with the inverse process $\nu_{\mu} + e^- \rightarrow$ $\mu^- + \nu_e$ underpin the SM assumption of lepton number conservation, the V - A interaction and universality [1]. All experiments up to now support the V - A structure of the weak interaction; however, substantial non-(V - A)components are not ruled out.

Complementary to these experiments, which are all based on observation of the charged leptons only, the Karlsruhe Rutherford Medium Energy Neutrino experiment (KARMEN) determines the energy spectrum of the ν_e emitted in the decay $\mu^+ \rightarrow e^+ + \nu_e + \overline{\nu}_{\mu}$ of unpolarized muons to draw conclusions on the Lorentz structure. In the well-known case of e^+ spectroscopy, it is the Michel parameter ρ which governs the shape of the e^+ energy spectrum. In an analogous way, the shape of the ν_e energy spectrum is determined by the parameter ω_L , which also depends on vector, scalar, and tensor components of the weak interaction, but in a different combination. In the SM all non-(V - A) components vanish, and ω_L is predicted to be 0. Thus an upper limit on ω_L derived from the analysis of the ν_e energy spectrum provides new limits on nonstandard couplings.

All features of muon decay are most generally described by a local, derivative-free, lepton-numberconserving, four-lepton point interaction with the matrix element given by [2]

$$M = \frac{4}{\sqrt{2}} G_F \sum_{\substack{\gamma = S, V, T\\ e, \mu = R, L}} g_{\epsilon\mu}^{\gamma} \langle \overline{e}_{\epsilon} | \Gamma^{\gamma} | (\nu_e)_n \rangle \langle (\overline{\nu}_{\mu})_m | \Gamma_{\gamma} | \mu_{\mu} \rangle.$$
(1)

The index γ labels the type of interaction $\Gamma(S = 4\text{-scalar}, V = 4\text{-vector}, T = 4\text{-tensor})$ and the indices ϵ and μ indicate the chirality (L = left-handed, R = right-handed) of electron and muon spinors, respectively. In this representation the chirality of the neutrino n or m is fixed to be equal to that of the associated charged lepton for the V interaction, but opposite for the S and T interactions. As G_F sets the absolute strength of the interaction, the ten coupling constants $g_{\epsilon\mu}^{\gamma}$ are dimensionless complex quantities normalized by

$$3|g_{RL}^{T}|^{2} + 3|g_{LR}^{T}|^{2} + \sum_{\epsilon,\mu=R,L} \left(\frac{1}{4}|g_{\epsilon\mu}^{S}|^{2} + |g_{\epsilon\mu}^{V}|^{2}\right) = 1$$
(2)

with $g_{RR}^T = g_{LL}^T \equiv 0$. In the SM, muon decay is a pure *V* interaction mediated between left-handed particles, so all coupling constants vanish except $g_{LL}^V \equiv 1$. Although this representation is elegant from the theoretical point of view, the individual coupling constants cannot be determined directly by experiment. However, the measurable parameters (ρ , η , ω_L , etc.) are expressible as positive semidefinite bilinear combinations of $g_{\epsilon\mu}^{\gamma}$ from which

upper or lower limits for the coupling constants can be derived.

The possibility of measuring ω_L with the KARMEN experiment was first pointed out by Fetscher [3]. More recently Greub *et al.* [4] have calculated the spectrum of left-handed ν_e including radiative corrections and effects of finite lepton masses. Taking significant terms only, the spectrum dN_L/dx can be described by

$$\frac{dN_L}{dx} = \frac{G_F^2 m_\mu^5}{16\pi^3} Q_L^{\nu} \{G_0(x) + G_1(x) + \omega_L G_2(x)\}, \quad (3)$$

where m_{μ} is the muon mass, $x = 2E_{\nu}/m_{\mu}$ is the reduced neutrino energy, and Q_L^{ν} denotes the probability of emission of a left-handed ν_e . The function $G_0(x)$ describes the pure V - A interaction, $G_1(x)$ takes into account radiative corrections, and $\omega_L G_2(x)$ includes the effect of scalar and tensor components according to

$$\omega_{L} = \frac{3}{4} \frac{|g_{RR}^{S}|^{2} + 4|g_{LR}^{V}|^{2} + |g_{RL}^{S} + 2g_{RL}^{T}|^{2}}{|g_{RL}^{S}|^{2} + |g_{RR}^{S}|^{2} + 4|g_{LL}^{V}|^{2} + 4|g_{LR}^{V}|^{2} + 12|g_{RL}^{T}|^{2}}.$$
(4)

The calculated ν_e energy spectra for different values of ω_L are shown in Fig. 1(a). Momentum conservation in the decay fixes the emission direction of ν_e near the kinematic end point to be opposite to that of the positron and the $\overline{\nu}_{\mu}$. Together with angular momentum conservation this implies suppression of emission of lefthanded ν_e in the case of vector coupling, while all other couplings enhance the decay rate at the end point. The total decay rate, and therefore the integral neutrino flux, is unchanged by nonstandard interactions.

The KARMEN experiment uses the pulsed spallation neutron facility ISIS at the Rutherford Appleton Laboratory to investigate neutrinos from μ^+ decay. The 800 MeV proton beam from ISIS is stopped in a Ta-D₂O target producing neutrons and pions. All charged pions are stopped inside the target within 10⁻¹⁰ s, the π^- being

absorbed by the heavy target material while the π^+/μ^+ decay chain $\pi^+ \rightarrow \mu^+ + \nu_{\mu}$, $\mu^+ \rightarrow e^+ + \nu_e + \overline{\nu}_{\mu}$ produces an intense burst of ν_{μ} , ν_e , and $\overline{\nu}_{\mu}$, emitted isotropically with equal intensity. Since both π^+ and μ^+ decay at rest, the energy spectra of the neutrinos are well defined. The π^+ decay produces monoenergetic ν_{μ} with $E_{\nu_{\mu}} = 29.8$ MeV; the ν_e and $\overline{\nu}_{\mu}$ from the μ^+ decay have continuous energy distributions up to $E_{\text{max}} = 52.8 \text{ MeV}.$ The time structure of ISIS-two 100 ns wide proton bunches 324 ns apart and recurring at 50 Hz-determines the production time of the different ν flavors: the short π^+ lifetime ($\tau_{\pi} = 26$ ns) leads to two ν_{μ} pulses within the first 500 ns after beam-on target. These pulses are well separated in time from the production of ν_e and $\overline{\nu}_{\mu}$, which follow with the much longer lifetime of the μ^+ $(\tau_{\mu^+} = 2.2 \ \mu s)$. This leads to a suppression factor of about 10⁴ for cosmic-ray background.

The neutrinos are detected in a segmented 56 ton liquid scintillation calorimeter consisting of 512 optical modules, each with a length of 3.53 m and a cross section of $18 \times 18 \text{ cm}^2$ [5]. The detector is an almost completely (96%) active calorimeter optimized for the measurement of electrons around 30 MeV and achieves resolutions of $\sigma(E)/E = 11.5\%/\sqrt{E(\text{MeV})}$ for energy, and $\sigma(X) \approx 7$ cm for position measurement. A 7000 ton shielding steel blockhouse together with two layers of active veto counters suppresses beam-correlated spallation neutrons and cosmic-ray muons.

The signature that unambiguously identifies a ν_e is a delayed coincidence consisting of an electron from the charged current reaction ${}^{12}C(\nu_e, e^-){}^{12}N_{g.s.}$ in the time window of ν_e production followed by a positron from the subsequent β decay of ${}^{12}N_{g.s.}$ ($\tau = 15.9$ ms) at the same location in the detector. Each event fully contained within the central detector with time 0.6–9.6 μ s after beam-on target and energy 10–36 MeV is identified as electron, provided it is followed by a positron event within 0.5–36 ms with energy 3.5–16.5 MeV. We demand the sequence to be detected in the same or adjacent module



FIG. 1. Influence of different values of $\omega_L = 0.0, 0.15$, and 0.3 on (a) the ν_e energy spectrum in μ^+ decay and on (b) the visible electron energy spectrum measured with the reaction ${}^{12}C(\nu_e, e^-){}^{12}N_{g.s.}$ (c) Experimental electron energy distribution together with MC expectation (solid line) and the subtracted background (shaded).

within a distance $\Delta X \leq 35$ cm along the module axis. Cuts used to reduce cosmic background are the same as used in previous data evaluations [6]. In data accumulated between June 1990 and December 1995—corresponding to 9122 C of protons or $2.51 \times 10^{21} \mu^+$ decays in the ISIS target—we find 513 e^-/e^+ sequences. Subtracting 13.3 ± 0.8 background events and taking into account an overall detection efficiency $\epsilon = 32.8\%$, the flux-averaged cross section is

$$\langle \sigma \rangle_{\rm exp} = [9.4 \pm 0.4({\rm stat}) \pm 0.8({\rm syst})] \times 10^{-42} {\rm ~cm}^2.$$
(5)

This is in good agreement with different theoretical calculations of $\langle \sigma \rangle_{\rm th}$ in the range of (9.1–9.4) $\times 10^{-42}$ cm² [7,8].

As the recoil energy of the ¹²N nucleus is negligible, the ν_e energy E_{ν} is determined from the measurement of the electron energy E_e via the kinematic relation $E_{\nu} = E_e + Q$, where Q = 17.3 MeV is the Q value of the detection reaction. The energy dependence of the cross section is dominated by the phase-space factor $(E_{\nu} - Q)^2$. Therefore, a low rate of additional ν_e at the kinematic end point $E_{\text{max}} = 52.8$ MeV due to nonstandard couplings is translated to the observation of a significantly higher rate of electrons and thus to a distortion of the visible energy spectrum of Fig. 1(b).

The KARMEN calorimeter allows a precise measurement of the energy E_e [see Fig. 1(c)]. The energy spectrum of ν_e from μ^+ decay is then determined in two steps. First, we apply the procedure of regularized unfolding described by Blobel [9] to derive the true electron energy. This method takes into account the detector response and minimizes inherent instabilities (oscillating solutions) by demanding a priori a certain degree of smoothness of the true electron distribution depending on statistical accuracy. The ν_e energy distribution is then calculated from the number of primary electrons, within a given interval ΔE from the unfolding procedure, divided by the corresponding mean cross section. This yields a ν_e energy spectrum with seven data points as shown in Fig. 2 and compared with the V - A expectation. This represents the first measurement of the neutrino energy spectrum from muon decay in addition to the well-known e^+ spectrum.

Because of the strong energy dependence of the detection cross section, the most detailed information on ω_L and Q_L^{ν} is obtained from the experimental electron spectrum of Fig. 1(c). The analysis is done by two independent methods: (1) the investigation of the measured decay rate on the basis of the flux-averaged cross section, and (2) the analysis of the spectral shape with a maximumlikelihood (ML) method.

As can be seen from Fig. 1(b), $\omega_L > 0$ would result in additional ${}^{12}C(\nu_e, e^-){}^{12}N_{g.s.}$ events; on the other hand, $Q_L^{\nu} < 1$ would reduce the number of events [see Eq. (3)]. In order to find allowed regions in the $Q_L^{\nu} - \omega_L$ parameter space, we compared measured and expected flux averaged cross sections. As theoretical cross section



FIG. 2. Energy spectrum of ν_e from μ^+ decay determined by an unfolding method compared with the standard model expectation (solid line).

 $\langle \sigma \rangle_{\rm th}$ with a realistic estimate of the systematic error we use $\langle \sigma \rangle_{\rm th} = (9.2 \pm 0.5) \times 10^{-42} \, {\rm cm}^2$. The experimental cross section is taken from Eq. (5) with statistical and systematic error added quadratically. The probability distribution of the ratio

$$R(Q_L^{\nu},\omega_L) = \frac{\langle \sigma \rangle_{\exp}}{\langle \sigma \rangle_{\text{th}}} = Q_L^{\nu}(1+S\omega_L) = \frac{9.4 \pm 0.9}{9.2 \pm 0.5}$$
(6)

incorporates a flux decrease by right-handed ν_e through Q_L^{ν} as well as an increase by nonzero ω_L values; S is the ratio of additional events in case of $\omega_L = 1$ relative to the expectation for $\omega_L = 0$. We have sampled the probability density function of the ratio R from Gaussian distributions of $\langle \sigma \rangle_{\rm exp}$ and $\langle \sigma \rangle_{\rm th}$ for 3 different energy ranges: (a) the range 10-36 MeV with the highest statistical accuracy, but only moderate sensitivity S = 0.81, (b) the range 28-36 MeV, where with S = 3.48 we are very sensitive to ω_L , and (c) the range 10–22.5 MeV, where the expected event number is almost independent of ω_L [S = 0.002; see Fig. 1(b)]. From range (c) we deduce a lower limit $Q_L^{\nu} \ge 0.796$. The shaded parameter space shown in Fig. 3 combines regions excluded at 90% confidence level of all three energy ranges. From inverse muon decay experiments it is known that $Q_L^{\nu} > 0.92$ [10,11]. Including this information in our analysis of range (b) restricts the allowed area and sets a 90% confidence upper limit $\omega_L \leq 0.12$.

In the second method we determine ω_L by analyzing the shape of the visible electron spectrum independent of Q_L^{ν} . In order to increase the energy resolution and to reduce the background level we applied more stringent cuts on the electron position along the module axis $|X| \leq 150$ cm and on the electron time 0.6–7.2 μ s. These cuts reduce the background to only 6.0 events in a sample of 441 events, thus nearly doubling the signal-to-background ratio.



FIG. 3. The $Q_L^{\nu} - \omega_L$ parameter space: The shaded regions are excluded at 90% confidence from the different analyses of the absolute flux in several energy ranges. The horizontal line is the result of the spectral shape analysis $\omega_L \le 0.113$ at 90% confidence. The vertical line is the current best limit $Q_L^{\nu} \ge 0.92$.

The theoretical ν_e energy spectrum of Eq. (3) was converted into a visible electron spectrum using the energy-dependent $\sigma(E_{\nu})$ taken from [8] folded with the detector response by a Monte Carlo (MC) calculation. The ML procedure was carried out on an event-by-event basis for several fit intervals all of which gave results compatible with $\omega_L = 0$ within a 1 σ error. The net result is

$$\omega_L = [2.7^{+3.8}_{-3.3}(\text{stat}) \pm 3.1(\text{syst}) \times 10^{-2}.$$
 (7)

Including the systematic error (energy shift of 0.25 MeV or 0.7% scaling error) we find, with the most conservative Bayesian approach, a 90% confidence upper limit $\omega_L \leq$ 0.113. This excludes the region above the horizontal line in Fig. 3. Combining Eqs. (2) and (4) the following relation between the shape parameter ω_L and nonstandard couplings is [10,12]

$$|g_{RL}^{S} + 2g_{RL}^{T}| \le \sqrt{\frac{16}{3}\omega_{L}}.$$
 (8)

The limit on ω_L thus results in an upper limit of $|g_{RL}^S + 2g_{RL}^T| \le 0.78$ for the interference term of scalar and tensor amplitudes.

In conclusion, the KARMEN experiment finds no evidence for nonstandard coupling constants in μ^+ decay at rest, either by a determination of the absolute ν_e flux or by analysis of the spectral shape. Our analysis excludes most of the Q_L^{ν} - ω_L parameter space and yields for the first time an upper limit on the neutrino Michel parameter ω_L .

During 1996 the experiment was upgraded by an additional active veto counter in order to increase the sensitivity of the search for neutrino oscillations in the channel $\overline{\nu}_{\mu} \rightarrow \overline{\nu}_{e}$ [13]. Since 1997 KARMEN has been taking data again. Up to the end of 1999 we expect about 400 further charged current events, which will reduce the statistical error by about a factor of 1.4. Considering also a reduction of the systematic error, this may result in a limit competitive with the present best limit $|g_{RL}^{S} + 2g_{RL}^{T}| \leq 0.45$ deduced from measurements of the positron polarization [10].

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