

Resonant Magnetic Field Control of Elastic Scattering in Cold ^{85}Rb

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A magnetic field-dependent Feshbach resonance has been observed in the elastic scattering collision rate between atoms in the $F = 2, M = -2$ state of ^{85}Rb . Changing the magnetic field by several Gauss caused the collision rate to vary by a factor of 10^4 , and the sign of the scattering length could be reversed. The resonance peak is at 155.2(4) G and its width is 11.6(5) G. From these results we extract much improved values for the three quantities that characterize the interaction potential: The van der Waals coefficient C_6 , the singlet scattering length a_S , and the triplet scattering length a_T . [S0031-9007(98)07904-6]

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Very-low-temperature collision phenomena can be quite sensitive to applied electromagnetic fields. Several groups have altered inelastic collision rates in optical traps by applying laser fields [1]. There have also been numerous proposals [2–6] for using laser and static electric or magnetic fields to influence the s -wave scattering length (a), and equivalently, the elastic collision cross sections ($\sigma = 8\pi a^2$) between cold atoms. Particularly notable is the prediction by Verhaar and co-workers [2] that as a function of magnetic field there should be Feshbach resonances in collisions between cold ($\sim \mu\text{K}$) alkali atoms. These resonances were predicted to cause dramatic changes in the cross section and to even allow the sign of the scattering length to be changed. Such resonances are very interesting collision physics, but they also offer a means to manipulate Bose-Einstein condensates (BEC), the properties of which are primarily determined by the scattering length. Some time ago we searched for such resonances in ^{133}Cs and ^{87}Rb , without success [7,8]. This was not surprising because there was enormous uncertainty as to the positions and widths of the predicted resonances. However, recent photoassociation spectroscopy has greatly improved the knowledge of the alkali interatomic interaction potentials, allowing these resonance parameters to be predicted with much less uncertainty [9]. As we show below, the fact that the widths and positions of these resonances are so sensitive to the interatomic potentials makes their measurement a good way to determine these potentials.

The improved predictions for the resonance positions facilitated their observation. In the past few months Feshbach resonances have been observed in both ^{23}Na [10] and ^{85}Rb [11]. In the sodium work, two magnetic field-dependent processes were observed: (1) A change in the expansion rate of BEC due to a change in the scattering length, and (2) an enormous, and as yet unexplained, increase in the loss rate. This loss rate precluded the study of collisions in the interesting field regime near the center of the resonance where the scattering length changes sign. In the rubidium work

[11] the resonance was detected as a magnetic field dependence of the photoassociation spectra. The resulting resonance width was measured to be far larger than originally predicted by theory [2].

Here we report the study of a Feshbach resonance in the elastic collision cross section between atoms in the $F = 2, M = -2$ state of ^{85}Rb . By changing the dc magnetic field we are able to change the collision rate by 4 orders of magnitude and explore regions of positive, negative, and essentially zero scattering length. We determine the width and position of the resonance about a factor of 10 more precisely than in Ref. [11], and from these data we improve the accuracy of the Rb interaction potential substantially. In contrast to Ref. [10], we do not observe two or three body loss at resonance because we work at much lower densities (10^9 atoms/cm 3).

We measured the collision rates using the technique of “cross-dimensional mixing,” as in our earlier work [7]. In this technique, a nonisotropic distribution of energy is created in a magnetically trapped cloud of atoms, and the time for the cloud to reequilibrate by elastic collisions is measured. The apparatus used is identical to that used in our previous work on BEC in ^{87}Rb [12]. It is a double magneto-optic trap (MOT) system in which multiple samples of atoms are trapped in a relatively high-pressure chamber and then transferred to a second MOT in a low-pressure chamber. The second MOT is then turned off and a “baseball coil” magnetic trap is turned on around them. The atoms are then cooled by forced evaporation.

After the atoms are evaporatively cooled to the desired temperature, the energy in the radial direction (and correspondingly the square of the width of the trapped cloud) is reduced to about 0.6 that of the axial direction. This is done by decreasing the frequency of the rf “knife” used for the evaporative cooling more rapidly than the cloud can equilibrate [13]. The bias magnetic field is then adiabatically ramped to the selected value, and the cloud is allowed to equilibrate for a fixed time. The shape of the cloud is then measured by absorptive imaging. This is repeated for different equilibration times, and the aspect

ratios of the cloud vs time are fit to an exponential to determine the equilibration time constant τ_{eq} . Although the energy dependence of the cross section can lead to small deviations from exponential equilibration, the relative changes in τ_{eq} with B are large enough to dominate any error in curve fitting. As discussed in Ref. [14], τ_{eq} is proportional to the inverse of the elastic collision rate [15]. From the measured trap oscillation frequencies and the measured shape and optical depth, we determine the temperature T and average density $\langle n \rangle$ of the cloud. The magnetic field value at the center of the cloud is determined to ± 0.1 G by finding the rf frequency that resonantly drives spin flip transitions for the atoms at that position and using the Breit-Rabi equation. The spread in the magnetic field across the cloud scales as $T^{1/2}$ and is 0.6 G FWHM for a 0.5 μK cloud.

The measured equilibration times depend on B and T and vary from 0.15 s to nearly 2000 s. To display data taken at several values of both density and temperature, we convert it to the normalized equilibration rate, $\Gamma_{\text{norm}} = 1/\langle n v \rangle \tau_{\text{eq}}$. Figure 1 displays the data versus mean magnetic field. Γ_{norm} has units of cross section and in fact is proportional to an empirical average over the field- and energy-dependent elastic scattering cross section $\sigma(B, E)$. To calculate this average is problematic because the magnetic field and energy are coupled—a detailed Monte Carlo simulation of particle trajectories [16] is probably required. However, in the low-temperature limit the spread in magnetic field across the cloud is small so the maximum and minimum values of Γ_{norm} will occur at the same magnetic field as the peak and zero of $\sigma(B, E = 0)$. For a rough qualitative comparison, in Fig. 1 we have also plotted the theoretically predicted values of $\sigma(B, E)$ for two temperatures. These predictions are calculated using the atomic potentials determined from the observed maximum and minimum values of the relaxation rate, as discussed below. Both the cross section curves and the equilibration data show the same qualitative features. For temperatures below a few μK , there is a slightly asymmetric peak near 155 G, and the width and height of this peak are strong functions of temperature. At 167 G there is a profound drop in the rate. This dip is quite asymmetric, but the shape is relatively insensitive to T , and at the bottom (field value B_{min}) the rate is essentially zero. The field value of the peak (B_{peak}) is customarily defined to be the position of the Feshbach resonance, and the resonance width Δ is the distance between B_{peak} and B_{min} [11]. The scattering length is positive for field values between B_{peak} and B_{min} and is negative for field values below B_{peak} or above B_{min} . This dependence of the sign is expected from previous theory. The observation that B_{peak} is at a lower field than B_{min} provides experimental confirmation that the scattering length is negative away from the resonance.

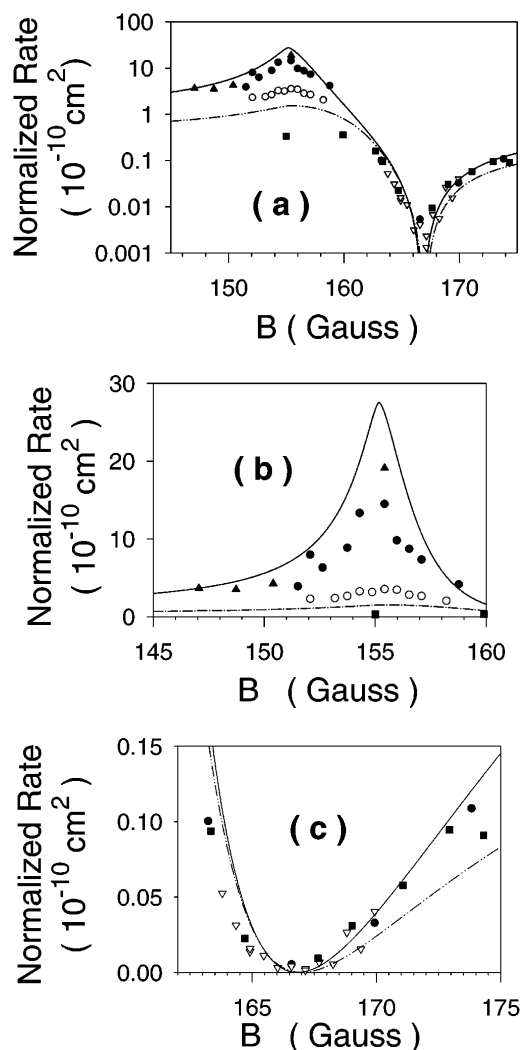


FIG. 1. (a) The data points show the measured equilibration rate divided by average density and velocity ($\Gamma_{\text{norm}} = 1/\langle n v \rangle \tau_{\text{eq}}$) vs magnetic field B . Data are shown for five different temperatures: 0.3 μK (\blacktriangle), 0.5 μK (\bullet), 1.0 μK (\circ), 3.5 μK (∇), and 9.0 μK (\blacksquare). The lines are calculations for the thermally averaged elastic cross sections (not equilibration rates) and hence are not expected to fit the data points. The solid line corresponds to 0.5 μK and the dot-dashed line to 9.0 μK . (b),(c) Expanded views of the regions of maximum and minimum cross section.

Because the functional form of the normalized equilibration rate is not known, the desired quantities B_{peak} and B_{min} cannot be found by a detailed fit to all of the data. We determine them by fitting a simple smooth curve to only the few highest (or lowest) points at each temperature below 5 μK , and assigning correspondingly conservative error bars that more than span the values determined for all temperatures [17]. We find $B_{\text{peak}} = 155.2(4)$ G and $B_{\text{min}} = 166.8(3)$ G, giving a resonance width $\Delta = 11.6(5)$ G. The values for B_{peak} and Δ are reasonably consistent with the less precise values of 164(7) G and 8.4(3.7) G measured in Ref. [11]. In our experiment,

the accuracy of the peak position is much higher primarily because of better field calibration. The better accuracy in the width is largely because the photoassociation measurement is a somewhat less direct way to observe the resonance, and so substantial and somewhat uncertain corrections are required to go from observed signal widths to the actual resonance width.

We can now use these measured quantities to determine the singlet and triplet Born-Oppenheimer potentials. The accuracy of predicted cold collision observables hinges on the quality of these potentials used in the radial Schrödinger equation for the nuclear motion. Because the scattering is purely s wave at these temperatures, accurate determination of these potentials in turn relies primarily on three parameters, the long range van der Waals coefficient C_6 and the zero-energy singlet and triplet phases. The phases are usually tabulated in terms of singlet a_S and triplet a_T scattering lengths calculated with hyperfine terms omitted from the radial equation.

The old “nominal” values of the scattering lengths and C_6 (optimized to achieve agreement with previous measurements) predicted the position of this Feshbach resonance reasonably well, but, as in Ref. [11] the predicted width was much smaller than observed. This width reflects the coupling of bound and continuum channels and is primarily controlled by the difference between the singlet and triplet scattering lengths. It is inconvenient to tabulate results as a function of singlet and triplet scattering lengths for ^{85}Rb , because both are very close to lying on top of a divergent pole. Accordingly, we present the results of our analysis in terms of a bound state phase (fractional bound state number) ν_D as in Ref. [11]. We define the corresponding singlet ν_{DS} or triplet ν_{DT} bound state phase in terms of the scattering length through the relation [18]: $\nu_D = (1/\pi) \cot^{-1}(a/a_{\text{ref}} - 1)$, where $a_{\text{ref}} = \Gamma(5/4)/[\sqrt{8} \Gamma(3/4)] (2mC_6/\hbar^2)^{1/4}$ [19] and m is the reduced mass of the atomic pair. The bound state phase is related to the short range quantum defect μ^{sr} presented in Ref. [20] by $\nu_D \cong \mu^{\text{sr}} - 1/4$. The quantum defect method developed in Ref. [20] was used to generate the theory curves in Fig. 1.

We have adjusted the singlet and triplet Rb-Rb potential curves at small distances until the theoretical calculation agrees simultaneously with the present measurements of B_{peak} and Δ , to within the experimental uncertainties. For a given C_6 , this severely constrains the values of both ν_{DS} and ν_{DT} (see Fig. 2). However, the position and width are not sufficient to totally constrain all three parameters C_6 , ν_{DS} , and ν_{DT} . In order to obtain values for all of the parameters, we also require that our triplet potentials predict the g -wave shape resonance in an energy range consistent with the measured value given in Ref. [21].

The results of this analysis are presented in Fig. 2. This illustrates that accurate measurements of Feshbach resonances are an extremely precise method for determining threshold properties of the interatomic potentials.

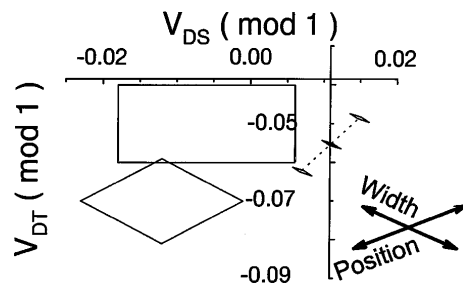


FIG. 2. Comparison of the allowed ν_{DS} , ν_{DT} parameter space based on this work and previous measurements for a constant value of C_6 . The rectangle is the allowed range from Ref. [24]. The large diamond is the range from Ref. [11]. The small filled diamond is the range from this work. Note that the rectangle and large diamond constraints were established using a different value of C_6 from the one used in this work (4550 instead of 4700 a.u.). As is discussed in the text, the position of the constrained regions depends linearly upon C_6 , which is indicated by the small connected diamonds on either side of the small filled diamond. These diamonds show the effect of 50 a.u. uncertainty in C_6 . The arrows indicate sensitivity to changes in resonance width and position in this parameter space.

Our measurement has allowed us to reduce the combined C_6 , ν_{DS} , and ν_{DT} parameter space by roughly a factor of 80. As discussed in Ref. [11], the position of the resonance depends mostly on C_6 and the sum $\nu_{DS} + \nu_{DT}$. The width depends mostly on the difference $\nu_{DS} - \nu_{DT}$. As Fig. 2 shows, the allowed ν_{DS} , ν_{DT} parameter region is an extremely correlated function of C_6 . In particular, we find that the area of the allowed parameter region is independent of both C_6 and C_8 . The uncertainty in these dispersion coefficients simply shifts the position of the “diamond.” Accordingly, we can represent the allowed parameter region in the following manner: $\nu_{DS} + \nu_{DT} = -0.0442 + 2.14(10^{-4})(C_6 - \bar{C}_6) - 4(10^{-8})(C_8 - \bar{C}_8) \pm 0.001$ and $\nu_{DS} - \nu_{DT} = 0.0652 - 6.6(10^{-5})(C_6 - \bar{C}_6) \pm 0.003$, where $\bar{C}_6 = 4700$ and $\bar{C}_8 = 550\,600$ [22]. Converting the bound state phases into scattering lengths, we find agreement with previous work [23–25]. We find for the nominal dispersion coefficients (in a.u.) $a_T(^{85}\text{Rb}) = -369 \pm 16$, $a_S(^{85}\text{Rb}) = 2400_{350}^{+600}$, $a_T(^{87}\text{Rb}) = 106 \pm 4$, $a_S(^{87}\text{Rb}) = 90 \pm 1$. Our value for C_6 , 4700(50) a.u., is slightly higher (and with smaller uncertainty) than a previous analysis based solely on the g -wave shape resonance [21]. We have confirmed that several scattering observables predicted by the new Born-Oppenheimer potentials are consistent with previous measurements. Specifically, the new potentials predict a broad d -wave shape resonance [26] in ^{87}Rb , the scattering length ratio $a_{2,1}/a_{1,-1}$ [27], and the thermally averaged $|2, 2\rangle + |1, -1\rangle$ inelastic rate constant [28] that are consistent with previous measurements. We also find 10 of the 12 measured d -wave bound states [24] within the 2σ error bars. The new potentials also permit us to predict additional ^{85}Rb

Feshbach resonances [6] at $B_{\text{peak}} = 226.0 \pm 4.0$ G with a width of ~ 0.01 G, and at $B_{\text{peak}} = 535.5 \pm 4.0$ G with a width of 2.2 ± 0.2 G.

We have shown that precise measurements of the Feshbach resonance in the elastic scattering cross section provide unprecedented knowledge of the Born-Oppenheimer potentials that control Rb-Rb scattering processes at sub-mK temperatures. We have also used this resonance to enhance evaporative cooling of ^{85}Rb in a low density magnetic trap and change the scattering length from negative to positive. These capabilities should allow novel studies of BEC in the future.

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