Sonin Replies: My paper [1] and the Comment by Gaifullin et al. [2] are in disagreement as to whether the magnetoabsorption resonances observed in Bi layered superconductors in strong magnetic fields were the Josephson-plasma resonance (JPR). Gaifullin et al. [2] insist that these were JPR because of a good agreement of the experiment with the JPR theory of Bulaevskii et al. [3,4] (I call it the strong-pinning theory, see below). But my analysis [1] (call it the revised JPR theory) did not confirm crucial predictions of the strong-pinning theory. If the revised theory is correct, the agreement is fictitious. I cannot discuss here alternative interpretations for resonances and new experiments in fields 0-15 mT presented by Gaifullin et al. [2] because my paper addressed only the JPR interpretation of *old* experiments in much higher magnetic fields 3-6 T. So the topic of discussion is as follows: What does the JPR theory predict in strong magnetic fields?

If the dc magnetic field *H* is normal or nearly normal to layers, the JPR frequency $\omega(H) = \omega(0)\sqrt{\langle \cos \varphi \rangle}$ depends on *H* because vortices produce space fluctuations of the phase difference φ between layers. In order to explain the experiment, $\langle \cos \varphi \rangle$ must be much smaller than unity (about 1/25) and the strong-pinning theory explained it by large fluctuations of φ due to strong pinning of pancakes. But the revised theory estimated $1 - \langle \cos \varphi \rangle$ to be about 1%, i.e., 100 times less than in the strong-pinning theory.

Where does the disagreement start from? The revised theory [1] estimates the φ fluctuation using the expression $\langle \varphi^2 \rangle \sim L_J^2/a^2 \ln(\lambda_J/L_J)$ which agrees with Eq. (9) of Bulaevskii *et al.* [3]. Here, L_J is the average length of a Josephson string connecting two pancakes in neighboring layers, λ_J is the Josephson penetration length, and $a = \sqrt{\Phi_0/H}$ is the intervortex distance. The factor $\ln(\lambda_J/L_J)$ originates from the φ variation far from the string. So the disagreement is not due to my assumption that "the phase is large only inside a Josephson string". In fact, implicitly Bulaevskii *et al.* also used it.

But the two theories disagree on the value of L_J . According to Ref. [1], $L_J \ll a$ and the phase fluctuation is small. But the strong-pinning theory assumes that $L_J \sim a$ due to strong pinning of pancakes. Using the formula $\langle \cos \varphi \rangle = \exp(-\langle \varphi^2 \rangle)$ (see below) yields $\langle \cos \varphi \rangle \propto 1/B^{\nu}$ ($\nu \sim 1$), in agreement with the observed anticyclotronic dependence. Neither the previous works [3,4], nor Gaifullin *et al.* [2] present any serious argument to prove $L_J \sim a$ in the whole interval of fields excepting that "correlations in vortex positions along the *c* axis are very weak", i.e., compared with pinning. But I shall explain that strong pinning is not enough for the condition $L_J \sim a$ to be valid for the whole interval of fields.

Let us introduce the average distance L_d between pinning sites. Suppose that $L_d \ll a$. Then pancakes may choose among the closest pinning sites, i.e., $L_J \sim L_d \ll$ *a* and the fluctuation $\langle \varphi^2 \rangle$ is small, in agreement with Ref. [1]. Now assume that $L_d \gg a$. In this case there are some long Josephson strings with length $\sim a$. But there are not enough pinning sites for all pancakes, and the number of long strings is by factor a^2/L_d^2 smaller than the number of vortices. This means that the average Josephson string length is about a^2/L_d . So the fluctuation is weak again, and $\omega(H) \propto \langle \cos \varphi \rangle \approx 1 - \alpha \Phi_0/HL_d^2$, where $\alpha \sim 1$. Thus the correct JPR theory disagrees with the experiment even if pinning is strong: It predicts a nonmonotonous dependence on *H*, and $L_J \sim a$ only for nearly equal densities of vortices and pinning sites.

Gaifullin *et al.* argued that one may use the formula $\langle \cos \varphi \rangle = \exp(-\langle \varphi^2 \rangle)$ even for large $\langle \varphi^2 \rangle$, stating that *sometimes* it is correct (for the Gaussian distribution, e.g.). They do not present any argument why it is correct for *the case under consideration*. However, the dispute over how to calculate $\langle \cos \varphi \rangle$ at large $\langle \varphi^2 \rangle$ is important if $\langle \varphi^2 \rangle$ is large but, as discussed above, $\langle \varphi^2 \rangle \leq 1$.

For fields parallel to layers, I pointed out [1] that the collective-mode spectrum in Ref. [4] was incorrect since it had a gap between "optic" and acoustic branches at the Brillouin-zone boundary, in contradiction with symmetry of a one-site vortex-lattice cell. Gaifullin *et al.* deny the fact that this gap was present in the spectrum of Ref. [4]. But then the discussion is beyond physics: Any reader may check himself that, according to Eqs. (54) and (57) in Ref. [4], frequencies of optic and acoustic branches are not equal at the Brillouin-zone boundary ($q = \pm \pi$ in notations of Ref. [4]) and the gap between them is proportional to k^2 . Here $\vec{k}(k_x, k_y, 0)$ is the wave vector in the layer plane. Gaifullin *et al.* stressed that spectra in my paper and in Ref. [4] coincide if k = 0. It is right, but the spectrum in Ref. [4] is incorrect for $k \neq 0$.

In summary, the Comment cannot challenge my conclusions on wrong predictions of the strong-pinning theory both in normal and parallel magnetic fields. Therefore the question in the title of my paper remains without an answer, at least for strong fields.

E. B. Sonin Racah Institute of Physics, Hebrew University of Jersulam Jerusalem 91904, Israel and Ioffe Physical Technical Institute St. Petersburg 194021, Russia

Received 2 April 1998 [S0031-9007(98)07306-2] PACS numbers: 72.30.+q, 74.60.Ec, 74.60.Ge

- [1] E.B. Sonin, Phys. Rev. Lett. 79, 3732 (1997).
- [2] M. D. Gaifullin, Y. Matsuda, and L. N. Bulaevskii, preceding Comment, Phys. Rev. Lett. 81, 3551 (1998).
- [3] L. N. Bulaevskii et al., Phys. Rev. Lett. 76, 1719 (1996).
- [4] L.N. Bulaevskii et al., Phys. Rev. B 53, 14601 (1996).