## Resonant X-Ray Fluorescence Spectroscopy of Correlated Systems: A Probe of Charge-Transfer Excitations

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X-ray fluorescence spectroscopy with monochromatic photon excitation is presented as a tool for studies of charge-transfer excitations in correlated systems, using CeO<sub>2</sub> and UO<sub>3</sub> as examples. Ce  $4f \rightarrow 3d$  and U  $5f \rightarrow 3d$  x-ray fluorescence, with excitation near the 3d thresholds, probes states as eigenvalues for the ground state Hamiltonian from the Anderson impurity model. Sweeping the excitation energy across 3d absorption edges enhances contributions of different electronic configurations to fluorescence so that observed resonances indicate the charge-transfer origin of the absorption satellites. [S0031-9007(96)00691-6]

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A degree of localization of the *f* states which can have both bandlike and localized character is a key point in the description of the electronic structure of rare-earth and actinide systems. While delocalized models based on band theory were proposed to explain ground state [1] and spectroscopic [2] properties, various transport and spectroscopic signs of electron correlation effects are observed even in "itinerant" materials and cannot be fully interpreted within one-electron formalism. As an alternative, localized approach the Anderson impurity model (AIM) is often used [3,4] which treats the f states of a rare-earth or actinide atom as a degenerate impurity level hybridized with the valence band, but neglects the interaction between flevels on different atoms. To describe the ground state, a model Hamiltonian is constructed which includes as parameters the energy of a localized state  $\varepsilon_f$ , delocalized states  $\varepsilon_v$ , the hybridization strength V, and the on-site f-f Coulomb interaction  $U_{ff}$ . The parameters are optimized by fitting both high-energy spectroscopic and low-energy thermodynamic data [3]. However, these parameters may be renormalized in a different way in high- versus lowenergy experiments [5,6]. In this case, those experimental techniques become attractive which can be related to both high-energy and low-energy scale techniques such as valence band resonant x-ray fluorescence spectroscopy (RXFS). In the localized, many-body approach, RXFS is expected to probe the states as eigenvalues of the ground state Hamiltonian via creation or annihilation of a core or hole due to site and symmetry selectivity of RXFS. In this Letter, we explore a potential of RXFS in studies of so-called mixed-valency compounds where the anion  $2p \rightarrow$  metal f charge-transfer excitations play an important role and the ground state can be described as a strong mixture of several electronic configurations. Using CeO<sub>2</sub> [4,7–11] and UO<sub>3</sub> [12–15] as systems with strong metal f–O 2p hybridization we demonstrate that the localized, many-body approach is appropriate for the description of the resonant x-ray fluorescence (RXF) process and that RXF spectra can be interpreted within the framework of the AIM. We show that it is essential to include the RXF data in spectroscopic analysis for these systems in order to derive a unique set of model parameters.

The measurements on  $CeO_2$  were performed at beam line BW3 of HASYLAB, Hamburg, with a modified SX-700 monochromator [16]. X-ray absorption (XA) spectra at the Ce 3d edge were obtained by measuring total electron yield. The Ce 4f-to-3d RXF spectra were recorded using a grazing-incidence grating spectrometer [17]. The emitted radiation was detected in the vertical plane in the direction perpendicular to that of the incident beam. During RXF (XA) measurements, the resolutions of the monochromator and spectrometer were about 1.6(0.4) and 2.1 eV, respectively. For UO<sub>3</sub>, the experiment was carried out at beam line X24A of NSLS, Brookhaven, using a double-crystal monochromator equipped with Si(111) crystals [18]. The U 5f-to-3d RXF spectra were obtained using a spectrometer based on a bent Si(111) crystal and a position sensitive proportional detector [19]. The spectrometer was placed in the horizontal plane perpendicular to the incident beam. The resolutions of the monochromator and spectrometer were about 0.5 and 1 eV, respectively. For both UO<sub>3</sub> and CeO<sub>2</sub>, the incident angle for the incoming beam was  $\sim 20^{\circ}$  from the sample surface in order to minimize self-absorption effects [20]. The  $UO_3$ 

sample was a pressed pellet prepared from 99.8% material obtained from Alpha Products. CeO<sub>2</sub> was pelletized using 5 kbar pressure and the pellet was then fired for 24 h in air at 1575 K.

In the one-electron approach, for the excitation tuned to the threshold or above it, the valence band RXF spectra are expected to appear at the same photon energies (except for the elastic peak) and to sample the partial density of states. Some changes in the spectral shape may be observed though if, for example, momentum conservation [21] is fulfilled. For  $CeO_2$  and  $UO_3$ , however, the Ce 4f-to-3dand U 5f-to-3d spectral variations (Figs. 1 and 2) occur on a larger energy scale than that (2-3 eV) expected based on one-electron calculations for the f density of states [8,12,22]. The energy of the measured RXF spectra follows the varying excitation energy, thus arguing in favor of a localized, many-body description of the RXF process (the emitted photon energy  $\omega$  must increase with increasing incident photon energy  $\Omega$  in accordance with the equation  $\Omega + E_g = \omega + E_f$ , where  $E_g$  and  $E_f$  are the energies of the ground and final states, respectively).

Ce 4f-to-3d RXF spectra of CeO<sub>2</sub> recorded at different  $\Omega$  near the Ce 3d threshold are displayed in Fig. 1. For purely tetravalent Ce, one would expect the RXF spectrum to be basically a single, recombination line appearing

at  $\omega$  equal to  $\Omega$  as a result of the  $4f^0 \rightarrow 3d^9 4f^1 \rightarrow 4f^0$ excitation-deexcitation process. However, due to strong Ce 4f-O 2p hybridization in CeO<sub>2</sub>, the initial and intermediate states of the RXF process (Fig. 3) can be mainly described as mixtures of  $4f^0$  and  $4f^1\underline{v}$  configurations and  $3d^94f^1$  and  $3d^94f^2\underline{v}$  configurations ( $\underline{v}$  stands for a hole in the valence band), respectively, so that there are radiative transitions back to the ground state and to charge-transfer excited states. This explains the double-line shape of the RXF spectra of  $CeO_2$ . One can notice, however, that the energy separation between these two lines as well as their relative intensities strongly depend on  $\Omega$ . The observed changes can be attributed to variations in the transition rates to different final states of the RXF process which are the bonding (ground state), nonbonding, and antibonding states between  $4f^0$  and  $4f^1\underline{v}$  configurations. For CeO<sub>2</sub>, the RXF process is schematically shown in Fig. 3. When  $\Omega$  is set to the main Ce  $3d_{5/2}$  and  $3d_{3/2}$  XA peaks (884.1) and 902.0 eV) the intermediate state which corresponds to the bonding combination between  $3d^94f^1$  and  $3d^94f^2\underline{v}$ configurations decays predominantly to the bonding and nonbonding final states while the transitions to the antibonding final state constitute only  $\sim 5\%$  of the total intensity of the inelastic component. The RXF spectrum consists of the intense recombination peak and the line  $\sim$ 4.5 eV lower in energy and broadened due to the finite



UO<sub>3</sub> 3672.0 eV  $UM_5$ Absorption Emission 3672.0 eV Normalized intensity 3569.5 eV 3563.9 eV 3557.8 eV 3555.7 eV 3553.5 eV 3520 3530 3540 3550 3560 3570 Energy [eV]

FIG. 1. For CeO<sub>2</sub>, the experimental (dots) and calculated (solid lines) Ce 3d XA spectrum, and Ce 4f-to-3d RXF spectra at different excitation energies. RXF data are normalized to the beam monitor reference signal. The arrows on the XA spectrum indicate the excitation energies used for the RXF measurements.

FIG. 2. For UO<sub>3</sub>, the U  $3d_{5/2}$  XA spectrum, and U 5f-to-3d RXF spectra recorded at different excitation energies. RXF data are normalized to the beam monitor reference signal. The arrows on the XA spectrum indicate the excitation energies used for the RXF spectra.



FIG. 3. A schematic representation of the Ce 4f-to-3d RXF process for CeO<sub>2</sub>. Energy levels which are labeled with electronic configurations correspond to the average energies of these configurations in the the limit of  $V \rightarrow 0$ . The shaded rectangles represent the nonbonding states for the  $4f^{1}\underline{v}$  and  $3d^{9}4f^{2}\underline{v}$  configurations.

width of the O 2*p* band. Setting  $\Omega$  to XA satellites (889.3) and 907.3 eV), which are the antibonding combinations between  $3d^94f^1$  and  $3d^94f^2\underline{v}$  configurations [4], gives rise to a remarkable enhancement in the relative intensity of the low-energy line which is now  $\sim 7 \text{ eV}$  below the weak recombination peak. This enhancement can be ascribed to a resonance of antibonding states [23], an indicator of the charge-transfer origin of the XA satellite structures. In this case the transitions to the antibonding final state constitute (80-90)% of the total intensity of the inelastic component. It has been suggested [24] that the XA satellite structures may reflect the extended f states (f continuum). In this case, one would expect a relaxation process leading to the core-hole ionized state and nonresonantlike fluorescence spectra at  $\Omega$  set to 889.3 and 907.3 eV. Indeed, the low-energy line in the corresponding RXF spectra appears at approximately the same  $\omega$  as normal fluorescence. However, this low-energy line does exhibit energy dispersion (not shown) with  $\Omega$  varying through the satellite energy range. The situation is clearer for UO<sub>3</sub> where, for  $\Omega$  set to a similar charge-transfer satellite (~3563 eV) in the U  $3d_{5/2}$  XA spectrum, the resonating low-energy line in the U 5f-to-3d RXF spectra appears at different  $\omega$  than that for normal fluorescence. The description of the RXF process in CeO<sub>2</sub>, shown in Fig. 3, and corresponding assignments are supported by the results of calculations (Fig. 1) using the AIM.

The RXF spectra were calculated using the expression for the coherent second order optical process given by

$$F(\Omega, \omega) = \sum_{f} \left| \sum_{m} \frac{\langle f | T_2 | m \rangle \langle m | T_1 | g \rangle}{E_g + \Omega - E_m - i \Gamma_m} \right|^2 \\ \times \delta(E_g + \Omega - E_f - \omega), \qquad (1)$$

where  $|g\rangle$  and  $|f\rangle$  are the ground and final states, and  $|m\rangle$  is the intermediate state with energy  $E_m$ .  $T_1$  and  $T_2$  represent the optical dipole 3d-to-4f and 4f-to-3d transitions,

respectively. We assumed nonpolarized incident light. The calculations were done including three configurations:  $4f^0$ ,  $4f^1\underline{v}$ ,  $4f^2\underline{v}^2$  for  $|g\rangle$  and  $|f\rangle$ , and  $3d^94f^1$ ,  $3d^94f^2\underline{v}$ ,  $3d^94f^3\underline{v}^2$  for  $|m\rangle$ . The effect of multiplet coupling was taken into account in the manner described in Ref. [4]. The values used for the main parameters are as follows:  $V = 1.0, \varepsilon_f - \varepsilon_v = 2.0, U_{ff} = 9.0, \text{ and } U_{fc} = 12.6$  (in units of eV), where  $\varepsilon_v$  is the center of the valence band assigned a width W = 2.0 eV and  $U_{fc}$  is the magnitude of the core-hole potential acting on the 4f electron. Furthermore, we took into account the configuration-dependent hybridization [25] using two reduction factors  $R_c = 0.6$ and  $R_v = 0.925$ ; the hybridization between  $3d^9 4f^1$  and  $3d^94f^2\underline{v}$  configurations is taken as  $R_cV$ , and that between  $4f^1\underline{v}$  and  $4f^2\underline{v}^2$  configurations as  $V/R_v$ . Since the integrated intensity of the experimental Ce 4f-to-3d RXF spectra of CeO<sub>2</sub> (Fig. 1) is not proportional to the absorption coefficient at the Ce 3d edge due to saturation [26] and self-absorption effects the calculated RXF spectra were simply scaled so as to compare with the experiment.

It is essential to introduce  $R_c$  in order to explain the difference in the energy separations between peaks in XA (~5 eV) and 889.3- and 907.3-eV excited RXF (~7 eV) spectra of CeO<sub>2</sub>. The actual value of  $R_c$  was estimated from a consistent fit of the Ce 3d x-ray photoemission, XA, and RXF data of  $CeO_2$  and is found to be in between those for  $\alpha$ -Ce and U, respectively, obtained from first principle calculations [25]. The value of  $R_v = 0.925$  is not essential for the description of Ce 4f-to-3d RXF spectra; however,  $R_v \sim 0.9$  gives a better fit for the Ce 3d x-ray photoemission spectrum. The values of the model parameters used to describe the ground state of  $CeO_2$ are in agreement with those obtained from first principle calculations [8]. The latter calculations reveal strong energy dependence for  $|V(\varepsilon)|^2$ , being the largest on the top of the O 2p valence band. Our  $V^2$  is close to  $|V(\varepsilon)|^2$ averaged over  $\sim 1.5$  eV on the top of the O 2p band and the effective width W is almost the same as the energy range which gives a predominant contribution to  $|V(\varepsilon)|^2$ . Furthermore, using the set of parameters derived in the present analysis the experimental Ce 4d XA, valence band photoemission, and bremsstrahlung isochromat spectra of  $CeO_2$  were successfully reproduced [27]. The size of the insulating gap was estimated to be  $\sim 4.5$  eV, which is in better agreement with experiment than the value of  $\sim 2 \text{ eV}$ obtained in an earlier analysis [28].

For UO<sub>3</sub>, due to the high oxidation state of U (6+), one can expect a significant  $5f^2\underline{v}^2$  admixture in the ground state of the system so that this state can be described as a mixture of  $5f^0$ ,  $5f^1\underline{v}$ , and  $5f^2\underline{v}^2$  configurations. In the limit of  $V \rightarrow 0$ , the difference between configurationaveraged energies can be written as  $E(5f^1\underline{v}) - E(5f^0) =$  $\Delta$  and  $E(5f^2\underline{v}^2) - E(5f^1\underline{v}) = \Delta + U_{ff}$ . The specific shape of U 5f-to-3d RXF spectra for UO<sub>3</sub> (Fig. 2), recorded at  $\Omega$  set near the U  $3d_{5/2}$  threshold, indeed indicates high covalency of chemical bonds in this oxide.

The resonant part of U 5*f*-to-3*d* fluorescence follows  $\Omega$ due to energy conservation while for  $\Omega$  set far above the threshold the contribution of nonresonant, normal fluorescence is clearly observed at a constant  $\omega$ . When  $\Omega$  is set to satellites in the U  $3d_{5/2}$  XA spectrum at about 2–4 and 10 eV above the main XA maximum, one can see some enhancement in the spectral weight of the U 5f-to-3dfluorescence at about 5, 8.5, and 14.5 eV, respectively, below the recombination line. As in the case of  $CeO_2$ , such resonances indicate the charge-transfer origin of satellites (antibonding combinations between  $3d^95f^1$ ,  $3d^95f^2\underline{v}$ , and  $3d^95f^3\underline{v}^2$  configurations) in the U  $3d_{5/2}$ XA spectrum of UO<sub>3</sub>. The most supportive evidence for the idea that the ground state of  $UO_3$  is a mixture of three configurations is the shape of the RXF spectrum recorded at the excitation energy of 3563.9 eV. This spectrum exhibits essentially three lines at about 3563.9, 3558.8, and 3549.4 eV which can be tentatively associated with  $5f^0$ ,  $5f^1\underline{v}$ , and  $5f^2\underline{v}^2$  configurations, respectively. The ability to determine from RXF spectra the energies of the states associated with the  $5f^2\underline{v}^2$  configuration puts additional restrictions on the value of  $U_{ff}$  for UO<sub>3</sub> because  $U_{ff}$  determines the energy of this configuration.

In conclusion, Ce 4f-to-3d and U 5f-to-3d RXF spectra of CeO<sub>2</sub> and UO<sub>3</sub> demonstrate that low-energy chargetransfer excitations in correlated systems can be studied by the corresponding core-hole excitation-deexcitation process when there are radiative transitions back to the ground state and low-energy charge-transfer excited states. Because of its site and symmetry selectivity, RXFS allows one to probe Ce 4f and U 5f states which are basically the eigenvalues of the ground state Hamiltonian in the AIM. The values of model parameters for the ground state of correlated systems can be derived by analyzing the corresponding resonant fluorescence spectra within this model.

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- [1] For example, B. Johansson and M.S.S. Brooks, in Handbook on the Physics and Chemistry of Rare-Earths (North-Holland, Amsterdam, 1993), Vol. 17, p. 149.
- [2] For example, M.R. Norman et al., Phys. Rev. Lett. 53,

1673 (1984); E. Weschke *et al.*, Phys. Rev. Lett. **69**, 1792 (1992).

- [3] O. Gunnarsson and K. Schönhammer, Phys. Rev. B 28, 4315 (1983); O. Gunnarsson *et al.*, J. Appl. Phys. 63, 3676 (1988).
- [4] A. Kotani and H. Ogasawara, J. Electron. Spectrosc. Relat. Phenom. 60, 257 (1992); Physica (Amsterdam) 186B– 188B, 16 (1993).
- [5] O. Gunnarsson and N.E. Christensen, Phys. Rev. B 42, 2363 (1990).
- [6] J. W. Allen and L. Z. Liu, Phys. Rev. B 46, 5047 (1992).
- [7] D. D. Koelling et al., Solid State Commun. 47, 227 (1983).
- [8] A.K. McMahan and R.M. Martin, in *Narrow-Band Phenomena*, edited by J.C. Fuggle, G.A. Sawatzky, and J.W. Allen (Plenum, New York, 1988), p. 133.
- [9] A. Fujimori, Phys. Rev. B 28, 2281 (1983); 28, 4489 (1983).
- [10] E. Wuilloud et al., Phys. Rev. Lett. 53, 202 (1984).
- [11] F. Marabelli and P. Wachter, Phys. Rev. B 36, 1238 (1987).
- [12] M. V. Ryzhkov and V. A. Gubanov, J. Radioanal. Nucl. Chem. 143, 85 (1990).
- [13] B. W. Veal and D. J. Lam, Phys. Lett. 49A, 466 (1974).
- [14] Yu. A. Teterin et al., Phys. Chem. Miner. 7, 151 (1981).
- [15] G. Kalkowski et al., Phys. Rev. B 35, 2667 (1987).
- [16] T. Möller, Synchrotron. Radiat. News 6, 16 (1993).
- [17] J. Nordgren et al., Rev. Sci. Instrum. 60, 1690 (1989).
- [18] P.L. Cowan *et al.*, Nucl. Instrum. Methods Phys. Res., Sect. A 246, 154 (1986).
- [19] S. Brennan et al., Rev. Sci. Instrum. 60, 2243 (1989).
- [20] RXF measurements, performed with two different incident angles (20° and 37°) at some excitation energies, reveal no significant angular dependence for the Ce 4f-to-3d spectra of CeO<sub>2</sub>. This may be connected with the poorly defined surface of the CeO<sub>2</sub> pellet and with specific x-ray detection geometry. In turn, analytical estimates made for the 20° incidence angle indicate that selfabsorption leads to the  $\sim 25\%$  reduction or to the  $\sim 35\%$ enhancement in the integrated-intensity ratio between elastic and inelastic components in the RXF spectra of CeO<sub>2</sub> when the excitation energy is set to the main line or to the satellite, respectively, in the Ce 3d XA edges. For UO<sub>3</sub>, changes in the shape of RXF spectra caused by selfabsorption are found to be less pronounced compared to those for  $CeO_2$  because of the smaller concentration of the f element in UO<sub>3</sub> and larger core-hole lifetime broadening for the U 3d edge.
- [21] Y. Ma et al., Phys. Rev. Lett. 69, 2598 (1992).
- [22] L. D. Finkelstein et al., Mater. Lett. 14, 115 (1992).
- [23] S. Tanaka et al., J. Phys. Soc. Jpn. 59, 1488 (1990).
- [24] R.C. Karnatak et al., Phys. Rev. B 36, 1745 (1987).
- [25] O. Gunnarsson and O. Jepsen, Phys. Rev. B 38, 3568 (1988).
- [26] For example, S. Eisebitt *et al.*, Phys. Rev. B **47**, 14103 (1993).
- [27] M. Nakazawa, S. Tanaka, T. Uozumi, and A. Kotani (unpublished).
- [28] T. Nakano et al., J. Phys. Soc. Jpn. 56, 2201 (1987).