

**Astrophysical S Factor of $^{12}\text{C}(\alpha, \gamma)^{16}\text{O}$ from the Beta-Delayed Alpha-Particle Emission of ^{16}N
[Phys. Rev. Lett. 70, 2066 (1993)]**

Z. Zhao, R. H. France III, K. S. Lai, S. L. Rugari, M. Gai, and E. L. Wilds

In the abstract it is stated that the statistical uncertainty of the measured S_{E1} factor is ± 16 keVb, where in fact it is clearly stated in the paper and in Fig. 1(b) that the statistical uncertainty is ± 6 keVb. Hence the final quoted S factor is $S_{E1}(300) = 95 \pm 6(\text{stat}) \pm 28(\text{syst})$ keVb.

In addition the uncertainty of the branching ratio of the beta decay to the broad 1^- state at 9.6 MeV measured at the Michigan University Radioactive Beam Facility is stated to be ± 2 but it should be ± 3 . Hence the final quoted branching ratio for beta decay to that state is $1.3(3) \times 10^{-5}$.

**The Galileo Solar Redshift Experiment
[Phys. Rev. Lett. 70, 2213 (1993)]**

Timothy P. Krisher, David D. Morabito, and John D. Anderson

In the second line of Eq. (3), the subscript of the first alpha should be sc/\odot (*not* st/\odot) and the subscript of the second U should be sc/\oplus (*not* st/\oplus).

**Wetting Dynamics in a Confined Symmetric Binary Mixture Undergoing Phase Separation
[Phys. Rev. Lett. 70, 2770 (1993)]**

Hajime Tanaka

J. Bodensohn and W. I. Goldberg [Phys. Rev. A 46, 5084 (1992), Ref. [13]] should have been cited in the discussion on the relation between the initial radius of disklike droplet $2a_0$ and the film thickness d ($2a_0 \sim d$). They first reported a similar relation in their paper. It should be noted that they derived the above relation in a different way from ours.