

## Noise-Sustained Structure in Taylor-Couette Flow with Through Flow

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We report experimental and theoretical results for the absolute and convective instability boundaries in Taylor-Couette flow with imposed axial flow as a function of the axial Reynolds number. Experiment and theory agree quantitatively. In the downstream region of a large-aspect-ratio system, noise-sustained structures of traveling vortices exist in much of the convectively unstable regime. These structures have a nearly time-independent amplitude, but a noisy phase. The phase noise ceases abruptly upon crossing the absolute instability boundary.

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The sensitivity of open flows to perturbations has been well documented since Leconte observed "musically inclined" jets responsive to acoustic excitation [1]. Numerous examples exist for open systems such as pipe and channel flows and obstacle wakes [2,3]. However, the role of noise in destabilizing laminar states and producing turbulence remains unclear.

In open systems it is important to distinguish between absolutely unstable flows, in which disturbances grow at any fixed point, and convectively unstable ones, in which fluctuations grow only in a comoving frame [4-6]. In the latter case, individual fluctuations grow as they are advected downstream, but are eventually "blown out" of the system so that no permanent structure is produced in the absence of noise or perturbations. However, in a study of a complex Ginzburg-Landau (CGL) equation which models many open flows, it was found that a weak but persistent noise source generated sustained structures in the downstream portion of the system [7].

In this Letter we describe an investigation of stability and noise in an open flow, namely, Taylor-Couette flow between concentric cylinders [8] with imposed axial flow. Experimentally we find that if the system is sufficiently long, noise-sustained structure predominates when the base flow is convectively unstable. The observed structures are patterns of traveling Taylor vortices whose phase varies irregularly with time. This irregularity vanishes within our resolution at the transition from convective to absolute instability. We presume that the patterns are sustained by noise originating in the inlet region. As shown below, noise effects can be observed near the onset of the centrifugal (Taylor) instability even at small axial Reynolds numbers  $R \lesssim 4$ . Our observations clarify the interplay between noise and stability, and permit a quantitative description by the appropriate CGL equation.

The experiments were performed in a modified Taylor-Couette apparatus with the inner cylinder rotating. A recirculating gravity feed generated axial flow whose rate was controlled to an accuracy of 0.2%. The radius ratio  $r_1/r_2$  was 0.738 and the aspect ratio  $L/d$  was 144. Here  $d=r_2-r_1$  and the total length  $L=98$  cm. Both cylinders were concentric and straight to  $\pm 0.04$  mm. We used water-glycerol mixtures seeded with 1% Kalliroscope suspension by volume for flow visualization. A computer-

interfaced charge-coupled-device camera recorded scan lines and time series of reflected light. The fluid temperature was controlled to better than  $\pm 30$  mK. Before entering the gap, the fluid passed through an annulus containing eight holes of diameter 1.1 mm, followed by two layers of 0.44-mm-mesh stainless-steel screen to produce azimuthally uniform axial flow.

This system has two control parameters. One is the axial Reynolds number  $R \equiv \langle w \rangle d / \nu$ , where  $\langle w \rangle$  is the mean axial flow velocity and  $\nu$  the kinematic viscosity. The second is the reduced angular rotation frequency of the inner cylinder  $\epsilon \equiv \omega / \omega_0 - 1$ , where  $\omega_0$  refers to the onset of Taylor-vortex flow at  $R=0$  [8]. The structureless base flow, comprised of circular Couette flow superimposed with axial Poiseuille flow, first becomes unstable to traveling vortices [9]. Recently, the distinction between convective and absolute instability has been examined for this system [10] and for thermal convection with through flow [11]. Below we describe high-resolution measurements of the convective and absolute instability regimes made possible by the long aspect ratio of our apparatus.

We found that convective instability could be directly observed by generating localized perturbations in the base flow at fixed  $\epsilon$  and  $R$ . This was done by rotating both cylinders about their axis back and forth once through a small angle [12]. As shown in the space-time plot of Fig. 1, a pulse localized to within two or three vortex pairs appeared at the inlet and traveled into the system. The pulse initially decayed over a fraction of the distance  $\langle w \rangle d^2 / \nu$ , which is approximately that required for the inlet flow to attain the angular momentum of the base flow. The pulse then spread, grew in amplitude, and was advected out of the system. To probe the onset of convective instability, time series of reflected light were recorded as pulses passed a point  $100d$  from the inlet. The amplitude of the fundamental peak in the discrete Fourier transform (DFT) was found for pulses at various  $\epsilon$  for given  $R$ . Reducing  $\epsilon$  caused slower pulse growth, and hence weaker signals at the observation point. The intersection of the DFT peak amplitude with the noise base line was taken as the onset of convective instability  $\epsilon_c$ . Figure 2 shows the results for  $R \lesssim 20$ . At larger  $R$ , spiral modes [13] grew and complicated the measurements.

We also calculated numerically the onset of convective

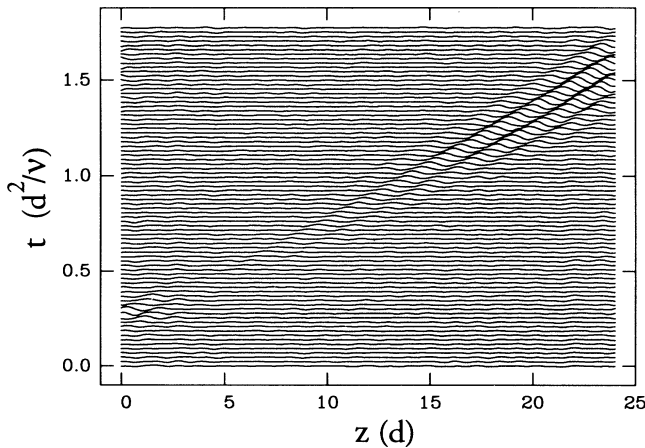


FIG. 1. Pulse growth in the convectively unstable regime;  $R=16.5$  and  $\epsilon=0.160$ . At time  $t=0.2$ , the apparatus was rocked back and forth once. The length shown is about  $L/6$ .

instability for our experimental geometry from the Navier-Stokes equations. Instability to modes of the form  $\exp[i(kz - \Omega t + m\Theta)]$  was examined, where  $t$  is the time, and  $z$  and  $\Theta$  are the axial and azimuthal coordinates, respectively. For axisymmetric modes ( $m=0$ ), the results are well fitted by  $\epsilon_c = 0.000381R^2(1 - 9.3 \times 10^{-5}R^2)$ , shown as the solid curve in Fig. 2. They are in excellent agreement with our measurements [14], and are consistent with previous calculations for other radius ratios [13].

For  $\epsilon$  sufficiently greater than  $\epsilon_c$ , structure was observed in the absence of applied perturbations. Figure 3 shows space-time plots of traveling vortex patterns at several  $\epsilon$  for  $R=3.0$ . The pattern amplitude was roughly time independent at a given position, but decreased to zero, or nearly zero, at the inlet because the flow entered with no azimuthal velocity. The spatial growth may be measured by the healing length  $l_h$  at which the pattern reaches half its maximum amplitude [11]. As can be seen in Fig. 3,  $l_h$  grew and the amplitude of the pattern at

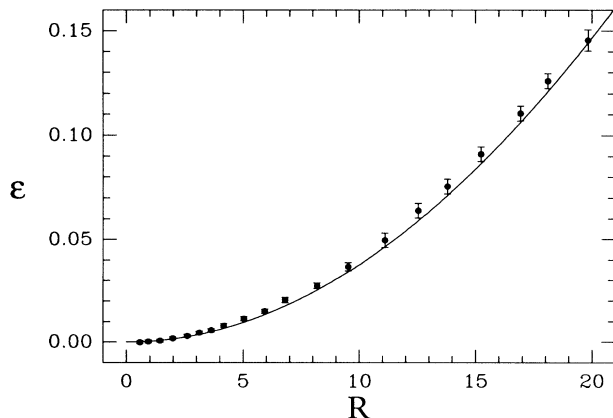


FIG. 2. Convective instability boundary  $\epsilon_c$  obtained by observing pulse growth. The curve is the linear stability result.

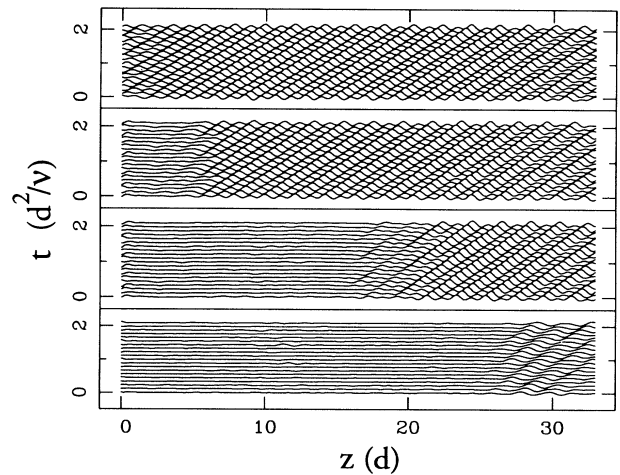


FIG. 3. Space-time plots of patterns at  $R=3.0$ , showing the increase in healing length  $l_h$  as  $\epsilon$  decreases. Top to bottom:  $\epsilon=0.1020, 0.0822, 0.0632,$  and  $0.0347$ . About  $\frac{1}{4}$  of the apparatus is shown.

large  $z$  decreased as  $\epsilon$  was reduced.

A recent numerical study of thermal convection with through flow found that, in the absence of noise, patterns resembling those in Fig. 3 arose only in the absolutely unstable regime [11]. The healing length  $l_h$  increased as  $\epsilon$  was reduced, and diverged at the boundary  $\epsilon_a$  between convective and absolute instability. However, as suggested by the CGL equation [7], one expects that noise-sustained structure with finite  $l_h$  can exist in the convectively unstable regimes of real systems. We shall see that the lower two plots in Fig. 3 are examples of such structures. A diagnostic other than the behavior of  $l_h$  is thus necessary to locate  $\epsilon_a$  in the presence of noise.

Detailed examination of the structures of Fig. 3 revealed a qualitative change as  $\epsilon$  was reduced. After

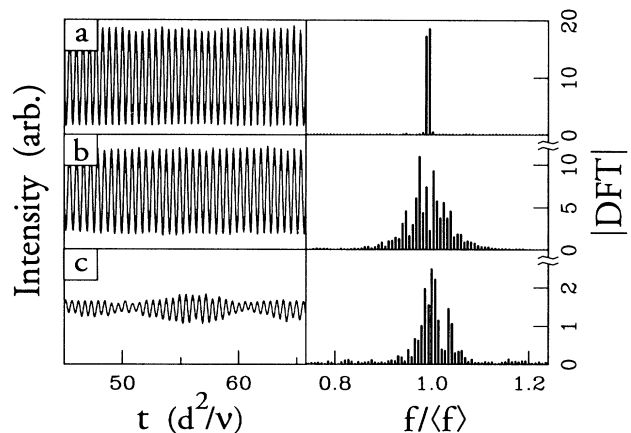


FIG. 4. Time series, and DFT moduli near the fundamental peak, at  $z=100d$  for  $R=3.0$ . About  $\frac{1}{4}$  of each series is shown. (a)  $\epsilon - \epsilon_c = 0.0896$ . (b)  $\epsilon - \epsilon_c = 0.0318$ . The DFT peak shows considerable broadening. In (c), structure persists at small  $\epsilon - \epsilon_c = 0.0077$ .

fixing  $\epsilon$  and  $R$  and waiting for transients to die away, time series of reflected light were recorded for  $z=100d$ . Figure 4 shows portions of time series and corresponding DFT moduli for various  $\epsilon$  at  $R=3.0$ . Each series covered about 160 periods and contained 2048 points. In Fig. 4(a),  $\epsilon - \epsilon_c = 0.0896$ , and the data appear periodic and of nearly uniform amplitude. The corresponding fundamental DFT peak is very sharp. At  $\epsilon - \epsilon_c = 0.0318$  in Fig. 4(b), the time series remains fairly uniform in amplitude, but the DFT peak has become much broader. Demodulation shows that this broadening is due to irregular time dependence of the phase. As shown for  $\epsilon - \epsilon_c = 0.0077$  in Fig. 4(c), sustained patterns persist to small  $\epsilon - \epsilon_c$ , but the amplitude shows time variation, and the DFT remains broad. The transition in spectral width can be characterized by the normalized variance of the fundamental peak in the power spectrum  $\sigma^2 = \langle (f - \langle f \rangle)^2 \rangle / \langle f \rangle^2$ , calculated by integrating over frequencies in the vicinity of the fundamental component. The experimental result for a range of  $\epsilon$  at  $R=3.0$  is shown in Fig. 5(a). The sharp increase in  $\sigma^2$  below  $\epsilon=0.065$  corresponds to the onset of phase noise.

Simulations of the CGL equation

$$\tau_0(\dot{A} + sA') = \tilde{\epsilon}(1 + ic_0)A + \xi_0^2(1 + ic_1)A'' - g(1 + ic_2)|A|^2A \quad (1)$$

showed that the transition in  $\sigma^2(\epsilon)$  is related to the boundary dividing absolute and convective instability. Equation (1) is derived from an expansion around  $\epsilon_c$ . Here  $\tilde{\epsilon} = \omega/\omega_c(R) - 1 = (\epsilon - \epsilon_c)/(1 + \epsilon_c)$ , and  $s \propto R$  is the group velocity of the critical mode at onset. Length and time are scaled by the gap  $d$  and diffusion time  $d^2/\nu$ , respectively. For the linear terms, coefficients specific to our geometry were obtained from the stability analysis described above [15]. A discrete version of Eq. (1) was

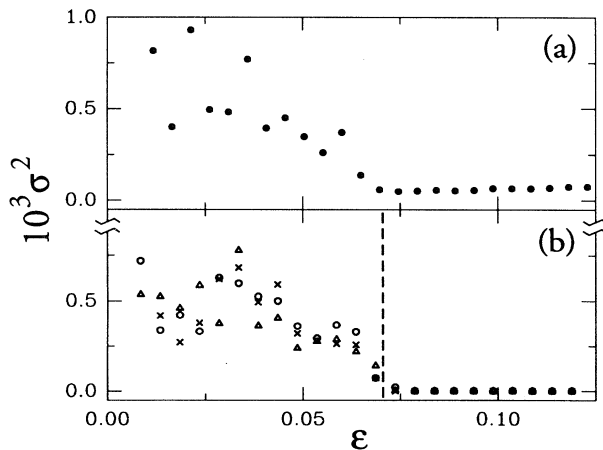


FIG. 5. (a) Normalized variance of the fundamental peak in the DFT power vs  $\epsilon$  at  $R=3.0$  for time series at  $z=100d$ . The transition at  $\epsilon=0.065$  indicates the onset of phase noise. (b) Corresponding results for Eq. (1) for noise levels  $10^{-6}$  ( $\times$ ),  $10^{-5}$  ( $\Delta$ ), and  $10^{-4}$  ( $\circ$ ), respectively. The dashed line locates  $\epsilon_a^{CGL} = 0.071$ .

numerically integrated. Noise was modeled by choosing random values for the real and imaginary parts of  $A(z=0)$  at each time step, a form chosen because the most relevant noise originates in the inlet region. Values were uniformly distributed between  $\pm n$ , where we call  $n$  the noise level. We mimicked the experiment by analyzing time series of the function  $\Psi = \text{Re}\{A(z,t)\exp[i(k_c z - \Omega_c t)]\}$  at  $z=100d$ . The time series and their DFTs closely resembled their experimental counterparts. As shown in Fig. 5(b), the peak variance  $\sigma^2$  undergoes a transition remarkably similar to that in the experiment. Demodulation shows that the broadening is again due to phase noise.

It is readily shown [7,11] that the boundary between convective and absolute instability in Eq. (1) occurs at  $\tilde{\epsilon}_a^{CGL} = (s\tau_0/2\xi_0)^2/(1+c_1^2)$ . The vertical dashed line in Fig. 5(b) corresponds to  $\tilde{\epsilon}_a^{CGL}$  for our geometry at  $R=3.0$ . The onset of phase noise is clearly identified with  $\tilde{\epsilon}_a^{CGL}$  and is also insensitive to the noise level. The onset varied less than  $\delta\epsilon = \pm 0.005$  for  $10^{-9} \leq n \leq 10^{-3}$ , where onset is taken to be the point where the variance  $\sigma^2$  first departs from the base line. Thus, even though the physical noise source and level are uncertain, we are confident that the experimental transition to phase noise corresponds to the boundary  $\epsilon_a$ .

Figure 6 shows the complete stability diagram for axisymmetric structures up to  $R \approx 4.0$ . The upper data points mark  $\epsilon_a$ , identified by the transition to phase noise. For  $R \gtrsim 4$ , spiral modes grew and interfered with measurements of  $\epsilon_a$ . We calculated the absolute stability boundary from the full equations [16], using methods outlined in Refs. [5] and [6]. The results are well fitted by  $\epsilon_a = 0.00782R^2(1 - 0.0043R^2)$ , and are shown as the upper

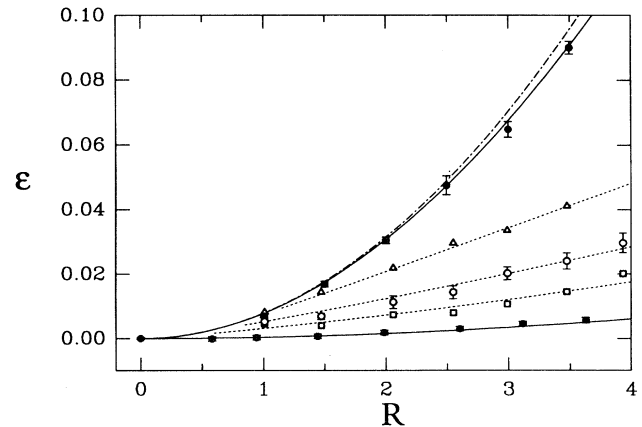


FIG. 6. Stability diagram for axisymmetric structures. The lower points and solid curve are for the onset of convective instability. The upper data points locate the experimental transition to phase noise, the upper solid curve is the linear stability result for  $\epsilon_a$ , and the dash-dotted line is  $\epsilon_a^{CGL}$ . Open symbols indicate the boundary  $\epsilon_s$  for noise-sustained structure at  $z=25d$  ( $\Delta$ ),  $z=50d$  ( $\circ$ ), and  $z=100d$  ( $\square$ ). The dotted lines are the estimates of  $\epsilon_s$ , given by Eq. (2) with  $A_{\min}/A_{\delta}^{\text{ff}} = 590$ .

solid curve. The lower set of data points is for the convective instability boundary  $\epsilon_c$ , and the corresponding solid curve is the linear stability result. Experiment and theory agree very well for both boundaries. The upper dash-dotted curve is the CGL result  $\epsilon_a^{\text{CGL}} = 0.00789R^2(1 - 0.0008R^2)$ . Its close agreement with the experiment and the full equations suggests that Eq. (1) applies rather well throughout the parameter range of Fig. 6.

With the stability boundaries precisely located, we can be sure that structures observed at  $\epsilon < \epsilon_a$ , such as the lower two plots of Fig. 3, are sustained by noise. We measured  $\epsilon_s$ , the minimum  $\epsilon$  for which structure was observed, at various distances from the inlet. A spectral method similar to that employed in measuring the convective instability boundary was used. The results for  $z = 25d$ ,  $50d$ , and  $100d$  are shown in Fig. 6 as open symbols. Further from the inlet, structure was observed at increasingly smaller  $\epsilon$ , because fluctuations had more time to grow. In the limit of very large aspect ratio, the boundary  $\epsilon_s$  approaches  $\epsilon_c$ , and noise-sustained structure predominates in the convectively unstable regime [17].

The effects of noise can be modeled by an effective inlet boundary condition  $A_0^{\text{eff}} \neq 0$  which gives rise to a time-independent amplitude

$$A(z) = A_0^{\text{eff}} \exp\{z\bar{\epsilon}_a^{1/2}[1 - (1 - \bar{\epsilon}/\bar{\epsilon}_a)^{1/2}]/\xi_0\} \quad (2)$$

obtained as the solution of the linear part of Eq. (1). We may invert this expression for  $\bar{\epsilon}$  to give  $\epsilon = \epsilon_c + (1 + \epsilon_c)\bar{\epsilon}$ . After inserting the linear stability values for  $\epsilon_c$ ,  $\epsilon_a$ , and  $\xi_0$ , this provides an estimate of the structure boundary  $\epsilon_s$  with the single free parameter  $A_{\text{min}}/A_0^{\text{eff}}$ . Here  $A_{\text{min}}$  is the minimum experimentally detectable amplitude which prevails at  $\epsilon_s$ . A fit to the data gives  $A_{\text{min}}/A_0^{\text{eff}} \approx 600$ . The corresponding estimate of  $\epsilon_s$  is shown as the dashed lines in Fig. 6. The estimate  $A_{\text{min}} = 10^{-3}$  determined in a similar experimental configuration [18,19] yields  $A_0^{\text{eff}} \approx 2 \times 10^{-6}$ . From simulations of Eq. (1), we find that a noise level about an order of magnitude larger than  $A_0^{\text{eff}}$ , or  $n \approx 2 \times 10^{-5}$ , produces amplitude profiles that match the data. These values motivated the choice of noise levels shown in Fig. 5(b).

In conclusion, we have shown that noise-sustained structure predominates in a convectively unstable open flow. The increase in phase noise at the transition from absolute to convective instability may serve as a useful diagnostic in other open flows (for example, object wakes [5]) which possess both instability regimes.

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- [15] We found

$$\begin{aligned} k_c &= 3.136 + 0.000104R^2(1 - 0.00014R^2), \\ \Omega_c &= 3.677R(1 + 0.000020R^2), \\ \tau_0 &= 0.0379 - 0.000024R^2(1 - 0.00046R^2), \\ s &= 1.230R(1 + 0.0000041R^2), \\ \xi_0^2 &= 0.0725 - 0.000048R^2(1 - 0.00053R^2), \\ c_0 &= 0.00720R(1 - 0.00038R^2), \\ c_1 &= 0.0235R(1 + 0.00026R^2). \end{aligned}$$

The amplitude scale was set by a nonlinear coefficient  $g$  of order 1. We used  $c_2 = 0.0113R$  as in Ref. [11] for convection with through flow. Our results were insensitive to the particular (small) value of  $c_2$ .

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- [17] In the experiment of Ref. [10], it was stated that no sustained structure was observed in the convectively unstable regime. That work used an aspect ratio  $L/d = 27$  and attempted to measure  $\epsilon_a$  by finding a divergence in the healing length  $l_h$ . The result deviated from the CGL value for  $R > 1.3$ . Note that our data for  $\epsilon_s$  at  $z = 25d$  in Fig. 6 intersects  $\epsilon_a$  near  $R = 1.2$ , suggesting that the deviation in Ref. [10] was due to noise-sustained structure with a comparable noise level.
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