

### Comment on "Space-Time as a Causal Set"

In a recent Letter by Bombelli, Lee, Meyer, and Sorkin,<sup>1</sup> the authors discuss the merits of viewing space-time not as a continuous manifold, but rather as a discrete set of points with connections between them defining the large-scale causal structure. This would remove the ultraviolet divergences plaguing present-day field theory, and the singularities of general relativity; it would provide a basis for making space-time, in addition to the fields existing on it, a dynamical variable<sup>2</sup>; and it would make the amount of information and computation going on in a finite volume of space-time finite as opposed to infinite.

This Comment is intended to explore the idea of a lattice which is Lorentz invariant, a property we know the Universe to have, at least at large scales. We know such a lattice would have to be random,<sup>3</sup> to avoid having preferred directions; then the local connections would have some probability distribution  $\rho(x_1 - x_2) = \rho(x)$ . Since space-time volume  $dV = d^3\mathbf{x} dt$  is invariant under Lorentz transformations,  $\rho$  must be a function only of  $\Delta s$ , the invariant interval:  $\rho = \rho(\Delta s)$ . Then the coordination number, or the average number of points each point is connected to, is

$$N = \int d^4x \rho(\Delta s). \quad (1)$$

But this integral diverges, since each hyperboloid shell of constant  $\Delta s$  has an infinite volume. Even if we restrict connections to the light cone, with  $\rho = \rho(r)\delta(ct - r)$ , Lorentz invariance requires that  $\rho(r) = \text{const}/r$ , and then

$$N = \int 4\pi r^2 dr \rho(r) \quad (2)$$

diverges again. Besides, the next-neighbor distribution would have the form (1). Connections outside the light cone, if they exist at all, similarly diverge.

So there is no normalizable probability distribution that is Lorentz invariant: More precisely, there is no Lorentz-invariant probability measure of  $\mathbb{R}^4$  besides  $\delta^4(x)$ .

This means that if the Universe is indeed a random lattice, the coordination number is infinite at the scale at which space-time has 3+1 dimensions and is Lorentz invariant—but of course one wants the real coordination

number, at the smallest scales, to be finite. Perhaps this puzzle can be solved in terms of the 3+1 space being a "coarse graining"<sup>1</sup> of a lattice whose dimensionality is higher at smaller scales; the coordination number will vary in some relation to the scale of the coarse graining. However, the coordination number would seem to *decrease* with the dimensionality as we increase the scale, rather than increasing to infinity as would be needed to resolve the paradox. In the absence of such a resolution there seem to be four possibilities:

(1) Space-time is not, in fact, discrete.

(2) Space-time is discrete, but it has the nasty property that every point is influenced by an infinity of "nearest neighbors" which, in a given frame, are arbitrarily far back in time.

(3) Physics is macroscopically Lorentz invariant and isotropic but the underlying lattice is not; this is somewhat plausible since similar phenomena can be seen in certain cellular automata with regard to space isotropy.<sup>4</sup> However, if there is a rigid structure then curvature must result from lattice defects: This is an ungainly way to do general relativity.

(4) Space-time is discrete, but the lattice can only be strictly defined in the nonrelativistic limit; relativistically (or quantum mechanically) the lattice points are ill defined in such a way that it is not even possible to construct an average  $\rho(x)$  in the flat-space state.

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<sup>1</sup>L. Bombelli, J. Lee, D. Meyer, and R. D. Sorkin, *Phys. Rev. Lett.* **59**, 521 (1987)

<sup>2</sup>See, for instance, T. D. Lee, *Phys. Lett.* **112B**, 217 (1983).

<sup>3</sup>N. H. Christ, R. Friedberg, and T. D. Lee, *Nucl. Phys.* **B202**, 89 (1982).

<sup>4</sup>For instance, lattice gases or any cellular automata based on diffusive processes such as diffusion-limited aggregation. See also N. Packard and S. Wolfram, *J. Stat. Phys.* **38**, 901 (1985).