

Comment on "Naked Singularities in Self-Similar Spherical Gravitational Collapse"

In a recent Letter,¹ Ori and Piran have shown that the self-similar spherical gravitational collapse of an adiabatic perfect fluid with a soft enough equation of state can give rise to a naked shell-focusing singularity.² This important work extends beyond dust³ known fluid-collapse histories which have nakedly singular end states. For dust it is known⁴ that the naked shell-focusing singularities are *not* strong-curvature singularities. The purpose of this Comment is to point out that the singularities produced by Ori and Piran *are* strong-curvature singularities.

Following the work of Clarke and Królak,⁵ consider a null geodesic (N), affinely parametrized by λ , with four-tangent l^a , and terminating at the shell-focusing singularity where $\lambda = 0$. If

$$\lim_{\lambda \rightarrow 0} \lambda^2 R_{\alpha\beta} l^\alpha l^\beta \neq 0, \quad (1)$$

where $R_{\alpha\beta}$ is the Ricci tensor, then N terminates in a strong-curvature singularity in the sense of Tipler.⁶

For the example given by Ori and Piran¹ (and in their notation) it follows from Einstein's equations, the comoving condition, the equation of state $p = (\Gamma - 1)\rho$, and the self-similar relations $\rho = D(y)/T^2$ and $e^\psi = (4\pi \times D)^{2(1-\Gamma)/\Gamma}$ that

$$\lambda^2 R_{\alpha\beta} l^\alpha l^\beta = 2\Gamma(4\pi D)^{(2-\Gamma)/\Gamma} \left[\frac{\lambda}{T} \frac{dT}{d\lambda} \right]^2. \quad (2)$$

Now the appropriate N here is the Cauchy horizon $y = y_1 = \text{const}$. The null geodesic equations for the Cauchy horizon integrate explicitly to give

$$T = R/y_1 \propto \lambda^\delta, \quad \delta = \text{const} > 0, \quad (3)$$

and so

$$\lambda^2 R_{\alpha\beta} l^\alpha l^\beta = 2\Gamma \delta^2 [4\pi D(y_1)]^{(2-\Gamma)/\Gamma}. \quad (4)$$

That is, the Cauchy horizon terminates in a *strong-curvature* singularity.

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¹A. Ori and T. Piran, Phys. Rev. Lett. **59**, 2137 (1987).

²D. M. Eardley and L. Smarr, Phys. Rev. D **19**, 2239 (1979).

³Reference 2; D. Christodoulou, Commun. Math. Phys. **93**, 171 (1984); R. P. A. C. Newman, Classical Quantum Gravity **3**, 527 (1986).

⁴See Newman (Ref. 3).

⁵C. J. S. Clarke and K. Królak, J. Geom. Phys. **2**, 127 (1986).

⁶See, for example, F. J. Tipler, C. J. S. Clarke, and G. F. R. Ellis, in *General Relativity and Gravitation*, edited by A. Held (Plenum, New York, 1980).