
ERRATA

Cosmic Quarkonium: A Probe of Dark Matter.
MARK SREDNICKI, STEFAN THEISEN, and JOSEPH SILK
[Phys. Rev. Lett. **56**, 263 (1986)].

The right-hand side of Eq. (2) should be multiplied by a factor of 3 for color. Correspondingly, the branching ratio in Fig. 2 should be reduced by a factor of 3. This reduces the predicted line fluxes by the same factor. We thank S. Rudaz for bringing this error to our attention.

Also, the interaction of Eq. (1) is to be interpreted as a vertex factor appropriate when one applies the usual Feynman rules, treating both λ and Q as *Dirac* spinors. If the Feynman rules are altered to take into account the Majorana nature of λ , the right-hand side of Eq. (1) should be multiplied by $\frac{1}{2}$. See H. E. Haber and G. L. Kane, Phys. Rep. **117**, 75 (1985), for a thorough discussion.

Precise Determination of $\sin^2\theta_w$ from Measurements of the Differential Cross Sections for $\nu_{\mu}p \rightarrow \nu_{\mu}p$ and $\bar{\nu}_{\mu}p \rightarrow \bar{\nu}_{\mu}p$. K. ABE, L. A. AHRENS, K. AMAKO, S. H. ARONSON, E. W. BEIER, J. L. CALLAS, P. L. CONNOLLY, D. CUTTS, D. C. DOUGHTY, L. S. DURKIN, B. G. GIBBARD, S. M. HEAGY, D. HEDIN, J. S. HOFTUN, M. HURLEY, S. KABE, Y. KURIHARA, R. E. LANOU, Y. MAEDA, A. K. MANN, M. D. MARX, T. MIYACHI, M. J. MURTAGH, S. MURTAGH, Y. NAGASHIMA, F. M. NEWCOMER, T. SHINKAWA, E. STERN, Y. SUZUKI, S. TATSUMI, S. TERADA, D. H. WHITE, H. H. WILLIAMS, Y. YAMAGUCHI, and T. YORK [Phys. Rev. Lett. **56**, 1107 (1986)].

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