Oscillatory Structures in GaAs/(AlGa)As Tunnel Junctions

Recently, Hickmott *et al.*¹ reported a remarkable oscillatory structure with period $e\Delta V = \hbar \omega_{LO}$, the longitudinal-optic phonon energy, in the lowtemperature reverse-bias J(V) characteristics of $n^+GaAs/(AlGa)As/n^-GaAs/n^+GaAs$ devices. The structure was observed only in magnetic fields (*B*) above 4 T which, they claimed, reduced ionizedimpurity scattering by inducing freezeout onto shallow donors. This permitted sequential ballistic acceleration of electrons up to the LO phonon emission threshold.

studied similar molecular-beam-We have epitaxy-grown, 200-µm-diam mesas with layer thicknesses and doping as follows: substrate and buffer, 200 μ m, 2×10¹⁸ cm⁻³; *n*⁻ layer, 1 μ m, 1×10¹⁵ cm⁻³; Al_{0.35}Ga_{0.65}As barrier, 168 ±10 Å, un-doped; top layer, 1 μ m, 2×10¹⁸ cm⁻³. Figure 1 shows typical J(V) plots. Mesas of type A have J(V) of similar overall form to those in Ref. 1, but of much larger amplitude ($\sim 10^3$ larger at V = -0.1 V). Mesas of type B have an additional low-bias (V > -0.2V) contribution to J which may arise from imperfections (e.g., microchannels) in the (AlGa)As barrier. Even at B=0, derivative plots of both mesa types show oscillatory structure (Fig. 2) of comparable relative amplitude (e.g., $\sim 2\%$ at -0.2 V) to that in Ref. 1. The periodicity of the main peaks is $\hbar \omega_{\rm LO}/e$. Structure is clearly visible (Fig. 2, curves c and d) even at temperatures (30 K) at which the majority of donors are ionized. The weaker intermediate extrema, seen only for T < 4 K, probably arise from higher harmonic structure in J(V); at 4 K some peaks in d^2J/dV^2 are only 2 mV wide. Note that our barriers are $\sim 20\%$ narrower than those in Ref. 1 and that J is exponentially sensitive² to barrier height and width. Calculation shows that J(V) of Fig. 1 (type A) is fully consistent with the barrier-tunneling characteristics. The



FIG. 1. $\log J vs - V$ for two of our devices.



FIG. 2. d^2J/dV^2 vs -V: Curves *a* and *d*, type-A mesas; curves *b* and *c*, type-B mesas.

larger J(V) passed by our devices may be crucial for observation of oscillatory structure even at B = 0 and up to 30 K and for understanding of its origin.^{3,4} The low reverse-bias shoulder on J(V) in type-B mesas is removed by modest (≈ 5 T) transverse magnetic fields $(J \perp B)$, suggesting that it may be due to microchannels in the barrier.⁵ However, oscillations for V > 0 are not observed.⁵ Oscillatory structure in both mesa types for V < -0.2 V is essentially unaffected by magnetic fields (\parallel and $\perp J$) up to 12 T.

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²The exponential argument is ~ -30 .

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