diagonalization of the matrix, Porter assigned the transitions provisionally and then made an adjustment of  $+8.9 \text{ cm}^{-1}$  in the vibrational frequencies and  $+1.6\%$  in the rotational energies to obtain the best fit of the experimental values. The close agreement  $(0.03 \text{ cm}^{-1} \text{ maximum error})$  on all four assigned transitions confirms the assignments, in our judgment. The remaining four transitions appear to involve higher vibrational levels, higher rotational levels, or both. Unequivocal assignment was not possible because of the high density of candidates.

This Letter reports preliminary data for a few transitions. We expect that additional resonances will be found soom. Only about  $4\%$  of the infrared spectral region presently accessible to our experimental apparatus has been explored thus far. Such additional data would greatly facilitate transition identification. We hope that these results, combined with the results of  $Oka<sub>1</sub><sup>1</sup>$  will lead to precise predictions of transitions observable in precise predictions of transitions observable<br>the interstellar medium in  $\text{H}_3^+$  and  $\text{H}_2\text{D}^+.$  We also plan to search for the  $H_2D^+$  spectrum in the laboratory.

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Ab Initio Prediction of the Rotation-Vibration Spectrum of  $H_3^+$  and  $D_3^+$ 

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The first few lines in the  $P$ ,  $Q$ , and  $R$  branches of the rotation-vibration spectra of the equilateral-triangle molecular ions  $H_3^+$  and  $D_3^+$ , obtained by ab initio nonperturbative calculations, are reported. Comparison with observations indicates an accuracy better than 1% was obtained for both the infrared-active fundamental vibration frequency and the equilibrium internuclear distances.

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In this Letter we report the first results of an ab initio nonperturbative calculation of the rotation-vibration spectrum of a highly anharmonic, nonrigid polyatomic molecule. Our choice of system is  $H_3^{\ \ +}$  (and its deuterium analogue  $D_3^{\ \ +}$ ), an ion thought to play an important role in the formation of molecules in the interstellar medium.<sup>1</sup> Although the calculations were carried out in the absence of observed spectra, during the preparation of our report of the results measurements of rotation-vibration spectra were achieved by Qka' tation-vibration spectra were achieved by Oka $^{\circ}$ <br>for H<sub>3</sub>  $^{\circ}$  and by Shy *et al* . $^3$  for D<sub>3</sub>  $^{\circ}$  , in two quite remarkable experiments.

The calculated spectra are sensitive to the potential-energy surface. $4$  By conventional standards (size of electronic basis set, low absolute electronic eigenenergies), the best calculation of electronic eigenenergies), the best calculation of<br>a potential-energy surface for H<sub>3</sub><sup>+</sup> is that of Dykstra et  $al.^5$  However, the most convenient surface available at the time these calculations were begun was a configuration-interaction (CI) surface' designed to be relatively consistent over the required range of nuclear coordinates. We therefor report here our results for our CI surface

and let its accuracy be assessed by comparison with the observations. Details of our use of the method of vibrational-configuration interaction to obtain the lowest pure-vibrational (i.e.,  $J=0$ ) eigenenergies have been reported. '

In the work reported here, we diagonalize the Born-Oppenheimer Hamiltonian for nuclear motion in the product basis of the pure-vibrational eigenfunctions and the symmetric-top eigenfunceigenfunctions and the symmetric-top eigenfunctions in real form.<sup>8</sup> If  $Q_i$   $(i = 1, 2, 3)$  are the normal-mode coordinates for an equilateral  $X_3$  molecule, the Hamiltonian for bound nuclear motion' can be expressed as a matrix in the pure-vibrational eigenfunction basis:

$$
H_{nm} = E_n S_{nm} + \frac{1}{2} A_{nm} J_x^2 + \frac{1}{2} B_{nm} J_y^2 + \frac{1}{2} C_{nm} J_z^2
$$

$$
+ \frac{1}{2} D_{nm} (J_x J_y + J_y J_x) + i F_{nm} J_z, \qquad (1)
$$

n	m	$\langle n   A   m \rangle$	$\langle n B m\rangle$	$\zeta n\, \,c\, \,m\>$ $>$	$\langle n D m\rangle$	$\langle n\, \,\mathbf{E}\, \,\mathbf{m}\,\rangle$	$\mathbf{Ln} \mid \mathbf{F} \mid \mathbf{m}$ $\mathbf{L}$
				$(H_2^+)$			
$\mathbf 1$	$\mathbf 1$	86.446	86.450	41.129	0.0	0.0	$0.0$
$\sqrt{2}$	$\,1$	$-15.197$	15.185	$-0.007$	0.096	0.0	0.0
3	$\mathbf 1$	0.096	$-0.096$	0.0	15.193	0.0	0.0
4	$1\hskip-3.6pt.$	$-12.484$	$-12.526$	$-6.223$	$0.0$	0.0	0.0
$\sqrt{2}$	$\sqrt{2}$	82.735	92.638	39.470	$-0.061$	2515.99	0.0
3	$\,2$	0.062	$-0.062$	$0.0$	$-4.846$	0.0	37.031
4	$\sqrt{2}$	$-2.353$	2.347	$-0.008$	0.015	0.0	0.0
3	3	92.509	82.833	39.471	0.062	2516.17	0.0
4	3	0.015	$-0.015$	0.0	2.360	0.0	0.051
4	4	84.308	84.384	39.984	0.0	3185.32	0.0
				$(D_3^+)$			
$\mathbf 1$	$\mathbf 1$	43.311	43.311	20.918	0.0	0.0	0.0
$\sqrt{2}$	$1\,$	$-6.417$	6.415	0.0	0.066	0.0	0.0
3	$\mathbf 1$	0.066	$-0.066$	0.0	6.416	0.0	0.0
$\overline{4}$	$\mathbf 1$	$-5.425$	$-5.429$	$-2.701$	0.0	0.0	0.0
$\sqrt{2}$	$\sqrt{2}$	41.939	45.353	20.308	$-0.035$	1826.39	0.0
3	$\sqrt{2}$	0.0351	$-0.0351$	0.000	$-1.695$	0.0	19.423
4	$\sqrt{2}$	$-0.795$	0.794	0.000	0.008	0.0	0.0
3	3	45.338	41.951	20.308	0.035	1826.42	0.0
4	3	0.008	$-0.008$	0.000	0.795	0.0	0.007
4	4	42.542	42.552	20.514	0.000	2307.47	0.0

TABLE I. Vibrational matrix for the rotational coefficients.<sup>a</sup>

<sup>a</sup>Values are given in cm<sup>-1</sup>. The matrices A, B, and C are symmetric; the matrix  $F$  is skew symmetric;  $E$  is the diagonal matrix of the pure vibration  $(J = 0)$  eigenenergies.

where  $\mathbf{\tilde{J}}$  is the rotational angular-momentum operator whose components are referred to a molecule-fixed coordinate system with origins at the center of mass,  $z$  axis perpendicular to the molecular plane, and  $x$  axis bisecting an  $X-X$  "bond." In Eq. (1),  $E_n$  is the *n*th pure-vibrational eigenenergy and the rotational coefficient matrices have the general form

$$
\langle n|f(Q_1,Q_2,Q_3)|m\rangle\,,
$$

where  $|m\rangle$  is the *m*th pure-vibrational eigenket and  $f$  is an appropriate function of the instantaneous moments of inertia given in terms of the normal-mode coordinates. Integration is effected by the generalized Gauss quadrature previously described.<sup>7</sup> The matrices  $\overline{A}$ ,  $\overline{B}$ ,  $\overline{C}$ , and  $\overline{D}$  are symmetric, and  $\overline{F}$  is skew symmetric. The straightforward construction of the matrix of Eq. (1) in the symmetric-top basis will be described in a full account to be published elsewhere.

As a minimal vibrational basis, we take the four lowest-energy states  $(0, 0, 0)$ ,  $(0, 1, 0)$ ,  $(0, 0, 0)$ 1), and  $(1,0,0)$ , where  $(v_1, v_2, v_3)$  gives the nominal<sup>10</sup> numbers of quanta in the  $Q_1$ ,  $Q_2$ , and  $Q_3$  vibrations, respectively. Numbering these states 1-4, we give the values of the vibrational matrix components of the rotational coefficients in Ta-

ble I. The  $D_{nm}$  term and a term in the difference  $A_{nm} - B_{nm}$  couple adjacent odd or adjacent even values of the molecular  $z$ -axis angular-momentum component quantum number K. The term in  $F_{nm}$ is the Coriolis coupling that splits levels with nonzero  $K$  in the presence of vibrational angular momentum.

We report our results in Tables II and III in the form of electric dipole-allowed transition frequencies (cm<sup>-1</sup>) for the first dozen or so lines in<br>the P,Q, and R branches of the  $\nu_F$  fundamental.<sup>11</sup> the P, Q, and R branches of the  $\nu_E$  fundamental.<sup>11</sup> A discussion of the term patterns and eigenfunctions will be given in the full report, but we mention here that for  $H_3^*$  (but not  $D_3^*$ ), the  $A_1$  breathing mode is rotationally coupled to the  $E$  infraredactive mode for  $J \ge 6$ . This may have implications for the relaxation of the metastable  $(1,0,0)$  vibration state to  $(0,1,0)$ ,  $(0,0,1)$ , and  $(0,0,0)$  at relatively low rotational excitation. Furthermore, a slight calculated splitting of some of the  $E$  levels at high J and low K in  $H_3^+$  appears to indicate that a larger vibrational basis is required for accurate results when  $J \gtrsim 5$  or 6. Calculations with a 10-member basis are in progress.

-member basis are in progress.<br>Since the D $_3^{+}$  amplitudes are smaller, the vibra tion-rotation interaction is less, and we expect our results to be more accurate than for  $H_3^*$ . We

P-Branch		$P_{Q-Branch}$		$r_{Q-Branch}$		R-Branch	
$P_{P_1}(1)E$	2452.277	$P_{Q_1}(1)E$	2539.963	$r_{Q_1(6)E}$	2526.808	$r_{R_1(1)E}$	2687.855
$P_{P_2}(2)E$	2374.398	$P_{Q_1}(2)E$	2548.592	$r_{Q_0(1)A_2}$	2525.417	$r_{R_0(1)A_2}$	2720.257
$P_{P_1}(2)E$	2367.786	$P_{Q_2}(2)E$	2550.053	$r_{Q_2(6)E}$	2518.993	$P_{R_1}(1) E$	2720.770
$P_{P_3}(3)$ A <sub>2</sub>	2295.215	$P_{Q_3}(3)$ A <sub>2</sub>	2561.062	$r_{Q_1(2)E}$	2515.678	$r_{R_2(2)E}$	2761.340
$P_{P_2(3)E}$	2292.806	$P_{Q_2}(3)E$	2563.684	$r_{Q_0(5)A_2}$	2512.940	$r_{R_1}(2)E$	2766.382
$P_{P_1}(3)E$	2292.550	$P_{Q_A}(4)E$	2568.927	$r_{Q_0(3)A_2}$	2512.467	$P_{R_2}(2)E$	2820.931
$r_{P_0(3)A_2}$	2292.696	$P_{Q_1}(3)E$	2568.698	$r_{Q_1}(3)E$	2510, 340	$P_{R_1}(2)E$	2824.740
$r_{P_1}(3)E$	2259.636	$P_{Q_5}(5)E$	2578.497	$r_{Q_1}(5)E$	2510.276	$r_{R_3(3)A_2}$	2833.144
$P_{P_1}(4)E$	2231.452	$P_{Q_3}(4)$ A <sub>2</sub>	2576.178	$r_{Q_1}(4)E$	2506.769	$r_{R_2}(3)E$	2840.291
$P_{P_2}(4)E$	2224.872	$P_{Q_6}(6)$ A <sub>2</sub>	2588.877	$r_{Q_2}(3)E$	2504,093	$r_{R_1}(3) E$	2844.015
$P_{P_3(4)A_2}$	2219.512	$P_{Q_2}(4)E$	2592.898	$r_{Q_2(5)E}$	2504.002	$r_{R_A(4)E}$	2904.899
$P_{P_A}(4)$ E	2215.838	$P_{Q_A}(5)E$	2597.411	$r_{Q_2(4)E}$	2501.479	${}^{r}R_{3}(4)A_{2}$	2915.697

TABLE II. Results of ab initio calculation of the  $H_3^+$  rotation-vibration spectrum.<sup>2</sup>

<sup>a</sup>Superscripts p, r denote  $\Delta K = -1$ , +1, respectively (in absorption) and P,Q,R denote  $\Delta J = -1,0, +1$  respectively; the numerical subscript is the initial (lower level)  $K$  value, and the number in parentheses is the initial  $J$  value. The letter beside the line symbol is the symmetry (in  $D_{3h}$ ) of the rotation-vibration species. Line positions are in  $cm^{-1}$ .

R-Branch 1879.503
1915.621
1926.891
1927.091
1954.145
1955.809
1974.883
1976.123
1976.690
1992.047
1992.249
1994.184

TABLE III. Results of ab *initio* calculation of the  $D_2$ <sup>+</sup> rotation-vibration spectrum.<sup>3</sup>

<sup>a</sup>See footnote to Table II.

therefore concentrate on  ${\rm D_3}^+$  in our compariso with observations. The principal errors in the calculation are in the rotationless vibrational transition frequency  $v_0$  (calculated to be 2515.990 transition frequency  $\nu_0$  (calculated to be 251.<br>cm<sup>-1</sup> for H<sub>3</sub><sup>+</sup>, 1826.391 cm<sup>-1</sup> for D<sub>3</sub><sup>+</sup>) and the equilibrium internuclear distance  $R_e$  (calculated to be 1.65737 $a_0$ ). The error in  $\nu_0$  is additive and the error in  $R_e$  enters as a factor of the rotational constants (roughly proportional to  $R_e$ <sup>-2</sup>). Thus we may adjust the *ab initio* line positions  $\nu$  by writing

 $\nu$ <sub>adj</sub> =  $\gamma \nu$  +  $q$ ,

where  $r = (R_{\text{calc}}/R_e)^2$  and  $q = (1 - r)\nu_0 + \Delta \nu$ . In Table IV we give adjusted values of the first four  $D_3^+$  line positions in the P branch. An excellent fit to observations<sup>12</sup> is obtained for  $\Delta v = 8.937 \text{ cm}^{-1}$ 

TABLE IV. Adjustment for errors in the pure vibrational transition frequency and equilibrium internuclear distance for  $D_3^+$ .

Transition	Ab Initio	Adjusted <sup>a</sup>	Observed <sup>b</sup>	
$P_{P_1}(1) E$ ${}^{p}P_{2}^{(2)}E$ $\overline{P_1(2)} E$ ${}^{r}P_{0}(2)$ $A_{1}$	1794.295 1753.937 1751.499 1749.826	1802.706 1761.686 1759.208 1757.508	1802.716 1761.709 1759.205 1757.478	
${}^{\text{a}}$ See Eq. (2).		<sup>b</sup> Reference 12.		

and  $r = 1.016394$  [corresponding to an estimated  $R_e$  of 1.644 $a_0$  (compare 1.650 $a_0$  obtained in Ref. 5 and in an "improved" ACI calculation reported in Refs.  $4$  and  $6$ ). The agreement of observed and adjusted line positions is remarkable in view of the fact that the  ${}^{r}P_{0}(2)$   $A_{1}$  line at 1757.478 cm<sup>-1</sup> has an upper state that is shifted  $\sim 1.4 \text{ cm}^{-1}$  by l. type doubling. In the case of  $H_3^+$ , the observed lines<sup>13</sup> are generally at positions 5 and 6  $cm^{-1}$ higher than the *ab initio* values. We conclude that our method for the rotation-vibration calculation is satisfactory and that future work should concentrate on refinement of methods for obtaining accurate and consistant potential-energy sur faces. accurate and consistant potential-energy surface.<br>Analysis of both the  $H_3^+$  and  $D_3^+$  theoretical and observed spectra for an accurate  $R_e$  value that will allow determination of the  $H_2D^+$  pure-rotation spectrum is in progress.

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