

TABLE II. Number n_c^\pm of $1p$ nucleons involved in the coherent excitation of Σ -hypernuclear 0^+ states in (K^-, π^+) reactions, respectively, at 0° on ^{16}O for $p_K = 720$ MeV/ c , and conversion width ratio $\Gamma/\Gamma_{\text{nm}}$. We show two representative cases, corresponding to weak or strong Σ spin-orbit coupling.

A_Z structure	I	n_c^-	n_c^+	$\Gamma/\Gamma_{\text{nm}}$
$(Np_{3/2}^{-1} \otimes \Sigma p_{3/2})_0^+$	1/2	4/3	0	1.3
	3/2	8/3	4	0.6
$(Np_{3/2}^{-1} \otimes \Sigma p_{1/2})_0^+$	1/2	2/3	0	1.2
	3/2	4/3	2	0.8
1S_0	1/2	2	0	1.4
	3/2	4	6	0.3
3P_0	1/2	0	0	1.1
	3/2	0	0	1.1

DYNAMIC CONFINEMENT FROM VELOCITY-DEPENDENT INTERACTIONS. M. King and F. Rohrlich [Phys. Rev. Lett. 44, 621 (1980)].

On page 622 in the top equation of the second column $(\vec{\xi} - \vec{\pi})^2$ should read $(\vec{\xi} \cdot \vec{\pi})^2$. On p. 624, Eq. (9b) $l - \frac{5}{2}$ should read $l + \frac{5}{2}$. In the equation preceding (9b) the argument of u_i should be $(\beta\xi)^2$ which corresponds to $t = (\beta\xi)^2$. In the last line of the left column on p. 624 "has been earlier" should read "has been noted earlier."

PHASE TRANSITIONS AND MAGNETIC MONOPOLE PRODUCTION IN THE VERY EARLY UNIVERSE. Alan H. Guth and S.-H. H. Tye [Phys. Rev. Lett. 44, 631 (1980)].

In the first paragraph, the time corresponding to $T = 10^{17}$ GeV should be $t \sim 10^{-41}$ sec. Equation (2) should read

$$p(t) = \exp \left[-\frac{4\pi}{3} \int_0^t dt_1 R^3(t_1) \lambda(t_1) \left(\int_{t_1}^t dt_2 \frac{v}{R(t_2)} \right)^3 \right].$$

Footnote 12 should read the following: If the density of monopoles exceeds this bound, they would noticeably influence the cosmological deceleration parameter. See Ref. 7.