al Academic Printing Co. Ltd., Tokyo, 1977). ⁵P. Braun-Munzinger, G. M. Berkowitz, T. M. Cormier, C. M. Jachcinski, J. W. Harris, J. Barrette,

and M. J. LeVine, Phys. Rev. Lett. 38, 944 (1977).

⁶M. R. Clover, R. M. DeVries, R. Ost, N. J. A. Rust, and H. E. Gove, Bull. Am. Phys. Soc. <u>23</u>, 615 (1978).

⁷J. C. Peng, J. V. Maher, W. Oelert, D. A. Sink, C. M. Cheng, and H. S. Song, Nucl. Phys. <u>A264</u>, 312 (1976).

⁸P. Braun-Munzinger, A. Gamp, C. K. Gelbke, and H. L. Harney, Z. Phys. A276, 107 (1976).

⁹In a recent publication [D. Dehnhard, V. Shkolnik,

and M. A. Franey, Phys. Rev. Lett. <u>40</u>, 1549 (1978)], the ¹⁶O +²⁸Si excitation function is reporduced by introducing a parity dependence into the optical model implying the importance of core exchange terms. This calculation, however, fails to reproduce the angular distribution at $E_{\rm c.m.}$ = 26.23 MeV reported here. ¹⁰K. W. McVoy, Phys. Rev. C <u>3</u>, 1104 (1971).

¹¹J. G. Cramer, R. M. DeVries, D. A. Goldberg, M. S. Zisman, and C. F. Maguire, Phys. Rev. C <u>14</u>, 2158 (1976).

¹²F. M. Strutinsky, Zh. Eksp. Teor. Fiz. <u>46</u>, 2078 (1964) [Sov. Phys. JETP <u>19</u>, 1401 (1964)].

Ambiguities in Pion-Nucleus Optical Potentials and the Determination of Neutron Radii

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The $\pi^{\pm}-{}^{48}$ Ca elastic scattering data are shown to be reproduced with a wide range of rms neutron radii if the optical-model parameters are suitably adjusted. The corresponding total cross sections vary significantly.

It has long been expected that any difference between the proton and neutron distributions in nuclei will lead to corresponding differences in π^{\pm} -nucleus differential¹ and total² cross sections in the energy region dominated by the (3, 3) πN resonance. This is because at these energies $\sigma(\pi^{\pm}p) \simeq 3\sigma(\pi^{\pm}n)$, so that the π^{\pm} mainly "sees" protons and, by symmetry, the π^{-} mainly sees neutrons. With the aid of recent data^{3,4} for ^{40,48}Ca, we have now found that if both differential- and total-cross-section data are fully analyzed, the ambiguities are reduced in extracting radial parameters.

Electron and proton scattering experiments⁵ indicate that in ⁴⁰Ca the rms proton and neutron radii, r_p and r_n , are nearly equal; also, r_p (⁴⁸Ca) $\simeq r_n$ (⁴⁰Ca). Hartree-Fock calculations⁶ give Δr_n $= r_n$ (⁴⁸Ca) - r_n (⁴⁰Ca) = 0.31 fm; however, scattering of 79-MeV α particles⁷ indicates Δr_n = 0.08 ± 0.08 fm, and 1-GeV proton scattering⁸ gives 0.16 ± 0.02 fm. Generally, in neutron-rich nuclei, measured neutron radii are smaller than the Hartree-Fock predictions.⁹

Two π^{\pm} experiments have been done on ^{40,48}Ca, yielding somewhat different conclusions about Δr_n . Jakobson *et al.*³ performed measurements at the Clinton P. Anderson Meson Physics Facility (LAMPF) at several energies of the differences in the total cross sections, $\Delta \sigma^{\pm} = \sigma_T^{\pm} ({}^{48}\text{Ca}) - \sigma_T^{\pm} ({}^{40}\text{Ca})$. They adjusted the radial parameters in an optical-model calculation to obtain a fit and concluded that the rms neutron radii differ by $\Delta r_n = 0.14 \pm 0.05$ fm. At the Swiss Institute for Nuclear Research, Egger *et al.*⁴ measured elastic scattering over a range of angles at 130 MeV. They associated the minima with nuclear radii via a black-disk model, and found that the π^- radius exceeded the π^+ radius by 0.21 fm for ⁴⁰Ca, and by 0.51 fm for ⁴⁸Ca. They ascribed the ⁴⁰Ca difference to Coulomb effects, and assumed that the additional radius difference of 0.3 fm in ⁴⁸Ca is due to a larger neutron radius. They state that these results are consistent with the Hartree-Fock predictions⁶ but not those of Jackson *et al.*³

One problem in interpreting both of these experiments is that some knowledge of the optical potential is required. Jakobson $et al.^3$ used optical-model parameters based on free- πN data; these parameters reproduce the general features of pion scattering in this energy region, and, as they noted, their analysis is relatively insensitive to the values of the parameters as long as they are the same for both ⁴⁰Ca and ⁴⁸Ca. In the elastic scattering experiment,⁴ the "blackness" of the nucleus is needed to associate the disk radius R with a neutron radius. If the medium is made more absorptive, the absorptive region will move further out into the nuclear surface; since the black-disk diffraction minimum occurs at $2kR\sin(\theta/2) = 3.83$, θ decreases as the imaginary part of the optical potential increases. The same effect occurs if the potential is made

more attractive, so that the pion wave number k is increased.¹⁰ By themselves, the positions of the minima yield little information about the nuclear radii.

Since we know the complete elastic-scattering distribution at 130 MeV, we can in fact check the optical-model assumptions at that energy. Using the electron-scattering density parameters¹¹ for both the protons and the neutrons and the free- πN phase shifts, we found that the differential cross sections calculated with the ABACUS-M code¹² agree qualitatively with the data, although they do not fit in detail. The minima are located rather well, suggesting that $r_n = r_b$ is consistent with the elastic data. However, the calculated minima are too deep; the predicted absorption is too large. Varying the off-shell behavior (Kisslinger¹³ versus Laplacian models¹⁴) and including the Lorentz-Lorenz effect¹⁵ or the c.m.system to laboratory correction¹⁶ does not change this general pattern (Fig. 1).

To see the effect of using phenomenological parameters in extracting neutron density information, we began with ⁴⁰Ca. We used a Kisslinger model with a full Lorentz-Lorenz correction. Initially, we set the proton and neutron half-density $(c_{p} \text{ and } c_{n})$ and thickness parameters $(a_{p} \text{ and } c_{n})$ and a_n) equal to the electron-scattering values¹¹; we held the small s-wave coefficient b_0 in the potential constant and varied the *p*-wave coefficient b_1 , which was taken¹⁷ to be pure isospin $\frac{3}{2}$. When we fitted to the elastic π^{\star} data simultaneously, we obtained an excellent fit; as anticipated, the fitted value of Imb_1 was significantly smaller than the free value (see Fig. 2 and Set I in Table I). Alternatively, we fitted the π^+ data alone, since these cross sections are rather insensitive to the neutron distribution. We then used the resulting b_1 parameter in π^- calculations, varying the neutron parameters c_n and a_n to optimize the fit; this gave $r_n < r_p$, with $c_n > c_p$ but $a_n < a_p$ (Set II in Table I). However, this procedure was not unique; when we changed the distribution used in fitting the π^+ data, the neutron parameters found from the π^- data also changed somewhat.

Turning to ⁴⁸Ca, we found that using the ⁴⁰Ca Set I or Set II best-fit b_1 values and the electronscattering ⁴⁸Ca density parameters¹¹ for both the protons and the neutrons gave a better agreement than found with the free- πN parameters. (Figure 2 shows the Set II calculations; results with Set I are similar.) Again the minima are located fairly well, so that there is no necessity to introduce

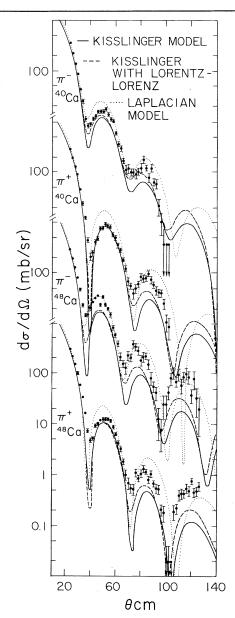


FIG. 1. Differential cross sections calculated with parameters obtained from free- πN phase shifts (Table I). Poorer results are obtained if the angle transformation is included. Data are from Ref. 4.

a larger neutron radius.

To see how much better agreement with the ⁴⁸Ca data is possible, we chose several sets of neutron density parameters and varied b_1 for each set (Table I). We found equally good fits for a range of density parameters, e.g., with $a_n = 0.52$ fm, for c_n from 3.4 to 4.4. In each case the minima are correctly located. However, as c_n is increased, the fitted values of Re b_1 and Im b_1 both decreases. In terms of a black-disk

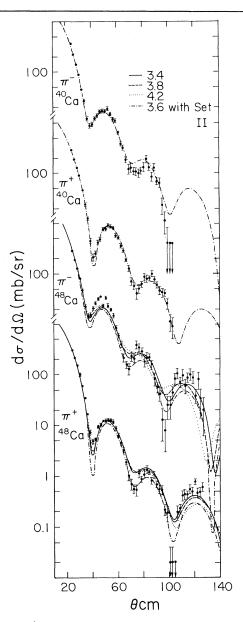


FIG. 2. ⁴⁰Ca best-fit calculations as discussed in the text. The ⁴⁸Ca curves are with fitted parameters for three τ_n values plus the Set II parameters obtained by fitting ⁴⁰Ca. See Table I for parameter values.

model, the wave number k and the opaqueness of the surface are both decreasing so as to offset the change in c_n .

Clearly the elastic scattering data cannot be used by themselves at this point to find a neutron radius, since we have no empirical or theoretical basis for preferring any one of the fitted sets of optical-model parameters. However, even though they all reproduce the elastic data equally well, they do predict rather different results for the

TABLE I. Optical-model parameters. Electron-scattering density parameters (Ref. 11) for 40 Ca are c = 3.31 fm, a = 0.60 fm; for 48 Ca, c = 3.60 fm, a = 0.52 fm. The last column gives the χ^2 per degree of freedom (d.f.).

	c _n (fm)	a _n (fm)	r _n (fm)	Reb ₁	Imb ₁	$\chi^2/d.f.$
Free πN				11.53	9.71	
⁴⁰ Ca Set I	3.31	0.60	3.40	11.32	6.86	3.8
Set II	3.60	0.46	3.27	11.46	6.61	3.3
⁴⁸ Ca	3.4	0.52	3.27	12.36	4.81	11.1
	3.6	0.52	3.39	11.41	4.65	10.4
	3.8	0.52	3.52	10.46	4.54	10.3
	4.0	0.52	3.65	9.63	4.38	10.4
	4.2	0.52	3.78	9.03	4.26	10.6
	4.4	0.52	3.92	8.51	4.08	11.0

total cross sections. This means that these nuclei are not opaque enough for the total cross sections and minima to vary with the same equivalent black-disk radius.

The Set I and Set II parameters found for ⁴⁰Ca

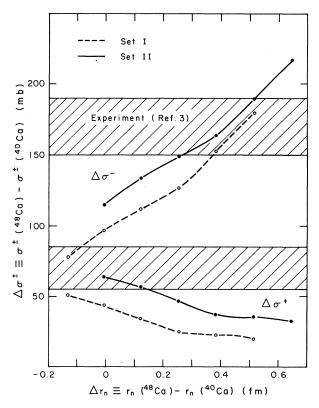


FIG. 3. Total-cross-section differences. The experimental data are taken from a smooth curve through the points at several energies in Ref. 3.

lead to different ^{40}Ca total cross sections; σ^{\star} are both smaller for Set II. These differences are reflected in the values of $\Delta \sigma^{\pm}$ shown in Fig. 3. In ⁴⁸Ca, as c_n increases, and the adjusted value of ${\rm Im} b_1 \; {\rm decreases}, \; \sigma^{\rm +(^{48}Ca)} \; {\rm decreases}.$ This happens because the π^+ interact mainly with the protons, whose radius is fixed, and the medium is becoming more transparent. However, at the same time $\sigma^{-}(^{48}Ca)$ increases; the π^{-} mainly interact with the neutrons, and the effect of increasing the neutron radius apparently more than offsets the change in Imb_1 .

It is clear from Fig. 3 that the variation in $\Delta \sigma^{\star}$ corresponding to the ⁴⁰Ca Set I and II parameters obscure their dependence on Δr_n . Most likely other curves could have been obtained by varying the ⁴⁰Ca neutron distribution. The fact that the Set II curves are both more or less consistent with a common value, $\Delta r_n \simeq 0.2$ fm, while the Set I curves lead to inconsistent Δr_n values might be interpreted as an argument in favor of the former fit. Note that the results in Fig. 3 are insensitive to the shape of the neutron distribution.

To summarize, we have shown that the elastic scattering data alone cannot be used to determine neutron density parameters, since changes in the density can be offset by changes in the opticalmodel parameters. Since a single set of opticalmodel parameters does not describe the scattering from ⁴⁰Ca and ⁴⁸Ca very well, it is also impossible to extract such information from the total-cross-section differences alone. However, the two kinds of information taken together may permit more definitive statements to be made.

We might also note that preliminary data are now available¹⁸ on the elastic scattering from ^{40,48}Ca at 291 MeV. The same kind of problems arise in extracting neutron parameters of this energy as at 130 MeV. It will be interesting to see whether the situation will be different at 180 MeV or so, where the nucleus is blackest.

One point we have not considered in this Letter is the possibility of an improved theoretical description of the pion-nucleus interaction. If sufficient progress is made so that the dependence on

the particular nucleus is understood, and the adjustable parameters are thereby eliminated or at least reduced in number, reliable extraction of a neutron radius may become feasible.

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¹L. S. Kisslinger, R. L. Burman, J. H. Koch, and M. M. Sternheim, Phys. Rev. C 6, 469 (1972).

²M. M. Sternheim and E. H. Auerbach, Phys. Rev. C 4, 1805 (1971). Earlier work is cited here. ³M. J. Jakobson *et al.*, Phys. Rev. Lett. <u>38</u>, 1201

(1977).

⁴J. P. Egger et al., Phys. Rev. Lett. 39, 1608 (1977). ⁵I. Ahmad, Nucl. Phys. <u>A247</u>, 418 (1975); G. D.

Alkhazov et al., Phys. Lett. 57B, 47 (1975).

⁶J. W. Negele, Phys. Rev. C <u>1</u>, 1260 (1970).

⁷G. M. Lerner *et al.*, Phys. Rev. C <u>12</u>, 778 (1975).

⁸G. D. Alkhazov et al., Nucl. Phys. <u>A274</u>, 443 (1976).

⁹S. Shlomo and E. Friedman, Phys. Rev. Lett. <u>39</u>, 1180 (1977).

¹⁰These remarks apply also to a more detailed blackdisk analysis given recently by M. B. Johnson and H. A. Bethe, Comments Nucl. Part. Phys. 8, 75 (1978).

¹¹H. R. Collard, L. R. B. Elton, and R. Hofstadter, in Nuclear Radii, Landolt-Börnstein, Numerical Data and Functional Relationships in Science and Technology. edited by H. Schopper (Springer, Berlin, 1967), Group I. Vol. 2.

¹²E. H. Auerbach and M. M. Sternheim, Brookhaven National Laboratory Report No. BNL-12696, 1968 (unpublished).

¹³L. S. Kisslinger, Phys. Rev. <u>98</u>, 761 (1955).

¹⁴H. K. Lee and H. McManus, Nucl. Phys. <u>A167</u>, 257 (1971); J. H. Koch and M. M. Sternheim, Phys. Rev. C 6, 1118 (1972); G. Fäldt, Phys. Rev. C 5, 400 (1972). ¹⁵M. Ericson and T. E. O. Ericson, Ann. Phys. (N.Y.)

<u>36</u>, 323 (1966). ¹⁶G. A. Miller, Phys. Rev. C <u>10</u>, 1242 (1974); L. S. Kisslinger and F. Tabakin, Phys. Rev. C 9, 188 (1974); Fäldt, Ref. 14.

¹⁷With this approximation, averaging over an equal number of protons and neutrons is equivalent to multiplying b_1 by $\frac{2}{3}$. This factor must be included in comparing the b_1 values in this paper with those in many earlier papers by the present author and others. The assumption of pure isospin $\frac{2}{3}$ is needed to minimize the number of free parameters for $Z \neq N$ or $\rho_n \neq \rho_p$.

¹⁸J. S. McCarthy, private communication.