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Theory of a Free-Electron Laser

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^A classical, linear theory is given for the gain of a short-wavelength free-electron laser. Waves propagating along the relativistic beam are unamplified when the beam is cold. Two of the six geometric optics modes have phase velocities $< c$. When thermal effects are included, the theory predicts the possibility of substantial amplification of these two modes for asymmetric electron distribution functions, due to wave-particle resonance.

Two recent experiments on the Stanford Linear Accelerator Center's superconductive linear accelerator have rekindled wide interest in the possibility of free-electron lasers.^{1,2} A classica luct
nte1
1,2 single-particle radiation analysis for this system was published in 1951 by Motz,³ who later carried out experiments' with 100-MeV electrons. Phillips⁵ posed the problem as an O -type travelingwave interaction, and successfully built and operated microwave devices (Ubitrons) of high efficiency ($\sim 10\%$) and high power output (~ 800 kW). Quantum-mechanical calculations, posing the interaction as stimulated Compton scattering-a nonlinear process—have been published by Manonfinear process—nave been published by Ma-
dey, 6 Sukhatme and Wolf, 7 and by Colson, 8 Hopf dey, Bushaline and word, and by Corson, \overline{H} have reproduced Madey's result using a classical theory based on the relativistic collisionless Boltzmann equation, and have also pubsionless Boltzmann equation, and have also pub-
lished a strong-signal theory.¹⁰ General feature of stimulated processes leading to gain have been of stimulated processes leading to gain have been
summarized by Granatstein and Sprangle.¹¹ Kwan Dawson, and Lin¹² have carried out a linear stability analysis for a cold beam employing peri-

odic boundary conditions and compared the predicted temporal growth rates with numerical simulations.

The linearized theory presented here invokes a mechanism not included in the above, namely wave-particle resonance, but neglects the freestreaming effects which underlie prior work. It indicates that within the geometric-optics approximation, a cold relativistic beam propagating along the axis of a helical static magnetic field does not amplify short-wavelength radiation propagating along the beam. Moreover, it suggests that sizable gain can be obtained using a beam with a finite momentum spread, but that this spread must be asymmetric. When the gain mechanism described here is operative it is likely that the above nonlinear theories must be modified so as to incorporate it.

We take the periodic transverse static magnetic field to be the curl of a vector potential

$$
\vec{A}(z) = (mc^2/e)\vec{V}(z), \qquad (1)
$$

where $\vec{V}(z) = \vec{e}_x V_x(z) + \vec{e}_y V_y(z)$ is the dimension-

less wiggle momentum with $V_x(z)$ and $V_y(z)$ periodic functions of z . This vector potential is valid near the axis of a system of oblate helically wound wires carrying zero net current. The choice V_{τ} $= V \cos k_0 z$ and $V_y = V \sin k_0 z$ corresponds to the experiments^{1,2}; the case $V_v = 0$ can be realized, for example, by an array of permanent magnets. As will be seen, the results do not depend critically upon the relative size of V_x and V_y .

The relativistic Lagrangian for an electron in

the field given by (1), neglecting self-electrostatic effects, yields three constants of the motion:

$$
\alpha = \gamma \dot{x}/c - V_x; \quad \beta = \gamma \dot{y}/c - V_y; \n\gamma = [1 - (\dot{x}^2 + \dot{y}^2 + \dot{z}^2)/c^2]^{-1/2}.
$$
\n(2)

Thus $f=f_0(\alpha, \beta, \gamma)$ is the corresponding general steady-state distribution function, without limit on the wiggle momentum.

The linearized kinetic equation for f_1 , ths small perturbation in the distribution function, is

$$
\partial f_1/\partial t + (c\vec{u}/\gamma) \cdot \nabla f_1 - (e/mc)(\vec{E}_1 + \vec{u} \times \vec{B}_1/\gamma) \cdot \nabla_u f_0 - (e/m\gamma)\vec{u} \times [\nabla \times (m\vec{V}c^2/e)] \cdot \nabla_u f_1 = 0,
$$
\n(3)

where \vec{u} is the electron momentum divided by mc, and \vec{E}_1 and \vec{B}_1 are the perturbed electromagnetic fields. In terms of the independent variables \vec{r} , t , α , β , and $u=(\gamma^2-1)^{1/2}$, (3) becomes

$$
\partial f_1/\partial t + (c\vec{u}/\gamma) \cdot \nabla f_1 = (e/mc)(\vec{E}_1 + \vec{u} \times \vec{B}_1/\gamma) \cdot [(\vec{u}/u)\partial f_0/\partial u + \vec{e}_y \partial f_0/\partial \beta + \vec{e}_x \partial f_0/\partial \alpha].
$$
\n(4)

Since we are interested in interactions with electromagnetic wavelengths much shorter than the periodicity length in $\tilde{V}(z)$, and propagation along the beam, we seek a steady-state geometric-optics¹³ soludicity length in $V(\vec{z})$, and propagation along the beam, we seek a steady-state geometric-optics⁻³ solu-
tion with $\vec{E}_1 = \vec{a} \exp[i(\Psi - \omega t)]$ and $\vec{B}_1 = (c\vec{k} \times \vec{a}/\omega) \exp[i(\Psi - \omega t)]$, where $\vec{k} = k(z)\vec{e}_z$ and $\Psi = \int dz k(z)$ tion with $E_1 = a \exp[i(\Psi - \omega t)]$ and $B_1 = (cK \times a/\omega) \exp[i(\Psi - \omega t)]$, where $K = R(z)e_z$ and $\Psi = jaz R(z)$. In addition one has $f_1 = g(z) \exp[i(\Psi - \omega t)]$, and current density $\vec{J}_1 = \vec{j}(z) \exp[i(\Psi - \omega t)]$, with $\vec{j}(z) = -ec\int d^3u \vec{u}g$. γ . The quantities $\vec{a}(z)$, $k(z)$, $g(z)$, and $\vec{j}(z)$ are all slowly varying in a wavelength $2\pi/k$.

We solve for the conductivity tensor $\vec{\sigma}$ for the electron gas, where $\vec{j} = \vec{\sigma} \cdot \vec{a}$, and find

$$
\overline{\sigma} = -\frac{ie^2}{m\omega} \int d\alpha \, d\beta du f_0(\alpha, \beta, u) \left\{ \frac{\partial}{\partial u} \left[\frac{\overline{u} \overline{u}}{u_z^2} \frac{\omega}{kc - \omega \gamma / u_z} \right] - \frac{\partial}{\partial \beta} \left[\frac{u \overline{u} \overline{e}_y}{\gamma u_z} \right] - \frac{\partial}{\partial \alpha} \left[\frac{u \overline{u} \overline{e}_x}{\gamma u_z} \right] \right\}.
$$
 (5)

First we consider a cold beam; e.g., $f_0(\alpha, \beta, u) = N_0 \delta(\alpha) \delta(\beta) \delta(u-v)$. Then (5) becomes $(\omega_p^2 = 4\pi N_0 e^2/m)$

$$
\overline{\sigma} = \frac{i\omega_{\rho}^{2}}{4\pi\omega} \left\{ \frac{v}{\gamma v_{z}^{3}} \left[u_{z}^{2}(\overline{\dot{I}} - \vec{e}_{z}\vec{e}_{z}) + \vec{V}\vec{V} \right] \right\}
$$
\n
$$
+ \frac{v}{v_{z}} \frac{\vec{V}(\vec{V} + \vec{e}_{z}v_{z}) + (\vec{V} + \vec{e}_{z}v_{z})\vec{V}}{kc/\omega - \gamma/v_{z}} + \frac{v(1 + V^{2})}{\gamma v_{z}^{3}} \frac{(\vec{V} + \vec{e}_{z}v_{z})(\vec{V} + \vec{e}_{z}v_{z})}{(kc/\omega - \gamma/v_{z})^{2}} \right\},
$$
\n(6)

where now $v_z^2 = v^2 - V^2$ and $\gamma = (1 + v^2)^{1/2}$. The dielectric tensor $\vec{\epsilon} = (1 - k^2 c^2/\omega^2)\vec{I} + c^2\vec{k}\vec{k}/\omega^2 + 4\pi i \vec{\sigma}/\omega$ may now be formed. The normal modes of the system obey $\vec{\epsilon} \cdot \vec{a} = 0$. It is convenient to use the right-handed orthonormal basis vectors $\vec{e}_1 = \vec{e}_z \times \vec{V}/V$, $\vec{e}_2 = \vec{V}/V$, and $\vec{e}_3 = \vec{e}_z$. Then

$$
\vec{\epsilon} = \epsilon_{11}\vec{e}_1\vec{e}_1 + \epsilon_{22}\vec{e}_2\vec{e}_2 + \epsilon_{23}(\vec{e}_2\vec{e}_3 + \vec{e}_3\vec{e}_2) + \epsilon_{33}\vec{e}_3\vec{e}_3,
$$
\n(7)

where

$$
\epsilon_{11} = 1 - (kc/\omega)^2 - (\omega_p^2/\omega^2)(v/\gamma v_2),
$$

\n
$$
\epsilon_{22} = 1 - (kc/\omega)^2 - (\omega_p^2/\omega^2)(v/\gamma v_z^3)[v^2 + 2\gamma V^2/v_z \lambda + V^2(1 + V^2)/\lambda^2 v_z^2],
$$

\n
$$
\epsilon_{23} = -(\omega_p^2/\omega^2)(v/\gamma v_z^3)[\gamma V/\lambda + (1 + V^2)V/\lambda^2 v_z],
$$

\n
$$
\epsilon_{33} = 1 - (\omega_p^2/\omega^2)(v/\gamma v_z^3)(1 + V^2)/\lambda^2,
$$
\n(8)

with $\lambda = kc/\omega - \gamma/v_z$. From det $\overline{\epsilon} = 0$ we see that the roots of the six geometric optics modes are given by

$$
\epsilon_{11} = 0 \text{ or } kc/\omega = \pm |1 - (\omega_p^2/\omega^2)(v/\gamma v_z)|^{1/2}
$$
\n(9)

(waves with $\omega/k > c$), and by $D = \epsilon_{22} \epsilon_{33} - \epsilon_{23}^2 = 0$, or, after some algebra,

$$
0 = [1 - (\omega_p^2/\omega^2)(v/\gamma v_z) - (k^2 c^2/\omega^2)][(kc/\omega - \gamma/v_z)^2 - (\omega_p^2/\omega^2)v/\gamma v_z^3].
$$
\n(10)

One pair of roots is identical to (9) . The other pair is

One part of 10000 to momentum to
$$
(0)
$$
. The other pair

$$
k\,c/\omega\,=\gamma/\,v_{\rm g}\,\pm\,(\omega_{\rm p}\,/\,\omega)(\upsilon/\gamma\,v_{\rm g}^{\phantom{\mathrm{u}}3})^{1/2}
$$

which are slow waves $(\omega/k < c)$. The forms of (9) and (11) are reminiscent of corresponding forms for conventional transverse electromagnetic and longitudinal space-charge waves on a cold plasma stream, but it must be noted that here the modes are neither purely transverse $(\vec{k} \cdot \vec{a} = 0)$ nor purely longitudinal $(k \times a = 0)$. In any case, it is apparent that (9) and (11) lead to purely real roots: The cold beam manifestly does not amplify.

We now consider the effect of a small thermal spread in the beam momentum. The most important change will be the addition of a small real component to (6) , on account of wave-particle coupling. The two modes given by (11) are thus the only candidates for growth. The possibility of amplification is governed by the steady-state transport equation for the wave energy U ,¹³ governed by the steady-state transport equation for the wave energy $U₁₃$

$$
\nabla \cdot (\omega_{\vec{\tau}} U) + \vec{\sigma}^{\text{H}}; \quad \vec{\dot{a}} \vec{a}^* = 0, \tag{12}
$$

where $U = (\omega/8\pi)(\partial \vec{\xi}/\partial \omega)$: $\vec{a}\vec{a}$, and $\vec{\sigma}^H$ is the Hermitian part of $\vec{\sigma}$. After some manipulation, one can show that when $v^2 \gg V^2 \approx 1$,

$$
U = \mp \frac{1}{4\pi} \frac{\omega}{\omega_p} \frac{1 + V^2}{V^2} \frac{a^2}{v^{1/2}},
$$
\n(13)

where the upper sign (negative energy) goes with the analog of the slow space-charge wave given by the upper sign in (ll), and where the lower sign (positive energy) goes with the analog of the fast spacecharge wave given by the lower sign in (11). The group velocity can be shown to be $\frac{\partial \omega}{\partial k} = c v_z / \gamma \approx c$, namely the z velocity of the beam. The lowest-order thermal contribution to $\overline{\sigma}^H$ is, from (5),

$$
\overline{\sigma}^{\mathrm{H}} = -\frac{\pi e^2}{m \omega} \int d\alpha \, d\beta \, du \, \frac{\partial f_0}{\partial u} \frac{\overline{\mathrm{uu}}}{u_z^2} \delta \left(\frac{kc}{\omega} - \frac{\gamma}{u_z} \right). \tag{14}
$$

It can be shown that $\overline{\tilde{u}\tilde{u}}$: $\overline{\tilde{a}\tilde{a}}^* \simeq a^2/V^2$. Thus if we define $N_0F(u) = \int d\alpha \, d\beta f_0(\alpha, \beta, u)$, we have from (12)

$$
\frac{\partial}{\partial z} \left[\mp \frac{c}{4\pi} \frac{\omega}{\omega_p} \frac{1+V^2}{V^2} \frac{a^2}{v^{1/2}} \right] - \frac{\omega_p^2}{4\omega} \frac{a^2 v}{(1+V^2)V^2} F'(w_\pm) = 0, \tag{15}
$$

where $w_{\pm} \simeq v^{\pm}(\omega_{p}/\omega)v^{3/2}(1+V^{2})^{-1}$ and the upper and lower signs correspond to those in (11). For the case $\vec{V} = V(\vec{e}_x \cos k_y z + \vec{e}_y \sin k_y z)$, Eq. (15) yields

$$
\frac{\partial}{\partial z} (\ln a^2) = \mp \pi \frac{\omega_p}{c} \left(\frac{\omega_p}{\omega}\right)^2 \frac{v^{3/2}}{(1 + V^2)^2} F'(w_\pm) \equiv G,
$$
\n(16)

and growth occurs if $F'(w_+) < 0$ for the slow spacecharge wave, and if $F'(w_-) > 0$ for the fast spacecharge wave.

The consequences of (16) can be readily understood with the aid of Fig. 1. Clearly the symmetric distribution $1(a)$ leads to spatial decay since $F'(w_+) > 0$ and $F'(w_-) < 0$. But the asymmetric distribution 1(b), which is qualitatively like that of many relativistic beams, can yield spatial amplification for the positive-energy wave for $v < w < u_0$. If we approximate $F'(w_+) \approx 1/\Delta^2$, where Δ is an effective width, then with $\omega = 10^{14}$, $\omega_{p} = 3 \times 10^{9}$, $v \approx \gamma = 50$, $V^{2} = 0.5$, and $\Delta = 10^{-2}$, (16) $\omega_p = 3 \times 10^{\circ}$, $v = \gamma = 30$, $v = 0.3$, and $\Delta = 10^{-6}$, (10)
yields $G = 10^{-4}$ cm⁻¹ = 1%/m.¹ A distribution function skewed in the opposite sense would amplify

! the negative-energy wave.

In conclusion, the gain of a free-electron laser described by the theory presented here depends critically on the axial momentum distribution of the beam. Asymmetry is required for gain, a low-energy tail leading to amplification of the positive-energy wave. Suitable tailoring of the distribution function could result in gains higher than those that have been observed. Clearly the results are very sensitive to the shape of $F(u)$ suggesting that appropriate experiments would be most valuable in establishing the predictions of this work.

We have benefitted from discussions with

 (11)

FIG. l. Axial momentum distributions. In (a) ^a symmetric distribution is shown which is stable against spatial amplification for both the fast and slow spacecharge waves, since $F'(w_+) > 0$ and $F'(w_-) < 0$. In (b) an . asymmetric distribution similar to that furnished by a linac is shown which can support amplification for the fast wave, when w_z is in the shaded region where $F'(w_z)$ $\geq 0.$

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Trapping of Cusp-Injected, Nonneutral, Electron Rings with Resistive Walls and Static Mirror Coils

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A hollow, rotating 2.3-MeV electron beam, moving with axial velocity of $\sim 0.25c$ after injection through a magnetic-field cusp, is slowed down and trapped in a small magnetic mirror well by interaction with a $30-60-\Omega$ /square resistive wall. The trapped-electron ring containing $\sim 10^{12}$ particles performs a damped oscillation about the well minimum and settles down into an equilibrium state.

The electron-ring accelerator (ERA) concept first proposed by Veksler et al , 1 uses the electrostatic potential well of a high-density, relativistic electron ring for collective acceleration of positive ions. Experiments in progress at several laboratories^{2–4} employ a pulsed magneti mirror field with azimuthal beam injection to obtain a compressed electron ring. In the ERA experiment at the University of Maryland, 5^{-6} on the other hand, the ring is being formed with the aid of a static, cusped magnetic field. The cusp

transforms a long hollow electron beam into a short rotating cylindrical electron layer $(E \text{ layer})$ which propagates in the axial direction at a fraction of the speed of light. The length and velocity of the E layer depend on the pulse shape of the injected electron beam and the axial magnetic field, B_e As B is increased, the axial velocity is reduced until a cutoff value, B_c , is reached where all particles are reflected in the cusp region. Experimentally, it is found that reproducible beams are obtained only when B is some-