

Is There a High- γ Anomaly in Antineutrino Interactions?

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We have analyzed data taken in the CERN narrow-band neutrino and antineutrino beams with regard to the "high- γ anomaly" observed by previous experiments at Fermilab. At neutrino energies between 30 and 200 GeV, the $\bar{\nu}$ and ν charged-current cross-section ratios and muon-inelasticity distributions disagree with the earlier results. In particular, there is no evidence for energy-dependent effects in the antineutrino data which constitute an important aspect of the alleged anomaly.

In recent years, low- and high-energy inelastic interactions of neutrinos with nuclei were believed to be consistent with $V-A$ coupling and Bjorken scale invariance in the framework of the spin- $\frac{1}{2}$ parton model¹⁻⁴. In 1974 evidence was reported for a significant departure from expectation of the antineutrino-muon inelasticity distributions at higher energy⁵ which was subsequently dubbed the "high- γ anomaly" after the Bjorken scaling variable $y = (E_\nu - E_\mu)/E_\nu$. Subsequent experimental results supported an anomalous behavior of $\bar{\nu}$ total and differential cross sections.⁶⁻⁹ This was considered as evidence for a deviation from charge symmetry, together with new particle production,⁵ the existence of right-handed currents, and a breaking of scale invariance.⁶ The experimental observations were the following¹⁰: An excess of events in the high- γ region of the antineutrino y distribution was observed for $x < 0.1$ ⁵; a strong energy dependence of the average value of y was found in the region $E_{\bar{\nu}} = 30-70$

GeV⁶; the ratio of antineutrino to neutrino charged-current cross sections was found to increase with energy above about 30 GeV⁷; the antiquark component in the antineutrino y distribution was observed to increase with energy.^{6,8,9}

We report here first results from an experiment which is currently running at the CERN 400-GeV proton synchrotron. The present analysis is addressed to the specific question of a high- γ anomaly in antineutrino interactions, and follows, therefore, in part the guidelines set out by the advocates of this anomaly. The experimental conditions, on the other hand, are quite different.

In this experiment, neutrinos and antineutrinos of known energy between 10 and 200 GeV are produced by selecting 200-GeV parents, pions and kaons of the appropriate charge, in 400-GeV proton-beryllium interactions. The parent beam with 0.2-mrad divergence and $\pm 5\%$ momentum spread is directed towards the detector, passing through a 300-m decay tunnel, and about 350-m

muon shielding. The average distance from the parent decay point to the detector is 620 m. An estimate of the neutrino energy can be obtained event by event from the decay kinematics and observed radial distance of a neutrino interaction from the central beam line. The uncertainty in the neutrino energy determined by this means is typically $\pm 20\%$, mostly due to the unknown decay position along the beam line. Although this measurement of the neutrino energy is not used here in particular, it is useful for a systematic study of the calibration of the detector.

The beam is instrumented by various devices in order to control the alignment, and to measure the beam composition and intensity. Apart from the positron detectors, there are, for flux monitoring, a differential Čerenkov counter separating pions, kaons, and protons; a beam current transformer to measure the parent flux; a total absorption calorimeter used as a beam stopper to integrate the total beam energy; and a set of detectors in the steel shield to map the ionization generated by the decay muons. The neutrino energy spectrum and the ratio of neutrino and antineutrino fluxes are accurately known at least for the purpose of the present investigation.

The basic feature of the detector¹¹ is the inte-

gration of target calorimeter and muon spectrometer in order to obtain a large acceptance. It is subdivided into nineteen magnetized iron modules, 3.75 m in diameter, with triple-gap drift chambers in between modules for muon-track measurement. Each magnet module weighs 65 tons. The average field is 1.65 T giving typically $\pm 10\%$ resolution on the muon momentum. Plastic scintillator is sandwiched with either 5- or 15-cm-thick iron plates. There are seven 15×5 cm plate modules and twelve 5×15 cm plate modules. The hadron energy resolution is measured to be $0.8(E^*)^{1/2}$ GeV^{1/2} in the 5-cm sampling section of the detector, $\sqrt{3}$ times larger in the 15-cm sample section. The energy calibration for hadronic showers generated by neutrino interactions has been checked with the help of the known neutrino energy as described above.

In the analysis, a 740-ton fiducial volume is defined by cuts in radius and depth, and hits in five consecutive chambers are required on the muon track. The angular region with good acceptance extends up to 400 mrad typically. This means that the geometric acceptance in y is near unity for all but the upper end of the y distribution, e.g., $y < 0.9$ for $E_\nu = 100$ GeV.

After fiducial and muon-track-length cuts there

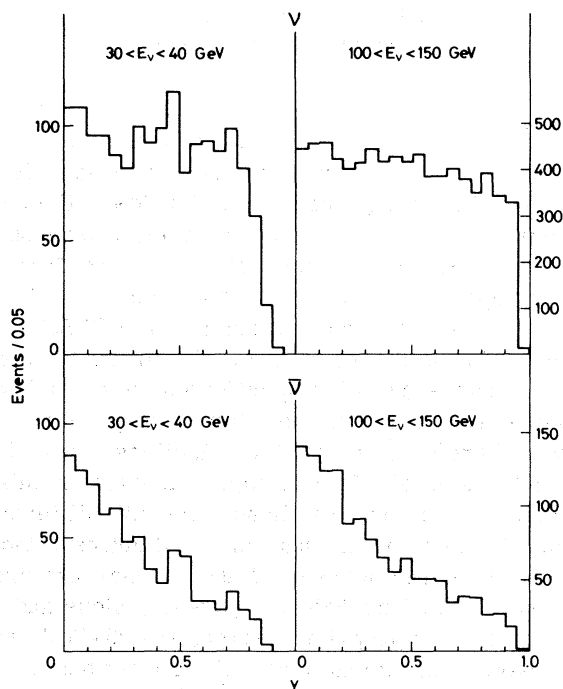


FIG. 1. The y distributions of neutrino and antineutrino events with $x < 0.6$ in the two energy bands $E_\nu = 30-40$ GeV and $E_\nu = 100-150$ GeV.

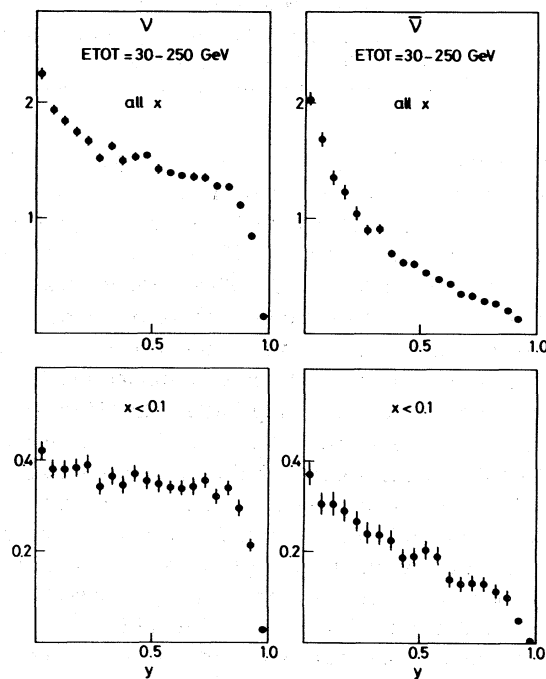


FIG. 2. Normalized y distributions for neutrino and antineutrino data, with and without a cut $x < 0.1$. The distributions are normalized to the same ν and $\bar{\nu}$ fluxes subject to a relative scale error of $\pm 10\%$.

remain 36 000 neutrino and 12 000 antineutrino events. For every event, the neutrino energy is computed from the sum of muon and hadron shower energy, $E_\nu = E_\mu + E_h$. The useful spectrum extends from about 20 to 200 GeV. About half the neutrino data and a quarter of the antineutrino data are at energies above 100 GeV. Further cuts are imposed on the data as required for comparison with previous results.

The y distributions of neutrino and antineutrino events are displayed in Fig. 1 for two different energy bins; in Fig. 2 for all energies $E_\nu > 30$ GeV, with and without a cut at $x < 0.1$. In Fig. 2 the neutrino and antineutrino distributions are normalized to the same neutrino flux in order to allow a test of charge symmetry in the limit $y \rightarrow 0$. We note the following: The y distributions at different energies are very similar. Charge symmetry is valid within our present uncertainties of the neutrino to antineutrino flux ratio ($\pm 10\%$), as is expected for an isospin-0 target when strangeness- and charm-changing currents are neglected. The y distribution of neutrino events

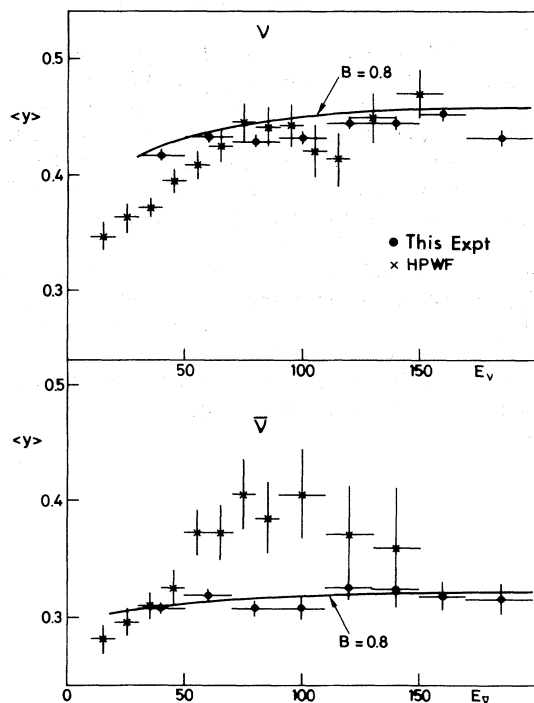


FIG. 3. The first moments of the y distributions as a function of energy, for neutrino and antineutrino data with $x < 0.6$. The curves marked $B = 0.8$ are calculated for the conditions of this experiment assuming a fixed antiquark component. The Harvard University–University of Pennsylvania–University of Wisconsin–Fermilab (HPWF) data are taken from Ref. 6.

is approximately flat up to $y \approx 0.9$, where the acceptance losses set in. The antineutrino y distribution falls rapidly with increasing y as expected for $do/dxdy \sim q(x)(1-y)^2 + \bar{q}(x)$ in the quark-parton model with a small amount of sea quarks $\int \bar{q}(x)dx \ll \int q(x)dx$. Even for $x < 0.1$, the y distributions for neutrinos and antineutrinos are very different, and are compatible with the picture that valence quarks dominate, but that the sea quarks concentrate at small x . These observations are in conflict with Ref. 5, where a flat y distribution was observed for both neutrinos and antineutrinos in the region $x < 0.1$.

The average values of y for accepted events are displayed in Fig. 3 as a function of neutrino energy. Here a cut $x < 0.6$ has been applied for a direct comparison with Ref. 6. Since the average y values are biased by the drop in acceptance at large y , we have indicated the expected dependence for a constant amount of sea quarks for the conditions of this experiment. The antineutrino data are compatible with an energy-independent fraction of sea quarks $B = \int [q(x) - \bar{q}(x)]dx / \int [q(y) + \bar{q}(x)]dx = 0.8$, in contrast to the conclusions obtained in Ref. 6.

From the antineutrino y distributions, we have calculated the B values as a function of neutrino energy, and compared them with those of Refs. 6, 8, and 9 (Fig. 4). The new data do not show any significant energy dependence of quark-sea component, contrary to what had been inferred^{6,8} from the earlier data.

The ratio of antineutrino and neutrino charged-current cross sections was computed from the observed number of events, and a Monte Carlo efficiency calculation. The Monte Carlo accounts for the small difference in detection efficiency

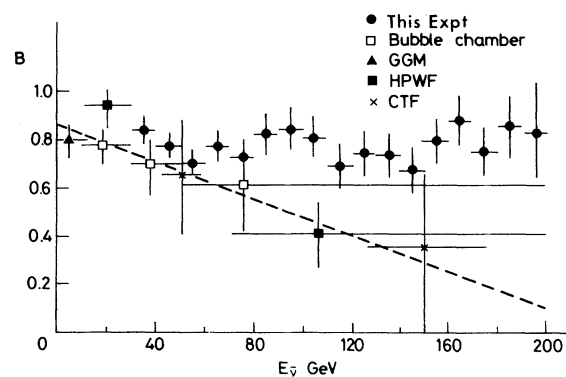


FIG. 4. The B values as a function of antineutrino energy, from various experiments [Refs. 4 ("GGM"), 6 ("HPWF"), 8 ("CTF"), and 9 ("Bubble Chamber")].

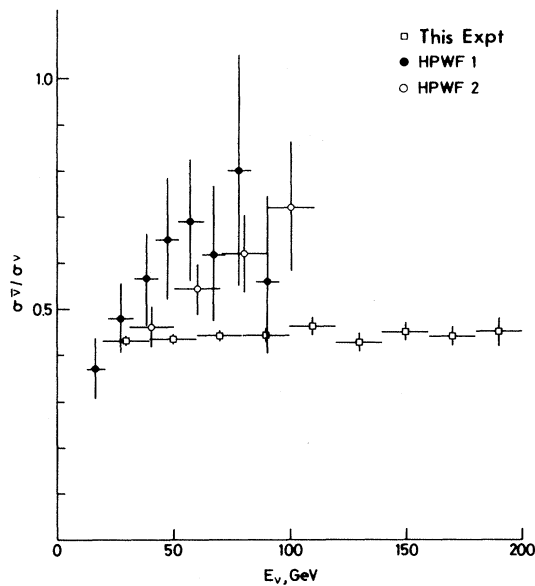


FIG. 5. The ratio between the antineutrino and neutrino charged-current total cross sections as a function of neutrino energy. The data marked HPWF 1/2 are taken from Ref. 7. The indicated errors are statistical only. Systematic errors are discussed in the text.

due to the different y distributions and the minimum muon momentum required in the analysis. The cross-section ratio is displayed in Fig. 5. We notice that the energy dependence of the antineutrino cross section follows the neutrino cross section quite well, in contradiction to the observation of Ref. 7. The present data are subject to two uncertainties: The relative $\bar{\nu}/\nu$ flux normalization has, for the moment, an uncertainty leading to a $\pm 10\%$ scale error common to all energies. In addition, below 90 GeV the events are mostly due to neutrinos from pion decay, and

above 90 GeV mostly kaon decay. The ratio of the K/π ratios in the negative and positive beams, as determined using the differential Čerenkov counter, may be in error by not more than $\pm 10\%$. This could lead to a corresponding discontinuity in the cross-section ratio around 90 GeV, which, however, is not observed.

The present data allow the following conclusion relevant to the "high- y anomaly": The experimental observations on which the evidence was based are not confirmed. In particular, the antineutrino y distribution and the ratio between the antineutrino and neutrino cross sections are found to be essentially independent of neutrino energy.

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¹⁰The scaling variables are defined as $x = Q^2/2M_\nu$ and $y = \nu/E_\nu$, where Q^2 is the square of the four-momentum transfer from the neutrino to the muon, and ν is the energy transfer to the hadrons.

¹¹M. Holder *et al.*, "A detector for high-energy neutrino interactions" (to be published).