Photoproton Cross Section for ^{18}O as a Measure of the Effect of the Ualence Neutrons on the '60 Core*

B. L. Herman, D. D. Faul, R. A. Alvarez, and P. Meyer Lawrence Livermore Laboratory, University of California, Livermore, California 94550 (Received 12 April 1976)

The ¹⁸O(γ , p)¹⁷N cross section has been measured with monoenergetic photons from threshold to 80 MeV. The results show striking similarities and differences with the $^{16}O(\gamma, p)^{15}N$ cross section.

We have measured an important photonuclear cross section on an unusual nucleus by a novel technique with surprising results. The 18 O nucleus, lying in the periodic table at the beginning of the $s-d$ shell, is unusual in that although it consists to first order simply of a pair of valence neutrons bound to the doubly magic 16 O core, it is reported to have a large value for $B(E2, 0-2)$, equal to 0.0048 b^2 .¹ If this nucleus were purely rotational, this value for $B(E2)$ would imply a large average ground-state deformation $\bar{\beta}_2$, equal to 0.37. Of course, this is not actually the case,¹ but possibly the ¹⁸O nucleus is described best by the "coexistence" of spherical and very deformed $(\beta_2 \ge 0.6)$ configurations. This makes it very interesting to measure the photoproton cross section for this nucleus, which, since the $^{16}O(\gamma, p)$ cross section has been studied extensively, should result in a direct measure of the effect of the presence of the valence neutrons on the deformation (or polarization) of the 16 O core. Furthermore, the 18 O nucleus is *unique* among stable nuclei in that its (γ, p) threshold (15.94 MeV) lies at a higher energy than its $(\gamma, 2n)$ threshold² (12.19) MeV); this means that the residual nucleus from the $^{18}O(\gamma, p)$ process, namely ^{17}N , can decay by β ⁻ emission to neutron-unstable states in ¹⁷O. Therefore, the ${}^{10}O(\gamma, p)$ cross section can be measured by observing the resulting delayed neutrons. In fact, nearly all $(95\%)^{3}$ ¹⁷N nuclei decay to such neutron-unstable states with a lifetime⁴ $(\tau_{1b} = 4.16 \text{ sec})$ which is ideal for the purposes of the measurement we have done. The results we obtained are surprising because of their striking similarities and differences with the $^{16}O(\gamma, p)$ cross section, as will be demonstrated below.

The few previous photonuclear measurements on 18 O all have been performed with continuous bremsstrahlung radiation sources. Prompt photoneutron measurements⁵ have heretofore been limited to the energy region below the giant resonance, delayed photoneutron measurements⁶ were not sufficiently detailed for any conclusions to be

drawn regarding structure in the giant resonance, and prompt photoproton measurements, ' done only at 90', suffer from high detection thresholds and the inability to distinguish between the groundand excited-state cross section. Also the only electron-scattering measurement⁸ was done at 180° , which selectively excites only magnetic transitions.

The source of radiation for our experiment was the monoenergetic photon beam produced by the annihilation in flight of fast positrons from the Lawrence Livermore Laboratory electron-positron linear accelerator. The photon energy resolution was approximately 1% . The technique we used to measure the $^{18}O(\gamma, p)^{17}N$ cross section was to count the delayed neutrons from $17N$ decay between beam bursts from the linac. This was made feasible by three conditions: (1) The 4.16 sec half-life of $17N$ is short compared with the length of a run necessary to acquire sufficient statistics at a given incident photon energy (10- 20 min), so that the activity was driven into a steady-state saturation condition quickly; (2) tbe balf-life is long compared with the time interval between beam bursts $(1/720 \text{ sec})$, so that there was no need to correct for the $17N$ decay lifetime in the time during which the delayed neutrons were detected; and (3) the mean decay time of the $\texttt{BF}_{\textup{3}}\text{-}$ tube –plus –paraffin 4π neutron detector 9,1 was sufficiently short¹⁰ (~90 μ sec) so that the number of prompt photoneutrons detected during the delayed gate (delayed with respect to the linac beam burst by 675 μ sec and 700 μ sec wide) was negligible. The monitoring and calibration of the photon beam, the subtraction of the neutron yield resulting from the positron bremsstrahlung, and the details of the neutron detector have been pubthe details of the neutron detector have been pub
lished elsewhere.⁹⁻¹¹ The sample used consiste of 120 g of H₂O enriched to 96.5 at. $\%$ ¹⁸O. It is worth noting that neutron backgrounds were negligible, consisting entirely of machine-off (cosmic-ray) background neutrons, since no material (other than 18 O) in or near the photon beam could

be a source of delayed neutrons. Sample-blank and annihilation-target-out runs confirmed this fact.

The $^{18}O(\gamma, p)^{17}N$ -cross-section results of this experiment are shown in Fig. 1, together with the cross sections for ${}^{16}O(\gamma,p)^{15}N$ and for ${}^{16}O(\gamma,n)^{15}O$
from the literature.¹² One immediately is strucl from the literature.¹² One immediately is strucl by the similarity (except for scale) of the cross sections below about 21 MeV (the appearance of the peaks at 17.3, 19.2, and 20.8 MeV) as well as by the dramatic difference between the appearance of the two broad and prominent structures at 23.4 and 27.5 MeV in the $^{18}O(\gamma,p)$ cross section and that of the four narrower and more tightly clustered peaks at 22.3, 23.1, 24.2, and 25.2 MeV in the $^{16}O(\gamma,b)$ cross section. (It should be noted that the 23.4-MeV peak has a shoulder on

FIG. 1. (a) The $^{18}O(\gamma, p)^{17}N$ cross section measured in the present experiment compared with (b) the $^{16}{\rm O}(\gamma\,,$ p ¹⁵N and (c) the ¹⁶O(γ , n¹⁵O cross sections taken from the literature (see text). Threshold energies (from Ref. 2) are indicated by the arrows. Solid lines are used to represent the 16 O cross sections because they have been synthesized from several experimental results; in any case, the relative precision of these results is good enough so that no important structure has been created or obscured by this procedure.

the low-energy side, at 22.4 MeV, and the 27.5- MeV structure perhaps is composed of two peaks, at about 26.6 and 28 MeV.)

The measured integrated cross section for the reaction $^{18}O(\gamma, p)$ is 29.8 MeV mb up to 30.6 MeV, compared with 81.2 MeV mb up to 29.0 MeV and 41.5 MeV mb up to 28.0 MeV for the reactions $^{16}O(\gamma, p)$ and $^{16}O(\gamma, n)$, respectively. It seems clear that the explanation of the difference in magnitude of a factor of 3 between the (γ, p) cross sections for 18 O and 16 O lies largely in the much greater number of competing reaction channels available in the ¹⁸O case, chiefly the (γ,n) , $(\gamma, 2n)$, and (γ, pn) channels (feeding states in ^{17}O , ^{16}O , and 16 N), whose threshold energies are 8.05, 12.19, and 21.83 MeV, respectively.² This is made even more plausible by the fact that the $^{16}O(\gamma, p)$ cross section (whose threshold energy is 12.13 MeV) is so much larger than the ${}^{16}O(\gamma,n)$ cross section (threshold at 15.67 MeV) [Figs 1(b) and 1(c)], while the $^{16}O(\gamma, p_0)$ and $^{16}O(\gamma, n_0)$ cross sections, each consisting of a single channel, are sections, each consisting of a single channel, are
comparable in magnitude.¹³ It should be noted as well that there also is a difference in the relative sizes of the low-energy peaks; the 19.2-MeV structure is more prominent for 18 O than for 16 O.

Likewise, the expected large neutron widths for the giant-resonance states in 18 O will broaden them, and the presence of the valence neutrons might cause them to shift in energy to some extent. But although we cannot rule out this possibility in the absence of a detailed calculation, it seems unlikely that at the same time that the states below 21.5 MeV are not broadened very much and hardly shifted at all, the states above this energy should be broadened and shifted to a large enough extent to explain the observed difference in shape between the two (γ, p) cross sections.

One might expect naively that the $27.5-MeV$ hump in the $^{18}O(\gamma\rho)$ cross section results from the deformation splitting of the giant resonance, but if one uses the hydrodynamic theory to compute the ground-state deformation from the 23.4to 27.5-MeV splitting, one obtains $\bar{\beta}_2 = 0.20$, which is much less than the value $\bar{\beta}_2$ = 0.37 mentioned above, and appears to yield no deep insight into the physical process involved. In any case, one should not use the (γ, p) cross section alone for this determination, but rather the total photonabsorption cross section.

If one is tempted, on the other hand, to ascribe this splitting of the giant resonance to isopin effects, one obtains equally unsatisfactory results.

First, one expects that nearly all the $^{18}O(\gamma, p)$ cross section represents T , $(T=2)$ strength because it lies in the same energy region as the 16 O cross sections (which must be T_z because of the $\Delta T = +1$ selection rule for dipole transitions in self-conjugate nuclei). Second, one expects that nearly all the T_{\leq} (T=1) strength is manifested in the photoneutron $[(\gamma,n)$ and $(\gamma,2n)]$ channels. Finally, the energy splitting should be given by something close to the empirical relation¹⁴

 $\Delta E = E(T) - E(T) = 55(T_0 + 1)/A$

which yields a value greater than 6 MeV for ^{18}O .

One can speculate that the $^{18}O(\gamma,p)$ cross section should look like some sort of superposition of the (γ, p) cross sections for ¹⁶O and ²⁰Ne, since the latter nucleus bears some resemblance to the deformed part of the ¹⁸O ground state in the coex-
istence picture.¹⁵ One gets no support for this istence picture.¹⁵ One gets no support for this point of view from existing measurements of the (γ, p) ¹⁶ and (γ, n) ¹⁷ cross sections for ²⁰Ne, however, which show no prominent structure in the 23- to 30-MeV energy region; but then again one expects that the large concentration of dipole strength characteristic of deformation splitting should lie above 30 MeV for 20 Ne, in an energy region where no measurements as yet have been carried out.

We are left with a substantial difference between the shapes of the (γ, p) cross sections for 18 O and 16 O which does not appear to result from any of the above considerations. Clearly, this difference must be a measure of the effect of the valence neutrons on the ¹⁶O core. One even can speculate that a large core-polarization effect in 18 O might result from an unexpectedly large ten- 18 O might result from an unexpectedly large te
sor component of the nuclear force.¹⁸ This, if true, would have important ramifications for many aspects of nuclear physics, and, in particular, would make necessary a fundamental reexamination of may microscopic nuclear calculations.

We have profited from conversations with Dr. J. E. E. Baglin, Professor G. E. Brown, Dr. J. T. Caldwell, Professor W. Greiner, Professor A. K. Kerman, Dr. T. W. Phillips, and Professor J. P. Vary.

Contract No. W-7405-Eng-48. A preliminary account of this work appeared in D. D. Faul et $al.$, Bull. Am. Phys. Soc. 21, 68 (1976).

¹A. M. Kleinfeld, K. P. Lieb, D. Werdecker, and U. Smilansky, Phys. Rev. Lett. 85, 1829 (1975), aud references therein.

 2 A. H. Wapstra and N. B. Gove, Nucl. Data Tables 9, $303(1971).$

 3 M. G. Silbert and J. C. Hopkins, Phys. Rev. 134, B16 (1964); J. Gilat, G. D. O'Kelley, and E. Eichler, Oak Ridge National Laboratory Report No. ORNL-3488,

1963 (unpublished), p. 4.

⁴I. Dostrovsky, R. Davis, Jr., A. M. Poskanzer, and P. L. Reeder, Phys. Rev. 139, B1513 (1965).

⁵H. Fuchs, Z. Phys. 171, 416 (1963); S. F. Mughabghab and W. E. Stephens, Phys. Rev. 133, B660 (1964); J. D. Allan, J.W. Jury, R. G. Johnson, K. G. McNeill, J. G. Woodworth, and Y. S. Horowitz, Can. J. Phys. 53, 786 (1975).

 6 W. E. Stephens, J. Halpern, and R. Sher, Phys. Rev. 82, 511 (1951); R. Montalbetti and L. Katz, Can. J. Phys. 81, ⁷⁹⁸ (1953)..

⁷R. Kosiek, K. Schlüpmann, H. W. Siebert, and R. Wendling, Z. Phys. 179, 9 (1964); W. Hofmann, R. Kosiek, G. Kraft, and R. Mundhenke, Z. Phys. 225, 808 (1969).

 ${}^{8}G.$ J. Vanpraet, Phys. Lett. 17, 120 (1965).

⁹S. C. Fultz, R. A. Alvarez, B. L. Berman, and P. Meyer, Phys. Rev. ^C 10, 608 (1974).

¹⁰B. L. Berman, J. T. Caldwell, R. R. Harvey, M. A. Kelly, R. L. Bramblett, and S.C. Fultz, Phys. Rev. $162, 1098$ (1967).

 $\overline{^{11}}$ B. L. Berman and S. C. Fultz, Rev. Mod. Phys. $\frac{47}{1}$, $713(1975).$

¹²Values for the ¹⁶O(γ , p_0) cross section have been taken from R. C. Morrison, thesis, Yale University, 1965 (unpublished); and from Z. E. E. Baglin and M. N. Thompson, Nucl. Phys. A138, 73 (1969), normalized to the mean of several other results in the literature; for $^{16}O(\gamma, p_{i>0})$, from J. T. Caldwell, thesis, Lawrence Livermore Laboratory Report No. UCRL-50287, 1967 (unpublished); and for $^{16}O(\gamma, n)$, from R. L. Bramblett, J. T. Caldwell, R. R. Harvey, and S. C. Fultz, Phys. Rev. 133, ^B ⁸⁶⁹ (1964), and J. T. Caldwell, R. L. Bramblett, B. L. Berman, R. R. Harvey, and S. C, Fultz, Phys. Rev. Lett. 15, 976 (1965).

 13 C.-P. Wu, F. W. K. Firk, and T. W. Phillips, Phys. Rev. Lett. 20, 1182 (1968).

 $14P$. Paul, in Proceedings of the International Conference on Photonuclear Reactions and Applications, edited by B. L. Herman (Lawrence Livermore Laboratory, Livermore, Calif., 1973), p. 407.

 15 G. E. Brown, in International Congress on Nuclear Physics, Paris, 1964, Proceedings, edited by P. Gugenberger {Centre National de la Recherche Scientifique, Paris, 1964), p. 129, and private communication.

 16 W. R. Dodge and W. C. Barber, Phys. Rev. 127 , 1746 $(1962); R. E. Segel, Z. Vager, L. Meyer-Schützmeister,$ P. P. Singh, and R. G. Allas, Nucl. Phys. A93, 31 (1967); W. Hofmann et al., $op. cit.$ (Ref. 7).

^{*}Work performed under the auspices of the U. S, Energy Research and Development Administration under

 17 A. Veyssière, H. Beil, R. Bergère, P. Carlos, A. Leprêtre, and A. de Miniac, Nucl. Phys. A227, 513 (1974).

 18 J. P. Vary, P. U. Sauer, and C. W. Wong, Phys. Hev. ^C 7, ¹⁷⁷⁶ (1978); J. P. Vary, private communication.

Evidence for New Minima in Photoionization Cross Section Obtained by Spin-Polarization Measurements

U. Heinzmann, H. Heuer, and J. Kessler Physikalisches Institut der Universität Münster, 44 Münster, Germany (Heceived 23 February 1976)

It has recently been predicted that Cooper-type minima of the photoionization cross section should also exist for $l \rightarrow l-1$ transitions. This theoretical conclusion is supported by measurements of the spin polarization of photoelectrons that have been ejected from thallium atoms by circularly polarized light.

Recently Msezane and Manson' predicted the existence of a new kind of minimum in the wavelength dependence of the photoionization cross section. Studying the example of the excited Cs 5d photoionization they found, apart from the _{ou} photoionization they iound, apart from the
well-known Cooper minimum,² a second minimum for the $l\rightarrow l+1$ photoionization channel and a minimum for the $l + l - 1$ channel. These minima arise when the matrix elements of the photoionization process vanish as a result of positive and negative contributions to their radial parts caused by details of the overlap between the discrete and continuum wave functions.

A zero of the matrix element for the $l - l - 1$ channel is a novel feature which has not been predicted before and which was thought not to exist. It is the purpose of the present Letter to report experimental evidence of such a zero which causes a, zero minimum of the partial cross section for the $l + l - 1$ channel.

The measurements have been made with thallium atoms in their ground state $6s^2p(^2P_{1/2})$ which were photoionized by circularly polarized light. A measurement of the wavelength dependence of the photoionization cross section would have hardly revealed the new minimum, since the effect to be studied is masked by the influence of the $l + l + 1$ channel. Consequently, instead of measuring the photoionization cross section the polarization of the photoelectrons produced by the circularly polarized light has been observed.

The relation between the polarization P and the photoionization cross section has been discussed in an earlier paper' which also gives a brief account of the experimental procedure (a detailed description of the apparatus will be given elsewhere⁴). The following facts have been

shown there: According to the selection rules $l + l + 1$, the outer p electron of a thallium atom can make a transition into the S or the D continuum. If circularly polarized light is used for photoionization, the photoelectrons in the S continuum have a polarization $P = 1$, whereas their polarization in the D continuum is $P = -0.5$. The resulting polarization of all the photoelectrons produced is therefore

$$
P = (1 \times Q_S - 0.5 \times Q_D) / (Q_S + Q_D), \qquad (1)
$$

where the polarizations of the two final states have been superimposed after weighting them with the cross sections Q_S and Q_D for reaching the S and D continuum, respectively.

In a conventional photoionization experiment one measures the cross section

$$
Q = Q_S + Q_D , \qquad (2)
$$

since one cannot distinguish between transitions into different continua. A measurement of the polarization yields, however, information on the individual channels as one can immediately see from Eq. (1). If, for example, Q_S or Q_D dominates, the polarization tends to $+1$ or -0.5 , respectively. If both P and Q are known, one has from Eqs. (1) and (2)

$$
Q_S = Q \frac{P + 0.5}{1.5}, \quad Q_D = Q \frac{1 - P}{1.5}, \tag{3}
$$

so that one can study the $l \rightarrow l - 1$ and the $l \rightarrow l + 1$ channels separately.

In the case of thallium discussed here, the situation is complicated by the fact that the continuum cannot be reached solely by the direct transitions on which our interest in focused. There are also transitions via autoionizing states