Laboratories in checking some of our source materials for radioactive contamination.

†Work supported in part by the National Science Foundation.

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¹T. D. Lee and C. S. Wu, Annu. Rev. Nucl. Sci. <u>15</u>, 381 (1965).

 2 T. Kirsten and H. W. Müller, Earth Planet. Sci. Lett. <u>6</u>, 271 (1969).

³B. Srinivasan, E. C. Alexander, Jr., R. D. Beaty, D. E. Sinclair, and O. K. Manuel, Econ. Geol. <u>68</u>, 252 (1973).

⁴R. K. Bardin, P. J. Gollon, J. D. Ullman, and C. S. Wu, Nucl. Phys. A 158, 337 (1970).

 5 For instance, Compton scattering of an ambient γ ray in the source followed by a second Compton scattering or a Møller scattering with detection of the scattered photon and at least one electron. See Ref. 4, p. 341, processes (F) and (G).

⁶E. Greuling and R. C. Whitten, Ann. Phys. (N. Y.) <u>11</u>, 510 (1960).

[†]The events that were found outside this energy region will be discussed elsewhere.

⁸The total half-life measured by Srinivasan *et al*. (Ref. 3) is about twice as large as that of Kirsten and Müller (Ref. 2). In order to set an upper limit on the branching ratio, the result reported by Srinivasan *et al*. is used.

⁹B. Pontecorvo, Phys. Lett. B <u>26</u>, 630 (1968). ¹⁰E. W. Hennecke, O. K. Manuel, and D. D. Sabu, Phys. Rev. C 11, 1378 (1975).

¹¹H. Primakoff and S. P. Rosen, Rep. Prog. Phys. <u>22</u>, 121 (1959).

 12 A much less reliable estimate can be obtained by assuming the N*(1236) mechanism proposed by Primakoff and Rosen. [See H. Primakoff and S. P. Rosen, Phys. Rev. 184, 1925 (1969).] By use of their theoretically calculated expression, we find $\alpha \lesssim 10^{-4}$, which is of the same order as obtained by Hennecke, Manuel, and Sabu (Ref. 10) from their recent measurements with tellurium.

¹³E. Fiorini, A. Pullia, G. Bertolini, F. Capellani, and G. Restelli, Nuovo Cimento Soc. Ital. Fis. A <u>13</u>, 747 (1973).

Exact Classical Solution for the 't Hooft Monopole and the Julia-Zee Dyon*

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We present an exact solution to the nonlinear field equations which describe a classical excitation possessing magnetic and electric charge. This solution has finite energy and exhibits explicitly those properties which have previously been found by numerical analysis.

Recently 't Hooft¹ has proposed a model for a magnetic monopole which arises as a static solution of the classical equations for the SU(2) Yang-Mills field coupled to an SU(2) Higgs field. The model has been extended by Julia and Zee² so that the monopole becomes a dyon, possessing both electric and magnetic charge. The purpose of this note is to present an exact analytic solution for a particular version of these models.

The Lagrangian density for the model is³

$$\mathcal{L} = -\frac{1}{4} F^{\mu\nu a} F_{\mu\nu}{}^{a} - \frac{1}{2} \Pi^{\mu a} \Pi_{\mu}{}^{a} + \frac{1}{2} \mu^{2} \varphi^{a} \varphi^{a} - \frac{1}{4} \lambda (\varphi^{a} \varphi^{a})^{2}, \tag{1}$$

where

$$F_{\mu\nu}{}^{a} = \partial_{\mu}A_{\nu}{}^{a} - \partial_{\nu}A_{\mu}{}^{a} + e \epsilon^{abc}A_{\mu}{}^{b}A_{\nu}{}^{c}, \qquad (2)$$

and

$$\Pi_{\mu}{}^{a} = \partial_{\mu}\varphi^{a} + e \epsilon^{abc} A_{\mu}{}^{b}\varphi^{c}. \tag{3}$$

The field equations in the static limit where all time derivatives are zero are

$$\partial_{\iota} F^{\mu i a} + e \epsilon^{abc} A_{\iota}{}^{b} F^{\mu i c} = e \epsilon^{abc} \Pi^{\mu b} \varphi^{c} , \qquad (4)$$

and

$$\partial_{i} \Pi^{ia} + e \epsilon^{abc} A_{\mu}{}^{b} \Pi^{\mu c} = -\mu^{2} \varphi^{a} + \lambda (\varphi^{b} \varphi^{b}) \varphi^{c} . \tag{5}$$

The Wu-Yang⁴-'t Hooft-Julia-Zee Ansatz is to seek a solution of the form

$$A_i^a = \epsilon_{aij} \hat{r}_j [1 - K(r)]/er, \tag{6}$$

$$A_0^a = \hat{\gamma}_a J(r)/er, \tag{7}$$

$$\varphi^a = \hat{r}_a H(r) / er. \tag{8}$$

Fields which satisfy Eqs. (4) and (5) also produce an extremum of the canonical Hamiltonian obtained from (1):

$$\mathcal{H} = \int d^3 \gamma \left[\frac{1}{4} F_{ij}^{\ a} F_{ij}^{\ a} - \frac{1}{2} F_{0i}^{\ a} F_{0i}^{\ a} + \frac{1}{2} \Pi_i^{\ a} \Pi_i^{\ a} - \frac{1}{2} \Pi_0^{\ a} \Pi_0^{\ a} - \frac{1}{2} \mu^2 \varphi^a \varphi^a + \frac{1}{4} \lambda (\varphi^a \varphi^a)^2 \right]. \tag{9}$$

When expressed in terms of K(r), J(r), and H(r) this becomes

$$3C = \frac{4\pi}{e^2} \int_0^\infty dr \left(K'^2 + \frac{(K^2 - 1)^2}{2r^2} - \frac{J^2 K^2}{r^2} - \frac{(rJ' - J)^2}{2r^2} + \frac{H^2 K^2}{r^2} + \frac{(rH' - H)^2}{2r^2} - \frac{\mu^2 H^2}{2} + \frac{\lambda H^4}{4e^2 r^2} \right), \tag{10}$$

and the field equations reduce to

$$\gamma^2 K'' = K(K^2 - 1) + K(H^2 - J^2), \tag{11}$$

$$\gamma^2 J'' = 2JK^2, \tag{12}$$

$$\gamma^2 H'' = 2HK^2 + (\lambda/e^2)(H^3 - C^2\gamma^2 H), \qquad (13)$$

where $C = \mu e/\sqrt{\lambda}$.

We seek solutions such that the square of the Higgs field goes to a constant as $r \to \infty$. Then from Eq. (13) we see that $H/r \to \pm C$. For definiteness we choose the positive sign. The particular version of these equations that we will consider is to take $\lambda \to 0$ with C fixed. Then the requirement $H/r \to \infty$ C is no longer forced upon us. Nevertheless, we shall look for solutions of Eqs. (11)-(13) for which it continues to be valid.

It is easy to verify that a solution of Eqs. (11)–(13) with $\lambda = 0$ and regular boundary conditions at r = 0 and ∞ is given by

$$K = Cr/\sinh(Cr), \tag{14}$$

$$J=0, (15)$$

$$H = Cr \coth(Cr) - 1, \tag{16}$$

which therefore constitutes an exact solution of 't Hooft's model for λ and μ^2 both set equal to 0.

Furthermore, by choosing

$$K = Cr/\sinh(Cr), \tag{17}$$

$$J = \sinh \gamma [Cr \coth(Cr) - 1], \qquad (18)$$

$$H = \cosh \gamma [Cr \coth(Cr) - 1], \tag{19}$$

we have a solution of our version of the model of Julia and Zee with $H/r \to \infty C \cosh \gamma$. Here γ is an arbitrary constant. Clearly the solution of 't Hooft's model corresponds to $\gamma = 0$.

The reader may wonder how such a solution was discovered. The answer is that if one seeks trial functions with good boundary conditions for a variational calculation, the hyperbolic sine and tangent arise quite naturally. Substitution of a few trial functions of this type in the equations revealed that they were almost satisfied. With a bit of "shimmying" the expressions (14)-(16) emerged.

The exact solutions leave unchanged most of the conclusions which have been deduced from the numerical analyses. Specifically, let us investigate the magnetic monopole strength $4\pi g$, the electric charge Q, and the mass M of the excitation we have found.

The Abelian electromagnetic field has been identified by 't Hooft^{1,5} as

$$\mathfrak{F}_{\mu\nu} = \partial_{\mu} (\hat{\varphi}^a A_{\nu}^a) - \partial_{\nu} (\hat{\varphi}^a A_{\mu}^a) - (1/e) \epsilon^{abc} \hat{\varphi}^a \partial_{\mu} \hat{\varphi}^b \partial_{\nu} \hat{\varphi}^c , \qquad (20)$$

where $\hat{\varphi}^a = \varphi^a (\varphi^b \varphi^b)^{1/2}$. Using Eqs. (6)-(8) we have

$$\mathfrak{B}_{i} = \frac{1}{2} \epsilon_{ijk} \mathfrak{F}_{jk} = -\hat{r}_{i}/er^{2}, \tag{21}$$

and

$$\mathcal{E}_{i} = -\mathfrak{F}_{0i} = \hat{r}_{i} \frac{d}{dr} \left(\frac{J}{r} \right) = \hat{r}_{i} \sinh \gamma \left(\frac{1 - K^{2}}{e r^{2}} \right). \tag{22}$$

We see that our solution represents a point monopole of strength $4\pi g = -4\pi/e$. In addition, since $K \to 1$ as $r \to 0$ there is a cloud of electric charge of amount $Q = (4\pi/e)\sinh\gamma$ with no pointlike core. Thus the solution represents a point monopole surrounded by a cloud of electric charge.

The energy or mass of the solution can be calculated from the Lagrangian density written in generally covariant form as $M = \int d^3r T^{00}(\vec{r})$, where

$$T^{\mu\nu} = (2/\sqrt{-g}) \partial (\mathcal{L}\sqrt{-g}) / \partial g_{\mu\nu} = F^{\mu\lambda a} F_{\lambda}^{\nu a} + \Pi^{\mu a} \Pi^{\nu a} + g^{\mu \nu} \mathcal{L}. \tag{23}$$

We find the same expression for M as for \mathcal{H} given by Eq. (10) except that the two terms with negative signs have those signs reversed. For the case $\lambda = 0$, Eqs. (12) and (13) allow us to write

$$\frac{1}{2r^2}(rJ'-J)^2 + \frac{J^2K^2}{r^2} = \frac{1}{2}\frac{d[J(J'-J/r)]}{dr},$$
 (24)

with a corresponding equation for the H terms so that

$$M = (4\pi/e^2) \left\{ \frac{1}{2} \left[H(H' - H/r) + J(J' - J/r) \right] \Big|_0^{\infty} + \int_0^{\infty} dr \left[K'^2 + \frac{1}{2} (K^2 - 1)^2 / r^2 \right] \right\}.$$
 (25)

The integral is easily evaluated with the result

$$M = (C/\alpha) \cosh^2 \gamma, \tag{26}$$

where $\alpha = e^2/4\pi$. The constant C governs the fall-off of the Yang-Mills field as $r \to \infty$ and is thus identified with its mass.

A comparison with the numerical results of 't Hooft and Julia and Zee is instructive. For $\gamma=0$ 't Hooft finds $M=(C/\alpha)f(\lambda/e^2)$, where f varies from 1.1 for $\lambda/e^2=0.1$ to 1.44 for $\lambda/e^2=10$. Julia and Zee compute f=1.42 for $\lambda/e^2=0.5$. We find f=1 for $\lambda/e^2=0$.

Secondly, Julia and Zee find for $\lambda/e^2 = 0.5$ solutions with $Q = 0.324e/\alpha$ which correspond to $\gamma = 0.319$ and 1.038, respectively. For such choices of γ we would find $M = 1.10C/\alpha$ and $2.56C/\alpha$, respectively, while they find $M = 2.62C/\alpha$ and $2.86C/\alpha$, respectively.

A few concluding remarks concerning stability are in order. Our solution appears to be unstable against changes in C since there is a continuum of solutions, each with mass proportional to C. But C is the only mass parameter in the theory and so it sets the scale of length. Thus solutions for different values of $C \neq 0$ are identical with respect to the appropriate length scale and hence can be considered stable, at least with regard to changes in C.

Concerning changes in γ it would appear on the classical level that since the mass depends con-

tinuously on γ , only $\gamma = 0$ can lead to stability unless some classical reason can be found to quantize $Q = (e/\alpha) \sinh \gamma$.

One of us (C.M.S.) would like to thank Loyal Durand, III, for the hospitality extend to him by the Aspen Center for Physics where part of this work was done.

*Research supported in part by the U.S. Energy Research and Development Administration under Contract No. AT(11-1)-3075.

¹G. 't Hooft, Nucl. Phys. B79, 276 (1974).

²B. Julia and A. Zee, Phys. Rev. D <u>11</u>, 2227 (1975).

³We use a metric $g^{11} = g^{22} = g^{33} = -g^{00} = 1$. Repeated Greek indices are summed from 0 to 3; Latin indices from 1 to 3. On the fields A, F, φ , and Π indices a, b, and c refer to isospin and i, j, and k to ordinary space. The symbol ϵ^{abc} is the usual completely antisymmetric object with $\epsilon^{123} = \epsilon_{123} = 1$. The symbol \hat{r}_i stands for a component of the unit radial vector and $\sqrt{-g}$ is the square root of the negative of the determinant of $g_{\mu\nu}$.

⁴T. T. Wu and C. N. Yang, in *Properties of Matter Under Unusual Conditions*, edited by H. Mark and S. Fernbach (Interscience, New York, 1969), pp. 349-354.

⁵J. Arafune, P. G. O. Freund, and C. J. Goebel, J. Math. Phys. (N. Y.) <u>16</u>, 433 (1975).