applied to the system. This wavelength is chosen to exactly connect the 5s level of Kr with the 7s of Hg. Equation (3) applies, with the quantity  $|\gamma|^2$ now given by  $(|\mu_{23}^A||\mu_{30}^A||\mu_{01}^B||E|/2\hbar^2\Delta\omega_1)^2$ .  $\Delta\omega_1 = 4054 \text{ cm}^{-1}$  is the difference between the applied optical frequency and the Hg 7s-8p transition. At an incident power density of  $10^9 \text{ W/cm}^2$ , an excited mercury concentration of  $10^{17}$  atoms/ cm<sup>3</sup>, and  $\overline{V} = 5 \times 10^4$  cm/sec, we find a Kr transition rate of  $2.6 \times 10^7$  transitions/sec. This transition rate corresponds to an effective cross section for inelastic collision  $[W(t) = N_A \sigma_{eff} \overline{V}]$  of  $\sigma_{eff}$ = 72  $Å^2$ . This cross section varies linearly with the power density of the applied 5886-Å field. One application of this type of pumping process may be selective rapid ionization of inner-shell electrons, thereby creating an inversion with regard to these levels.<sup>8</sup>

From a somewhat different point of view, we have described a process where by using one or more photons of an external electromagnetic field, large effective cross sections for inelastic collision can be attained without the usual requirement of energy resonance of the atomic species. Experiments to demonstrate these ideas will be undertaken in our laboratory.

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<sup>1</sup>J. F. Young, G. C. Bjorklund, A. H. Kung, R. B. Miles, and S. E. Harris, Phys. Rev. Lett. <u>27</u>, 1551

(1971); R. B. Miles and S. E. Harris, IEEE J. Quantum Electron. 9, 470 (1973).

<sup>2</sup>R. T. Hodgson, P. P. Sorokin, and J. J. Wynne, Phys. Rev. Lett. <u>32</u>, 343 (1974).

<sup>3</sup>S. E. Harris and D. M. Bloom, Appl. Phys. Lett. <u>24</u>, 229 (1974); D. M. Bloom, James T. Yardley, J. F.

Young, and S. E. Harris, Appl. Phys. Lett. <u>24</u>, 427 (1974). <sup>4</sup>P. R. Berman, Phys. Rev. A <u>6</u>, 2157 (1972), and <u>5</u>,

927 (1972). <sup>5</sup>L. I. Schiff, *Quantum Mechanics* (McGraw-Hill, New York, 1968) 3rd ed., p. 260.

<sup>6</sup>Progress in Quantum Electronics, edited by J. H. Sanders and S. Stenholm (Pergamon, New York, 1972), Vol. 2, Pt. 3.

<sup>7</sup>C. E. Moore, *Atomic Energy Levels*, U. S. National Bureau of Standards, National Standard Reference Data Systems—35 (U. S. GPO, Washington, D. C., 1971), Vols. 2 and 3.

<sup>8</sup>M. A. Duguay and P. M. Rentzepis, Appl. Phys. Lett. <u>10</u>, 350 (1967).

## Radiative Decay of the $2^{3}S_{1}$ and $2^{3}P_{2}$ States of Heliumlike Vanadium (Z = 23) and Iron $(Z = 26)^{*}$

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Lifetimes of the M1 decay  $2^3S_1 \rightarrow 1^1S_0$  and of the decay  $2^3P_2 \rightarrow 1^1S_0$  have been measured in the two-electron ions V<sup>+21</sup> and Fe<sup>+24</sup>. The measured lifetimes are  $\tau(2^3S_1) = 16.9(7)$  nsec for V<sup>+21</sup> and  $\tau(2^3S_1) = 4.8(6)$  nsec for Fe<sup>+24</sup>. The  $2^3P_2$  lifetimes are compared with a calculation that considers relativistic corrections and hyperfine-structure effects. It is found that for V<sup>+21</sup>, hyperfine effects contribute appreciably to the lifetime. For Fe<sup>+24</sup> we obtain  $\tau(2^3P_2) = 0.11(2)$  nsec.

The study of radiative decay from the  $2^3S_1$  and  $2^3P_2$  levels of the two-electron system offers an opportunity to test the theory of forbidden decay in a system where precise, unambiguous calculations of decay rates can be made. In this paper, we describe some measurements designed to extend existing experimental information on the rates of these decays to the two-electron atoms V<sup>+21</sup> and Fe<sup>+24</sup>. We develop a theory for the  $2^3P_2$  rates which takes into account leading-order relativistic corrections and hyperfine effects. Our results on vanadium show the first evidence for the influence of the hyperfine inter-

action on the radiative decay of an energy level belonging to the two-electron system.

Experimental work on the  $2^3S_1$  levels of the twoelectron system has established the single-photon nature of the decay in  $\operatorname{Ar}^{+16\,1}$  and ordinary helium.<sup>2</sup> The theory of this decay has been examined by several authors<sup>3</sup> and detailed calculations of the rates have been made by Drake<sup>4</sup> and by Johnson and Lin<sup>5</sup> using somewhat different starting points. Measurements of this decay rate have been made on  $\operatorname{Ar}^{+16}$ ,  $\operatorname{Ti}^{+20}$ ,<sup>6</sup> and  $\operatorname{Cl}^{+15}$ .<sup>7</sup> The measured rates in Cl and Ar differ from theory by several times the quoted error, whereas the



FIG. 1. Sample decay curve obtained with  $V^{+21}$  beam. The points beyond 22 cm are fitted by a single exponential and ascribed to decay of the  $2^{3}S_{1}$  level. The small-separation points are fitted by a composite of four exponentials as described in the text and are ascribed to decay of the  $2^{3}P_{2}$  level.

result in Ti is in agreement. In order to establish the Z dependence of this discrepancy we report here measurements on  $V^{+21}$  and  $Fe^{+24}$ .

Interest in the  $2^{3}P_{2}$  levels centers mainly on the M2 decay mode to the  $1^{1}S_{0}$  ground level which has a rate comparable with the E1 rate to  $2^{3}S_{1}$ for ions with  $Z \simeq 20$ . The M2 decay was first observed<sup>8</sup> in Ar<sup>+16</sup> and rates have now been measured in S<sup>+14</sup>, <sup>9</sup> Cl<sup>+15</sup>, <sup>10</sup> and Ar<sup>+16</sup>. <sup>11</sup> In this paper we present evidence indicating that the decay of this level is strongly influenced by the hyperfine interaction in V<sup>+21</sup> and give results for Fe<sup>+24</sup>.

The lifetimes were measured by the beam-foil time-of-flight method. Our apparatus has been described previously<sup>6</sup> and the details are not repeated here. The vanadium (Z = 23) and iron (Z = 26) ions were obtained from the Lawrence Berkeley Laboratory SuperHILAC at an energy of 7.2 MeV/amu. Excitation of the beams into the metastable states was done with a 50- $\mu$ g/cm<sup>2</sup> carbon foil. Decay curves are taken by varying the foil-detector separation. The total number of counts under the peak is integrated and normalized to the integrated beam current collected in a Faraday cup. This quantity is plotted as a function of foil-detector separation. A sample decay curve

is shown in Fig. 1.

Since the  $2^{3}P_{2}$  and  $2^{3}S_{1}$  energy separations are small compared with the detector resolution, the decay curves are composites, exhibiting fast components and slow components. The slow components are ascribed to the  $2^{3}S_{1}$  decay and the fast components to decay from  $2^{3}P_{2}$ . As discussed below, the slow and fast components are probably composites of two or more exponentials as the result of hyperfine effects in V<sup>+21</sup> and cascading effects in Fe<sup>+24</sup>.

The isotope <sup>51</sup>V has a nuclear spin  $I = \frac{7}{2}$  and a magnetic moment  $\mu = 5.15\mu_N$ .<sup>12</sup> The resulting hyperfine structure influences our decay curves by admixing  $2^3P_2$  with  $2^3P_1$  and  $2^1P_1$ . The  $2^3P_2$  level is split into five components with total angular momentum between  $F = \frac{3}{2}$  and  $F = \frac{11}{2}$  (Fig. 1). Because the hyperfine interaction is diagonal in F, the rates from both  $F = \frac{3}{2}$  and  $F = \frac{11}{2}$  are unaffected by hyperfine structure. However, the rates from  $F = \frac{5}{2}$ ,  $\frac{7}{2}$ , and  $\frac{9}{2}$  will all be altered. Hence, the observed decay curve will be a composite of four exponentials weighted according to the initial populations of each of the hyperfine levels.

The most abundant stable iron isotope has zero

spin and hyperfine effects are not present. Hence the decay of  $2^{3}P_{2}$  is exclusively by E1 decay to  $2^{3}S_{1}$  and M2 to  $1^{1}S_{0}$ .

A problem arises in the interpretation of the  $2^{3}S_{1}$  decay because of cascading from  $2^{3}P_{0}$ . The lifetime of  $2^{3}P_{0}$  is calculated to be  $\tau = 2.7$  nsec which is close to that of  $2^{3}S_{1}$ . Hence cascading effects from this level may be important. The shape of the resultant curve will depend upon the relative population  $R \equiv N_0(2^3P_0)/N_0(2^3S_1)$  of the two states, and the decay rates. It was found that the data could be fitted by two exponentials for  $0 \le R$  $\leq$  0.8. Varying the lifetimes of the two exponentials, we find that the "best fit" values for  $\tau(2^3S_1)$ vary between 4.2 and 5.3 nsec, where 5.3 nsec results from a single exponential fit (R = 0) and. hence, is an upper limit to the  $2^{3}S_{1}$  lifetime. Similarly, the values for  $\tau(2^{3}P_{0})$  vary between 1.3 and 3.3 nsec. We note that the theoretical value of  $\tau(2^3P_0) = 2.7$  nsec corresponds to  $\tau(2^3S_1) = 4.8$ nsec. We have chosen to take the mean value as the experimental value and to take cascading into account by using an increased error.

In Fig. 1, an experimental decay curve is shown for the V<sup>+21</sup> beam. The decay curve is fitted by identifying the data points beyond 22 cm with the  $2^3S_1$  decay. These yield a lifetime  $\tau(2^3S_1) = 16.9(7)$ nsec. When this decay is subtracted from the experimental data it yields the points for small detector-foil separation as shown. These are fitted by a composite decay curve constructed by assuming that each of the *F* states associated with  $2^3P_2$ has the theoretical lifetime shown in Table I. These lifetimes include effects due to hyperfine quenching. Moreover, the initial populations of the *F* states are taken as proportional to 2F + 1. TABLE I. Theoretical transition rates and lifetimes of the  $2^{3}P_{2}$  state in heliumlike ions.

Z	$\begin{array}{c} A_{E1} \\ (\text{nsec}^{-1}) \end{array}$	$A_{M^2} $ (nsec <sup>-1</sup> )	F	$\begin{array}{c} A_{E1}^{\text{hfs}} \\ (\text{nsec}^{-1}) \end{array}$	au (nsec)
16 17 18 22	0.259 0.301 0.352 0.687	0.117 0.194 0.312 1.64	A11	< 0.007	$2.66 \\ \simeq 2.01 \\ 1.51 \\ 0.429$
23 26	0.820	2.37 6.50	$\begin{pmatrix} 3/2 \\ 5/2 \\ 7/2 \\ 9/2 \\ 11/2 \end{pmatrix}$	0 0.99 1.75 1.69 0	$\begin{array}{c} 0.313 \\ 0.239 \\ 0.202 \\ 0.205 \\ 0.313 \\ 0.126 \end{array}$

The resultant fit to the experimental points is as shown and is seen to be quite good.

To calculate accurate theoretical rates for the  $2^{3}P_{2}$  level, the transition matrix elements are evaluated to zeroth and first order in the quantities  $Z^{-1}$  and  $(Z\alpha)^{2}$ . The zeroth-order term is the nonrelativistic hydrogenic approximation. The term of order  $Z^{-1}$  is the first correction in the nonrelativistic Z expansion of the matrix element. The leading relativistic correction, of order  $(Z\alpha)^{2}$ , is obtained by evaluating the relativistic transition operator between states formed from properly symmetrized products of hydrogenic Dirac wave functions.

For the M2 transition, the coefficient of  $Z^{-1}$  has been determined by Drake<sup>13</sup> from the Z-expansion calculation of Dalgarno and Parkinson.<sup>14</sup> Taking his result together with the leading relativistic and finite-wavelength (retardation) corrections yields

(1)

 $A_{M2}(2^{3}P_{2}) = \alpha k^{5}(2^{15}/5\times 3^{10})(Z\alpha)^{-2}[1+0.147Z^{-1}-0.640(Z\alpha)^{2}]^{2}$ 

(in units where  $m_e = c = \hbar = 1$ ), where k is the transition energy.

Because the nonrelativistic-velocity form of the matrix element for  $n \rightarrow n E1$  transitions vanishes to lowest order in  $Z^{-1}$ , the evaluation of the E1 transition rate  $2^{3}P_{2} \rightarrow 2^{3}S_{1}$  is simplified by making the dipole approximation and converting the matrix element to the length form. The error in making this approximation is negligible because of the smallness of the transition energy:  $\langle \vec{k} \cdot \vec{r} \rangle = O(\alpha)$ . The conversion to length form is valid relativistically and so relativistic corrections may be obtained as described above. For this transition matrix element, the coefficient of  $Z^{-1}$  has been obtained by Cohen and Dalgarno.<sup>15</sup> We thus have

$$A_{E1}(2^{3}P_{2}) = \alpha k^{3} 12 (Z\alpha)^{-2} [1 + 0.759Z^{-1} - 0.167(Z\alpha)^{2}]^{2}.$$
(2)

The transition probability for the E1 decay  $2^{3}P_{0} + 2^{3}S_{1}$ , obtained as for the  $2^{3}P_{2}$  decay, is given by

$$\boldsymbol{A}_{E1}(2^{3}\boldsymbol{P}_{0}) = \alpha k^{3} 12 (\boldsymbol{Z} \alpha)^{-2} [1 + 0.759 \boldsymbol{Z}^{-1} - 0.417 (\boldsymbol{Z} \alpha)^{2}]^{2}.$$
(3)

We note that the theoretical uncertainty in the expressions for the transition rates listed above is expected to be of the order of 1% or less for Z in the range 10-40.



FIG. 2. Comparison between measured and calculated decay rates for the  $2^{3}S_{1}$  level.

In the case of vanadium, which has a nonzero nuclear spin, the  $2^{3}P_{2}$  and  $2^{3}P_{0}$  states undergo an E1 transition to the  $1^{i}S_{0}$  state due to hyperfine mixing.<sup>16</sup> An estimate of the transition rate is obtained in the following way. Nonrelativistic intermediate-coupling wave functions are taken as the unperturbed basis. The hyperfine interaction which is diagonal in F = J + I, is treated in firstorder perturbation theory. Only the effect of mixing of the  $2^{3}P_{2}$  and  $2^{3}P_{0}$  states with the nearby  $2^{3}P_{1}$  and  $2^{1}P_{1}$  states is included. The dipole transition operator then has a nonvanishing matrix element between the perturbed  $2^{3}P_{2}$  and  $2^{3}P_{0}$ states and the  $1^{1}S_{0}$  state which is proportional to the dipole matrix element between the LS coupled  $2^{1}P_{1}$  state and the  $1^{1}S_{0}$  state. The latter is evaluated with the aid of the Z expansion of Dalgarno and Parkinson.<sup>14</sup> The hyperfine matrix elements are approximated by evaluating the contact-interaction term between hydrogenic product wave functions. Energy differences are evaluated by means of the Z expansion of the nonrelativistic energies,<sup>17</sup> together with the Z expansion of the order- $\alpha^4$  corrections.<sup>18</sup> The transition rates thus obtained are added to the rates discussed above and the resulting values for the lifetime of the  $2^{3}P_{2}$  state are listed in Table I.

In Fig. 2 we compare all of the measured decay rates of  $2^{3}P_{2}$  with the corresponding theoretical rates. Agreement is seen to be very good over a wide range of Z. For vanadium, the theoretical rate includes the contribution from hyperfine quenching. It is seen that the experimental and theoretical rates would be in serious disagreement without contributions from this mechanism. We take this as very strong evidence that hyperfine quenching is indeed present.

In Fig. 3 we compare the measured  $2^3S_1$  decay rates with the calculated rates. Agreement between theory and experiment is satisfactory for Z = 22, 23, and 26, but puzzling discrepancies exist at Z = 17 and 18.



FIG. 3. Comparison between measured and calculated M2 rates for decay from the  $2^{3}P_{2}$  level. The point at Z = 23 labeled "without HFS" is obtained by making a best fit to our vanadium data using a single exponential.

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<sup>1</sup>R. Marrus and R. W. Schmieder, Phys. Lett. <u>32A</u>, 431 (1970).

<sup>2</sup>H. W. Moos and J. R. Woodworth, Phys. Rev. Lett. <u>30</u>, 775 (1973).

<sup>3</sup>G. Feinberg and J. Sucher, Phys. Rev. Lett. <u>26</u>, 681 (1971); H. R. Griem, Astrophys. J. <u>156</u>, 1103 (1969); I. L. Beigman and V. I. Safronova, Zh. Eksp. Teor. Phys. <u>60</u>, 2045 (1971) [Sov. Phys. JETP <u>33</u>, 1102 (1971)]; S. Feneuille and E. Koenig, C. R. Acad. Sci., Ser. B <u>274</u>, 46 (1972); G. W. F. Drake, Phys. Rev. A <u>5</u>, 1979 (1972).

<sup>4</sup>G. W. F. Drake, Phys. Rev. A 3, 908 (1971).

<sup>5</sup>W. R. Johnson and C. Lin, Phys. Rev. A <u>9</u>, 1486 (1974).

<sup>6</sup>H. Gould, R. Marrus, and R. W. Schmieder, Phys. Rev. Lett. <u>31</u>, 504 (1973).

<sup>7</sup>C. L. Cocke, B. Curnutte, and R. Randall, Phys. Rev. Lett. 31, 507 (1973).

<sup>8</sup>R. Marrus and R. W. Schmieder, Phys. Rev. Lett. <u>25</u>, 1689 (1970).

<sup>9</sup>C. L. Cocke, B. Curnutte, and R. Randall, Phys. Rev. A 9, 1823 (1974).

<sup>10</sup>C. L. Cocke, B. Curnutte, J. R. MacDonald, and

R. Randall, Phys. Rev. A 9, 57 (1974).

<sup>11</sup>R. Marrus and R. W. Schmieder, Phys. Rev. A <u>5</u>, 1160 (1972).

<sup>12</sup>W. F. Ramsey, *Molecular Beams* (Oxford Univ.

Press, Oxford, England, 1963), p. 173.

<sup>13</sup>G. W. F. Drake, Astrophys. J. <u>158</u>, 1199 (1969).

<sup>14</sup>A. Dalgarno and E. M. Parkinson, Proc. Roy. Soc., Ser. A <u>301</u>, 253 (1967).

<sup>15</sup>M. Cohen and A. Dalgarno, Proc. Roy. Soc., Ser. A

293, 359 (1966).

 $^{16}$ A discussion of the history and calculations of this effect is given in R. H. Garstang, J. Opt. Soc. Amer. 52, 845 (1962).

<sup>17</sup>R. E. Knight and C. W. Sherr, Rev. Mod. Phys. <u>35</u>, 431 (1963).

<sup>18</sup>H. T. Doyle, in *Advances in Atomic and Molecular Physics*, edited by D. R. Bates and I. Esterman (Academic, New York, 1969), Vol. 5, p. 337.

## **Reverse Temperature Gradient in the Kinetic Theory of Evaporation**

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The singular-eigenfunction-expansion method for solving the transport equation is applied to a modeled kinetic-theory Boltzmann equation to solve the problem of evaporation and condensation between parallel surfaces. A remarkable result is obtained for the temperature profile which seems to confirm previous predictions made on the basis of asymptotic approximations.

In a recent paper, Pao<sup>1</sup> suggested that the slope of the temperature profile of a saturated vapor between two parallel surfaces at different temperatures could be in opposition to the imposed gradient for certain values of  $\beta$ , where  $\Delta N = \beta \Delta T$  in a saturated liquid-vapor system. Here N represents the molecular density and T the temperature. Pao's prediction was based on an asymptotic analysis for large gap width, and used results obtained for the half-space problem<sup>2</sup> by applying the Wiener-Hopf technique to a modeled Boltzmann equation.

We have approached the same problem through the singular-eigenfunction analysis of Kriese, Chang, and Siewert,<sup>3</sup> also used by Thomas and co-workers,<sup>4-6</sup> and have avoided any asymptotic approximations. Thus for the first time, inverted temperature profiles and critical values of  $\beta$  are calculated explicitly. The surprising result is that our analysis seems to confirm Pao's novel prediction.

We consider the problem of a vapor between two interphase (vapor-liquid or vapor-solid) surfaces maintained at  $x = \pm d/2$ ; we assume that the condensed phase in x < -d/2 is kept at temperature  $T_0 - \frac{1}{2}\Delta T$ and that the condensed phase in x > d/2 is kept at temperature  $T_0 + \frac{1}{2}\Delta T$ . As a result of evaporation and condensation there is a flow of mass and energy from x = + d/2 to x = -d/2. Within the vapor, we assume the state of the fluid to be described by the linearized single-relaxation model of the Boltzmann equation<sup>7</sup>:

$$c_x \partial f(x, \vec{c}) / \partial x + f(x, \vec{c}) = \pi^{-3/2} [N(x) + (c^2 - \frac{3}{2})T(x) + 2c_x U(x)].$$
(1)

Here  $f(x, \vec{c})$  is the perturbation of the particle-distribution function, x is the spatial variable,  $\vec{c}$  is the molecular velocity, and  $c_x$  is the x component of the velocity, all in dimensionless units. In addition, N(x), T(x), and U(x) represent perturbations of the number density, the temperature, and the x com-