Static Nuclear Deformation and the Total Reaction Cross Section*

John M. Alexander, Louis C. Vaz, † and S. Y. Lin

Department of Chemistry, State University of New York at Stony Brook, Stony Brook, New York 11794

(Received 9 October 1974)

An empirical study is made of the total reaction cross sections for 4 He reactions with ^{208}Pb , ^{209}Bi , ^{233}U , and ^{237}Np . Systematic differences are revealed that are attributed to the effects of nuclear deformation. This behavior can be accounted for by a reduction of the average interaction barrier by 0.8 to 0.6 MeV or by an increase in the width of the spectrum of barriers by 0.4 to 0.⁷ MeV.

Calculated values of the total reaction cross section $(\sigma_{\mathbf{z}})$ depend on the deformation of each of the colliding nuclei.¹⁻⁴ The calculations, from several models of varying sophistication, indicate that the sensitivity of σ_{p} to deformation is very small for energies greater than the interaction barrier, but appreciable for lower energies.¹⁻⁴ One or both reaction partners may exhibit static nuclear deformation or, in addition, dynamic shape changes can be induced during the collision act. Several theoretical investigations lead to the expectation that dynamic effects are 'very small.^{5,6} Static deformation, however, is expected to lead to decidedly enhanced reaction cross sections at low energies because of the lower interaction barrier for collisions near thepolar $axis.^{1-3}$

Previous studies of the available experimental data have not been able to isolate this effect. $1 - 3.7$ In this paper we present an empirical approach to this problem that does indicate significant effects of static deformation on σ_p . We use the experimental data of Barnett and Lilley' for the near-spherical nuclei ²⁰⁸Pb and ²⁰⁹Bi, those of Freiesleben and Huizenga³ and Lin⁹ and Fleury et al.¹⁰ for the statically deformed nuclei ^{233}U and 237 Np. The results are, of course, completely dependent on the accuracy of the experimental data—in particular, the effective beam energies. The relevant experimental data from Refs. 3 and 8-10 were obtained with modern techniques for control of beam energy at Van de Graaff accelerators.

We have used the equations for penetration of a real parabolic barrier to fit measured total reaction cross sections. The transmission coefficient $T(l, E)$ for the partial wave l is given by the Hill-Wheeler¹¹ formula

$$
T(l, E) = [1 + \exp\{2\pi(E_l - E)/\hbar\omega_l\}^{-1}, \qquad (1)
$$

where E is the incident energy $(c.m.)$. The effec-

tive potential barrier height E_i is given by

$$
E_{l} = V_{N}(R_{l}) + V_{C}(R_{l}) + \hbar^{2}l(l+1)/2\,\mu R_{l}^{2}, \tag{2}
$$

with R_i the radial distance corresponding to the maximum and with V_N and V_C the nuclear and Coulomb potentials. Wong' has derived a very simple expression for the total reaction cross section:

$$
\sigma_R = (R_0^2/2)(\hbar\omega_0/E)
$$

×ln[1+exp{2\pi(E-E_0)/\hbar\omega_0}], (3)

with the approximation that the curvature $\hbar\omega$, and radius R_i are independent of l (i.e., $\hbar\omega_1 = \hbar\omega_0$ and $R_i = R_0$.

Vaz and Alexander' have modified this expression by consideration of a uniform distribution of sion by consideration of a uniform distribution c
barrier heights between $\overline{E}_0 - \Delta$ and $\overline{E}_0 + \Delta$.¹² With this modification they have analyzed and systematized all the available reaction cross sections in terms of the four parameters \overline{E}_0 , R_0 , $\hbar\omega_0$, and Δ . No clear variations could be determined for the empirical parameters $\hbar\omega_0$ and Δ ; values of \approx 4 MeV and \approx 3 MeV, respectively, provide good fits to σ_R for $E \ge \overline{E}_0$. For ⁴He projectiles \overline{E} and R_0 are well fitted by the following expressions:

$$
E_0 = 2Ze^2/R_e \tag{4}
$$

with

$$
R_e = r_e A^{1/3} + 2.53 \text{ fm},\tag{5}
$$

and

$$
R_0 = r_0 A^{1/3} + 2.53 \text{ fm.}
$$
 (6)

The empirical value of r_0 (1.41 fm) seems to be essentially independent of Z. The empirical values of r_e vary rather slowly with Z^7 .

Our knowledge of the shape and height of the interaction barrier from theory or from other measurements (i.e., elastic scattering) is not yet well enough developed to give us precise values

FIG. 1. The total reaction cross section σ_R versus laboratory energy for 4 He reactions with four targets. Data are from Hefs. 8, 8, 9, and 10. Data points are indicated as follows: ^{208}Pb , \triangleright ; ^{209}Bi , ∇ ; ^{233}U , + ²³⁷Np, \times . Calculated values of σ_R are given by the lines with "best-fit" parameters as indicated. In these fits r_0 was fixed at 1.41 fm and \overline{E}_0 , $\hbar\omega_0$, and Δ were free parameters.

of r_e , r_0 , $\hbar\omega_0$, and Δ . Freiesleben and Huizenga³ have employed independently estimated potential parameters in several formulations, in an attempt to describe $\sigma_R(E)$ and identify the effect of static deformation. They were unable to obtain a good fit to σ_R at low energies (with no free parameters). Even with barrier heights fitted to experiment no effect of deformation was apparent.

We have taken the empirical route in the examination of the reactions of 4 He with 208 Pb, 209 Bi, ^{233}U , and ^{237}Np . (The data for the first three systems were also examined in Ref. 3. New data for $\mathrm{^{37}Np}$ are now available.⁹) First we fitted the $\sigma_{_{\bm{R}}}$ data for all systems with the three free parameters r_e (or \overline{E}_0), $\hbar\omega_0$, and Δ ; r_0 was fixed at 1.41 fm.⁷ The quality of fit and the parameters used are shown in Fig. 1. These values of \overline{E}_0 are somewhat larger than those obtained from analysis of elastic scattering measurements.¹³ This

FIG. 2. σ_R versus $E - \overline{E}_0$ (c.m.) with \overline{E}_0 for all systems fixed to those indicated by the reference systems ²⁰⁸Pb and ²⁰⁹Bi. (r_e = 1.47 for ²⁰⁸Pb and ²⁰⁹Bi; r_e = 1.46 for 233 U and 237 Np). The solid, dashed, and dot-dashed lines were calculated with \triangle values as indicated ($\hbar \omega_0$) and r_0 fixed at 3.1 MeV and 1.41 fm). Symbols for data points are as in Fig. l.

difference may well be due to the models used namely, the use of a single potential barrier $(\Delta = 0)$ in Ref. 13 and a spectrum of barriers here (finite Δ).

At first glance there is no clear systematic difference between the parameters for the nearspherical nuclei ^{208}Pb and ^{209}Bi and the deformed nuclei ^{233}U and ^{237}Np . However, the picture changes if we take the near-spherical nuclei as reference systems and then calculate σ_R for ²³³U and ²³⁷Np. Values of r_0 , r_e , Δ , and $\hbar\omega$ from the reference systems are 1.41 fm, 1.47 fm, 3.0 MeV, and 3. 1 MeV, respectively. The systematic variation of r_e with Z leads us to expect a change of -0.01 fm for r_e from the reference Z of 82, 83 to Z of 92, 93. The systematics give no indication of a change for $\hbar\omega_0$ or Δ . In Fig. 2 we replot σ_R against $E - E_0$ for the four systems. As R_0 is very nearly the same for all four, the data

points should follow the solid line calculated with parameters for the reference systems Pb and Bi [see Eq. (3)]. The observation from Fig. 2 is that σ_R for Np and U is 1.5 to 2 times as great as that for Pb and Bi $(E - E_0 < -2 \text{ MeV}).$

In this formulation these systematic differences can be described either by a change in Δ (the width of the spectrum of barriers) or in \overline{E}_0 (the average barrier height). In Fig. ² the dashed curves give fits to the data for U and Np with Δ allowed to vary from the reference value. Increases in Δ of 0.7 and 0.4 MeV give good fits for U and Np, respectively. Alternatively \overline{E}_0 (or r_e) can be allowed to vary with Δ , $\hbar\omega_0$, and r_0 fixed at the reference values. In Fig. 3 we show the eonvergenee of the data resulting from decreases in \overline{E}_0 of 0.6 and 0.3 MeV, respectively, for U and Np. Wong³ has estimated that target deformation in this region will raise r_e by ≈ 0.1 fm; Fig. 3 indicates an increase of 0.03 to 0.05 fm.

Our conclusions are as follows: Empirical comparison reveals systematic differences in the reaction cross sections for 4 He reactions with ${}^{208}Pb$,

FIG. 3. σ_R versus $E-\overline{E}_0$ (c.m.) with \overline{E}_0 and r_e values as indicated. Values of r_0 , $\hbar\omega_0$, and Δ fixed as 1.41 fm, 8.1 MeV, and B.^O MeV, respectively.

 209 Bi, 233 U, and 237 Np. These differences are exhibited if potential parameters are obtained from the Pb and Bi systems and then used to calculate expected values of σ_R for U and Np. The values of σ_R for statically deformed Np and U seem to be enhanced by 1.⁵ to 2 times compared to the reference systems, Pb and Bi. This enhancement can be accounted for by a change in the halfwidth of the spectrum of barrier heights by 0.4 to 0.7 MeV or by a reduction of the average barrier by 0.3 to 0.6 MeV (or some combination).

The conclusion that we reach is that static deformation of one collision partner does affect the reaction cross section at low energies. The conclusion of Ref. 3 was that such an effect is not evident if one uses independently obtained parameters to describe the potential. We agree that our present models and our knowledge of the internuclear potential are not well enough developed to reveal such small effects. Thus, we conclude that semiempirical methods are most useful for systemization and characterization of the phenomena.

*Work supported in part by the U. S. Atomic Energy Commission.

)Present address: Radiochemistry Division, Bhabha Atomic Research Centre Trombay, Bombay 85, India. $1J$. O. Rasmussen and K. Sugawura-Tanabe, Nucl.

Phys. A171, 703 (1971); J.O. Rasmussen, Lawrence Berkeley Laboratory Report No. LBL 2866, 1978 (unpublished) .

 ${}^{2}C$. Y. Wong, Phys. Lett. $42B$, 186 (1972), and Phys. Rev. Lett. 31, 766 (1973).

 3 H. Freiesleben and J. R. Huizenga, Nucl. Phys. A224. $503(1974)$.

 4 T. Tamura, Rev. Mod. Phys. 37, 679 (1965).

 5 H. Holm, W. Scheid, and W. Greiner, Phys. Lett. 29B, 473 (1969).

 $\overline{{}^6P}$. W. Riesenfeld and T. D. Thomas, Phys. Rev. C $\overline{2}$, 711 (1970).

 7 L. C. Vaz and J. M. Alexander, Phys. Rev. C 10, 464 (1974).

 8A . R. Barnett and J. S. Lilley, Phys. Rev. C 9, 2010 (1974).

 9 S. Y. Lin, M. S. thesis, State University of New York at Stony Brook, 1974 (unpublished).

 10 A. Fleury, F. H. Ruddy, M. N. Namboodiri, and J. M. Alexander, Phys. Rev. ^C 7, ¹²⁸¹ (1978).

 11 D. L. Hill and J. A. Wheeler, Phys. Rev. 89, 1102 (1958). It is clear that a parabola cannot provide an accurate representation of the real potential for energies well below the barrier height. Nevertheless it is possible that Eq. (1) may provide a useful empirical representation of the reaction probability (see Ref. 1). It is in this spirit that we obtain and use best-fit parameters for $E \leq \overline{E}_0$.

 12 A spectrum of barriers resulting from random orientations of a deformed nucleus is discussed in Refs. 1-8. The possibility of a spectrum of barriers due to nuclear

vibrations has been put forward by J.R. Nix (private communication).

 13 G. Goldring, M. Samuel, B. A. Watson, M. C. Bertin, and S.I. Tabor, Phys. Lett. 32B, 465 (1970).

Element 106⁺

A. Ghiorso, J. M. Nitschke, J. R. Alonso, C. T. Alonso, M. Nurmia, and G. T. Seaborg Lawrence Berkeley Laboratory, University of California, Berketey, California 94720

and

E. K. Hulet and R. W. Lougheed Lawrence Liverrnore Laboratory, University of California, Liverrnore, California 94550 (Received 9 September 1974)

We have produced element 106 by bombarding 249 Cf with 18 O ions accelerated by the SuperHILAC. The new nuclide $^{263}106$, produced by the $(^{18}O, 4n)$ reaction, is shown to decay by α emission with a half-life of 0.9 ± 0.2 sec and a principal α energy of 9.06 ± 0.04 MeV to the known nuclide 259 Rf, which in turn is shown to decay to the known nuclide 255 No.

The identification of new elements at the upper end of the periodic table is especially difficult because of extremely low production rates and because there are large uncertainties in predicting their nuclear properties. For these reasons, positive identification requires some means of determining the atomic number directly. Among the proven methods are (1) measurement of distinctive K x rays^{1,2} following α decay and (2) establishment of a genetic link between an α emitter of a new element and a previously identified daughter nuclide. Our identification of element 106 is based on the latter method because of its higher sensitivity. This method was also used in discovering α -emitting isotopes of rutherfordium (element 104 ³ and hahnium (element 105).⁴ For element 106, we have carried this method one step further by demonstrating that the granddaughter $[{}^{255}$ No, $t_{1/2}$ = 3 min, E_{α} = 8.11 MeV (57%)]^{5,6,1} is in the chain of α decay of ²⁶³106. Thus, our proof for the atomic number of element 106 comes from demonstrating the following decay sequence: $^{263}106 \stackrel{\alpha}{\rightarrow}$ 259 Rf $\stackrel{\alpha}{\rightarrow}$ 255 No $\stackrel{\alpha}{\rightarrow}$.

These genetic relationships were established in two ways depending on whether $^{263}106$ α particles escaped from their backing surface. (1) When these particles were detected leaving the surface, we observed with a certain probability in a time interval of 12 sec the α 's of the 3-sec daughter ²⁵⁹Rf (E_α = 8.77 and 8.86 MeV) that also were directed outward; i,e., we observed the decay

sequence $^{263}106 \frac{\alpha}{4}$ 259 Rf $\frac{\alpha}{4}$. (2) When $^{263}106 \alpha' s$ were directed into the backing surface (and hence were not detected), the recoil energy imparted to the daughter nucleus allowed it to escape from the surface and to be implanted in the face of an opposing detector. Upon periodically moving these detectors away from the original sources, the α decay of daughter ²⁵⁹Rf and the subsequent α decay of the granddaughter were observed: i.e., we detected the decay sequence $^{259}Rf \overset{\alpha}{\rightarrow} ^{255}No$ $\frac{\alpha}{\tau}$. Considering the finite thickness (~1 μ g/cm²) of the NaCl deposits containing the 263106 atoms. the considerable recoil energy required to transfer the observed number of daughter 259 Rf atoms to the detector faces could be furnished only by a preceding α emitter. We thus were provided with a second genetic linkage to ²⁶³106 by α decay.

The ¹⁸O beam from the SuperHILAC (average 3×10^{12} ions/sec) was wobbled electromagnetically to prevent localized overheating of the target, which was both edge-cooled by contact with a water-cooled copper block and gas-cooled by helium impinging on the aluminum backing (Fig. 1). The energy of the ^{18}O ions emerging from the target was determined by measuring the energy of these ions scattered from the target into a Si(Au) surface barrier detector placed at 30° to the beam axis.

The target was prepared by subliming 259 μ g of ²⁴⁹Cf as CfF₃ onto a 27- μ g/mm² substrate of