



However one can show (by symmetry about the reflection surface) that the phase of the horizontal ray passing through the beam splitter at  $E'$  will correspond (at that point) to an equal reduction in length  $\Delta L'$  relative to a ray at  $E$ . As a consequence there will be no contribution (to first order in  $g$ ) from this effect to the interference.

A similar path distortion occurs because the ray  $BC$  arrives at  $C$  at a Bragg angle  $\theta - \epsilon$  instead of  $\theta$ , with

$$\epsilon = gr \cos(2\theta)/v^2 \sin(2\theta). \quad (6)$$

The ray  $CE$  will then have an angle  $\epsilon$  above the horizontal. Consequently it will emerge from the beam splitter at a point  $E''$  (instead of  $E$ ), which is displaced upwards *along the surface* by an amount

$$\Delta L'' \cong 2s\epsilon \cos\theta/\sin(2\theta). \quad (7)$$

However the ray from (near)  $D$  which reflects at  $E''$  will have a phase (at  $E''$ ) equivalent to an increased length  $\Delta L''$ . Again there will be no contribution to the interference.

Shifts of the reflection points  $C$ ,  $D$ , and  $E$  caused by the gravitational distortion of the rays into parabolas are also  $\sim 10 \text{ \AA}$  and lead to phase changes that cancel to first order in  $g$ .

The total phase difference between the two paths is that given by Eqs. (2) and (3). If the interferometer is rotated about the axis  $AD$ , the relative roles of the alternate paths are interchanged. Consequently the total fringe shift  $\Delta N$  will be twice  $\Delta\phi/2\pi$ :

$$\Delta N = 2M^2 grs\lambda/h^2. \quad (8)$$

Even if the neutron  $\lambda$  is expressed in terms of its classical momentum, the fringe shift  $\Delta N$  still depends on the ratio  $g/h$ .  $\Delta N$  is proportional to the area,  $rs$ , enclosed by the two paths. If this area is  $6 \text{ cm}^2$  and  $\lambda = 1.42 \text{ \AA}$ , the expected fringe shift is  $\Delta N = 10.7$ .

Elastic distortion of the interferometer during the  $180^\circ$  rotation will contribute a shift of about 1 fringe. This can be measured independently by repeating the experiment with x rays. Design of an interferometer that will function for both neutrons and x rays involves a compromise on the thickness of the beam splitter (slab  $BE$  of Fig. 1). We have found a suitable design based on the use of  $0.71\text{-\AA}$  ( $\text{Mo } K\alpha$ ) x rays and  $1.42\text{-\AA}$  neutrons. Similar beam paths result when the former undergo (440) and the latter (220) Bragg reflections.

Theoretical performance of our design has been computed<sup>5</sup> by using the dynamical theory appro-

priate<sup>1</sup> to an interferometer of the type shown in Fig. 1. The beam splitter was taken to be 1 mm thick. To optimize x-ray transmission the reflecting planes form an angle of  $8^\circ$  with the surface of the beam splitter. The integrated intensity<sup>6</sup> (for x rays) at  $F$ , when the two beams interfere constructively, is  $2.3 \times 10^{-8}$ . Since the dynamical index of refraction for x rays differs from unity by  $\sim 5 \times 10^{-6}$ , it follows that the crystal surfaces must be smooth to  $\sim \pm 10^{-4} \text{ cm}$  over the areas where the beams interact with the crystal. This condition is not difficult to achieve.

For the case of two  $1.42\text{-\AA}$  neutron beams interfering constructively, the integrated intensity is  $2.5 \times 10^{-7}$ . With neutron beams currently available at high-flux reactors, the counting rate at the detector should be tens of counts per second. The dynamic index of refraction for neutrons differs from unity by  $\sim 10^{-7}$ . Consequently the requirements on surface flatness are less stringent ( $\sim \pm 50 \times 10^{-4} \text{ cm}$ ) than that for x rays.

We conclude that the proposed experiment is feasible. It is not obvious (to us) that all conceivable theories of gravitation will lead to the fringe shift, Eq. (8),<sup>7</sup> predicted by a scalar Newtonian potential. An experimental result in disagreement with Eq. (10) would question the validity of the principle of equivalence<sup>8</sup> in the quantum limit. Agreement could of course be regarded as a check of the equivalence principle away from the classical limit,  $\hbar \rightarrow 0$ .

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<sup>1</sup>U. Bonse and M. Hart, Z. Phys. **194**, 1 (1966).

<sup>2</sup>Successful construction of a neutron interferometer has been reported recently: H. Rauch, W. Treimer, and U. Bonse, Phys. Lett. **47A**, 369 (1974).

<sup>3</sup>The central peak of the Bragg reflection (which has a Darwin width of 0.5 sec) is totally reflected at  $B$ ,  $C$ , and  $E$  and emerges from the interferometer along a line parallel to  $DE$ . The rocking curve of the interferometer consists of two 0.08-sec wide peaks separated by the 0.5-sec Darwin gap.

<sup>4</sup>A. W. McReynolds, Phys. Rev. **83**, 172 (1951); J. W. T. Dabbs, J. A. Harvey, D. Paya, and H. Horstmann, Phys. Rev. **139**, B756 (1965).

<sup>5</sup>The existence of multiply reflected beams within the beam splitter has been neglected. For the case of neutrons these may have appreciable intensity because absorption is small. They can be eliminated (for narrow beams) by suitable shielding. An important point in the theory (Ref. 1) is that the phase difference between the two beams at  $F$  does not depend on the precise angle of incidence (within the 0.2-sec angular window). This implies that the incident wave at  $A$  can be spheri-

cal.

<sup>6</sup>Unit integrated intensity is defined to be 100% of the incident beam arriving at the detector throughout an angular rotation of the crystal of 1 rad.

<sup>7</sup>We emphasize that Eq. (8) applies to an interferometer for which the beam splitter is very thin. Rays are refracted as they enter and leave the beam splitter (near *B* and *E*). Phase differences caused by this phe-

nomenon could not arise were the angle  $\epsilon - \delta$ , given by Eqs. (4) and (6), equal to zero. Since  $\epsilon - \delta \sim 10^{-8}$ , a small correction to Eq. (8) may result for a beam splitter of finite thickness, depending on design.

<sup>8</sup>Equation (8) can also be derived when no gravitational field is present provided the nuclear reactor, beam port, etc., and the interferometer have a uniform acceleration *g*.

## Multiplicity Distributions and Multiplicity Scaling in $p + p \rightarrow p + \text{MM}$ at 28.5 GeV/c\*

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The charged particles associated with a fast forward trigger proton of four-momentum transfer *t* in the reaction  $p + p \rightarrow p + \text{MM}$  at 28.5 GeV/c were measured using the Multiparticle Argo Spectrometer System. Charged-particle multiplicity distributions show multiplicity scaling with respect to the missing mass, MM, to ~10% within each of two *t* regions,  $0 < |t| < 2$  (GeV/c)<sup>2</sup> and  $3 < |t| < 5$  (GeV/c)<sup>2</sup>. The width of the scaling curve shrinks between regions.

We previously reported the behavior of the average number of charged particles,  $\langle n_c \rangle$ , produced in the reaction

$$p_1 + p_2 \rightarrow p_3 + \text{MM} \quad (1)$$

at 28.5 GeV/c.<sup>1</sup> We observed a sharp rise in  $\langle n_c \rangle$  with increasing *t*, the four-momentum transfer from *p*<sub>1</sub> to *p*<sub>3</sub>, for fixed mass MM recoiling against *p*<sub>3</sub>. Here we describe certain regularities in the shapes of the charge multiplicity distributions,  $P_{n_c}(\text{MM}, t)$ , and examine the scaling properties in MM and *t* and their correlation with the rise in  $\langle n_c \rangle$ . Such observations may be helpful in understanding the dynamical mechanism of particle production in high-energy *pp* collisions.

Our data base consists of ~200 000 events of Reaction (1). The data were taken at Brookhaven National Laboratory by the Multiparticle Argo Spectrometer System (MASS).<sup>2</sup> The high-momentum spectrometer (HMS) momentum analyzed and identified *p*<sub>3</sub>, and the vertex spectrometer (VS)<sup>3</sup> momentum analyzed the remaining charged particles. The VS tracks were reconstructed by the computer code PITRACK<sup>4</sup> and fitted to a common vertex. Corrections were applied<sup>1</sup> to the multiplicity distributions to compensate for loss of charged particles due to limited solid angle (~10%), low-momentum cutoff (~1%), and software inefficiencies (~6%).

The probability for a given charged multiplicity

associated with a trigger proton of given MM and *t* is

$$P_{n_c}(\text{MM}, t) \equiv \frac{d^2 \sigma_{n_c} / d(\text{MM}) dt}{\sum_i d^2 \sigma_i / d(\text{MM}) dt} \quad (2)$$

The data, which cover the range  $1.2 < \text{MM} < 6.5$  GeV and  $0.2 < |t| < 6.5$  (GeV/c)<sup>2</sup>, have been divided into 31 bins of MM and *t* by taking six intervals in MM and seven in *t*. For each bin we calculated the average charged multiplicity in the final state,  $\langle n_c \rangle$ , the dispersion,  $D^2 \equiv \langle n_c^2 \rangle - \langle n_c \rangle^2$ , and two higher normalized moments,  $C_i \equiv \langle n_c^i \rangle / \langle n_c \rangle^i$ ,  $i = 3, 4$ , of  $P_{n_c}(\text{MM}, t)$ ; the results are given in Table I. In analyzing the data with finer binning<sup>5</sup> one sees that for fixed MM > 2 GeV, the average charge multiplicity is approximately constant for  $|t| < 2$  (GeV/c)<sup>2</sup>, then rises abruptly by  $\Delta n_c \sim 0.6$  charged particles and again becomes approximately constant for  $|t| > 3$  (GeV/c)<sup>2</sup>. This constancy of  $\langle n_c \rangle$  for fixed MM in the two *t* regions is also observed<sup>5</sup> in the individual  $P_{n_c}$ . The magnitude of the increase and the *t* value at which it is centered are approximately constant for all intervals of MM > 2 GeV. We found that the rise is centered at  $|t| \sim 2.5$  (GeV/c)<sup>2</sup> and occurs in a characteristic interval  $\Delta t \sim 1$  (GeV/c)<sup>2</sup>.

Figure 1 is a plot of  $P_{n_c}(\text{MM}, t)$  versus MM averaged over two *t* intervals on either side of the rise in  $\langle n_c \rangle$ : Region I has  $1 < |t| < 2$  (GeV/c)<sup>2</sup> and region II has  $3 < |t| < 5$  (GeV/c)<sup>2</sup>. Compari-