Attenuation of the Coriolis Interaction within the Cranking Model*

P. Ringt

Lawrence Berkeley Laboratory, University of California, Berkeley, California 94720

and

H. J. Mang

Physikdepartment der Technischen Universität München, 8046 Garching, West Germany (Received 19 August 1974)

The description of strongly distorted rotational bands within the cranking model allows an interpretation of the attenuation factors used in the particle-plus-rotor model. It turns out that they are not very much influenced by the residual interaction, but are strongly dependent on the angular momentum. A simple model is proposed to calculate distorted spectra which is in rather good agreement with the experimental data and with the fully self-consistent calculation.

The description of very distorted rotational bands of odd-mass deformed nuclei is possible within the particle-plus-rotor model' (PRM) by the coupling of particles to the collective rotation using a Coriolis interaction. Practical calculations, however, allow a reproduction of the experimental data only by reducing the strength of this interaction. The attenuation factors R used for this purpose lie between 0.4 and 0.9.² There exist attempts to give an interpretation of these factors by taking into account the coupling of the outside particle to collective vibrations of the core.³ Recently⁴ it has been shown that the application of the cranking model within the framework of the Hartree-Pock-Bogolyubov (HFB) theory, which can be derived from a projection of the angular momentum before the variation,⁵ allows a quantitative description of these bands without any fit parameter. In particular, no extra attenuation of the Coriolis term has to be introduced. In the present paper we show to what extent this description can be compared with the PRM and why in the latter model the Coriolis term has to be attenuated. A very simple model is introduced, where the outside particle is coupled to a cranked core. It is justified by the self-consistent calculation and agrees very well with the experiment. As a numerical example the very distorted band with positive parity in 159 Dy is investigated.

Within the cranking model the internal wave function φ_{α} of the odd nucleus is calculated by the variational equation

$$
\langle \delta \varphi_{\alpha} | \hat{H} - \omega \hat{J}_x - E^{\alpha} | \varphi_{\alpha} \rangle = 0. \tag{1}
$$

If one restricts φ_{α} to the HFB functions, it corresponds to the blocked HFB equations' in the rotating frame. Qne has to look for solutions of

this system which have odd particle-number parity' and which are eigenfunctions of a rotation about 180 $^{\circ}$ around the x axis:

$$
\exp(i\pi J_x)\varphi_\alpha = i(-1)^{I-1/2}\varphi_\alpha. \tag{2}
$$

 I is the total angular momentum and the cranking frequency ω is determined by the subsidiary condition⁸

$$
\langle \varphi_{\alpha} | J_{x} | \varphi_{\alpha} \rangle^{2} + \langle \varphi_{\alpha} | J_{z}^{2} | \varphi_{\alpha} \rangle = I(I+1). \tag{3}
$$

Equation (1) is solved directly in Ref. 4 for 159 Dy. For comparison with the particle-plus-rotor model, however, it is useful to decompose φ_{α} ,

$$
\varphi_{\alpha} = \gamma_{\alpha}^{\dagger} |\varphi_0\rangle = \sum_{K} C_K^{\alpha} \beta_K^{\dagger} |\varphi_0\rangle, \tag{4}
$$

where φ_0 is the underlying HFB wave function of the even core and β_K^{\dagger} are the quasiparticle operators corresponding to this core, which diagonalize the Hamiltonian ($H¹¹$ diagonal). For $\omega = 0$, K is a good quantum number (it corresponds to the eigenvalue of J_s) because of the axial symmetry of the core. For higher ω this is not true. In the numerical calculation of 159 Dy (see Fig. 1) however, it turns out that for a large region of spin values $(I<\frac{21}{2})$ the core stays nearly axially symmetric and K is a rather good quantum number. The variation (1) is therefore decomposed into a variation of the core function φ_0 and a variation of the mixing coefficients C_K^{α} ,

$$
\langle \delta \varphi_0 | \gamma_\alpha (H - \omega J_x) \gamma_\alpha^{-\dagger} | \varphi_0 \rangle = 0, \tag{5}
$$

$$
\sum_{K} \left\{ \left[\langle \varphi_{0} | H - \omega J_{x} | \varphi_{0} \rangle + E_{K}(\omega) \right] \delta_{K K'} - \omega (j_{x}^{11})_{K K'} \right\} C_{K'}^{\alpha} = E^{\alpha} C_{K}^{\alpha}; \qquad (6)
$$

 j_x^{11} is the one-quasiparticle part of J_x corresponding to the operators β_K^{\dagger} . Equation (5) corresponds to blocked HFB equations for the determination of the core wave function φ_0 with even

FIG. 1. The dependence of (a) the quasiparticle energies E_K , and (b) the matrix elements j_{KK+1} ¹¹ of Eq. (6) on the angular momentum I; $K \approx -\frac{1}{2}$ corresponds to the decoupling parameter. Full lines correspond to the favored solutions, dashed lines to the unfavored solutions (see Ref. 9).

number parity. It is coupled by the blocking of γ_{α} to Eq. (6) which determines the mixing amplitudes C_{κ}^{α} .

The PRM replaces the calculation of φ_0 by assuming a rotor with a fixed moment of inertia. Equation (6) corresponds to the diagonalization of the PRM for the calculation of the mixing amplitudes.

Besides the fact that the cranking model gives energies in the rotating frame, there is a close analogy between Eq. (6) and the PRM concerning the amplitudes C_K^{α} . (a) With the neglect of constants, the diagonal elements are in both cases essentially the quasiparticle energies E_{κ} . In the cranking model they depend on ω , but only very weakly, as shown in Fig. 1(a). (b) The nondiagonal elements vanish exactly for $K \neq K' \pm 1$ in the PRM and approximately in the cranking model. In the latter model the frequency is ω $=\langle \varphi_{\alpha} | J_{x} | \varphi_{\alpha} \rangle / \mathcal{G}_{sc}$. \mathcal{G}_{sc} is the self-consistently determined moment of inertia. With regard to Eq. (3) the elements $K' = K + 1$ are

$$
-\frac{\left[I(I+1)-\langle J_z^2\rangle\right]^{1/2}}{g_{\rm sc}}\left[j_x^{11}(\omega)\right]_{K_{K+1}}
$$
(7a)

in the cranking model, and

$$
-\frac{\left[I(I+1)-K(K+1)\right]^{1/2}}{g_{\text{rotor}}} \left[j_x^{11}(\omega=0)\right]_{KK+1} \tag{7b}
$$

in the PRM.¹ Both expressions are very similar. If one neglects the small ω dependence of the matrix elements j_{x}^{11} in the cranking model [see Fig. 1(b)] and the fact that $[I(I+1) - K(K+1)]^{1/2}$ is replaced by $[I(I+1)-\langle J_z^2\rangle]^{1/2}$ in the cranking model, there remains only one big difference between both Coriolis interactions, explaining why one needs attenuation factors in the PRM but not in the cranking model: The cranking model uses a self-consistently determined moment of inertia \mathfrak{g}_{sc} , which includes the effect of the decoupling particle and which is strongly I dependent (see Fig. 2). For small I values, where the particle is coupled to the core, it is very easy to gain angular momentum in the x direction by decoupling the particle. Therefore, the value of g_{sc} is large and the Coriolis interaction is strongly attenuated. This effect can also be seen in the simple Inglis formula for the odd nucleus in the state α .

$$
g_{\alpha}^{\text{Ing}} = \sum_{K,K' \neq \alpha} \frac{\left| \left(J_x^{20} \right)_{KK'} \right|^2}{E_K + E_{K'}} + \sum_{K \neq \alpha} \frac{\left| \left(j_x^{11} \right)_{K\alpha} \right|^2}{E_K - E_{\alpha}} \,. \tag{8}
$$

The first part comes from the core. The second part describes the particle. Because of the small energy denominator it can become much larger than the first part. In the case of 159 Dy, we found $\mathbf{g}_{sc} = 26.82 + 96.53 = 123.35 \text{ MeV}^{-1}$.

However, for higher spin values a perturbationtheoretic treatment is no longer possible. The

FIG. 2. Moments of inertia dependent on the angular momentum. $g_{sc} = \langle \varphi_{\alpha} | J_x | \varphi_{\alpha} \rangle / \omega$ and $g_{c \text{ore}} = \langle \varphi_0 | J_x | \varphi_0 \rangle$ correspond to the many-body wave function $(K0_c$ in Fig. 3). \mathcal{I}_0 and $\mathcal{I}_0 + \mathcal{I}_b$ correspond to the particle-plus-cranking model [see Eq. (10)].

exact solution (see Fig. 2) shows that the particle is more and more aligned and its contribution to the moment of inertia becomes smaller and smaller. Therefore the self-consistent moment of inertia \mathfrak{g}_{sc} diminishes with increasing spin. Only for very high spin values should it increase again because of the antipairing and the stretching effect of the core.

Figure 3 shows the experimental spectrum of the positive-parity band in ¹⁵⁹Dy and different calculations. $K1_{sc}$ is the fully self-consistent solution of Eq. (1) as described in Ref. 4. It uses a pairing-plus-quadrupole force including the exchange term of the quadrupole-quadrupole force, its contribution to the pairing potential, and the contributions of the pairing force to the self-consistent field. $K2_{sc}$ uses a similar force, which does not include the latter three terms and which is adjusted to reproduce the same energy gap and the same deformation:

Units and details are given in Ref. 4.

In the column KO_c the influence of the residual interaction is neglected; i.e., the calculation is

FIG. 3. The positive-parity band in 159 Dy: Experiment (see Ref. 10) and different calculations as described in the text.

done within constant fields Γ and Δ taken from $K2_{sc}$ at $\omega = 0$. This procedure changes the behavior of the spectrum at very high spin values. However, the attenuation of the Coriolis interaction is only very little influenced by the residual interaction.

We studied it in the following simple model (column labeled crank+part in Fig. 3) suggested by Eq. (6) . One outside particle is coupled to a rotor with moment of inertia g_0 :

$$
\langle \varphi_0 | H - \omega J_x | \varphi_0 \rangle = \frac{1}{2} \mathbf{g}_0 \omega^2.
$$
 (9)

Neglecting the ω dependence of E_K and $j_{KK'}$ ¹¹ one has to diagonalize

$$
\frac{1}{2}g_0\omega^2 + E_K - \omega(j_x^{11})_{KK'}.
$$
 (10)

The subsidiary condition for ω is

$$
\mathcal{G}_0 \omega + \langle j_x \rangle = [I(I+1) - \langle j_z^2 \rangle]^{1/2}.
$$
 (11)

Therefore the Coriolis interaction can be written as

$$
H_{\text{Cor}} = -\omega j_x^{11} = -\frac{[I(I+1) - \langle j_z^2 \rangle]^{1/2} j_x^{11} R}{\vartheta_0}.
$$
 (12)

Compared to the PRM it is attenuated by a factor '

$$
R = \frac{g_0}{g_0 + g_p} = 1 - \frac{\langle j_x \rangle}{[I(I+1) - \langle j_z^2 \rangle]^{1/2}}.
$$
 (13)

 $\mathcal{G}_b = \langle j_x \rangle / \omega$ is the contribution of the outside particle (see Fig. 2). Taking into account that there was no fit parameter used $(\mathcal{G}_0$ is taken from KO_c), we find the agreement of this simple model with the experiment and with the fully self-consistent calculation surprisingly good. The last two co1 umns in Fig. 3 are calculations within the PRM, without attenuation (column 7) and with a fit over four parameters (see Ref. 5).

The attenuation R is caused by the decoupling of the outside particle, which can be described very easily within the cranking model. It is strongly spin dependent (see Table I) and approaches 1 for high spin values. The "favored" states $I = \frac{5}{2}, \frac{9}{2}, \ldots$ are more attenuated than the unfavored ones $I = \frac{7}{2}, \frac{11}{2}, \ldots$. It should be emphasized that the attenuation [see Eq. (13)] is contained within the solution of the cranking model and there is no further parameter needed.

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)On leave of absence from Technische Universitat, München, West Germany.

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Thermal Fluctuations and Experiments on the Free Fall of Electrons

Humphrey J. Maris*

Physics Department, Brown University, Providence, Rhode Island 02912 (Received 29 July 1974)

In order to study the gravitational force on an electron Witteborn and Fairbank have measured the acceleration of electrons moving along the axis of a vertical copper tube. I show that in experiments of this type there are large corrections to the motion of the electrons due to thermal fluctuations in the electromagnetic fields in the tube.

In a remarkable series of experiments Witteborn and $Fairbank¹⁻⁵$ (WF) have attempted to measure the gravitational free fall of electrons in vacuum. They studied the motion of electrons along the axis;of a vertical "drift tube" of copper. Apart from known or applied electric fields, it was believed that all vertical electric or magnetic potential. energy gradients had been reduced to below about 10^{-25} erg cm⁻¹. This was necessary to observe the effects of gravity, since for an

electron in the earth's gravitational field the gradient of the gravitational potential energy is only 10^{-24} erg cm⁻¹. It had earlier been predicted by Schiff and Barnhill⁶ that within a metal enclosure, itself in the earth's gravitational field, there should be an induced electric field of magnitude mg/e (*m* is the electron mass) directed downwards. This field arises because the electrons in a metal mill be redistributed as a result of gravity. The Schiff-Barnhill field (SB) exerts a