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ϵ Expansion in Semi-infinite Ising Systems

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The Wilson-Fisher ϵ expansion is used to calculate critical exponents in a semi-infinite Ising model to first order in ϵ . It is found that potentials renormalize just as in the bulk, and all surface information is contained in the wave-function renormalization. ν is $\frac{1}{2} + \frac{1}{12}\epsilon$ and η is 0, just as in infinite systems. η_{\perp} and η_{\parallel} are, respectively, $1 - \frac{1}{6}\epsilon$ and $2 - \frac{1}{3}\epsilon$ and the surface gap exponent Δ_1 is $\frac{1}{2} - \frac{1}{12}\epsilon$ if scaling is assumed.

The Wilson renormalization procedure has been highly successful in calculating critical exponents in bulk or infinite systems.¹⁻³ In this paper we will outline a calculation of critical exponents in a semi-infinite Ising system using the ϵ expansion.^{2,3} Particular emphasis will be placed on the calculation of the exponents η_{\parallel} and η_{\perp} introduced by Binder and Hohenberg⁴ and the gap exponent Δ_1 introduced by Fisher.⁵ The major purpose of this paper is to present results. Computational details will appear in a subsequent publication.

We start with the standard S^4 Hamiltonian

$$\mathcal{H} = \frac{1}{2}b^2 \sum_{\vec{x}} S^2(\vec{x}) - \frac{1}{2}K \sum_{\vec{x}, \vec{\delta}} S(\vec{x})S(\vec{x} + \vec{\delta}) + u \sum_{\vec{x}} S^4(\vec{x}), \quad (1)$$

where $\vec{x} = (x_{\perp}, x_1, \dots, x_{d-1}) \equiv (x_{\perp}, \vec{x}_{\parallel})$ labels the sites on a semi-infinite cubic lattice and $\vec{\delta}$ is the nearest-neighbor position. \vec{x}_{\parallel} is the $(d-1)$ -dimensional vector parallel to the surface, and x_{\perp} the component of \vec{x} perpendicular to the surface. x_{\perp} takes on values $1, 2, 3, \dots$. All other components take on all positive and negative values. It is convenient to extend the sums in Eq. (1) to include the plane $x_{\perp}=0$ and to require that $S(0, \vec{x}_{\parallel})$ be zero. Then $S(\vec{x})$ can be expressed as a Fourier sine integral

$$S(\vec{x}) = \int_{-\pi}^{\pi} \frac{d^d \vec{p}}{(2\pi)^d} \sigma(\vec{p}) \exp(i\vec{p}_{\parallel} \cdot \vec{x}_{\parallel}) \sqrt{2} \sin p_{\perp} x_{\perp}, \quad (2)$$

where $\vec{p} = (p_{\perp}, \vec{p}_{\parallel})$. [The $\sqrt{2}$ appearing in Eq. (2) puts p_{\perp} and \vec{p}_{\parallel} on an equal footing in the Hamiltonian below.] Following Wilson, we introduce a modified Hamiltonian in which \vec{p} is restricted to a *unit cyl-*

inder D of radius 1 in the parallel direction and height 2 in the perpendicular direction:

$$\mathcal{H} = \frac{1}{2} \int \frac{d^d p}{(2\pi)^d} (r + p^2) \sigma(\vec{p}) \sigma(-\nu\vec{p}) + u \int_D \frac{d^d p_1 d^d p_2 d^d p_3 d^d p_4}{(2\pi)^{3d}} \sigma(\vec{p}_1) \sigma(\vec{p}_2) \sigma(\vec{p}_3) \sigma(\vec{p}_4) \delta^d(\vec{p}_1, \vec{p}_2, \vec{p}_3, \vec{p}_4), \quad (3)$$

where $\nu\vec{p} = (-p_\perp, \vec{p}_\parallel)$,

$$\delta^d(\vec{p}_1, \vec{p}_2, \vec{p}_3, \vec{p}_4) = \delta^{d-1}(\vec{p}_{1\parallel} + \vec{p}_{2\parallel} + \vec{p}_{3\parallel} + \vec{p}_{4\parallel}) \Delta(p_{1\perp}, p_{2\perp}, p_{3\perp}, p_{4\perp}), \quad (4)$$

and

$$\Delta(p_{1\perp}, p_{2\perp}, p_{3\perp}, p_{4\perp}) = \frac{4}{2\pi} \sum_{x_\perp > 0} \sin p_{1\perp} x_\perp \sin p_{2\perp} x_\perp \sin p_{3\perp} x_\perp \sin p_{4\perp} x_\perp. \quad (5)$$

The renormalization procedure on the Hamiltonian (3) is essentially the same as for infinite systems. Let D_1 be the cylinder of radius b^{-1} in the $d-1$ parallel directions and of height $2b^{-1}$ in the perpendicular direction, and let $D_2 = D - D_1$. Taking the trace over all $\sigma(\vec{p})$ with $\vec{p} \in D_2$, we obtain a new Hamiltonian, $\hat{\mathcal{H}}$, which is a function only of $\hat{\sigma}(\vec{p}) \equiv \sigma(\vec{p})$ for $\vec{p} \in D_1$. $\hat{\mathcal{H}}$ is given to second order in u by

$$\begin{aligned} \hat{\mathcal{H}}[\hat{\sigma}(\vec{p})] = & \frac{1}{2} \int \frac{d^d p}{(2\pi)^d} [p^2 + r + 12u C_1(r)] \hat{\sigma}(\vec{p}) \hat{\sigma}(-\nu\vec{p}) - \frac{1}{2} \int \frac{d^d p_1}{(2\pi)^d} \int \frac{d^d p_2}{(2\pi)^d} v(\vec{p}_1, \vec{p}_2) \hat{\sigma}(\vec{p}_1) \hat{\sigma}(\vec{p}_2) \\ & + [u - 36u^2 C_2(r)] \int \frac{d^d p_1 d^d p_2 d^d p_3 d^d p_4}{(2\pi)^{3d}} \hat{\sigma}(\vec{p}_1) \hat{\sigma}(\vec{p}_2) \hat{\sigma}(\vec{p}_3) \hat{\sigma}(\vec{p}_4) \delta^d(\vec{p}_1, \vec{p}_2, \vec{p}_3, \vec{p}_4) \\ & + \int \frac{d^d p_1 d^d p_2 d^d p_3 d^d p_4}{(2\pi)^{4d}} v_4(\vec{p}_1, \vec{p}_2, \vec{p}_3, \vec{p}_4) \hat{\sigma}(\vec{p}_1) \hat{\sigma}(\vec{p}_2) \hat{\sigma}(\vec{p}_3) \hat{\sigma}(\vec{p}_4), \end{aligned} \quad (6)$$

where all explicit integrations are over D_1 ,

$$v(\vec{p}, \vec{p}') = 6u \delta^{d-1}(\vec{p}_\parallel + \vec{p}'_\parallel) \int_{D_2} \frac{d^d p_3}{p_3^2 + r} [\delta(p_\perp + p'_\perp + 2p_{3\perp}) - \delta(p_\perp - p'_\perp + 2p_{3\perp})], \quad (7)$$

and

$$C_1(r) = \int_{D_2} \frac{d^d p}{(2\pi)^d} \frac{1}{p^2 + r}, \quad C_2(r) = \int_{D_2} \frac{d^d p}{(2\pi)^d} \frac{1}{(p^2 + r)^2}; \quad (8)$$

$v(\vec{p}_1, \vec{p}_2)$ is a nonlocal two-spin potential, and v_4 contains all four-spin interactions except for the $\vec{p} = 0$ local part.

We now perform a series of three transformations on $\hat{\mathcal{H}}$ and $\hat{\sigma}$ to produce a new Hamiltonian \mathcal{H}' and a new spin variable σ' such that $\mathcal{H}'(\sigma')$ looks as much as possible like the original Hamiltonian. The first transformation eliminates the nonlocal two-spin interaction from $\hat{\mathcal{H}}$. Let $\hat{\sigma}(\vec{p}) = \hat{\sigma}'(\vec{p}) + \delta\hat{\sigma}(\vec{p})$. If $\delta\hat{\sigma}(\vec{p})$ satisfies

$$\delta\hat{\sigma}(\vec{p}) = \frac{1}{2p^2} \int_{D_1} \frac{d^d p'}{(2\pi)^d} v(\vec{p}, \vec{p}') \hat{\sigma}'(-\nu\vec{p}'), \quad (9)$$

then $\hat{\mathcal{H}}[\hat{\sigma}'(\vec{p})]$ contains no nonlocal two-spin interaction to first order in u . The second transformation takes \vec{p} to $b\vec{p}$ so that integrations are over D rather than D_1 . Finally, the third transformation renormalizes $\hat{\sigma}'$ so that the coefficient of the p^2 term is one half:

$$\hat{\sigma}'(b^{-1}\vec{p}) = \zeta \sigma'(\vec{p}). \quad (10)$$

To order ϵ , ζ is $b^{1+d/2}$ which is the same as for an infinite system. The final transformed Hamiltonian has the same form as Eq. (3) with new potentials

$$r' = b^2[r + 12u C_1(r)], \quad (11a)$$

$$u' = b^\epsilon[u - 36u^2 C_2(r)], \quad (11b)$$

where $\epsilon = 4 - d$. These are the same as the bulk renormalization equations.² Equation (11) yields a fixed-point value of u to first order in ϵ of $u^* = (2\pi^2/9)\epsilon$ and a coherence-length exponent ν equal to $\frac{1}{2} + \frac{1}{12}\epsilon$ which is the same as in infinite systems. Hence, there is a single divergent correlation length in a semi-infinite system which diverges in the same way as in an infinite system. This is consistent

with previous assertions.^{4,5,7-11} The nonlocal four-spin potential v_4 can be shown to be irrelevant⁶ to first order in ϵ .

We have just seen that potentials in a semi-infinite system renormalize in exactly the same way as potentials in a bulk system, at least to first order in ϵ . All information about the surface is lumped into the wave-function renormalization. In bulk systems, wave-function renormalization is characterized completely by a single number ζ . In the semi-infinite system, ζ becomes an operator $\zeta(\vec{p}\vec{p}')$ resulting from transformations one and three, given by

$$\zeta(\vec{p}, \vec{p}') = \zeta[(2\pi)^d \delta(\vec{p} - \vec{p}') + (1/2\rho^2)v(-\nu\vec{p}, \vec{p}')] \tag{12}$$

At the fixed point, the correlation function $\Gamma^*(\vec{p}\vec{p}') = \langle \sigma(\vec{p})\sigma(\vec{p}') \rangle$ satisfies

$$\Gamma^*(\vec{p}, \vec{p}') = \int \frac{d^d p_1}{(2\pi)^d} \int \frac{d^d p_2}{(2\pi)^d} \zeta^*(\vec{p}, \vec{p}_1) \zeta^*(\vec{p}', \vec{p}_2) \Gamma^*(b\vec{p}_1, b\vec{p}_2), \tag{13}$$

where the star indicates the fixed-point values of $\zeta(\vec{p}\vec{p}')$ and $\Gamma(\vec{p}\vec{p}')$. The solution to Eq. (13) for $d = 4 - \epsilon$ to first order in ϵ is

$$\Gamma^*(\vec{p}\vec{p}') = \Gamma_0^*(\vec{p}\vec{p}') + \frac{1}{\rho\rho'} (2\pi)^{d-1} \delta^{d-1}(\vec{p}_\parallel, \vec{p}'_\parallel) \gamma^*(\vec{p}, \vec{p}'), \tag{14}$$

where $p = (p_\perp^2 + p_\parallel^2)^{1/2}$, $\Gamma_0^*(\vec{p}\vec{p}')$ is the Gaussian fixed-point function

$$\Gamma_0^*(\vec{p}\vec{p}') = \frac{(2\pi)^d}{2} \frac{1}{\rho^2} [\delta(\vec{p} + \vec{p}') - \delta(\vec{p} + \nu\vec{p}')], \tag{15}$$

and

$$\gamma^*(\vec{p}\vec{p}') = \frac{3}{4\pi^2} u^* \frac{1}{\rho\rho'} \left\{ (p_\perp + p_\perp') \tan^{-1} \frac{2}{p_\perp + p_\perp'} - (p_\perp - p_\perp') \tan^{-1} \frac{2}{p_\perp - p_\perp'} \right\}. \tag{16}$$

It should be stressed that the factors of 2 appearing in \tan^{-1} in Eq. (16) result from reflections off of the surface and are unrelated to any cutoff. Fourier transformation of Eq. (14) yields the spatial correlation function

$$\Gamma^*(\vec{x}, \vec{x}') = \Gamma_0^*(\vec{x}, \vec{x}') + \Gamma_1^*(\vec{x}, \vec{x}'), \tag{17}$$

where Γ_0^* is the Gaussian half-space propagator,

$$\Gamma_0^*(\vec{x}, \vec{x}') = \frac{\Gamma(\frac{1}{2}d - 1)}{4\pi^{d/2}} \left(\frac{1}{|\vec{x} - \vec{x}'|^{d-2}} - \frac{1}{|\vec{x} - \nu\vec{x}'|^{d-2}} \right), \tag{18}$$

and

$$\Gamma_1^*(\vec{x}, \vec{x}') = \frac{3u^*}{8\pi^3 \rho} \{ J_+(x_\perp - x_\perp' + i\rho, x_\perp') - J_-(x_\perp + x_\perp' + i\rho, x_\perp') + J_+(x_\perp' - x_\perp + i\rho, x_\perp) - J_-(x_\perp + x_\perp' + i\rho, x_\perp) - J_+(x_\perp + x_\perp' + i\rho, 0) + J_-(x_\perp x_\perp' + i\rho, 0) \}, \tag{19}$$

where $\rho = |\vec{x}_\parallel - \vec{x}'_\parallel|$ and

$$J_\pm(\eta, x_\perp) = \text{Re} \int_0^2 dy \int_0^\infty (dp_\perp/2\pi) [\exp(ip_\perp \eta) \exp(-yx_\perp)/(y \pm 2ip_\perp)]. \tag{20}$$

As both x_\perp and x_\perp' go into the bulk, i.e., $x_\perp, x_\perp' \rightarrow \infty$ with $x_\perp - x_\perp'$ finite, $\Gamma_0^*(xx')$ becomes the bulk Gaussian (or mean field) propagator^{4,9} $(1/4\pi^2)|\vec{x} - \vec{x}'|^{-d+2}$, and Γ_1^* dies off as

$$\frac{1}{(x_\perp + x_\perp')^2} \ln \frac{x_\perp + x_\perp'}{|\vec{x} - \vec{x}'|}.$$

Hence, the fixed-point correlation function in the bulk becomes the same as for infinite systems with $\eta = 0$ to first order in ϵ . Note, however, that the effect of the surface dies off very slowly at

the fixed point. This corroborates results obtained by high-temperature expansions.⁴

Two other limits are of interest. In the first, \vec{x}' is fixed on the surface and \vec{x} goes to infinity in the bulk at an angle θ to the normal to the surface. In this case, Γ^* behaves like

$$\Gamma^*(\vec{x}, \vec{x}') \sim A(\theta)/|\vec{x} - \vec{x}'|^{d-2+\eta_\perp}. \tag{21}$$

By comparing $|\vec{x} - \vec{x}'|^{-3} \ln |\vec{x} - \vec{x}'|$ terms in Eqs.

(19) and (21), we obtain

$$\eta_{\perp} = (1 - \frac{1}{8}\epsilon). \quad (22)$$

In the Gaussian limit, $A(\theta)$ is $\pi^{-2} \cos\theta$. Corrections to $A(\theta)$ to order ϵ can be obtained from $|\vec{x} - \vec{x}'|^{-3}$ terms in Eq. (21). They will be presented in a more detailed subsequent publication. In the second limit of interest, both \vec{x} and \vec{x}' are on the surface and

$$\Gamma^*(\vec{x}, \vec{x}') \sim B/\rho^{d-2+\eta_{\parallel}}. \quad (23)$$

Again comparing $\rho^{-4} \ln\rho$ terms in Eqs. (19) and (23), we obtain

$$\eta_{\parallel} = (2 - \frac{1}{3}\epsilon). \quad (24)$$

These are to be compared with the mean-field and spherical-model⁸ values of $\eta_{\parallel} = 2$ and $\eta_{\perp} = 1$.¹²⁻¹⁴ They are consistent with the numerical calculations of Binder and Hohenberg⁴ on the three-dimensional Ising model in which η_{\parallel} and η_{\perp} are less than their mean-field values.

Using the scaling relation derived by Binder and Hohenberg,⁴ we can calculate the exponent γ_1 governing the variation of the surface magnetization m_s with bulk magnetic field h , $\chi_1 = \partial m_s / \partial h \sim t^{-\gamma_1}$, where $t = |T - T_c|$:

$$\gamma_1 = \nu(2 - \eta_{\perp}) = \frac{1}{2} + \frac{1}{8}\epsilon. \quad (25)$$

γ_1 can also be calculated from a scaling form for the surface free energy^{5, 9, 10}

$$F^X(t, h_s, h) = |t|^{2-\alpha^X} Q^X(h_s/t^{\Delta_1}, h/t^{\Delta}), \quad (26)$$

where h_s is the surface magnetic field, $\alpha^X = \alpha + \nu$ is the surface specific-heat exponent, and $\Delta = \beta + \gamma = \beta\delta$ and Δ_1 are, respectively, the bulk and surface gap exponents. We have

$$\chi_1 = \partial^2 F / \partial h_s \partial h \sim |t|^{-\gamma_1},$$

where

$$\gamma_1 = \Delta_1 + \nu - \beta. \quad (27)$$

Comparing Eqs. (25) and (27), we obtain

$$\Delta_1 = \frac{1}{2} - \frac{1}{12}\epsilon. \quad (28)$$

Extrapolation of this result to $\epsilon = 1$ gives $\Delta_1 = 0.416$. Binder and Hohenberg^{4, 9} estimate γ_1 for the three-dimensional Ising model to be about $\frac{7}{8}$ from high-temperature expansions. This gives $\Delta_1 \approx 0.5$ in disagreement with our result evaluat-

ed at $\epsilon = 1$. Clearly, near $d = 3$, ϵ^2 terms are important. At the moment, it is difficult to estimate the magnitude of these higher order terms within the context of the ϵ expansion. If we terminate the ϵ expansion for η_{\perp} at ϵ^2 and write $\eta_{\perp} = 1 - \frac{1}{8}\epsilon + c\epsilon^2$ and equate this to 0.64 for $\epsilon = 1$, we find $c \sim -0.2$. A coefficient of ϵ^2 of this size is perfectly reasonable within the ϵ expansion.

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