comes very close to producing an accidental tricritical point. Near the triple point F(m) is a very flat function of m, and it is likely that in this region of the phase diagram large fluctuations will make the Hartree-Fock approximation used here very inaccurate. Associated with the first-order transition is a discontinuous entropy change ΔS and a latent heat $l = T \Delta S$. Since in this model the moments disappear rather than disorder, the entropy change has no particular relation to $N \ln 2.5$ Because we have chosen a narrow conduction band, the entropy change over most of the first-order phase transition is close to $N \ln 4$ since there are 4N single electron states in the conduction band. For a more realistic model, ΔS would probably be considerably less. Similarly, if we had used a conduction band of finite width, the susceptibility above and below the first-order transition would be roughly independent of temperature, but a discontinuity in the susceptibility of either sign could occur at the transition temperature.

This simplified model, in which only the Heisenberg interaction was included (to demonstrate its effect) and in which both bandwidth and hybridization effects have been neglected, yields qualitatively good agreement with experiments on NiS. An arbitrary density of states for the itinerant electrons, an interband electron-electron interaction (in the Hartree-Fock approximation), and more complicated localized states $(S \neq \frac{1}{2})$, can be incorporated into the model merely at the expense of computer time.⁶ Hybridization may be a small effect for rare-earth compounds exhibiting MNM transitions, but for the transitionmetal compounds its inclusion may be essential for a realistic model. The difficulty in including hybridization is a major stumbling block in attaining an improved two-band model of the MNM transition.

The author wishes to thank Professor R. G. Goodrich and Professor A. K. Rajagopal for their discussions and helpful suggestions.

¹Two good comprehensive reviews of this subject are D. Adler, Rev. Mod. Phys. <u>40</u>, 714 (1968); N. F. Mott, Rep. Progr. Phys. <u>33</u>, 881 (1970).

²N. F. Mott, Proc. Phys. Soc., London, Sect. A <u>62</u>, 416 (1949).

⁴S. V. Tyablikov, *Methods in the Quantum Theory of Magnetism* (Plenum, New York, 1967).

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Germanium-64

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The new isotope ⁶⁴Ge has been produced via the reaction ⁶⁴Zn(³He,3*n*)⁶⁴Ge, chemically isolated, and its decay studied. Since ⁶⁴Ge is therefore nucleon stable, it can in principle have an important role in the synthesis of mass 64 by α capture during explosive stellar events, as proposed by Arnett, Truran, and Woosley. However, from the measured half-life (62.3±2.0 sec) and β -decay systematics, it appears that this mechanism is probably responsible for a negligible fraction of the observed abundance of mass 64.

Recently Arnett, Truran, and Woosley¹ (ATW) have shown that the elements on the high-mass side of the iron peak in the elemental abundance curve can be synthesized in an explosive stellar event. The process involves, for example, a quasiequilibrium among ⁴He, ⁵⁶Ni, ⁶⁰Zn, and ⁶⁴Ge, and the resulting abundances for mass 60 and 64 depend primarily on the corresponding binding energies. Since ⁶⁴Ge has never been observed ex-

perimentally, ATW used for that isotope a value of the binding energy calculated by Garvey *et al.*² The abundances obtained¹ for masses near 64 are typically 2 orders of magnitude smaller than those observed in nature.

This discrepancy could possibly be removed if 64 Ge were more tightly bound, since production of that nuclide by radiative α capture would then be enhanced. In that hope, a number of intensive

³J. C. Slater, Phys. Rev. 82, 538 (1951).

searches for ⁶⁴Ge were undertaken³⁺⁷ with a view to obtaining an experimental estimate of its mass. The fact that none of these searches was successful raised doubts about the nucleon stability of ⁶⁴Ge and its possible contribution to the synthesis of mass 64.

This Letter reports the detection of ⁶⁴Ge produced via the reaction ⁶⁴Zn(³He, 3n)⁶⁴Ge near 50 MeV. Conclusive identification has been made by (a) chemical separation of germanium, (b) observation of γ rays from the decay of ⁶⁴Ge to known excited states in the daughter nucleus ⁶⁴Ga, and (c) observation of the growth of the daughter activity.

In view of the expected low yield from the reaction ${}^{64}\text{Zn}({}^{3}\text{He},3n){}^{64}\text{Ge}$, it is fortunate that the ground state of ${}^{64}\text{Ga}$ is 0^+ , 8 since the isospin selection rule⁹ then requires most of the β decay of ${}^{64}\text{Ge}$ to lead to excited states in ${}^{64}\text{Ga}$. Furthermore, one expects the reaction ${}^{64}\text{Zn}(p,n){}^{64}\text{Ga}$ to populate many of the same states. γ rays from that reaction have been studied in beam by the present authors, 10 by Davids, Matthews, and Whitmire, 11 and by Hansen, Gregory, and Dietrich.¹² The latter two investigations included threshold measurements, and a level scheme for ⁶⁴Ga has been constructed^{11,12} on that basis.

Sources of ⁶⁴Ge were prepared by irradiating 10-mg/cm² targets of 99.66%-enriched ⁶⁴Zn on thick copper backings with the ³He beam from the Michigan State University cyclotron. The beam energy was degraded from 70 to 50 MeV with a Zn absorber. Following a 2-min irradiation the target was transported by pneumatic "rabbit" to a laboratory where (with the adaption of the method of Porile¹³) the ⁶⁴Zn was dissolved in concentrated HCl containing KClO₃. The volatile GeCl, was vacuum distilled at room temperature into a cold trap in front of a Ge(Li) detector. No evidence of activities other than germanium isotopes and their daughters was seen in the spectra except for ¹⁰C, which also forms a volatile tetrachloride. Counting was begun approximately 25 sec after the end of irradiation and continued for eight periods of 50.0 sec each.

Figure 1 shows portions of the 0-50-sec and 100-150-sec spectra resulting from a total of twelve irradiations. The strongest lines decay



FIG. 1. Low-energy portions of γ -ray spectra accumulated at two different times after the chemical separation of germanium. The rapid decay of ⁶⁵Ge ($t_{1/2}$ =30 sec) and the slower decay of the new isotope ⁶⁴Ge ($t_{1/2}$ =62 sec) are apparent. The growth of the 115-keV line from ⁶⁵Ga and the 992-keV line from ⁶⁴Ga, the daughter activities, can also be seen.

with a short half-life $(30 \pm 2 \text{ sec})$ and are attributed to ⁶⁵Ge (despite the large disagreement with the previous half-life¹³), on the basis of the rapid growth of the ⁶⁵Ga daughter and the good energy fit with levels observed¹⁴ in ⁶⁴Zn(³He,*d*)⁶⁵Ga.

Other lines, at 128.2 ± 0.2 , 384.1 ± 0.3 , 427.0 ± 0.3 , 667.1 ± 0.3 , and 774.5 ± 0.3 keV, decay with a (weighted average) half-life of 62.3 ± 2.0 sec and are assigned to the decay of the new isotope ⁶⁴Ge. All of these lines have been seen in the ⁶⁴Zn($p, n\gamma$)⁶⁴Ga experiments. A relative excitation function for ⁶⁴Ge and ⁶⁵Ge shows that lines from ⁶⁴Ge are weak at 29 MeV and undetectable at 20 MeV (the calculated threshold² is 21 MeV), while the ⁶⁵Ge γ rays remain strong. From the known excitation frunction for the reaction ⁶⁴Zn(³He,n)-⁶⁶Ge,⁵ the cross section for ⁶⁴Zn(³He,3n)⁶⁴Ge at 50 MeV is estimated to be 50 μ b.



FIG. 2. Upper graph, decay of the 427-keV γ ray from ⁶⁴Ge. The straight line represents the adopted half-life, 62.3 sec, fitted by least squares for intensity. Lower graph, growth and decay of the 992-keV line from the daughter activity ⁶⁴Ga. The solid curve is a calculation assuming the measured half-lives of ⁶⁴Ge and ⁶⁴Ga, fitted by least squares for the initial activities of the two isotopes. The initial activity of ⁶⁴Ga found in this way is, within error, zero; thus essentially all the observed ⁶⁴Ga results from ⁶⁴Ge decay.

In Fig. 2 are shown the decay of the 427-keV line from ⁶⁴Ge, and the growth and decay of the 992-keV line from the daughter ⁶⁴Ga, whose half-life is $159 \pm 2 \text{ sec.}^{15}$ Correction has been made for weak γ rays of 427 keV in the decays of ⁶⁶Ge and ⁶⁴Ga.

With the help of the 64 Zn $(p,n\gamma)$ 64 Ga results, a decay scheme for ⁶⁴Ge has been deduced as shown in Fig. 3. Absolute γ -ray intensities have been obtained by normalizing to the measured growth of ⁶⁴Ga daughter activity. Provided none of the observed γ rays is in cascade or appreciably converted, $(79 \pm 10)\%$ of the decay of ⁶⁴Ge has been accounted for. The 86-keV line is masked in these experiments by Pb x rays, and the 43-keV line lies below the threshold of the analog-to-digital converter. The 667-keV γ ray originates below 930-keV excitation¹² in ⁶⁴Ga, and the 775keV one below 1460 keV, but their exact locations are not known. Spin and parity assignments of 1⁺ have been made where $\log ft$ values indicate an allowed β transition. The log ft values shown have been calculated assuming mass excesses ΔM of - 58.83 MeV¹¹ for ⁶⁴Ga and - 54.03 MeV² for ⁶⁴Ge. This mass excess for ⁶⁴Ge seems quite realistic because the *ft* values so obtained are similar to those in neighboring nuclei. One can set an approximate lower limit on the mass excess by assuming that the $\log ft$ value for the transition to the 427-keV state is unlikely to be



FIG. 3. Decay scheme for 64 Ge. The notation and conventions of Nuclear Data Sheets have been adopted. The log*ft* values have been calculated assuming the mass excess for 64 Ge derived by Garvey *et al.* (Ref. 2). No direct measurement of the mass excess has been made.

less than 4.5, the smallest found in a survey of 134 transitions in the decays of ${}^{60-63}$ Zn, ${}^{64-68}$ Ga, and ${}^{66-69}$ Ge. This leads to $\Delta M({}^{64}$ Ge) ≥ -54.5 MeV and, correspondingly, an α -separation energy $S_{\alpha} \leq 2.7$ MeV.

This limiting value of S_{α} for ⁶⁴Ge is only 0.5 MeV greater than that assumed in the calculations of ATW; its use would cause only minor changes in the predicted abundances.¹⁶ Thus it appears that an (α, γ) capture chain proceeding along the N=Z line cannot synthesize the observed abundance of mass 64.

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Two-Neutron Transfer Reactions on Rare-Earth Nuclei

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The angular-momentum projection method is applied to the description of (t, p) and (p, t) reactions on deformed nuclei. The difference between the deformations of a target nucleus and that of a residual nucleus is taken into account in the calculation of the spectroscopic amplitude. The calculation is especially focused on the ground-state to ground-state (p, t) transitions. The neutron-number dependence of the calculated distorted-wave Born-approximation cross sections is in good agreement with the experimental results.

Two-neutron transfer reactions have been studied in transition and deformed nuclei in the rare-earth region, and a great deal of experimental data have been recently accumulated.¹⁻⁴ The experimental data have been analyzed by several authors⁵⁻¹¹ from theoretical viewpoints. In the framework of the plane-wave Born approximation, Yoshida⁵ has predicted that the cross section of the ground-state to ground-state transition for nuclei in superconducting states is proportional to $(\Delta/G)^2$, where Δ and G are the gap energy and pairing interaction strength, respectively. The enhancement of the cross section in the ground-state (g.s.) transition has been successfully explained by this theory. The calculation of the cross section in the distorted-wave Born approximation (DWBA)¹⁰ gave almost the same results as did the Yoshida prediction. However, those theories did not explain successfully the neutron-number dependence of the experimen-