

mania, Australia, August 1971 (to be published).

¹²W. R. Webber, in *Proceedings of the Eleventh International Conference on Cosmic Rays, Budapest, 1969*, edited by T. Gémesy *et al.* (Akademiai Kiado, Budapest, 1970), p. 275.

¹³N. L. Grigorov *et al.*, *Acta Phys.*, Suppl. 1, **29**, 518 (1970), and papers presented at the Twelfth International Conference on Cosmic Rays, Hobart, Tasmania, August 1971 (to be published). The Grigorov *et al.* results on proton, α -particle, and all particle fluxes are unusual and should lead to several inconsistencies with other cosmic-ray experiments. If the proton component has a "cutoff" below 1000 GeV then the charge composition of cosmic rays at 10^6 GeV would be mainly heavy nuclei. The energy fluctuations in the hadronic component of air showers at $E \geq 10^6$ GeV would be expected to be much narrower than observed [see, for

example, H. V. Bradt and S. A. Rappaport, *Phys. Rev. Lett.*, **22**, 960 (1969)], and in addition the extrapolated flux should be considerably less than that observed by air-shower experiments. It would also be difficult to account for deserved muon flux above TeV energies if the composition were mainly heavy nuclei.

¹⁴M. J. Ryan *et al.*, in *Proceedings of the Twelfth International Conference on Cosmic Rays, Hobart, Tasmania, Australia, August 1971* (to be published).

¹⁵K. Pinkau *et al.*, *Acta Phys.*, Suppl. 3, **29**, 291 (1970).

¹⁶V. A. Barger and R. J. N. Phillips, *Phys. Rev. Lett.*, **24**, 291 (1970).

¹⁷R. R. Amann, *Phys. Rev. D* **3**, 2861 (1971).

¹⁸H. Cheng and T. T. Wu, *Phys. Rev. Lett.*, **24**, 1456 (1970).

¹⁹M. Holder *et al.*, *Phys. Lett.*, **35B**, 361 (1971).

Neutrinos with Mass and the Decay $K_L^0 \rightarrow \bar{\nu}_l + \bar{\nu}_l$

Saul Barshay

The Niels Bohr Institute, University of Copenhagen, Copenhagen, Denmark

(Received 14 February 1972)

One of the few exact results that would be vitiated by the recently discussed possibility of neutrinos with mass, would be the statement of the forbidden nature of the decay $K_L^0 \rightarrow \bar{\nu}_l + \nu_l$ because of angular-momentum conservation in the usual neutrino theory. We note that the decay $K_L^0 \rightarrow \bar{\nu}_\mu + \nu_\mu$ could easily exist with a rate comparable to that for $K_L^0 \rightarrow 2\gamma$ if $m_{\nu_\mu} \approx 100$ eV. The processes $\nu_l + p \rightarrow \nu_l + \Sigma^+$, $\nu_l + n \rightarrow \nu_l + \Lambda$ and $\Sigma^+ \rightarrow p + \bar{\nu}_l + \nu_l$, $\Lambda \rightarrow n + \bar{\nu}_l + \nu_l$ would then occur in lowest order, but at minute rates.

It has recently been suggested,¹ in connection with the unexpectedly low counting rate in the solar-neutrino experiment,² that neutrinos with a finite mass could be unstable. Since the empirical limit³ on the mass of ν_μ is only $m_{\nu_\mu} < 1.6$ MeV, whereas that for ν_e is $m_{\nu_e} < 60$ eV, the question of finite mass and possible instability is surely also relevant for ν_μ .¹ Apart from their possible instability, neutrinos with mass would vitiate the exactness of the statement that the decay $K_L^0 \rightarrow \bar{\nu}_l + \nu_l$ is forbidden by angular-momentum conservation in the usual neutrino theory.⁴ This decay could be mediated by an effective Lagrangian density^{5,6}

$$L_{K^0}^{\text{eff}} = (-\lambda m_{\nu_l} \sqrt{G_F}) [K^0 \bar{\psi}_{\nu_l} (1 - \gamma_5) \psi_{\nu_l} + \bar{K}^0 \bar{\psi}_{\nu_l} (1 + \gamma_5) \psi_{\nu_l}], \quad (1a)$$

$$\Rightarrow L_{K_L^0}^{\text{eff}} = (\sqrt{2} \lambda m_{\nu_l} \sqrt{G_F}) K_L^0 \bar{\psi}_{\nu_l} \gamma_5 \psi_{\nu_l}, \quad (1b)$$

where $G_F \approx 10^{-5} m_N^{-2}$ is the Fermi constant in terms of the nucleon mass m_N , and λ is a dimensionless number. In fact an effective interaction of the form (1b) would be generated in perturbation theory by a "standard" (neutral) intermediate-boson Lagrangian⁷

$$L_W = [ig m_W W_\mu (\partial_\mu K^0) + g W_\mu \bar{\psi}_{\nu_l} \gamma^\mu (1 - \gamma_5) \psi_{\nu_l}] + \text{H.c.}, \quad (2)$$

where m_W denotes the boson mass and g is the dimensionless semiweak coupling, $g^2/m_W^2 = G_F/\sqrt{2}$. The two terms in (2) generate (1b) with $\lambda = 2^{3/4}g$.⁷ From (1b) we compute a rate for $K_L^0 \rightarrow \bar{\nu}_l + \nu_l$ in terms of the effective coupling constant $f^2 = \lambda^2 (m_{\nu_l}/m_N)^2 \times 10^{-5}$,

$$R(K_L^0 \rightarrow \bar{\nu}_l + \nu_l) = f^2 (6 \times 10^{22} \text{ sec}^{-1}) \Rightarrow 6.8 \times 10^3 \text{ sec}^{-1} \quad (3)$$

for $m_{\nu_l} = m_{\nu_\mu} = 100$ eV and $\lambda \approx 1$.⁷ For comparison we note the empirical $R(K_L^0 \rightarrow 2\gamma) \approx 10^4 \text{ sec}^{-1}$. Thus K_L^0 could easily be decaying into neutrino pairs at a rate which is a fraction of a percent of $(\tau_{K_L^0})^{-1}$!

Together with strong interactions, (1b) generates the following processes in lowest order⁸:

$$\nu_l + p \rightarrow \nu_l + \Sigma^+, \quad (4a)$$

$$\nu_l + n \rightarrow \nu_l + \Lambda, \quad (4b)$$

and

$$\Sigma^+ \rightarrow \bar{\nu}_i + \nu_i + p, \quad (4c)$$

$$\Lambda \rightarrow \bar{\nu}_i + \nu_i + n. \quad (4d)$$

However, one immediately calculates that these effects are minute *because* of the smallness of m_{ν_i} . For process (4a)

$$\begin{aligned} \frac{d\sigma}{dt} &= f^2 \left(\frac{G_{K^0 \Sigma^+ p}}{4\pi} \right)^2 \frac{F(t)}{4} \frac{t[t - (m_\Sigma - m_N)^2]}{[t - m_K^2]^2 (s - m_N^2)^2} \\ &\Rightarrow 0.8 \times 10^{-50} \text{ cm}^2 (\text{GeV}/c)^{-2} \end{aligned} \quad (5)$$

for a 10-GeV ν_μ with $t \cong - (m_\Sigma - m_N)^2 = - (0.250 \text{ GeV}/c)^2$ [$F(t) \cong F(0) = 1$] and $(G_{K^0 \Sigma^+ p}/4\pi) \cong 3,^9$ and for the above value of $f^2 \cong 10^{-19}$ corresponding to $m_{\nu_\mu} = 100 \text{ eV}$.⁷ This is to be compared with the corresponding $d\sigma/dt$ for $\nu_\mu + n \rightarrow \mu^- + p$ of about $10^{-38} \text{ cm}^2 (\text{GeV}/c)^2$. Similarly for process (4c) one computes a rate of about 0.06 sec^{-1} which may be compared with the rate for $\Sigma^- \rightarrow e^- + \bar{\nu}_e + n$ of about $0.7 \times 10^7 \text{ sec}^{-1}$.

Since we have noted that $R(K_L^0 \rightarrow \bar{\nu}_\mu + \nu_\mu)$ could be comparable to $R(K_L^0 \rightarrow 2\gamma)$, it is amusing to inquire into the contribution of the neutrino-pair intermediate state to the imaginary part of the amplitude for $K_L^0 \rightarrow \mu^+ + \mu^-$, given the problems surrounding this amplitude at the moment.¹⁰ Apart from the weakness of the "diagonal"¹¹ amplitude for $\bar{\nu}_\mu + \nu_\mu \rightarrow \mu^+ + \mu^-$ in the conventional theory, the exact $V-A$ structure causes the contribution from this intermediate state to vanish. However, remembering the recent admonition of Gell-Mann *et al.*,¹¹ "No connection between this (diagonal) process and the weak interactions should be assumed" *a priori*, we may parametrize an anomalous interaction, $z G_F \bar{\psi}_\mu \gamma_\lambda \psi_{\nu_\mu} \bar{\psi}_{\nu_\mu} \gamma^\lambda \psi_\mu$, and compute the value of z which makes the magnitude of the neutrino-pair contribution equal to that of the two-photon contribution. We find $|z| \cong 1.6 \times 10^4$. Such an anomaly can presumably be ruled out by current or future experiments on the process $\nu_\mu + Z \rightarrow \nu_\mu + \mu^+ + \mu^- + Z$ in the Coulomb field of the nucleus Z .^{12,13} However, note that since the product fz occurs, increasing f by 10 ($m_{\nu_\mu} = 1 \text{ keV}$) reduces $|z|$ to $\approx 1.6 \times 10^3$.

Clearly, the transition⁷ $K_L^0 \rightarrow \bar{\nu}_i + \nu_i + K_S^0$, with virtual neutrino-pairs predominating over the mass-shell intermediate state, represents a realization of an effective superweak model of CP noninvariance,¹⁴ *except* for the possible detectability of the decay $K_L^0 \rightarrow \bar{\nu}_i + \nu_i$.

In summary, we note that if one entertains the

notion of neutrinos with mass of the order of a few tens of electron volts, then the decay $K_L^0 \rightarrow \bar{\nu}_i + \nu_i$ could be occurring. How would one exclude this possibility? Given the surprises that neutral- K -meson decays have already held for physics, it might be amusing to consider this question.

I thank P. Oleson for discussion.

¹J. N. Bahcall, N. Cabibbo, and A. Yahil, Phys. Rev. Lett. **28**, 316 (1972).

²R. Davis, Jr., L. C. Rogers, and V. Rodeka, Bull. Amer. Phys. Soc. **16**, 631 (1971).

³A. Rittenberg *et al.*, Rev. Mod. Phys. Suppl. **43**, 1 (1971).

⁴T. D. Lee and C. N. Yang, Phys. Rev. **105**, 1671 (1957); A. Salam, Nuovo Cimento **5**, 299 (1957); L. Landau, Nucl. Phys. **3**, 127 (1957).

⁵This interaction is CP and T invariant, whereas terms like $iaK_L^0 \bar{\psi}_{\nu_i} \psi_{\nu_i}$ or $ibK_S^0 \bar{\psi}_{\nu_i} \gamma_5 \psi_{\nu_i}$ would break this invariance. [N. Brene and J. Dethlefsen (private communication) have pointed out that the neutral W_μ may of course have both Hermitian and anti-Hermitian parts, which can even mutually couple with CP and T noninvariance, thus generating an effective interaction $ibK_S^0 \bar{\psi}_{\nu_i} \gamma_5 \psi_{\nu_i}$. Note that the terms in (1a) without γ_5 are not generated by (2). If one replaces m_W in the first term of (2) by the decay constant for $K^+ \rightarrow \mu^+ + \bar{\nu}_\mu$, f_K , then λ becomes $(f_K/m_N)(4.5 \times 10^{-3}) \approx 10^{-3}$. A ν_μ mass of $\approx 100 \text{ keV}$ then gives (3).]

⁶Effective or primary neutral-lepton current couplings to hadrons are severely limited in strength by the empirical absence of $K^+ \rightarrow \pi^+ + \bar{l} + l$ (see Ref. 3). The interaction (1) does not contribute to these because $K^+ + K^0 + \pi^+$ strongly. The distinct feature of the interaction (1) is, of course, that the smallness of the effective coupling is related to the smallness of m_{ν_i} , rather than being *ad hoc*.

⁷Brene and Dethlefsen, Ref. 5.

⁸If the π^0 participates in an interaction similar to (1b) then the process $\nu_i + p \rightarrow \nu_i + p$ would also be generated at a minute level.

⁹Private communication from Y. A. Chao; Y. A. Chao and E. Pietarinen, Phys. Rev. Lett. **26**, 1060 (1971).

¹⁰A. R. Clark, T. Elioff, R. C. Field, H. J. Frisch, R. P. Johnson, L. T. Kerth, and W. A. Wenzel, Phys. Rev. Lett. **26**, 1667 (1971).

¹¹M. Gell-Mann, M. L. Goldberger, N. M. Kroll, and F. E. Low, Phys. Rev. **179**, 1518 (1969).

¹²W. Czyz, G. C. Sheppey, and J. D. Walecka, Nuovo Cimento **34**, 404 (1964).

¹³D. H. Perkins, in *Topical Conference on Weak Interactions*, CERN, Geneva, Switzerland, 14-17 January, 1969 (CERN Scientific Information Service, Geneva, Switzerland, 1969), p. 6.

¹⁴L. Wolfenstein, Phys. Rev. Lett. **13**, 562 (1964).