## Anomalous Viscosity as a Possible Explanation for an Anomalous Plasma Skin Effect

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Anomalous viscosity of electrons across the magnetic field in plasmas is calculated, assuming that the electrostatic ion acoustic wave is excited by the drift velocity of electrons parallel to the magnetic field lines. The mean free path of the wave perpendicular to the magnetic field is estimated. The momentum transfer rate between waves and electrons,  $\tau^{-1}$ , also was calculated from the quasilinear theory. The anomalous (kinematic) viscosity was then estimated as  $l^2\tau^{-1}$ . The theory appears to explain the experimentally observed results of the anomalous skin effect.

It has remained a mystery why there appears to be no detectable skin effect in Tokamaks; that is, the laser electron-temperature-profile measurement failed to show the increase in the temperature at the periphery. This obviously can be explained in two ways. One is that there is an enhanced heat conductivity to the center which makes the temperature profile almost flat, in spite of the fact that current distribution is restricted to the surface. The second is that there is an anomalous skin effect which allows the current to penetrate to the center. Although this was not experimentally checked in Tokamak devices, it is known that the  $\theta$ -pinch and collisionless shock waves have an anomalous skin effect. Here, we propose that an anomalous viscosity may give rise to the anomalous skin effect. This also implies that the anomalous skin effect need not be equivalent to anomalous resistivity. Experimentally, the resistivity found after the Tokamak plasmas are established cannot explain the seemingly rapid penetration of the current across the magnetic field. This concept of anomalous viscosity was introduced by Morse and Stovall<sup>1</sup> to explain the observed field reversal in  $\theta$  pinches.

The physical model we propose here is very simple. We assume that the plasma density  $n_0$  is uniform, that the electron temperature  $T_e$  is constant, and that the ion temperature is negligible. The current density, which is carried by electrons, has a gradient in the x direction and a strong, constant magnetic field in the z direction. The ion-plasma frequency  $\omega_{pi}$  is considered to be much higher than the ion cyclotron frequency  $\Omega_i$ . The system is unstable against many modes, such as the two-stream instability and ion acoustic waves. We assume that the nonlinear limit has set in. Thus, the plasma no longer is quiescent, but has a finite density of phonons. Since ion temperature is assumed to be cold, no strong interaction appears between phonons and ions, although the interaction may play a part in setting the nonlinear limitation of phonon density. The electrons then interact with the phonons and exchange momentum. The momentum is then carried by phonons across the magnetic field and imparted to the electrons whose parallel velocity is smaller than the phase velocity of the phonons. In this way the momentum is carried out from one electron to another electron via phonons. Therefore, if we know the density of the phonons,  $n_{\rm h}$  (m<sup>-3</sup>), the collision rate between a phonon and an electron, S (m<sup>3</sup>/sec), and the mean free path of the phonon perpendicular to the magnetic field, l (m), then the kinematic viscosity perpendicular to B is given by  $n_{p}Sl^{2}$ . We now try to estimate S and l. Unfortunately,  $n_{b}$  cannot be estimated unless the full nonlinear equation is solved. Instead, as will be explained, we shall express the



FIG. 1. (a) Drift velocity  $v_d$  and growth rate  $\gamma$  as a function of x (perpendicular to the magnetic field). (b) Wave propagation, +x direction and absorbed for x > 0. The momentum parallel to B is transferred from a point x < 0 to another point, x > 0.

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phonon density in terms of the plasma density fluctuation.

In practice, instead of phonons we calculate the wave-electron interaction, which is the classical picture of the phonon interaction.

As waves that have perpendicular group velocity, we chose electrostatic ion cyclotron waves.<sup>2</sup> For the wave with an  $\exp[i(\omega t + k_x x + k_z z)]$  dependence,<sup>3</sup> the dispersion relation is

$$\lambda_{\rm D}^2 k_x^2 \left[ 1 + \frac{\omega_{pi}^2}{\Omega_i^2 - \omega^2} \right] + k_z^2 \lambda_{\rm D}^2 \left( 1 - \frac{\omega_{pi}^2}{\omega^2} \right) + 1 = 0.$$
 (1)

Here,  $\lambda_D$  is the Debye length of electrons. We look for the solution for which the electron drift frequency is of the order of  $|\omega/k_z|$ . Then, pro-

vided that the drift frequency is higher than the plasma sound velocity (which we assume),

$$k_{x}^{2} \approx \frac{\omega^{2} - \Omega_{i}^{2}}{\Omega_{i}^{2} + \omega_{pi}^{2} - \omega^{2}} \frac{\omega_{pi}^{2}}{v_{s}^{2}}, \quad v_{s}^{2} \equiv \frac{kT_{e}}{M}.$$
 (2)

Waves propagate in this frequency range:  $\Omega_i^2 < \omega^2 < \omega_{pi}^2$ . We restrict to ourselves waves for which  $\omega^2 \ll \omega_{pi}^2$ . Then,

$$k_{x}^{2} \approx \frac{\omega^{2} - \Omega_{i}^{2}}{v_{s}^{2}} \equiv k_{0}^{2} \quad (k_{0} > 0).$$
(3)

The growth (or damping) rate of this wave can be easily calculated from the standard Landau damping calculation. We take the zeroth-order distribution function as a displaced Maxwellian. The result is

$$k_{x}^{2} v_{s}^{2} = (\omega^{2} - \Omega_{i}^{2}) \left[ 1 + \left( -\frac{\omega}{k_{z}} - v_{d} \right) \pi i f_{0e} \left( -\frac{\omega}{k_{z}} \right) \frac{1}{n_{0}} \frac{k_{z}}{|k_{z}|} \right];$$

$$\tag{4}$$

here  $f_{0e}$  is the zeroth-order electron distribution function. The growth rate  $\gamma$  is given by [for  $|\omega/k_z| \ll (kT_e/m)^{1/2}$ ]

$$\gamma = \frac{k_x^2}{2|k_z|} \frac{(kT_em)^{1/2}}{M} \left(\frac{\pi}{2}\right)^{1/2} \left(-\frac{v_d k_z}{\omega} - 1\right).$$
(5)

Thus, if  $v_d k_z / \omega < -1$ , the plasma is unstable.

We first estimate the mean free path of the wave perpendicular to the magnetic field. There, noting that  $v_d$  is a function of x (Fig. 1), we take  $\varphi = \varphi(x) \exp[i(\omega \tau + k_z z)]$ . Then we write, instead of Eq. (4) (we take  $k_z > 0$ ),

$$\frac{d^2\varphi(x)}{dx^2} + \frac{\omega^2 - \Omega_i^2}{v_s^2} \left[ 1 + \left( -\frac{\omega}{k_z} - v_d \right) \left( \frac{\pi}{2} \frac{m}{kT_e} \right)^{1/2} i \right] \varphi(x) = 0.$$
(6)

Assuming that the second term in the bracket is small, we apply the eikonal approximation, obtaining

$$\varphi(x) = \exp(ik_0 x) \exp\left[-\int \left(-\frac{\omega}{k_z} - v_d\right) \left(\frac{\pi}{2} \frac{m}{kT_e}\right)^{1/2} \frac{dx k_0}{2}\right].$$
(7)

In order to estimate the integral, we choose the coordinates so that  $v_d = v_{d1} - (x/L)v_{d0}$ , with  $v_{d1}$  and L positive. We expect that  $v_{d0}$  and  $v_{d1}$  are comparable to the typical drift velocity of electrons. Near x = 0,  $v_d > 0$ , so only the  $\omega/k_z < 0$  wave is unstable. We calculate the penetration depth for the phase velocity  $\omega/k_z = -v_{d1}$ . Then  $(-\omega/k_z - v_d) = v_{d0}x/L$ . Equation (7) then becomes

$$\varphi(x) = \exp(ik_0 x) \exp\left[-\frac{1}{4} \frac{k_0 x^2 v_{d0}}{L} \left(\frac{\pi}{2} \frac{m}{kT_e}\right)^{1/2}\right], \quad x \ge 0.$$
(8)

The depth *l* for which the amplitude decreases by a factor  $e^{-1/2}$  is

$$l^{2} = \left(\frac{2kT_{e}}{\pi m}\right)^{1/2} \frac{1}{v_{d0}} \frac{2L}{k_{0}}, \quad v_{s}^{2} k_{0}^{2} = \omega^{2} - \Omega_{i}^{2}.$$
(9)

The interaction of the wave with the electron distribution function can be estimated from the quasilinear theory.<sup>4, 5</sup> In the absence of the external electric field and collision, and if  $\vec{E}$  is not too large (for more detail, see Sec. 6 of Ref. 5),

$$\frac{\partial f_2}{\partial t} = \frac{1}{2} \operatorname{Re} \left( \frac{\partial f_1}{\partial v_z} E_z^* \right) = \frac{1}{2} \frac{e^2}{mkT_e} \operatorname{Re} \left\{ E_z^* \varphi \frac{\partial}{\partial v_z} \left[ \frac{(-\omega - k_z v_d)}{\omega + k_z v} f_{0e}(v) \right] \right\},\tag{10}$$

or

$$\frac{\partial}{\partial t}\Gamma \equiv \frac{\partial}{\partial t}\int v_{z}f_{2}dv = \frac{1}{2}\frac{e^{2}}{mkT_{e}}|\varphi|^{2}\left(\frac{\pi m}{2kT_{e}}\right)^{1/2}(-\omega-k_{z}v_{a})n_{0}\frac{k_{z}}{|k_{z}|} = \frac{1}{2}\frac{e^{2}}{mkT_{e}}|\varphi|^{2}\left(\frac{\pi m}{2kT_{e}}\right)^{1/2}\left(-\frac{\omega}{k_{z}v_{d}}-1\right)\Gamma_{0}|k_{z}|, \quad (11)$$

with  $\Gamma_0 \equiv \int v_z f_0 dv$ . Thus, noting that  $\gamma > 0$  corresponds to

$$-\frac{\omega}{k_z v_d} - 1 < 0, \tag{12}$$

we find that  $\Gamma$  changes by  $\Gamma_{\scriptscriptstyle 0}$  in the time  $\tau,$  with

$$\tau^{-1} = \frac{1}{2} \left( \frac{e^2}{mkT_e} \right) |\varphi|^2 \left( \frac{\pi m}{2kT_e} \right)^{1/2} \left( 1 + \frac{\omega}{k_z v_d} \right) |k_z|.$$
(13)

Therefore, the kinematic viscosity is estimated to be

$$\mu = \frac{l^2}{\tau} = \sum \frac{2kT_e}{\pi m} \frac{1}{v_{d0}} \frac{L}{k_0} \left| \frac{e\varphi}{kT_e} \right|^2 \left( 1 + \frac{\omega}{k_z v_d} \right) |k_z|, \qquad (14)$$

where the summation sign indicates the sum over all the modes. We may rewrite Eq. (14) as

$$\mu \approx \frac{2kT_{e}}{\pi m} \frac{L}{v_{d0}^{2}} \sum \left| \frac{k_{z}v_{d0}}{\omega} \right| \left| \frac{e\varphi}{kT_{e}} \right|^{2} \frac{\omega}{(\omega^{2} - \Omega_{i}^{2})^{1/2}} \left( \frac{kT_{e}}{M} \right)^{1/2} = 2L \left( \frac{kT_{e}}{M} \right)^{1/2} \frac{1}{\pi} \frac{kT_{e}}{mv_{d0}^{2}} \left\langle \left| \frac{k_{z}v_{d0}}{\omega} \right| \frac{\omega}{(\omega^{2} - \Omega_{i}^{2})^{1/2}} \right\rangle_{\text{av}} \sum \left( \frac{e\varphi}{kT_{e}} \right)^{2}, \quad (15)$$

where the average is taken over all the modes. This average is taken of the order of unity and denoted by C. We also note that  $e\varphi/kT_e \approx \tilde{n}/n_0$  from the dispersion relation where  $\tilde{n}$  is the density perturbation. Hence,

$$\mu = 2L \left(\frac{kT_e}{M}\right)^{1/2} \frac{1}{\pi} \frac{kT_e}{mv_{d0}^2} C \left\langle \frac{(n-n_0)^2}{n_0^2} \right\rangle_{\rm av},\tag{16}$$

with the assumption that the  $(n-n_0)^2$  average is taken over the frequency region between  $\Omega_i$  and  $\omega_{pi}$ . L is related to the drift velocity density gradient, and may be approximated as  $1/L = |d(\ln v_d)/dx|$ .

With this kinematic viscosity given, it is possible to estimate the skin effect. Equations to govern the skin effect are given by

$$(\nabla \times \vec{B})_z = \mu_0 J_z = -e \mu_0 \Gamma_z, \tag{17}$$

and ( $\nu_{\rm C}$  is the Coulomb collision frequency)

$$e n_0 E_z = +m \,\mu \nabla^2 \Gamma_z - m \,\nu_C \Gamma_z - m \,\partial \Gamma_z / \partial t. \tag{18}$$

Introducing vector potential  $\vec{A}$ , and applied constant electric field  $E_0$  in the z direction, we get

$$(m\mu\nabla^2 - m\nu_C)\nabla^2 A_z = m\nabla^2 \frac{\partial A_z}{\partial t} - e^2\mu_0 n_0 \frac{\partial A_z}{\partial t} + e^2\mu_0 n_0 E_0.$$
<sup>(19)</sup>

We now determine the time constant for a slab with the thickness of  $2x_0$ . (Asymptotically, the solution is  $A_z = -x^2 e^2 \mu_0 n_0 E_0 / 2m \nu_{\rm C}$ .) The normal-mode solution for the lowest eigenmode is then

$$A_{z} = A_{0} \cos\left(\frac{\pi}{2} \frac{x}{x_{0}}\right) \exp\left(\frac{-t}{\tau_{s}}\right), \tag{20}$$

$$\frac{1}{\tau_s} = \frac{1}{e^2 \mu_0 n_0} \frac{\pi^2}{4} \frac{m}{x_0^2} \left( \frac{\pi^2}{4} \frac{1}{x_0^2} \mu + \nu_C \right) \frac{1}{1 + \pi^2 c^2 / 4 x_0^2 \omega_{pe}^2}.$$
(21)

Here,  $\omega_{pe}/2\pi$  is the electron plasma frequency. The gradient distance L is estimated to be  $2x_0/\pi$ . Then

$$\tau_{s} = \frac{4e^{2}\mu_{0}n_{0}x_{0}^{2}}{\pi^{2}m} \left[ \frac{C}{x_{0}} \left( \frac{kT_{e}}{M} \right)^{1/2} \frac{kT_{e}}{mv_{d0}^{2}} \left\langle \frac{(n-n_{0})^{2}}{n_{0}^{2}} \right\rangle + \nu_{C} \right]^{-1} \left( 1 + \frac{\pi^{2}C^{2}}{4x_{0}^{2}\omega_{pe}^{2}} \right).$$
(22)

Now the estimate of the new term is in order. For the Tokamak,<sup>6</sup> we assume that  $\langle (n-n_0)^2/n_0^2 \rangle = 10^{-2}$ ,  $x_0 = 10$  cm, C = 1, M/m = 1800,  $n_0 = 2 \times 10^{13}$  cm<sup>-3</sup>,  $kT_e/mv_{d0}^2 = 10^2$ ; we then have the figures shown in Ta-

ble I.

In the table,  $\tau_s^*$  is the skin time without the viscosity correction.

It thus appears that the anomalous skin effect, as observed in Tokamaks, may be traceable to this effect. An experimental check should be relatively easy, especially if the time of current penetration is measured after the plasma is heated to a relatively high temperature. Oscillation over the ion cyclotron frequency should also be measurable.

Somewhat paradoxical results are indicated in Eq. (22), because it may appear that increasing the drift velocity increases  $\tau_s$ . However, in reality  $\langle (n-n_0)^2/n_0^2 \rangle_{\rm av}$  should be a function of  $v_{d0}$ , and presumably an increasing function. This probably compensates for the inverse-square dependence on  $v_{d0}$ .

Aside from the application to the Tokamak, this mechanism may also be responsible for the anomalously thick skin of collisionless shocks perpendicular to the magnetic field, as well as the anomalous viscosity hypothesized by Morse and Stovall.<sup>1</sup>

Furthermore, we have introduced here a new concept for the mechanism of transport of various plasma parameters. This technique may be useful for calculating heat conductivity and diffusion coefficients.

In the absence of both this viscosity effect and ion Landau damping, the electron and wave momenta would reach equilibrium values such as predicted by Drummond and Pines<sup>7, 8</sup> and the resistivity due to the phonon-electron interaction would disappear. Thus either ion Landau damping (which also leads to anomalous resistivity) or Table I. Skin times calculated for different electron temperatures.

<i>kT<sub>e</sub></i> (eV)	$ au_s$ (msec)	$ au_{s}^{*}$ (msec)
100	2.1	7.2
400	1.45	57.6
1000	0.95	220
2000	0.65	640

anomalous viscosity is essential in explaining the anomalous skin effect.

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<sup>3</sup>T. Stix, *The Theory of Plasma Waves* (McGraw-Hill, New York, 1962), p. 41.

<sup>4</sup>Stix, Ref. 1, p. 154.

<sup>5</sup>S. Yoshikawa and H. Yamato, Phys. Fluids <u>9</u>, 1814 (1966).

<sup>6</sup>See, for example, L. A. Artsimovich, A. M. Ana-

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## Comment on the Proposed Magnetic Cooling of He<sup>3</sup>

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We examine a recent proposal by Goldstein for producing very low temperatures by freezing  $\text{He}^3$  in the presence of magnetic fields. We conclude that the highly isotropic nature of solid  $\text{He}^3$  will completely suppress the particular cooling mechanism envisaged, but a related magnetothermal effect should be observable. At readily accessible temperatures this effect will be small, but below about 1 mK a potentially useful amount of magnetic cooling is predicted. The field required for the maximum cooling is about 70 kOe.

In a recent Letter,<sup>1</sup> Goldstein proposed a new method for obtaining extremely low temperatures by freezing He<sup>3</sup> in the presence of a magnetic field. It was implied that the method depends on the peculiar nature of the anisotropy of solid He<sup>3</sup>, but the precise role of the anisotropy was not discussed explicitly.<sup>2</sup> In this note we wish to point out that (1) the particular mean-field calcu-