

EVIDENCE FOR PHOTOFISSION OF IRON*

C. B. Fulmer, I. R. Williams, and T. H. Handley
Oak Ridge National Laboratory, Oak Ridge, Tennessee

and

G. F. Dell and L. N. Blumberg†
Cambridge Electron Accelerator, Cambridge, Massachusetts
(Received 4 August 1967)

Studies of proton-induced reactions in the GeV energy region^{1,2} have given evidence that fission occurs in nuclei at least as light as silver. It has been pointed out that any nucleus can be made to undergo fission provided it is supplied with sufficient excitation energy.^{3,4} In this note we present evidence of photofission in iron foils that were bombarded with high-energy electrons.

In a study⁵ of residual radioactivity produced in targets exposed to electrons in the GeV energy region, radionuclides were identified over a wide range of mass number. It was demonstrated that the residual nuclei are produced predominantly by photonuclear reactions that are induced by the electron bremsstrahlung photons.

The experiments are done by exposing targets to monitored beams of electrons at the Cambridge Electron Accelerator and subsequently measuring the gamma spectra from the exposed targets with a Ge(Li) gamma spectrometer. Radionuclides are identified by gamma energy and half-life.

The yields of radionuclides identified in iron foils exposed to electron beams are plotted in Fig. 1 as a function of the number of nucleons removed from the target nucleus. The yield of ³²P + ³³P, which are pure beta emitters, was identified by radiochemistry. (Analysis of the decay curve indicates that the two isotopes are produced with about equal intensity.) All of the other nuclides were identified in gamma spectra obtained by the Ge(Li) spectrometer. Additional identifications of ²⁴Na and ⁴³K were made by radiochemistry.

In the mass region $A \geq 43$ in Fig. 1 the yield of radionuclides exhibits an exponential decrease with the number of nucleons emitted from the target nucleus. The yields of ²⁴Na and ³²P + ³³P are, however, much larger than would be expected, unless the mechanism for production of these nuclei is different from that of the nuclides with $A \geq 43$.

In spectra obtained from aluminum targets

that were bombarded with 3-GeV electrons an appreciable yield of ⁷Be was observed.⁵ Careful examination of the gamma spectra obtained from the iron targets yielded no evidence for ⁷Be. A radiochemical separation also yielded no evidence for ⁷Be in an iron foil that was bombarded with 3-GeV electrons. Studies of proton- and alpha-induced reactions⁶ have shown that emission of ⁷Be is enhanced by rotational motion of the compound nucleus. Photon-induced reactions produce little rotational motion of the target nucleus.

The curve of Fig. 1 is an exponential drawn through a plot of calculated cross sections for production of residual nuclei by nucleon-induced reactions. The cross sections were calculated

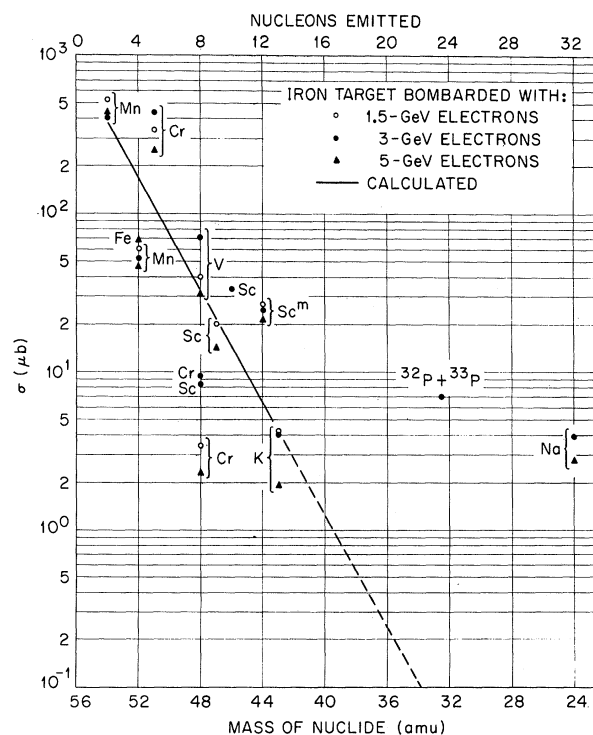


FIG. 1. Yields of radionuclides identified in iron samples that were exposed to a beam of 3-GeV electrons. The values of σ are based on the integrated electron beam intensity.

ed from the cascade-evaporation results of Bertini.⁷ Calculations were made at 50-MeV intervals between 50 and 400 MeV; the cross sections thus obtained were multiplied by the intensity of the bremsstrahlung spectrum⁸ at each energy. These products were summed for each mass number to yield the plot through which the curve of Fig. 1 was drawn. This curve was then normalized to the experimental data, for mass 43 and larger, in Fig. 1. The calculations included ⁴³K and residual nuclei of larger mass. Although the absolute magnitude of the calculated cross section for nucleon-induced reactions is almost three orders of magnitude larger than that of the measured cross sections in Fig. 1, the fall-off of cross section with intensity is in reasonable agreement with that of the data down to mass 43.

The yield of ²⁴Na is a factor of between 10³ and 10⁴ larger than would be expected from extrapolation of the cascade-evaporation theory. The yield of ³²P + ³³P is also much higher than would be predicted by the theory. Apparently another reaction mechanism is responsible for the yield of ²⁴Na that is observed.

A spectrochemical analysis of a piece of the iron foil that was used for the bombardments indicated the presence of impurities that can account for less than 1% of the ²⁴Na production, if it is assumed that the impurities, near mass 24, have cross sections for production of ²⁴Na that are about the same as that of aluminum.⁵ It was also observed that the yield of ²⁴Na in a thick target varies with depth in the material in the same manner as other radionuclides.

One may attempt to attribute the relatively large yield of ²⁴Na to nuclear fragmentation, i.e., the emission of fragments of mass 5-10. Emission of several fragments would be required, and this process would not yield similar amounts of ²⁴Na and ³²P.

The most reasonable explanation for the relatively large yields of ²⁴Na and ³²P appears to be that photofission of iron nuclei is induced by high-energy bremsstrahlung photons. The *Q* value for this reaction is -28 MeV and the Coulomb barrier is ~22 MeV, therefore the fission barrier is ~50 MeV.⁹ Many of the photons in the bremsstrahlung spectrum from 3-

GeV electrons have enough energy to produce fission in iron.

One would expect that photofission of iron would result in production of other residual nuclei in the same mass region. A small yield of ²²Na was observed. Many of the radioactive nuclei in the mass region between 24 and 32 have half-lives that are too long or too short to be identified in the experiment reported here and in Ref. 5. Some evidence of ²⁸Mg was observed; however, it has a beta decay chain length of two, compared with one for ²⁴Na and ³²P, and thus the yield of ²⁸Mg would be appreciably lower than that of ²⁸Al or ²⁸Si.

Fission is energetically favored over other reaction mechanisms such as multiple-nucleon (32 for production of ²⁴Na in ⁵⁶Fe) or alpha-particle emission. Both cascade-evaporation theory⁷ and nuclear evaporation theory¹⁰ predict much lower relative yields of ²⁴Na than is observed. Any mechanism other than fission would result in a much lower yield of ²⁴Na than of ³²P, whereas the experimentally observed yields are approximately equal.

*Research sponsored by the U. S. Atomic Energy Commission.

†Present address: Accelerator Department, Brookhaven National Laboratory, Upton, New York.

¹A. A. Caretto, J. Hudis, and G. Friedlander, *Phys. Rev.* **110**, 1130 (1958).

²E. W. Baker and S. Katcoff, *Phys. Rev.* **123**, 641 (1961).

³J. M. Miller and J. Hudis, *Ann. Rev. Nucl. Sci.* **9**, 159 (1959).

⁴R. W. Spence and G. P. Ford, *Ann. Rev. Nucl. Sci.* **2**, 411 (1953).

⁵C. B. Fulmer, I. R. Williams, G. F. Dell, and L. N. Blumberg, *Bull. Am. Phys. Soc.* **12**, 499 (1967); and to be published.

⁶R. H. Lindsay, *Phys. Rev.* **147**, 782 (1966).

⁷H. W. Bertini, *Phys. Rev.* **131**, 801 (1963); **138**, AB2(E) (1965). Also Oak Ridge National Laboratory Report No. ORNL-TM-1225, 10 September 1965 (unpublished).

⁸Y. S. Tsai and Van Whitis, *Phys. Rev.* **148**, 124 (1966).

⁹W. D. Myers and W. J. Swiatecki, *Nucl. Phys.* **81**, 1 (1966).

¹⁰I. Dostrovsky, Z. Frankel, and G. Friedlander, *Phys. Rev.* **116**, 683 (1959).