

a charged particle passes through its anticoincidence scintillator simultaneously. This correction is quite large for the forward telescopes.

(2) Isotropic correction to  $C_0$ . (a) Geometric correction,  $\delta$  ray, and Dalitz pair conversion. (b) Cross-section correction for the processes  $\pi^+p\eta^0$ ;  $\pi^+p(k\pi^0)$ ;  $\pi^0p(k\pi^\pm)$ ;  $\pi^+n\pi^0(k\pi^\pm)$ . We used bubble chamber results to carry out these corrections.<sup>2,3</sup> The contribution of the process  $\pi^+ + p \rightarrow \pi^+ + p + \gamma$  is negligible.<sup>4</sup>

Since the data consists of five points the least-squares method does not allow us to calculate the fourth-order fit. The statistical errors are so large that the value calculated by solving a linear system for five equations is not significant at all energies.

The experimental results are shown in Table I and Fig. 2. Bubble chamber data<sup>2,3,5,6</sup> are included in Fig. 2. The third-order fit is not statistically significant. The second-order fit gives the dominant features of the distributions up to 1070 MeV. The dominant features of the results are (a) the rise of  $C_0$  between 500 and 800 MeV with a shoulder near 800 MeV; (b) the large forward-backward asymmetry ( $C_1$  coefficient) which has a maximum in the region of about 1 GeV; (c) the steep rise of  $C_2$  coefficient between 700 and 1100 MeV.

In the following Letter a comparison is made between counters and bubble chamber results.

Above 1070 MeV sixth-order fits are necessary and counter results are not significant.<sup>2,7</sup>

We wish to express our gratitude to H. Kraybill for communication of unpublished results and to G. Valladas for helpful discussion. We wish to thank Professor A. Berthelot for his constant interest and support.

\*Present address: Lawrence Radiation Laboratory, University of California, Berkeley, California.

†Chercheur Institut Interuniversitaire des Sciences Nucléaires, Laboratoire des Hautes Energies, Brussels, Belgium.

<sup>1</sup>J. F. Detoeuf, Y. Ducros, J. P. Merlo, A. Stirling, B. Thevenet, L. van Rossum, and J. Zsembery, Phys. Rev. **134**, B228 (1964).

<sup>2</sup>D. Stonehill, C. Baltay, H. Courant, W. Fickinger, E. C. Fowler, H. Kraybill, J. Sandweiss, J. Sanford, and H. Taft, Phys. Rev. Letters **6**, 624 (1961).

<sup>3</sup>C. Gensollen, P. Granet, R. Barloutaud, A. Leveque, and J. Meyer, Proceedings of the Sienna International Conference on Elementary Particles (Società Italiana di Fisica, Bologna, Italy, 1963), p. 84.

<sup>4</sup>B. Thevenet, thesis, CEA Report No. R 2306, p. 49.

<sup>5</sup>P. C. A. Newcomb, Phys. Rev. **132**, 1283 (1963).

<sup>6</sup>C. A. Tilger, C. P. Poirier, E. D. Alyea, H. J. Martin, and J. H. Scandrett, Phys. Rev. **142**, 972 (1966).

<sup>7</sup>H. Kraybill, D. L. Stonehill, B. Deler, W. Laskar, J. P. Merlo, G. W. Tautfest, and G. Valladas, following Letter, [Phys. Rev. Letters **16**, 863 (1966)].

## ANGULAR DISTRIBUTION OF THE $\pi^0$ IN THE REACTION $\pi^+ + p \rightarrow N_{1238}^{*++} + \pi^0$ FROM 0.5 TO 1.46 GeV

H. L. Kraybill and D. L. Stonehill\*

Yale University, New Haven, Connecticut

and

B. Deler, W. Laskar, J. P. Merlo, G. Valladas, and G. W. Tautfest†

Département de Physique des Particules Élémentaires, Centre d'Études Nucléaires de Saclay,  
Gif-sur-Yvette, Seine et Oise, France

(Received 1 March 1966)

It is well known that in the neighborhood of 1 GeV, single-pion production is characterized by a strong production of the isobar  $N_{1238}^{*}(\frac{3}{2})$ .<sup>1,2</sup> Figures 1(a) and 1(b) give the behavior of the  $\pi^+p$  cross sections up to 1.6 GeV.<sup>3-8</sup> Figure 1(c) shows the ratio  $P$  of the  $\pi^+p\pi^0$  events in the isobar mass range  $p\pi^+(1.23 \pm 0.07)$  GeV/ $c^2$  to the total number of events  $\pi^+p\pi^0$ . The values of  $P$  up to 1 GeV are in good agreement with the values predicted by the Lindenbaum-

Sternheimer model<sup>9</sup> [dashed line on Fig. 1(c)]. Above 1 GeV the disagreement between experimental and predicted  $P$  values may be related to observed small values of the ratio  $R = \sigma_{\pi^+p\pi^0} / \sigma_{\pi^+\pi^+\eta}$  [Fig. 1(d)], a value inconsistent with the isobar-model prediction of  $R = 6.5$  (dashed line). This fact could be explained by the increase of  $T = \frac{1}{2}$  attractive interactions in the  $\pi^+\pi^+n$  channel.<sup>10,11</sup> The effect of this interaction is small in the  $\pi^+p\pi^0$  channel and the con-

sidered isobar ( $\pi^+p$ ) mass range, representing about 45% of the single-pion production under 1 GeV. In view of a two-body partial-wave analysis of the reaction  $\pi^+ + p \rightarrow \pi^0 + N_{1238}^*$  we report in this paper new bubble-chamber results from Purdue, Saclay, and Yale concerning angular distribution of the  $\pi^0$  associated with a  $\pi^+p$  mass  $1.23 \pm 0.07$  GeV/ $c^2$  at incident energies of 0.81, 0.91, 0.98, 1.09, 1.13, 1.26, and 1.30 GeV. Some of these experiments have been already described: 0.91, 1.09, and 1.26 GeV (Yale<sup>2</sup>), 1.09 and 1.13 GeV (Purdue<sup>12</sup>). The experiments at 0.81 and 1.30 GeV are new (Saclay<sup>10</sup>). We have also used published results<sup>11,13,14</sup> at 0.5, 0.6, and 1.46 GeV to plot Fig. 1(c) and 1(d).

The results of each experiment are expressed by a distribution of the production cross section into 20 intervals of the cosine of the angle  $\theta$  between the incident pion and the  $\pi^0$  in the over-all center of mass. The  $\pi^0$  angular distributions corresponding to the isobar  $\pi^+p$  region are expanded in Legendre polynomials of  $\cos\theta$ ,

$$\frac{2}{\pi\lambda^2} \frac{d\sigma}{d(\cos\theta)} (1.16 \leq M_{\pi^+p} \leq 1.30) \\ = C_0 + \sum_{l=1}^L C_l P_l(\cos\theta).$$

The values of the coefficients  $C_l$  obtained by the sixth-order fit are given in Table I and plotted in Fig. 2. The  $C_l$  values obtained by the fit of distributions in 40 intervals at 0.81, 1.13, and 1.3 GeV are compatible with the  $C_l$  values for 20 interval distributions within statistical errors. For the data below 1 GeV a fourth-order fit is satisfactory. Above 1 GeV polynomials of the sixth order are required. Seventh-order fits are given above 1.1 GeV. There is not a large improvement of the  $\chi^2$  by inclusion of a  $P_7$  term except at 1.13 and 1.3 GeV. This may be related to elastic distributions that give small values for  $C_7$  and  $C_8$ .<sup>6,7</sup>

To compare  $C_l$  coefficients for  $\pi^0$  associated with the  $\pi^+p$  isobar to the ones obtained for all  $\pi^0\pi^+p$  events, we have plotted in Fig. 2 the values of  $C_0 = P\sigma_{\pi^+p\pi^0}/\pi\lambda^2$ ,  $C_1$  and  $C_2$  of the counter experiment of Detoeuf *et al.*<sup>15</sup> We have used interpolated values of  $P$  to determine the coefficients  $C_0$  of the counter experiment. The

value  $B_0$  for the differential cross section at  $90^\circ$  for counter and bubble-chamber data is plotted on Fig. 2.  $B_0$  is related to the coefficients  $C_l$  by the relation

$$B_0 = C_0 - \frac{C_2}{2} + \frac{3}{8}C_4 - \frac{5}{16}C_6 \cdots$$

The dominant features of the variation of  $C_l$  coefficients versus the incident-pion kinetic energy (Fig. 2) are the following: (1) The  $C_0$  coefficient rises from 0.5 to 0.81 GeV with a shoulder near 0.8 GeV. It presents a weak maximum near 1.3 GeV. (2) The  $\pi^0$  is mainly emitted in the forward direction. The coefficient  $C_1$  has a maximum in the region about 1 GeV. (3) The coefficient  $C_2$  shows a steep rise between 0.7 and 1.1 GeV with a maximum near 1.2 GeV. (4) The coefficient  $C_3$  rises between 0.9 and 1.25 GeV from negative values to positive values. (5) The coefficient  $C_4$  does not exhibit any strong variation and is significantly negative below 1.3 GeV. It reaches a positive value at 1.4 GeV. (6) The coefficient  $C_5$  is small at all the energies. (7) The coefficient  $C_6$  becomes significant above 1 GeV and presents a strong minimum near 1.3 GeV. (8) The differential cross section at  $90^\circ$  shows a maximum at 0.81 GeV which is a consequence of the rise of  $C_2$ .

Below 1 GeV small negative values of  $C_5$  and  $C_6$  could be explained by a scanning bias in the forward direction. At higher energies  $\rho$  production modifies the  $\pi^0$  angular distribution in the forward direction. We report in Table I the result of two fits at 0.81, 1.13, and 1.30 GeV: (a) Fit of 39 points of the distribution without the forward point ( $\cos\theta < 0.95$ ). The results of these fits are compatible with the previous values. (b) Fit of the  $\pi^0$  distribution after a tentative correction of the  $\rho$  contamination. The behavior of  $C_2$ ,  $C_4$ , and  $C_6$  is not modified by the  $\rho$  correction. The coefficient  $C_1$  is decreased by about 30% at 1.13 GeV and 60% at 1.30 GeV;  $C_2$  and  $C_3$  are somewhat smaller than the previous values.

The variations of the coefficients  $C_l$  as deduced from this analysis for  $\pi^0$  associated with  $\pi^+p$  isobar are larger than the ones obtained for all the  $\pi^0\pi^+p$  events. This confirms that  $\pi^0$  angular effects characterize chiefly the  $N_{1238}^*$  production.

The differential cross section for the process  $\pi^+ + p \rightarrow \pi^0 + N_{(1.23 \pm 0.07)}^{*++}$  has the following

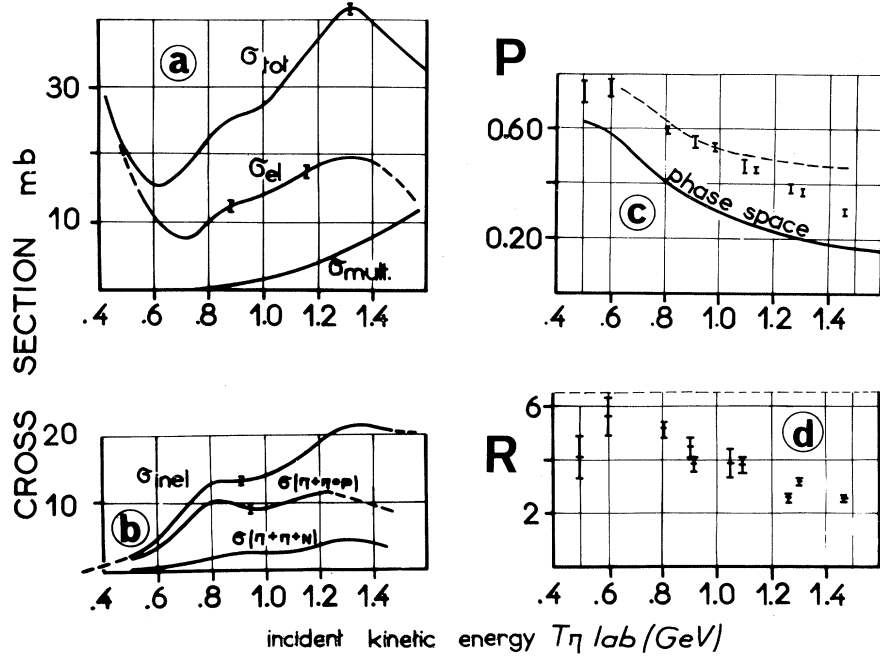


FIG. 1. (a) Total ( $\sigma_{\text{tot}}$ ), elastic ( $\sigma_{\text{el}}$ ), and multipion-production ( $\sigma_{\text{mult}}$ ) cross sections for  $\pi^+p$  interaction from 0.4 to 1.4 GeV. (b) Total inelastic ( $\sigma_{\text{inel}}$ ) and single-pion-production ( $\sigma_{\pi^+p\pi^0}$  and  $\sigma_{\pi^+\pi^+n}$ ) cross sections. (c) Ratio  $P$  of the number of events in the isobar mass ( $p\pi^+$ ) region ( $1.23 \pm 0.07$  GeV/c<sup>2</sup>) to the total number of  $\pi^+p\pi^0$  events compared with phase space (continuous line) and isobar-model (dashed line) predictions. (d) Ratio  $R$  of the cross section  $\sigma_{\pi^+p\pi^0}$  to the cross section  $\sigma_{\pi^+\pi^+n}$ , and the isobar-model prediction (dashed line).

form<sup>16</sup>:

$$\frac{1}{2\pi} \frac{\partial \sigma}{\partial (\cos \theta)} = \frac{9}{15} \frac{\lambda^2}{4} r \sum_{\alpha \leq \beta} \text{Re} R_{\alpha}^* R_{\beta} (J_{\alpha} + \frac{1}{2})^{1/2} \times (J_{\beta} + \frac{1}{2})^{1/2} (\delta_{\alpha\beta} + \sum_{l \geq 1} C_l^{\alpha\beta} P_l(\cos \theta)),$$

where  $R_{\alpha}$  is a reduced matrix element relative to the channel  $N_{(1238)}^* + \pi$ ;  $\alpha$  denotes a partial wave of the isobar characterized by the set  $(l_{\alpha}^i, l_{\alpha}, 2J_{\alpha})$ , where  $l_{\alpha}^i$  is the orbital angular momentum of the incident pion,  $l_{\alpha}$  is the orbital angular momentum of the isobar, and  $J_{\alpha}$  is the total angular momentum. We use the notations  $S, P, D, \dots$  for  $l=0, 1, 2, \dots$ .  $9/15$  is the probability of getting the charge state  $N_{(1238)}^{*++}$ , and  $r$  is the fraction of  $N_{3/2,3/2}^*$  in the mass range 1.16-1.30 GeV/c<sup>2</sup>. Near 1 GeV  $r$  is about 0.75. ( $r$  is calculated using the pole formulation of Bergia, Bonsignori, and Stanghellini<sup>17</sup> and the empirical expression for the width given by Klepikov, Mescheryakov, and Sokolov.<sup>18</sup> The values of the coefficients  $C_l^{\alpha\beta}$  given in Table II are related to the exper-

imental  $C_l$  values by

$$C_0 = (9/15)r \sum_{\alpha} |R_{\alpha}|^2 (J_{\alpha} + \frac{1}{2}),$$

$$C_l = (9/15)r \sum_{\alpha < \beta} \text{Re} R_{\alpha}^* \times R_{\beta} (J_{\alpha} + \frac{1}{2})^{1/2} (J_{\beta} + \frac{1}{2})^{1/2} C_l^{\alpha\beta}.$$

(a) The 0.80-GeV region.  $-F$ -wave production of the isobar  $N_{(1238)}^*$  is certainly small at these energies, so the negative  $C_4$  value can be explained by the production through the  $DD5$  wave, where the notation  $L_1 L_2 (2J)$  denotes the  $\pi$ - $p$  incident orbital angular momentum, the  $\pi$ - $N^*$  orbital angular momentum, and twice the total angular momentum. The interference between  $PP3$  and  $DD5$  could explain  $C_3$  and in part  $C_1$ . The resultant value of  $C_2$  is then small. We have made an estimate of the amplitudes of  $DD5$  and  $PP3$ , consistent with the observed  $C_l$ 's, which are compatible with the values of the absorption parameters deduced from recent phase-shift analyses.<sup>19-21</sup> The relative importance of the isotropic waves  $SD1, PP1, DD3$ , and  $DS3$  cannot be deduced from the angular

Table I. Numerical values of the coefficients of the expansion in Legendre polynomials of the  $\pi^0$  angular distribution corresponding to isobar mass region. Type of fit: 20, 40, fit of distributions in 20, 40, intervals; 39, fit of distributions without the forward point ( $\cos\theta < 0.95$ );  $\rho$ , fit of distributions with a tentative correction of the  $\rho$  contamination.

$T_\pi$ GeV	Fit type	order	$\chi^2/n$	$C_0$	$C_1$	$C_2$	$C_3$	$C_4$	$C_5$	$C_6$	$C_7$
0.81	20	6	1.11	1.48 ± 0.05	0.85 ± 0.08	0.04 ± 0.10	-0.40 ± 0.11	-0.48 ± 0.14	-0.17 ± 0.16	-0.26 ± 0.18	
	40	6	1		0.85 ± 0.08	0.04 ± 0.09	-0.41 ± 0.11	-0.46 ± 0.13	-0.05 ± 0.14	-0.21 ± 0.15	
	39	6	0.96		0.82 ± 0.08	-0.04 ± 0.11	-0.55 ± 0.14	-0.63 ± 0.17	-0.22 ± 0.19	-0.36 ± 0.18	
0.91	20	6	0.33	1.50 ± 0.07	1.02 ± 0.14	0.68 ± 0.17	-0.54 ± 0.20	-0.58 ± 0.24	-0.52 ± 0.24	-0.16 ± 0.28	
0.98	20	6	1.38	1.52 ± 0.03	1.34 ± 0.07	1.18 ± 0.08	-0.21 ± 0.08	-0.67 ± 0.11	-0.23 ± 0.12	-0.52 ± 0.13	
1.09	20	6	1.24	1.67 ± 0.08	1.04 ± 0.18	2.18 ± 0.21	0.04 ± 0.25	-0.91 ± 0.29	-0.04 ± 0.31	-0.61 ± 0.30	
	7	7	1.33		1.05 ± 0.18	2.20 ± 0.21	0.03 ± 0.25	-0.92 ± 0.29	-0.12 ± 0.38	-0.64 ± 0.34	-0.16 ± 0.40
1.13	20	6	1.03	1.75 ± 0.04	1.26 ± 0.10	2.54 ± 0.11	0.18 ± 0.14	-0.53 ± 0.16	-0.36 ± 0.15	-0.78 ± 0.17	
	20	7	0.88		1.25 ± 0.10	2.54 ± 0.11	0.14 ± 0.13	-0.55 ± 0.16	-0.55 ± 0.19	-0.85 ± 0.17	-0.32 ± 0.19
	40	6	1.07		1.25 ± 0.09	2.52 ± 0.11	0.10 ± 0.12	-0.69 ± 0.14	-0.44 ± 0.14	-0.95 ± 0.16	
	39	6	0.86		1.38 ± 0.11	2.73 ± 0.15	0.48 ± 0.18	-0.24 ± 0.21	-0.09 ± 0.18	-0.65 ± 0.18	
	$\rho$	6	1.14		0.83 ± 0.10	2.33 ± 0.12	-0.14 ± 0.14	-0.68 ± 0.17	-0.36 ± 0.17	-0.80 ± 0.19	
1.26	20	6	2.65	1.80 ± 0.09	0.88 ± 0.22	3.0 ± 0.23	1.11 ± 0.26	-0.77 ± 0.33	-0.18 ± 0.31	-2.02 ± 0.32	0.22 ± 0.43
	20	7	2.85		0.87 ± 0.22	3.02 ± 0.23	1.15 ± 0.26	-0.75 ± 0.33	-0.05 ± 0.39	-1.97 ± 0.32	
1.30	20	6	1.25	1.77 ± 0.05	0.76 ± 0.11	2.05 ± 0.13	1.10 ± 0.14	-0.43 ± 0.18	-0.28 ± 0.18	-1.78 ± 0.21	
	20	7	0.96		0.78 ± 0.11	2.06 ± 0.13	1.03 ± 0.15	-0.50 ± 0.22	-0.53 ± 0.22	-1.84 ± 0.21	
	40	6	1.35		0.79 ± 0.09	2.22 ± 0.13	1.11 ± 0.14	-0.23 ± 0.17	-0.34 ± 0.17	-1.56 ± 0.19	
	39	6	1.38		0.82 ± 0.13	2.22 ± 0.17	1.17 ± 0.20	-0.16 ± 0.22	-0.28 ± 0.22	-1.50 ± 0.22	
	$\rho$	6	1.05		0.30 ± 0.11	1.70 ± 0.13	0.81 ± 0.15	-0.60 ± 0.19	-0.25 ± 0.20	-1.83 ± 0.22	
1.46	20	6	1.33	1.39 ± 0.05	0.57 ± 0.07	1.27 ± 0.13	0.89 ± 0.14	0.65 ± 0.17	-0.13 ± 0.18	-1.16 ± 0.20	
	20	7	1.43		0.57 ± 0.10	1.28 ± 0.13	0.88 ± 0.15	0.65 ± 0.17	-0.17 ± 0.20	-1.16 ± 0.20	

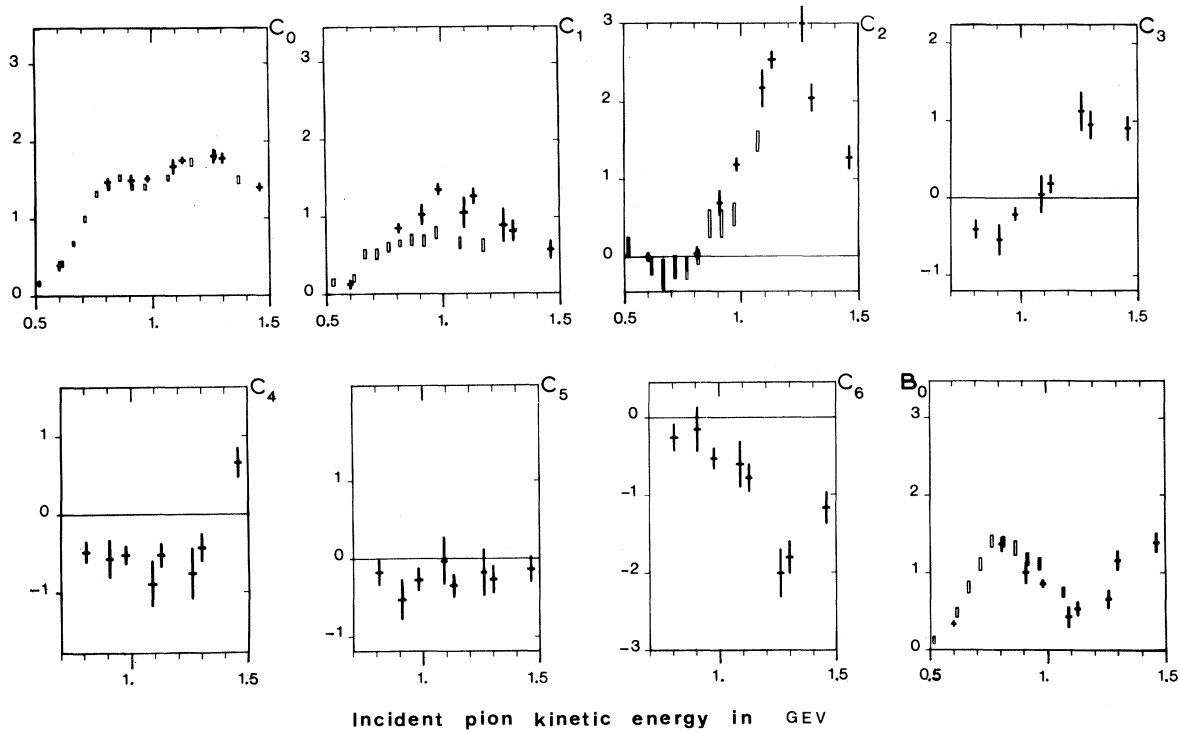


FIG. 2. Coefficients of the expansion in Legendre polynomials ( $C_l$ ) of the  $\pi^0$  angular distribution and differential cross section at  $90^\circ$  ( $B_0$ ). Solid points correspond to bubble-chamber results for events in the isobar mass region ( $1.23 \pm 0.07 \text{ GeV}/c^2$ ). Open rectangles correspond to counter results for all  $\pi^0$ .<sup>15</sup>

distribution of the  $\pi^0$  alone.

(b) The 0.90- to 1.30-GeV energy range. — At the energy of 1.30 GeV, the dominant wave is known to have a total angular momentum  $J = \frac{7}{2}$ .<sup>6,7,22</sup> The large value of  $C_6$  would correspond to the presence of the wave  $FF7$  (or  $GG7$ ). The interference between  $FF7$  and  $FF5$  (or  $GG7$  and  $DD5$ ) could also contribute to a large  $C_6$ . The wave  $GD7$  would not give any  $P_6$  term.  $C_6$  becomes statistically significant near 1 GeV; therefore, it is more probable that the wave responsible for the large  $C_6$  at 1.3 GeV is  $FF7$ . This is in agreement with the absorption of  $F_{37}$  in the elastic channel.

The only explanation of the large value of the ratio  $C_2/C_0 \approx 1.70$  and the small value of  $C_4/C_2 \approx -0.26$  is an interference between isobar waves of the same parity. These waves could be  $PP3$ , already present near 0.8 GeV, and  $FF7$ . The presence  $PP1$  in interference with  $FF7$  could give a positive coefficient  $C_4$  to attenuate the large negative  $C_4$  value given by the other waves. The waves  $SD1$ ,  $DS3$ , and  $DD5$  could also contribute to  $C_2$ . Above 1 GeV the amplitude of  $DD5$  might decrease because its interference with  $FF7$  would give a large

coefficient  $C_5$  which is not observed.

From this survey of the main features of the  $\pi^0$  angular distribution, we can conclude as follows: (1) In agreement with the phase-shift analysis of inelastic  $\pi^+p$  scattering up to 0.9 GeV, there is a strong indication of an important isobar production via the waves  $PP3$  and  $DD5$ . (2) The maximum at 1.3 GeV can be associated with a strong isobar production through the  $FF7$  wave, in good agreement with the analysis of the elastic scattering. (3) The simplest explanation of the rise of  $C_2$  and  $C_3$  between 0.9 and 1.2 GeV is an interference effect of the  $FF7$  wave with, respectively, the  $PP3, PP1$  waves, and the  $SD1, DS3$  waves.

In a future paper we shall examine the angular effects of the decay of the isobar in its center of mass, and we shall compare more accurately inelastic with elastic results.

We wish to thank Professor A. Berthelot for constant encouragement and support.

\*Now at Brookhaven National Laboratory, Upton, New York.

†On leave from Purdue University, Lafayette, Indiana.

Table II. Calculated distributions  $C_{\alpha\beta} = \delta_{\alpha\beta} + \sum_{l>1} C^{\alpha\beta} P_l(\cos\theta)$  in the pure  $(\frac{3}{2}, \frac{3}{2})$  isobar model<sup>a</sup> for two sets  $(l_{\alpha}^i, l_{\alpha}, 2J_{\alpha}), (l_{\beta}^i, l_{\beta}, 2J_{\beta})$  of incident, isobar, and total angular momenta.

Waves	$C^{\alpha\beta}$	Waves	$C^{\alpha\beta}$
SD1 SD1	1.	DD3 DS3	$-2P_2$
PF3	$2.68P_1$	FP5	$-0.537P_1 - 2.15P_3$
PP1	$2P_1$	GD7	$-0.916P_2 - 2.29P_4$
DD3	$2P_2$	FF5 FF5	$1 + 0.628P_2 - 0.428P_4$
FF5	$2.19P_3$	GG7	$2.81P_1 + 0.873P_3 - 1.06P_5$
GG7	$2.39P_4$	PP3	$1.12P_2 - 2.10P_4$
PP3	$-0.894P_1$	DD5	$0.328P_1 + 1.08P_3 - 2.84P_5$
DD5	$-1.31P_2$	FF7	$0.554P_2 + 1.05P_4 - 3.39P_6$
FF7	$-1.63P_3$	DS3	$-2.19P_3$
DS3	$-2P_2$	FP5	$-0.840P_2 - 2.10P_4$
FP5	$-2.68P_3$	GD7	$-0.251P_1 - 1.17P_3 - 2.09P_5$
GD7	$-3.21P_4$	GG7 GG7	$1. + 0.884P_2 + 0.15P_4 - 0.606P_6$
PF3 PF3	$1 + 0.80P_2$	PP3	$1.43P_3 - 2.50P_5$
PP1	$2.68P_2$	DD5	$0.596P_2 + 1.16P_4 - 3.32P_6$
DD3	$0.537P_1 + 2.15P_3$	FF7	$0.186P_1 + 0.8P_3 + 0.965P_5 - 3.9P_7$
FF5	$0.840P_2 + 2.10P_4$	DS3	$-2.39P_4$
GG7	$1.07P_3 + 2.14P_5$	FP5	$-1.07P_3 - 2.14P_5$
PP3	$-1.20P_2$	GD7	$-0.456P_2 - 1.34P_4 - 2.03P_6$
DD5	$-0.350P_1 - 1.40P_3$	PP3 PP3	$1. - 0.80P_2$
FF7	$-0.626P_2 - 1.57P_4$	DD5	$2.46P_1 - 1.87P_3$
DS3	$-2.68P_3$	FF7	$2.82P_2 - 2.09P_4$
FP5	$-0.172P_2 - 3.43P_4$	DS3	$0.895P_1$
GD7	$-0.319P_3 - 3.98P_5$	FP5	$1.20P_2$
PP1 PP1	1.	GD7	$1.43P_3$
DD3	$2P_1$	DD5 DD5	$1. + 0.408P_2 - 0.980P_4$
FF5	$2.19P_2$	FF7	$2.75P_1 + 0.356P_3 - 2.04P_5$
GG7	$2.39P_3$	DS3	$1.31P_2$
PP3	$-0.894P_2$	FP5	$0.351P_1 + 1.40P_3$
DD5	$-1.31P_3$	GD7	$0.60P_2 + 1.50P_4$
FF7	$-1.63P_4$	FF7 FF7	$1 + 0.794P_2 - 0.117P_4 - 1.01P_6$
DS3	$-2P_1$	DS3	$1.63P_3$
FP5	$2.68P_2$	FP5	$0.626P_2 + 1.56P_4$
GD7	$-3.21P_3$	GD7	$0.187P_1 + 0.873P_3 + 1.56P_5$
DD3 DD3	1.	DS3 DS3	1.
FF5	$2.63P_1 - 0.438P_3$	FP5	$2.68P_1$
GG7	$3.07P_2 - 0.683P_4$	GD7	$3.21P_2$
PP3	$0.716P_1 - 1.61P_3$	FP5 FP5	$1 + 0.80P_2$
DD5	$0.935P_2 - 2.24P_4$	GD7	$2.87P_1 + 1.43P_3$
FF7	$1.09P_3 - 2.72P_5$	GD7 GD7	$1 + 1.02P_2 + 0.55P_4$

<sup>a</sup>See Ref. 9.

<sup>1</sup>C. Gensollen, P. Granet, R. Barloutaud, A. Leveque, J. Meyer, in Proceedings of the Sienna International Conference on Elementary Particles (Società Italiana di Fisica, Bologna, Italy, 1963), p. 84.

<sup>2</sup>H. Foelsche, E. C. Fowler, H. L. Kraybill, J. R. Sanford, D. Stonehill, Proceedings of the International Conference on Instrumentation for High-Energy Physics at CERN, edited by F. M. Farley (North-Holland Publishing Company, Amsterdam, 1963), p. 36; and Rev. Mod. Phys. **34**, 503 (1962).

<sup>3</sup>A. Stirling, thesis, R-CEA Report No. 2938 (unpublished).

<sup>4</sup>T. J. Devlin, J. Solomon, and G. Bertsch, Phys. Rev. Letters **14**, 1031 (1965).

<sup>5</sup>B. Thevenet, thesis, R-CEA Report No. 2306 (unpublished).

lished).

<sup>6</sup>J. A. Helland, T. J. Devlin, D. E. Hagge, M. J. Longo, B. J. Moyer, and C. D. Wood, Phys. Rev. Letters **10**, 27 (1963).

<sup>7</sup>P. J. Duke, D. P. Jones, M. A. R. Kemp, P. G. Murphy, J. D. Prentice, J. J. Thresher, and H. H. Atkinson, Phys. Rev. Letters **15**, 468 (1965).

<sup>8</sup>J. F. Detoef, Y. Ducros, J. P. Merlo, A. Stirling, B. Thevenet, L. van Rossum, and J. Zsembery, Phys. Rev. **134**, B228 (1964).

<sup>9</sup>R. M. Sternheimer and S. J. Lindenbaum, Phys. Rev. **123**, 333 (1961).

<sup>10</sup>B. Deler, to be published.

<sup>11</sup>P. Daronian, A. Daudin, M. A. Jabiol, C. Lewin, C. Kochowski, B. Ghidini, S. Mongelli, and V. Pic-

ciarelli, to be published.

<sup>12</sup>G. W. Tautfest and R. B. Willmann, in Proceedings of the Second Topical Conference on Resonant Particles, Ohio University, Athens, Ohio, 10-12 June 1965 (unpublished), p. 421.

<sup>13</sup>J. Debaisieux, F. Grard, J. Heughebaert, R. Sevranckx, and R. T. van de Walle, Nucl. Phys. **63**, 273 (1965).

<sup>14</sup>P. C. A. Newcomb, Phys. Rev. **132**, 1283 (1963).

<sup>15</sup>J. F. Detoeuf, Y. Ducros, J. P. Merlo, A. Stirling, B. Thevenet, L. van Rossum, and J. Zsembery, preceding Letter [Phys. Rev. Letters **16**, 860 (1966)].

<sup>16</sup>The effect of the interference between the two isobar creation modes is neglected. At 800 MeV the ef-

fect of this interference on the coefficients  $C_l$  is 10 to 15% and it decreases quickly at higher energies.

<sup>17</sup>S. Bergia, F. Bonsignori, and A. Stanghellini, Nuovo Cimento **16**, 1073 (1960).

<sup>18</sup>N. P. Klepikov, V. A. Mescheryakov, and S. N. Sokolov, Dubna Report No. D-584, 1960 (unpublished).

<sup>19</sup>P. Bareyre, C. Brieman, A. V. Stirling, and G. Villet, Phys. Letters **18**, 342 (1965).

<sup>20</sup>A. Donnachie, A. T. Lea, and C. Lovelace, Phys. Letters **19**, 146 (1965).

<sup>21</sup>B. H. Brandsen, P. J. O'Donnell, and R. G. Moorhouse, Phys. Letters **19**, 420 (1965).

<sup>22</sup>P. Auvil and C. Lovelace, Nuovo Cimento **33**, 473 (1964).

## SUBTRACTIONS IN PARTIAL-WAVE DISPERSION RELATIONS\*

Toichiro Kinoshita

Laboratory of Nuclear Studies, Cornell University, Ithaca, New York  
(Received 8 April 1966; revised manuscript received 20 April 1966)

In the application of partial-wave dispersion relations in strong interaction physics, it is customary to assume that the number of subtractions does not exceed one.<sup>1</sup> Thus far no justification has been given for such an assumption. In fact it seems impossible to prove it if we restrict ourselves to the usual set of postulates for partial-wave amplitudes. The purpose of this note is to examine the nature of the supplementary assumptions one has to make in order to obtain the desired result, and to present the weakest assumption that has been found thus far.

For simplicity, we shall restrict our consideration to the elastic scattering of spinless particles of equal mass. The  $l$ th partial-wave amplitude  $f_l(s)$ , where  $s$  is the square of the center-of-mass energy, is assumed to have the following properties: (A) It is regular in the cut  $s$  plane with two cuts  $(-\infty, s_0)$  and  $(4\mu^2, +\infty)$ , real in the interval  $(s_0, 4\mu^2)$ , and continuous on the cuts. It has no essential singularity at any finite point on the cuts. (B) It has the threshold behavior  $f_l(s) \equiv (s-4\mu^2)^l F_l(s)$  where  $F_l(s)$  has a finite limit as  $s \rightarrow 4\mu^2$ . (C) It satisfies the unitarity condition on the right-hand cut. There is a considerable amount of freedom in choosing assumptions on the asymptotic behavior of  $f_l(s)$  for large  $s$ . The following is found to be the weakest that is still useful: (D) For sufficiently large  $s$ ,  $f_l(s)$  satisfies

$$|f_l(s)| < \exp[C(\ln|s|)^{2-\epsilon}] \text{ for any } \epsilon > 0.$$

Obviously the assumption (D) does not rule out infinitely many subtractions in general. To restrict the subtraction to a finite number, it is found necessary to impose a further assumption on the behavior of  $f_l(s)$  as  $|s| \rightarrow \infty$ . There are several more or less equivalent ways of stating this assumption. The following is perhaps the simplest: (E) The number of times that the left-hand cut discontinuity  $\text{Im}f_l(s+i0)$  changes its sign as  $s \rightarrow -\infty$  does not exceed  $C'(\ln|s|)^{1-\epsilon}$ , where  $\epsilon$  is the same as in (D). At present it is not known whether such an assumption can be justified on the basis of more fundamental properties of  $f_l(s)$ .

Since a general discussion requires some mathematics which might obscure the essential feature, we shall present here an explicit proof only for the case where (E) is replaced by a stronger assumption that  $\text{Im}f_l(s)$  changes its sign a finite number of times on the left-hand cut.<sup>2</sup> An outline of the problems encountered in the general case is given at the end of this note. Details will be discussed in a separate paper.<sup>3</sup>

Our method of proof is based on the technique introduced by Jin and Martin<sup>4</sup> which relates  $f_l(s)$  to a Herglotz function whose asymptotic behavior is well known.<sup>5</sup> Suppose  $f_l(s)$  has  $p$  real zeros  $r_1, r_2, \dots, r_p$  in the interval  $(s_0, 4\mu^2)$  and  $2q$  complex zeros  $c_1, c_1^*, \dots, c_q, c_q^*$  in the cut  $s$  plane. (In general  $q$  may be infinite. But it turns out to be finite under our