

MONTE CARLO INTRANUCLEAR CASCADE CALCULATIONS ON C^{12}
WITH MEDIUM-ENERGY PHOTONS

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It is generally considered that the Monte Carlo cascade calculations cease to be valid when the mass number of the target nucleus is too low. However, it has been shown by Abate *et al.*¹ and Bertini² that results compatible with experiments may be obtained down to $A = 12$. The present calculations were performed as an attempt to extend further the comparison with experimental results including, as Abate *et al.* did, the assumption of α clusters in light nuclei. The presence of α clusters is taken into account through the probability P for protons to collide with such clusters. A comparison is made between results obtained with $P = 40\%$, $P = 30\%$, and $P = 0\%$.

Our program is essentially the one used formerly³ with the following values for the nucleon-nucleon scattering cross sections⁴:

$$\sigma_{nm} = \sigma_{pp} = (10.63/\beta^2 - 29.92/\beta + 42.9) \text{ mb},$$

$$\sigma_{np} = \sigma_{pn} = (34.10/\beta^2 - 82.2/\beta + 82.2) \text{ mb};$$

and the differential cross section⁴:

$$d\sigma/d\Omega = K(A \cos^4\theta + B \cos^3\theta + 1),$$

where A and B vary with incident nucleon energy.

For (p, α) collisions we used experimental values of $d\sigma/d\Omega$ given by Selove and Teem⁵ and Cormack *et al.*⁶ For (n, α) collisions, $d\sigma/d\Omega$ was taken equal to $K \exp(-5.0\theta^2)$ after Tannenwald.⁷ The methods employed to correlate the above distributions with our uniform distribution of random numbers $0 \leq N \leq 1$ are described in reference 1 and by Gradsztajn.⁸

The potential square well is taken as in reference 3. The nucleons have a Fermi momentum distribution, whereas the α clusters are all considered as having an energy $E_0 = 5$ MeV. The binding energies taken for nucleons and α clusters are, respectively, 10 MeV and 7 MeV; and Coulomb barrier values are 3.5 MeV and 7 MeV, respectively.

In the present calculations we have postulated that an α cluster has only two possibilities after being knocked on: It can be emitted out of the nucleus if its kinetic energy is greater than the sum of its binding energy, Coulomb energy, and

E_0 . If not, it is absorbed but cannot cause other collisions.

All geometrical cross sections are calculated on the basis of $r_0 = 1.3$ fermis. The following results have been obtained:

(1) $C^{12}(p, p\alpha)Be^8$.—If $P = 40\%$, the calculated cross section is 37 mb for an incident proton energy of 156 MeV. The experimental value given by Samman and Cürer⁹ is 33 mb for 180-MeV protons and 22 mb for 340-MeV protons. The excitation energy distribution is not incompatible with the values given in reference 9.

(2) $C^{12}(p, 3pxn)Be$.—Our results are compared with the experimental ones of Yuasa, Bowden, and Bowman, which give $\sigma \approx 4$ mb for a proton energy of 110–135 MeV.¹⁰ The calculated cross sections with $P = 40\%$ and $P = 0\%$ are, respectively, for 156-MeV protons, 2.4 mb and 8.4 mb; for 135 MeV, 1.9 mb and 5.8 mb. The number of calculated events lies between 5000 and 9000. In Fig. 1 the number of events is plotted against $\Delta E = E^* + (n-1)B$, where E^* is the excitation energy of the residual nucleus, B the binding energy, and n the number of outgoing particles.

(3) $C^{12}(p, 2pxn)B$.—The experimental cross section, taken from Yuasa and Poulet¹¹ for 110-MeV protons, is 29 ± 5 mb. The calculated cross sections are the following (in mb), E_p being the energy of the incoming proton:

E_p (MeV)	P		
	40%	30%	0%
156	28	...	57
110	20	25	41

The ΔE spectrum is shown in Fig. 2.

(4) Total inelastic cross section.—On C^{12} , calculations on 1000 events lead, as expected, to the same results ($\sigma = 180$ mb) regardless of P values, in agreement with the experimental value of Muirhead and Rosser¹²: 210 ± 30 mb for 140-MeV protons.

The above calculations show that the Monte Carlo method applied for intranuclear cascade in light nuclei can give a good agreement with ex-

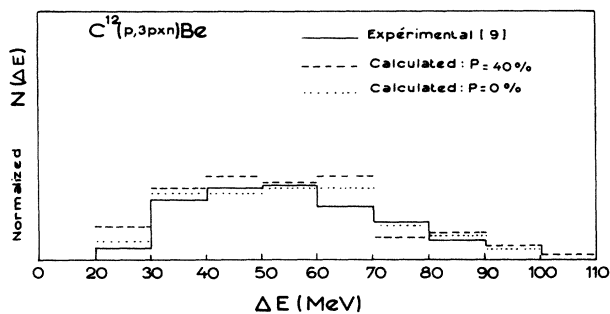


FIG. 1. Number of $(p, 3p\alpha n)$ events against $\Delta E = E^* + (n-1)B$.

periment.

The calculations of $(p, p\alpha)$, $(p, 2p\alpha n)$ and $(p, 3p\alpha n)$ cascades, taking into account the p - α collisions, give good values for the cross sections and ΔE spectra, but results still compatible with experiment can be obtained with $P=0\%$ for the last two.

Other Monte Carlo calculations are in progress for several other light nuclei and various reactions.

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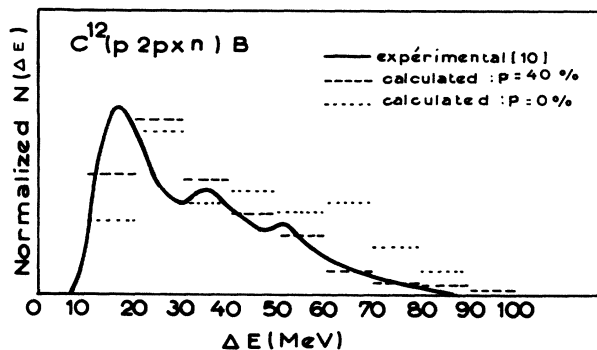


FIG. 2. Number of $(p, 2p\alpha n)$ events against $\Delta E = E^* + (n-1)B$.

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LIFETIMES OF $d_{3/2}$ HOLE STATES IN SCANDIUM ISOTOPES*

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The location of $d_{3/2}$ hole states in the odd- A isotopes of scandium has been inferred from the angular distributions of pickup reactions by Yntema and Satchler.¹ They observed such states in Sc^{45} , Sc^{47} , and Sc^{49} , and on this basis French² predicted that such a $d_{3/2}$ "hole" state (formed by promoting a $d_{3/2}$ proton to the $f_{7/2}$ shell) should also occur in Sc^{48} near the ground state. We therefore concluded that the first excited state of Sc^{48} , which we had previously shown to be metastable with a lifetime 190 times the single-particle estimate for an $M2$ transition,³ is the predicted $d_{3/2}$ "hole" state and that the transition is $M2$. The "hole" level in Sc^{45} was observed by Yntema and Erskine⁴ with the reaction $Ti^{48}(p,$

$\alpha)Sc^{45}$ and the excitation energy was found to be 13 keV. In Sc^{47} , Yntema and Satchler¹ found the hole state at about 800-keV excitation with the reaction $Ti^{48}(d, He^3)Sc^{47}$; and in Sc^{49} , which we have not investigated, they found the hole state at 2.4 MeV.

We have now observed the gamma rays (which were not previously seen) from these levels in Sc^{48} , Sc^{45} , and Sc^{47} and measured their lifetimes. Our results show that the first excited states of all three isotopes are indeed metastable (corresponding to $M2$ transitions) so they may be explained as $d_{3/2}$ hole states formed by promoting a $d_{3/2}$ proton to the $f_{7/2}$ shell. The level diagrams for these nuclei are given in Fig. 1.