## Measurements of ${}^{3}_{\Lambda}$ H and ${}^{4}_{\Lambda}$ H Lifetimes and Yields in Au + Au Collisions in the High Baryon Density Region

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We report precision measurements of hypernuclei  ${}^{3}_{\Lambda}$ H and  ${}^{4}_{\Lambda}$ H lifetimes obtained from Au + Au collisions at  $\sqrt{s_{NN}} = 3.0$  GeV and 7.2 GeV collected by the STAR experiment at the Relativistic Heavy Ion Collider, and the first measurement of  ${}^{3}_{\Lambda}$ H and  ${}^{4}_{\Lambda}$ H midrapidity yields in Au + Au collisions at  $\sqrt{s_{NN}} = 3.0$  GeV.  ${}^{3}_{\Lambda}$ H and  ${}^{4}_{\Lambda}$ H, being the two simplest bound states composed of hyperons and nucleons, are cornerstones in the field of hypernuclear physics. Their lifetimes are measured to be  $221 \pm 15(\text{stat}) \pm 19(\text{syst})$  ps for  ${}^{3}_{\Lambda}$ H and  $218 \pm 6(\text{stat}) \pm 13(\text{syst})$  ps for  ${}^{4}_{\Lambda}$ H. The  $p_{T}$ -integrated yields of  ${}^{3}_{\Lambda}$ H and  ${}^{4}_{\Lambda}$ H are presented in different centrality and rapidity intervals. It is observed that the shape of the rapidity distribution of  ${}^{4}_{\Lambda}$ H is different for 0%–10% and 10%–50% centrality collisions. Thermal model calculations, using the canonical ensemble for strangeness, describes the  ${}^{3}_{\Lambda}$ H yield well, while underestimating the  ${}^{4}_{\Lambda}$ H yield. Transport models, combining baryonic mean-field and coalescence (JAM) or utilizing dynamical cluster formation via baryonic interactions (PHQMD) for light nuclei and hypernuclei production, approximately describe the measured  ${}^{3}_{\Lambda}$ H and  ${}^{4}_{\Lambda}$ H yields. Our measurements provide means to precisely assess our understanding of the fundamental baryonic interactions with strange quarks, which can impact our understanding of more complicated systems involving hyperons, such as the interior of neutron stars or exotic hypernuclei.

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Hypernuclei are nuclei containing at least one hyperon. As such, they are excellent experimental probes to study the hyperon-nucleon (Y-N) interaction. The Y-N interaction is an important ingredient, not only in the equation-of-state (EOS) of astrophysical objects such as neutron stars, but also in the description of the hadronic phase of a heavy-ion collision [1]. Heavy-ion collisions provide a unique laboratory to investigate the Y-N interaction in finite temperature and density regions through the measurements of hypernuclei lifetimes, production yields, etc.

The lifetimes of hypernuclei ranging from A = 3 to 56 have previously been reported [2–11]. The light hypernuclei (A = 3, 4), being simple hyperon-nucleon bound states, serve as cornerstones of our understanding of the *Y-N* interaction [12,13]. For example, their binding energies  $B_{\Lambda}$  are often utilized to deduce the strength of the Y-N potential [14-16], which is estimated to be roughly 2/3 of the nucleon-nucleon potential. In particular, the hypertriton  ${}^{3}_{\Lambda}$ H, a bound state of  $\Lambda pn$ , has a very small  $B_{\Lambda}$  of several hundred keV [17,18], suggesting that the  ${}^{3}_{\Lambda}$ H lifetime is close to the free- $\Lambda$  lifetime  $\tau_{\Lambda}$ . Recently, STAR [10,11], ALICE [7,8] and HypHI [9] have reported  ${}^{3}_{\Lambda}$ H lifetimes with large uncertainties ranging from  $\sim 50\%$  to  $\sim 100\%\tau_{\Lambda}$ . The tension between the measurements has led to debate [19]. In addition, recent experimental observations of twosolar-mass neutron stars [20-22] are incompatible with model calculations of the EOS of high baryon density matter, which predict hyperons to be a major ingredient in neutron star cores [20–22]. These observations challenge our understanding of the *Y*-*N* interaction, and call for more precise measurements [12].

In heavy-ion collisions, particle production models such as statistical thermal hadronization [23] and coalescence [1] have been proposed to describe hypernuclei formation. While thermal model calculations primarily depend only on the freeze-out temperature and the baryo-chemical potential, the Y-N interaction plays an important role in the coalescence approach, through its influence on the dynamics of hyperon transportation in nuclear medium [24], as well as its connection to the coalescence criterion for hypernuclei formation from hyperons and nucleons [1]. At high collision energies, the  ${}^{3}_{\Lambda}$ H yields have been measured by ALICE [8] and STAR [10]. ALICE results from Pb + Pb collisions at  $\sqrt{s_{NN}} = 2.76$  TeV are consistent with statistical thermal model predictions [23] and coalescence calculations [25]. At low collision energies  $(\sqrt{s_{NN}} < 20 \text{ GeV})$ , an enhancement in the hypernuclei yield is generally expected due to the higher baryon density [1,23], although this has not been verified experimentally. The E864 and HypHI collaborations have reported hypernuclei cross sections at low collision energies [26,27], however both measurements suffered from low statistics and lack of midrapidity coverage. Precise measurements of hypernuclei yields at low collision energies are thus critical to advance our understanding in their production mechanisms in heavy-ion collisions and to establish the role of hyperons and strangeness in the EOS in the high-baryondensity region [28]. In addition, such measurements provide guidance on searches for exotic strange matter such as double- $\Lambda$  hypernuclei and strange dibaryons in low energy heavy-ion experiments, which could lead to broad implications [29–31].

In this Letter, we report  ${}^{3}_{\Lambda}$ H and  ${}^{4}_{\Lambda}$ H lifetimes obtained from data samples of Au + Au collisions at  $\sqrt{s_{NN}} =$ 3.0 GeV and 7.2 GeV, as well as the first measurement of  ${}^{3}_{\Lambda}$ H and  ${}^{4}_{\Lambda}$ H differential yields at  $\sqrt{s_{NN}} = 3.0$  GeV. We focus on the yields at midrapidity in order to investigate hypernuclear production in the high-baryon-density region. The yields at  $\sqrt{s_{NN}} = 7.2$  GeV are not presented here due to the lack of midrapidity coverage. The data were collected by the Solenoidal Tracker at the Relativistic Heavy Ion Collider (STAR) [32] in 2018, using the fixed-target (FXT) configuration. In the FXT configuration a single beam provided by the Relativistic Heavy Ion Collider (RHIC) impinges on a gold target of thickness 0.25 mm (corresponding to a 1% interaction probability) located at 201 cm away from the center of the STAR detector. The minimum bias trigger condition is provided by the beam-beam counters [33] and the time of flight detector [34]. The reconstructed primary-vertex position along the beam direction is required to be within  $\pm 2$  cm of the nominal target position. The primary-vertex position in the radial plane is required to lie within a radius of 1.5 cm from the center of the target to eliminate possible backgrounds arising from interactions with the vacuum pipe. In total,  $2.8 \times 10^8$  (1.5 × 10<sup>8</sup>) qualified events at  $\sqrt{s_{NN}} = 3.0$ (7.2) GeV are used in this analysis. The  $\sqrt{s_{NN}} =$ 3.0 GeV analysis and  $\sqrt{s_{NN}} = 7.2$  GeV analysis are similar. In the following, we describe the former; details related to the latter can be found in Supplemental Material [35].

The centrality of the collision is determined using the number of reconstructed charged tracks in the time projection chamber (TPC) [36] compared to a Monte Carlo Glauber model simulation [37]. Details are given in [38]. The top 0%-50% most central events are selected for our analysis.  ${}^{3}_{\Lambda}$ H and  ${}^{4}_{\Lambda}$ H are reconstructed via the two-body decay channels  ${}^{A}_{\Lambda}H \rightarrow \pi^{-} + {}^{A}He$ , where A = 3, 4. Charged tracks are reconstructed using the TPC in a 0.5 Tesla uniform magnetic field. We require the reconstructed tracks to have at least 15 measured space points in the TPC (out of 45) and a minimum reconstructed transverse momentum of 150 MeV/c to ensure good track quality. Particle identification for  $\pi^{-}$ , <sup>3</sup>He, and <sup>4</sup>He is achieved by the measured ionization energy loss in the TPC. The KFParticle package [39], a particle reconstruction package based on the Kalman filter utilizing the error matrices, is used for the reconstruction of the mother particle. Various topological variables such as the decay length of the mother particle, the distances of closest approach (DCA) between the



FIG. 1. Top row: Invariant mass distributions of (a)  ${}^{3}\text{He}\pi^{-}$  and (b)  ${}^{4}\text{He}\pi^{-}$  pairs. In the insets, black open circles represent the data, blue histograms represent the background constructed by using rotated pion tracks. In the main panels, black solid circles represent the rotational background subtracted data, and the red dashed lines describe the residual background. Bottom row: the transverse momentum  $(p_T)$  versus the rapidity (y) for reconstructed (c)  ${}^{3}_{\Lambda}\text{H}$  and (d)  ${}^{4}_{\Lambda}\text{H}$ . The target is located at the y = -1.05.

mother-daughter particles to the primary vertex, and the DCA between the two daughters, are examined. Cuts on these topological variables are applied to the hypernuclei candidates in order to maximize the signal significance. In addition, we place fiducial cuts on the reconstructed particles to minimize edge effects.

Figures 1(a) and 1(b) show invariant mass distributions of  ${}^{3}\text{He}\pi^{-}$  pairs and  ${}^{4}\text{He}\pi^{-}$  pairs in the  $p_{T}$  region (1.0-4.0) GeV/c for the 50% most central collisions. The combinatorial background is estimated using a rotational technique, in which all  $\pi^-$  tracks in a single event are rotated with a fixed angle multiple times and then normalized in the sideband region. The background shape is reasonably reproduced using this rotation technique for both  ${}^{3}_{\Lambda}$ H and  ${}^{4}_{\Lambda}$ H as shown in Figs. 1(a) and 1(b). The combinatorial background is subtracted from the data in 2D phase space  $(p_T \text{ and rapidity } y)$  in the collision center-ofmass frame. In addition to subtracting the rotational background, we perform a linear fit using the sideband region to remove any residual background. The subtracted distributions are shown in Figs. 1(c) and 1(d). The target is located at y = -1.05, and the sign of the rapidity y is chosen such that the beam travels in the positive y direction. The mass resolution is 1.5 and 1.8 MeV/ $c^2$  for  ${}^3_{\Lambda}$ H and  ${}^4_{\Lambda}$ H, respectively.

The reconstructed  ${}^{3}_{\Lambda}$ H and  ${}^{4}_{\Lambda}$ H candidates are further divided into different  $L/\beta\gamma$  intervals, where L is the decay

length,  $\beta$  and  $\gamma$  are particle velocity divided by the speed of light and Lorentz factor, respectively. The raw signal counts,  $N^{\rm raw}$ , for each  $L/\beta\gamma$  interval are corrected for the TPC acceptance, tracking, and particle identification efficiency, using an embedding technique in which the TPC response to Monte Carlo (MC) hypernuclei and their decay daughters is simulated in the STAR detector described in GEANT3 [40]. Simulated signals are embedded into the real data and processed through the same reconstruction algorithm as in real data. The simulated hypernuclei, used for determining the efficiency correction, need to be reweighted in 2D phase space  $(p_T-y)$  such that the MC hypernuclei are distributed in a realistic manner. This can be constrained by comparing the reconstructed kinematic distributions  $(p_T, y)$  between simulation and real data. The corrected hypernuclei yield as a function of  $L/\beta\gamma$  is fitted with an exponential function (see Supplemental Material [35]) and the decay lifetime is determined as the negative inverse of the slope divided by the speed of light.

We consider four major sources of systematic uncertainties in the lifetime result: imperfect description of topological variables in the simulations, imperfect knowledge of the true kinematic distribution of the hypernuclei, the TPC tracking efficiency, and the signal extraction technique. Their contributions are estimated by varying the topological cuts, the MC hypernuclei  $p_T$ -y distributions, the TPC track quality selection cuts, and the background subtraction method. The possible contamination of the signal due to multibody decays of A > 3 hypernuclei is estimated using MC simulations and found to be negligible (< 0.1%) within our reconstructed hypernuclei mass window. The systematic uncertainties due to different sources are tabulated in Table I. They are assumed to be uncorrelated with each other and added in quadrature in the total systematic uncertainty. As a cross-check, we conducted the measurement of  $\Lambda$  lifetime from the same data and the result is consistent with the Particle Data Group value [41] (see Supplemental Material [35]).

The lifetime results measured at  $\sqrt{s_{NN}} = 3.0$  GeV and  $\sqrt{s_{NN}} = 7.2$  GeV are found to agree well with each other.

TABLE I. Summary of systematic uncertainties for the lifetime and top 10% most central dN/dy (|y| < 0.5) measurements using  $\sqrt{s_{NN}} = 3.0$  GeV data.

Source	Lifetime		dN/dy	
	$^{3}_{\Lambda}$ H	$^4_{\Lambda}{ m H}$	$^{3}_{\Lambda}$ H	$^4_{\Lambda}{ m H}$
Analysis cuts	5.5%	5.1%	15.1%	6.9%
Input MC	3.1%	1.8%	8.8%	3.8%
Tracking efficiency	5.0%	2.4%	14.1%	5.2%
Signal extraction	1.5%	0.7%	14.3%	7.7%
Extrapolation			13.6%	10.9%
Detector material	< 1%	< 1%	4.0%	2.0%
Total	8.2%	6.0%	31.9%	16.6%



FIG. 2.  ${}^{3}_{\Lambda}$ H (a) and  ${}^{4}_{\Lambda}$ H (b) measured lifetime, compared to previous measurements [3–5,7–11,42–48], theoretical calculations [49–54], and  $\tau_{\Lambda}$  [41]. Horizontal lines represent statistical uncertainties, while boxes represent systematic uncertainties. The experimental average lifetimes and the corresponding uncertainty of  ${}^{3}_{\Lambda}$ H and  ${}^{4}_{\Lambda}$ H are also shown as vertical blue shaded bands.

The combined results are 221 + 15(stat) + 19(syst) for  ${}^{3}_{\Lambda}\text{H}$ and 218 + 6(stat) + 13(syst) for  ${}^{4}_{\Lambda}\text{H}$ . As shown in Fig. 2, they are consistent with previous measurements from ALICE [7,8], STAR [10,11], HypHI [9], and early experiments using imaging techniques [3–5,10,42–48]. Using all the available experimental data, the average lifetimes of  ${}^{3}_{\Lambda}\text{H}$  and  ${}^{4}_{\Lambda}\text{H}$  are  $200 \pm 13$  ps and  $208 \pm 12$  ps, respectively, corresponding to  $(76 \pm 5)\%$  and  $(79 \pm 5)\%$  of  $\tau_{\Lambda}$ . All data from ALICE, STAR, and HypHI lie within  $1.5\sigma$  of the global averages. These precise data clearly indicate that the  ${}^{3}_{\Lambda}\text{H}$  and  ${}^{4}_{\Lambda}\text{H}$  lifetimes are considerably lower than  $\tau_{\Lambda}$ .

Early theoretical calculations of the  ${}^{\Lambda}_{\Lambda}$ H lifetime typically give values within 15% of  $\tau_{\Lambda}$  [50–52]. This can be explained by the loose binding of  $\Lambda$  in the  ${}^{\Lambda}_{\Lambda}$ H. A recent calculation [49] using a pionless effective field theory approach with  $\Lambda d$ degrees of freedom gives a  ${}^{\Lambda}_{\Lambda}$ H lifetime of  $\approx 98\%\tau_{\Lambda}$ . Meanwhile, it is shown in recent studies that incorporating attractive pion final state interactions, which has been previously disregarded, decreases the  ${}^{\Lambda}_{\Lambda}$ H lifetime by  $\sim 15\%$  [19,53]. This leads to a prediction of the  ${}^{\Lambda}_{\Lambda}$ H lifetime to be  $(81 \pm 2)\%$  of  $\tau_{\Lambda}$ , consistent with the world average.

For  ${}^{4}_{\Lambda}$ H, a recent estimation [54] based on the empirical isospin rule [55] agrees with the data within  $1\sigma$ . The isospin rule is based on the experimental ratio  $\Gamma(\Lambda \rightarrow n + \pi^{0})/\Gamma(\Lambda \rightarrow p + \pi^{-}) \approx 0.5$ , which leads to the prediction  $\tau({}^{4}_{\Lambda}\text{H})/\tau({}^{4}_{\Lambda}\text{He}) = (74 \pm 4)\%$  [54]. Combining the average value reported here and the previous  ${}^{4}_{\Lambda}$ He lifetime measurement [56,57], the measured ratio  $\tau({}^{4}_{\Lambda}\text{H})/\tau({}^{4}_{\Lambda}\text{He})$  is  $(83 \pm 6)\%$ , consistent with the expectation.

Previous measurements on light nuclei suggest that their production yields in heavy-ion collisions may be related to their internal nuclear structure [58]. Similar relations for hypernuclei are suggested by theoretical models [1]. To further examine the hypernuclear structure and its production mechanism in heavy-ion collisions, we report the first measurement of hypernuclei dN/dy in two centrality selections: top 0%-10% most central and 10%-50% midcentral collisions. The  $p_T$  spectra can be found in Supplemental Material [35], and are extrapolated down to zero  $p_T$  to obtain the  $p_T$ -integrated dN/dy. Different functions [59] are used to estimate the systematic uncertainties in the unmeasured region, which correspond to 32%-60% of the p<sub>T</sub>-integrated yield in various rapidity intervals, and introduce 8%-14% systematic uncertainties. Systematic uncertainties associated with analysis cuts, tracking efficiency, and signal extraction are estimated using the same method as for the lifetime measurement. We further consider the effect of the uncertainty in the simulated hypernuclei lifetime on the calculated reconstruction efficiency by varying the simulation's lifetime assumption within a  $1\sigma$  window of the average experimental lifetime, which leads to 8% and 4% uncertainty for  ${}^{3}_{\Lambda}$ H and  ${}^{4}_{\Lambda}$ H, respectively. Finally, hypernuclei may encounter Coulomb dissociation when traversing the gold target. The survival probability is estimated using a Monte Carlo method according to [60]. The results show the survival probability > 96(99)% for  ${}^{3}_{\Lambda}H$  ( ${}^{4}_{\Lambda}H$ ) in the kinematic regions considered for the analysis. The dissociation has a strong dependence on  $B_{\Lambda}$  of the hypernuclei. Systematic uncertainties are estimated by varying the  $B_{\Lambda}$  of the  ${}^{3}_{\Lambda}$ H and  ${}^{4}_{\Lambda}$ H, which are equal to  $0.27 \pm 0.08$  MeV and  $2.53 \pm 0.04$  MeV, respectively [61]. As a conservative estimate, we assign the systematic uncertainty by comparing the calculation using the central values of  $B_{\Lambda}$  and its  $2.5\sigma$  limits. A summary of the systematic uncertainties for the dN/dy measurement is listed in Table I.

The  $p_T$ -integrated yields of  ${}^3_{\Lambda}$ H and  ${}^4_{\Lambda}$ H times the branching ratio (BR) as a function of y are shown in Fig. 3. For  ${}^4_{\Lambda}$ H, we can see that the midrapidity distribution changes from convex to concave from 0%–10% to 10%–50% centrality. This change in shape is likely related to the change in the collision geometry, such as spectators playing a larger role in noncentral collisions.

Also shown in Fig. 3 are calculations from the transport model, JET AA Microscopic Transportation Model (JAM) [62] coupled with a coalescence prescription to all produced hadrons as an afterburner [63]. In this model, deuterons and tritons are formed through the coalescence of nucleons, and subsequently,  ${}_{\Lambda}^{3}$ H and  ${}_{\Lambda}^{4}$ H are formed through the coalescence of  $\Lambda$  baryons with deuterons or tritons. Coalescence takes place if the spatial coordinates and the relative momenta of the constituents are within a sphere of radius ( $r_C$ ,  $p_C$ ). It is found that calculations using coalescence parameters ( $r_C$ ,  $p_C$ ) of (4.5 fm, 0.3 GeV/c),



FIG. 3. BR × dN/dy as a function of rapidity y for  ${}^{3}_{\Lambda}$ H (black circles) and  ${}^{4}_{\Lambda}$ H (red circles) for (a) 0%–10% centrality and (b) 10%–50% centrality Au + Au collisions at  $\sqrt{s_{NN}}$  = 3.0 GeV. Vertical lines represent statistical uncertainties, while boxes represent systematic uncertainties. The dot-dashed lines represent coalescence (JAM) calculations. The coalescence parameters used are indicated in the text.

(4 fm, 0.3 GeV/*c*), (4 fm, 0.12 GeV/*c*), and (4 fm, 0.3 GeV/*c*) for *d*, *t*,  ${}^{3}_{\Lambda}$ H, and  ${}^{4}_{\Lambda}$ H, respectively, can qualitatively reproduce the centrality and rapidity dependence of the measured yields. The smaller *p*<sub>*C*</sub> parameter used for  ${}^{3}_{\Lambda}$ H formation is motivated by its much smaller *B*<sub> $\Lambda$ </sub> (~0.3 MeV) compared to  ${}^{4}_{\Lambda}$ H (~2.6 MeV). The data offer first quantitative input on the coalescence parameters for hypernuclei formation in the high baryon density region, enabling more accurate estimations of the production yields of exotic strange objects, such as strange dibaryons [1].

The decay BR of  ${}^{3}_{\Lambda}$ H  $\rightarrow$   ${}^{3}$ He +  $\pi^{-}$  was not directly measured. A variation in the range 15%–35% for the BR [11,51,52] is considered when calculating the total dN/dy. For  ${}^{4}_{\Lambda}$ H  $\rightarrow$   ${}^{4}$ He +  $\pi^{-}$ , a variation of 40%–60% based on [17,56] is considered in this analysis.

The  ${}^{3}_{\Lambda}$ H and  ${}^{4}_{\Lambda}$ H midrapidity yields for central collisions as a function of center-of-mass energy are shown in Fig. 4. The uncertainties on the BRs are not shown in the main panels. Instead, the insets show the  $dN/dy \times$  BR as a function of BR. We observe that the  ${}^{3}_{\Lambda}$ H yield at  $\sqrt{s_{NN}} =$ 3.0 GeV is significantly enhanced compared to the yield at  $\sqrt{s_{NN}} = 2.76$  TeV [8], likely driven by the increase in baryon density at low energies.

Calculations from the thermal model, which adopts the canonical ensemble for strangeness [64] that is mandatory at low beam energies [65] are compared to data. Uncertainties arising from the strangeness canonical volume are indicated by the shaded red bands.  $\gamma$  decay of the excited state  ${}^{4}_{\Lambda}$ H(1<sup>+</sup>) to the ground state is accounted for in this calculation. Interestingly, while the  ${}^{3}_{\Lambda}$ H yields at  $\sqrt{s_{NN}} = 3.0$  GeV and 2.76 TeV are well described by the model, the  ${}^{4}_{\Lambda}$ H yield is underestimated by approximately a factor of 4. Coalescence calculations using DCM, an intranuclear cascade model to describe the dynamical stage of the reaction [1], are consistent with the  ${}^{3}_{\Lambda}$ H yield while underestimating the  ${}^{4}_{\Lambda}$ H yield, whereas the



FIG. 4. (a)  ${}^{3}_{\Lambda}$ H and (b)  ${}^{4}_{\Lambda}$ H yields at |y| < 0.5 as a function of beam energy in central heavy-ion collisions. The symbols represent measurements [8] while the lines represent different theoretical calculations. The data points assume a BR of 25(50)% for  ${}^{3}_{\Lambda}$ H( ${}^{4}_{\Lambda}$ H)  $\rightarrow$   ${}^{3}$ He( ${}^{4}$ He) +  $\pi^{-}$ . The insets show the (a)  ${}^{3}_{\Lambda}$ H and (b)  ${}^{4}_{\Lambda}$ H yields at |y| < 0.5 times the BR as a function of the BR. Vertical lines represent statistical uncertainties, while boxes represent systematic uncertainties.

coalescence (JAM) calculations are consistent with both. We note that in the DCM model, the same coalescence parameters are assumed for  ${}^{3}_{\Lambda}H$  and  ${}^{4}_{\Lambda}H$ , while in the JAM model, parameters are tuned separately for  ${}^{3}_{\Lambda}$  H and  ${}^{4}_{\Lambda}$  H to fit the data. It is expected that the calculated hypernuclei yields depend on the choice of the coalescence parameters [1]. Recent calculations from PHQMD [66,67], a microscopic transport model which utilizes a dynamical description of hypernuclei formation, is consistent with the measured yields within uncertainties. Compared to the JAM model which adopts a baryonic mean-field approach, baryonic interactions in PHQMD are modeled by density dependent two-body baryonic potentials. Meanwhile, the UrQMD-hydro hybrid model overestimates the yields at  $\sqrt{s_{NN}} = 3.0 \text{ GeV}$  by an order of magnitude. Our measurements possess distinguishing power between different production models, and provide new baselines for the strangeness canonical volume in thermal models and coalescence parameters in transport-coalescence models. Such constraints can be utilized to improve model estimations on the production of exotic strange matter in the high baryon density region.

In summary, precise measurements of  ${}^{3}_{\Lambda}$ H and  ${}^{4}_{\Lambda}$ H lifetimes have been obtained using the data samples of Au + Au collisions at  $\sqrt{s_{NN}} = 3.0$  and 7.2 GeV. The lifetimes are measured to be 221 + 15(stat) + 19(syst) for  ${}^{3}_{\Lambda}$ H and 218 + 6(stat) + 13(syst) for  ${}^{4}_{\Lambda}$ H. The averaged  ${}^{3}_{\Lambda}$ H and  ${}^{4}_{\Lambda}$ H lifetimes combining all existing measurements

are both smaller than  $\tau_{\Lambda}$  by ~20%. The precise  ${}^{\Lambda}_{\Lambda}$ H lifetime reported here resolves the tension between STAR and ALICE. We also present the first measurement of rapidity density of  ${}^{\Lambda}_{\Lambda}$ H and  ${}^{\Lambda}_{\Lambda}$ H in 0%–10% and 10%–50%  $\sqrt{s_{NN}} =$ 3.0 GeV Au + Au collisions. Hadronic transport models JAM and PHQMD calculations reproduce the measured midrapidity  ${}^{\Lambda}_{\Lambda}$ H and  ${}^{A}_{\Lambda}$ H yields reasonably well. Thermal model predictions are consistent with the  ${}^{\Lambda}_{\Lambda}$ H yield. Meanwhile, the same model underestimates the  ${}^{A}_{\Lambda}$ H yield. We observe that the  ${}^{\Lambda}_{\Lambda}$ H yield at this energy is significantly higher compared to those at  $\sqrt{s_{NN}} = 2.76$  TeV. This observation establishes low energy collision experiments as a promising tool to study exotic strange matter.

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\*Deceased.

- [1] J. Steinheimer, K. Gudima, A. Botvina, I. Mishustin, M. Bleicher, and H. Stöcker, Phys. Lett. B **714**, 85 (2012).
- [2] H. Bhang, S. Ajimura, K. Aoki, T. Hasegawa, O. Hashimoto *et al.*, Phys. Rev. Lett. **81**, 4321 (1998).
- [3] R. J. Prem and P. H. Steinberg, Phys. Rev. **136**, B1803 (1964).
- [4] R. E. Phillips and J. Schneps, Phys. Rev. 180, 1307 (1969).
- [5] Y. W. Kang, N. Kwak, J. Schneps, and P. A. Smith, Phys. Rev. 139, B401 (1965).
- [6] G. Bohm, J. Klabuhn, U. Krecker, F. Wysotzki, G. Bertrand-Coremans, J. Sacton, J. Wickens, D. H. Davis, J. E. Allen, and K. Garbowska-Pniewska, Nucl. Phys. B23, 93 (1970).
- [7] S. Acharya *et al.* (ALICE Collaboration), Phys. Lett. B **797**, 134905 (2019).

- [8] J. Adam *et al.* (ALICE Collaboration), Phys. Lett. B 754, 360 (2016).
- [9] C. Rappold et al., Nucl. Phys. A913, 170 (2013).
- [10] B. I. Abelev *et al.* (STAR Collaboration), Science **328**, 58 (2010).
- [11] L. Adamczyk *et al.* (STAR Collaboration), Phys. Rev. C 97, 054909 (2018).
- [12] T. Saito et al., Nat. Rev. Phys. 3, 803 (2021).
- [13] A. Gal, E. V. Hungerford, and D. J. Millener, Rev. Mod. Phys. 88, 035004 (2016).
- [14] R. H. Dalitz, R. C. Herndon, and Y. C. Tang, Nucl. Phys. B47, 109 (1972).
- [15] R. C. Herndon, Y. C. Tang, and E. W. Schmid, Phys. Rev. 137, B294 (1965).
- [16] N. N. Kolesnikov and S. A. Kalachev, Phys. Part. Nucl. Lett. 3, 341 (2006).
- [17] M. Juric et al., Nucl. Phys. B52, 1 (1973).
- [18] J. Adam et al. (STAR Collaboration), Nat. Phys. 16, 409 (2020).
- [19] A. Pérez-Obiol, D. Gazda, E. Friedman, and A. Gal, Phys. Lett. B 811, 135916 (2020).
- [20] P. Demorest, T. Pennucci, S. Ransom, M. Roberts, and J. Hessels, Nature (London) 467, 1081 (2010).
- [21] J. Antoniadis et al., Science 340, 6131 (2013).
- [22] E. D. Barr, P. C. C. Freire, M. Kramer, D. J. Champion, M. Berezina, C. G. Bassa, A. G. Lyne, and B. W. Stappers, Mon. Not. R. Astron. Soc. 465, 1711 (2017).
- [23] A. Andronic, P. Braun-Munzinger, J. Stachel, and H. Stocker, Phys. Lett. B 697, 203 (2011).
- [24] Z.-Q. Feng, Phys. Rev. C 102, 044604 (2020).
- [25] K.-J. Sun, C. M. Ko, and B. Dönigus, Phys. Lett. B 792, 132 (2019).
- [26] T. A. Armstrong, K. N. Barish, S. Batsouli, S. J. Bennett, M. Bertaina *et al.* (E864 Collaboration), Phys. Rev. C 70, 024902 (2004).
- [27] C. Rappold et al., Phys. Lett. B 747, 129 (2015).
- [28] J. Chen, D. Keane, Y.-G. Ma, A. Tang, and Z. Xu, Phys. Rep. 760, 1 (2018).
- [29] T. Ablyazimov *et al.* (CBM Collaboration), Eur. Phys. J. A 53, 60 (2017).
- [30] J. Ahn *et al.* (J-PARC-HI Collaboration), Letter of Intent for J-PARC Heavy-Ion Program (J-PARC-HI) (2016), https:// j-parc.jp/researcher/Hadron/en/Proposal\_e.html#1707.
- [31] R. L. Jaffe, Phys. Rev. Lett. 38, 195 (1977); 38, 617(E) (1977).
- [32] K. H. Ackermann *et al.* (STAR Collaboration), Nucl. Instrum. Methods Phys. Res., Sect. A 499, 624 (2003).
- [33] C. A. Whitten (STAR Collaboration), AIP Conf. Proc. 980, 390 (2008).
- [34] W. J. Llope (for STAR Collaboration), Nucl. Instrum. Methods Phys. Res., Sect. A 661, S110 (2012).
- [35] See Supplemental Material at http://link.aps.org/supplemental/ 10.1103/PhysRevLett.128.202301 for additional details on the data analysis.
- [36] M. Anderson *et al.*, Nucl. Instrum. Methods Phys. Res., Sect. A **499**, 659 (2003).
- [37] R. L. Ray and M. Daugherity, J. Phys. G 35, 125106 (2008).
- [38] J. Adam *et al.* (STAR Collaboration), Phys. Rev. C 103, 034908 (2021).
- [39] M. Zyzak *et al.*, The KFParticle package for the fast particle reconstruction in ALICE and CBM, Verhandlungen der Deutschen Physikalischen Gesellschaft (2013).

- [40] R. Brun et al., Report No. CERN-DD-EE-84-1, 1987.
- [41] P. Zyla *et al.* (Particle Data Group), Prog. Theor. Exp. Phys. 2020, 083C01 (2020).
- [42] G. Keyes, J. Sacton, J. H. Wickens, and M. M. Block, Nucl. Phys. B67, 269 (1973).
- [43] G. Keyes, M. Derrick, T. Fields, L. G. Hyman, J. G. Fetkovich, J. McKenzie, B. Riley, and I.-T. Wang, Phys. Rev. D 1, 66 (1970).
- [44] G. Bohm et al., Nucl. Phys. B16, 46 (1970); B16, 523(E) (1970).
- [45] G. Keyes, M. Derrick, T. Fields, L. G. Hyman, J. G. Fetkovich, J. McKenzie, B. Riley, and I.-T. Wang, Phys. Rev. Lett. 20, 819 (1968).
- [46] S. A. Avramenko et al., Nucl. Phys. A547, 95 (1992).
- [47] H. Outa, M. Aoki, R. S. Hayano, T. Ishikawa, M. Iwasaki, A. Sakaguchi, E. Takada, H. Tamura, and T. Yamazaki, Nucl. Phys. A585, 109 (1995).
- [48] N. Crayton et al., in Proceedings of the11th International Conference on High-Energy Physics (CERN, Geneva, 1962), pp. 460–462, https://cds.cern.ch/record/108582? ln=en or https://inspirehep.net/literature/1187848.
- [49] F. Hildenbrand and H. W. Hammer, Phys. Rev. C 102, 064002 (2020).
- [50] M. Rayet and R. H. Dalitz, Nuovo Cimento A **46**, 786 (1966).
- [51] J.G. Congleton, J. Phys. G 18, 339 (1992).
- [52] H. Kamada, J. Golak, K. Miyagawa, H. Witala, and W. Glockle, Phys. Rev. C 57, 1595 (1998).
- [53] A. Gal and H. Garcilazo, Phys. Lett. B 791, 48 (2019).
- [54] A. Gal, EPJ Web Conf. 259, 08002 (2022).
- [55] J. Cohen, Phys. Rev. C 42, 2724 (1990).
- [56] H. Outa, M. Aoki, R. S. Hayano, T. Ishikawa, M. Iwasaki, A. Sakaguchi, E. Takada, H. Tamura, and T. Yamazaki, Nucl. Phys. A639, 251c (1998).
- [57] J. D. Parker, M. J. Athanas, P. D. Barnes, S. Bart, B. Bassalleck *et al.*, Phys. Rev. C **76**, 035501 (2007); **76**, 039904(E) (2007).
- [58] T. A. Armstrong, K. N. Barish, S. Batsouli, S. J. Bennett, A. Chikanian *et al.* (E864 Collaboration), Phys. Rev. Lett. 83, 5431 (1999).
- [59] B. I. Abelev *et al.* (STAR Collaboration), Phys. Rev. C 79, 034909 (2009).
- [60] V. L. Lyuboshitz and V. V. Lyuboshitz, Phys. At. Nucl. 70, 1617 (2007).
- [61] P. Liu, J. Chen, D. Keane, Z. Xu, and Y.-G. Ma, Chin. Phys. C 43, 124001 (2019).
- [62] Y. Nara, N. Otuka, A. Ohnishi, K. Niita, and S. Chiba, Phys. Rev. C 61, 024901 (1999).
- [63] H. Liu, D. Zhang, S. He, K.-j. Sun, N. Yu, and X. Luo, Phys. Lett. B 805, 135452 (2020).
- [64] A. Andronic, P. Braun-Munzinger, J. Stachel, and H. Stöcker, Phys. Lett. B 697, 203 (2011).
- [65] M. Abdallah et al. (STAR Collaboration), arXiv:2108.00924.
- [66] S. Gläßel, V. Kireyeu, V. Voronyuk, J. Aichelin, C. Blume, E. Bratkovskaya, G. Coci, V. Kolesnikov, and M. Winn, Phys. Rev. C 105, 014908 (2022).
- [67] V. Kireyeu, J. Aichelin, E. Bratkovskaya, A. Le Fèvre, V. Lenivenko, V. Kolesnikov, Y. Leifels, and V. Voronyuk, Bull. Russ. Acad. Sci. Phys. 84, 957 (2020).