

Universal AdS Black Holes in Theories with 16 Supercharges and Their Microstates

Seyed Morteza Hosseini^{1,*} and Alberto Zaffaroni^{2,3,†}

¹*Kavli IPMU (WPI), UTIAS, The University of Tokyo, Kashiwa, Chiba 277-8583, Japan*

²*Dipartimento di Fisica, Università di Milano-Bicocca, I-20126 Milano, Italy*

³*INFN, sezione di Milano-Bicocca, I-20126 Milano, Italy*



(Received 21 December 2020; accepted 2 April 2021; published 28 April 2021)

We provide a universal microscopic counting for the microstates of the asymptotically AdS black holes and black strings that arise as solutions of the half-maximal gauged supergravity in 4 and 5 dimensions. These solutions can be embedded in all M-theory and type II string backgrounds with an AdS vacuum and 16 supercharges and provide an infinite set of examples dual to $\mathcal{N} = 2$ and $\mathcal{N} = 4$ conformal field theories in four and three dimensions, respectively. The counting is universal and it is performed by either studying the large N limit of the relevant supersymmetric index of the dual field theory or by using the charged Cardy formula.

DOI: [10.1103/PhysRevLett.126.171604](https://doi.org/10.1103/PhysRevLett.126.171604)

Introduction.—The microscopic counting of black hole microstates is a fundamental question for all theories of quantum gravity. String theory has provided a microscopic explanation for the entropy of a class of asymptotically flat black holes [1]. Despite the AdS/CFT correspondence [2], the analogous question for asymptotically anti-de Sitter (AdS) black holes has remained elusive until recently, except for AdS₃. In the last few years there has been some progress, first for a class of magnetically charged supersymmetric black holes in AdS₄ [3], and later for a class of supersymmetric Kerr-Newman (KN) black holes in AdS₅ [4,5]. The AdS/CFT correspondence provides a nonperturbative definition of quantum gravity in asymptotically AdS space in terms of a dual boundary quantum field theory (QFT) and the black hole microstates appear as particular states in the boundary description. The entropy of a supersymmetric black hole with angular momentum J and a set of conserved electric and magnetic charges is reproduced by counting the states with spin J and the same quantum numbers in the dual QFT. Computations in a strongly coupled supersymmetric QFT are difficult but the numbers of states of interest can be extracted from supersymmetric indices that can be often evaluated using exact nonperturbative techniques. Supersymmetric localization [6], for example, allows one to reduce the indices to matrix models that can be evaluated in a saddle-point approximation when the number of colors N is large, which is the regime where holography applies.

Unfortunately, asymptotically AdS black holes are difficult to find, and not so many 4- and 5-dimensional examples are known besides those that can be embedded in AdS₄ × S⁷ or AdS₅ × S⁵ and some universal examples that arise from embedding minimal gauged supergravity into string compactifications [7–10]. Some progress has been made instead in constructing infinite classes of supersymmetric AdS₂ and AdS₃ solutions that can arise as the near horizon limit of AdS black objects [11,12]. In this Letter, in the spirit of the above-mentioned universal examples, we consider a large class of black holes and black strings that arise as solutions of the half-maximal supergravity in AdS₄ and AdS₅. Such solutions can be embedded in all AdS₄ and AdS₅ type II or M-theory backgrounds with 16 supercharges. Indeed, for any supersymmetric solution of 10- or 11-dimensional supergravity of the warped product form AdS_D ×_w M, there is a consistent truncation to pure gauged supergravity in D dimensions containing that solution and having the same supersymmetry [13–16]. This observation calls for a *universal* large N formula for the number of supersymmetric states in conformal field theories (CFT) with 8 supercharges (16 including conformal supersymmetries). We will indeed perform a universal counting of states in CFTs with 8 supercharges using the superconformal index and the charged Cardy formula. Extrapolating from known results about the large N behavior of the index in various limits, we precisely reproduce the Bekenstein-Hawking entropy of the black objects. We will check the results for the known classes of CFTs with 8 supercharges that admit a holographic dual. We will also provide a conjecture for the R -symmetry charge dependence of the S³ free energy of 3D CFTs with 8 supercharges. A similar universal computation was done in [7] and in [8–10] for theories with 4 supercharges. For other related universal results see [17–19].

Published by the American Physical Society under the terms of the Creative Commons Attribution 4.0 International license. Further distribution of this work must maintain attribution to the author(s) and the published article's title, journal citation, and DOI. Funded by SCOAP³.

$\mathcal{N} = 4$ gauged supergravity in 5 dimensions.—The bosonic part of minimal 5D $\mathcal{N} = 4$ gauged supergravity [20,21] consists of the metric $g_{\mu\nu}$, a $U(1)_R$ gauge field a_μ , an $SU(2)_R$ Yang-Mills gauge field A_μ^I , $I = 1, 2, 3$, two antisymmetric tensor fields $B_{\mu\nu}^\alpha$, $\alpha = 4, 5$, and one real scalar ϕ . The fermionic components are four gravitini $\psi_{\mu i}$, $i = 1, 2, 3, 4$ and four spin-1/2 fermions χ_i . These fields form the $\mathcal{N} = 4$ gauged supergravity multiplet $(g_{\mu\nu}, \psi_{\mu i}, a_\mu, A_\mu^I, B_{\mu\nu}^\alpha, \chi_i, \phi)$. The bosonic Lagrangian is given by

$$\begin{aligned} \frac{1}{\sqrt{-g}}\mathcal{L} = & -\frac{R}{4} + \frac{1}{2}(\partial^\mu\phi)(\partial_\mu\phi) + \frac{g_2}{8}(g_2\xi^{-2} + 2\sqrt{2}g_1\xi) \\ & - \frac{\xi^{-4}}{4}f^{\mu\nu}f_{\mu\nu} - \frac{\xi^2}{4}(F^{\mu\nu I}F_{\mu\nu}^I + B^{\mu\nu\alpha}B_{\mu\nu}^\alpha) \\ & + \frac{1}{4\sqrt{-g}}\varepsilon^{\mu\nu\rho\sigma\tau}\left(\frac{1}{g_1}\varepsilon_{\alpha\beta}B_{\mu\nu}^\alpha D_\rho B_{\sigma\tau}^\beta - F_{\mu\nu}^I F_{\rho\sigma}^I a_\tau\right), \end{aligned}$$

where $\xi = \exp(\sqrt{2/3}\phi)$, and the field strengths are $f_\mu = \partial_\nu a_\mu - \partial_\mu a_\nu$, $F_{\mu\nu}^I = \partial_\nu A_\mu^I - \partial_\mu A_\nu^I + g_2\varepsilon^{IJK}A_\mu^J A_\nu^K$. The theory with $g_2 = \sqrt{2}g_1 = 2\sqrt{2}$ has an AdS vacuum with radius $\ell_5 = 1$ that preserves all of the 16 real supercharges.

The universal KN black hole in AdS₅.—There is a universal solution of $\mathcal{N} = 4$ gauged supergravity corresponding to a supersymmetric, asymptotically AdS black hole with two electric charges Q_1 and Q_2 under $U(1)^2 \subset U(1)_R \times SU(2)_R$ and two angular momenta J_1 and J_2 . It can be obtained as a particular case of the KN black holes in $AdS_5 \times S^5$ with angular momenta (J_1, J_2) and electric charges (Q_1, Q_2, Q_2) under the Cartan subgroup of $SO(6)$ [22–26]. The latter are solutions of the $U(1)^3$ truncation of $AdS_5 \times S^5$ and one can easily check that they become solutions of the minimal $\mathcal{N} = 4$ gauged supergravity when two $U(1)$ gauge fields are identified. The entropy can be compactly written as [27]

$$S(J_I, Q_i) = 2\pi\sqrt{Q_2^2 + 2Q_1Q_2 - 2a(J_1 + J_2)}, \quad (1)$$

where $a = \{[\pi\ell_5^3]/[8G_N^{(5)}]\}$, with ℓ_5 being the radius of AdS_5 and $G_N^{(5)}$ the Newton constant is the central charge of the dual $\mathcal{N} = 2$ CFT [28] at leading order in N . In order for the black hole to have a smooth horizon, the charges must satisfy the nonlinear constraint

$$\begin{aligned} 0 = & 2Q_2(Q_1 + Q_2)^2 + 2a(Q_2^2 + 2Q_1Q_2) \\ & - 2a(J_1 + J_2)(Q_1 + 2Q_2) - 2aJ_1J_2 - 4a^2(J_1 + J_2). \end{aligned} \quad (2)$$

The entropy (1) can be written as the constrained Legendre transform of the quantity [29]

$$\log \mathcal{Z}(X_i, \omega_i) = -4\pi i a \frac{X_1 X_2^2}{\omega_1 \omega_2}. \quad (3)$$

The entropy is obtained indeed by extremizing

$$S = \log \mathcal{Z}(X_i, \omega_i) - 2\pi i(\omega_1 J_1 + \omega_2 J_2 + X_1 Q_1 + 2X_2 Q_2),$$

with respect to X_i and ω_i with the constraint $X_1 + 2X_2 - \omega_1 - \omega_2 = \pm 1$. Both signs lead to the same critical value (1) [30], which is real precisely when (2) is satisfied. The solution can be embedded in all AdS_5 type II or M-theory backgrounds preserving 16 supercharges.

Examples of $\mathcal{N} = 2$ quiver theories with a holographic dual and central charges of order $\mathcal{O}(N^2)$ are provided by orbifolds of $AdS_5 \times S^5$. Another large class of $\mathcal{N} = 2$ theories with a holographic dual can be obtained by compactifying N M5-branes on a Riemann surface Σ_g of genus g with regular punctures [31]. We will refer to these theories as holographic class \mathcal{S} theories. The theories are generically non-Lagrangian and their central charge is of order $\mathcal{O}(N^3)$. On general grounds [14,15], one knows that the effective 5D dual gravitational theory can be consistently truncated to the minimal $\mathcal{N} = 4$ gauged supergravity. As an example, one can check explicitly that the universal KN solution can be embedded in the compactification with no punctures, corresponding to the M-theory AdS_5 solution originally found in [32]. To this purpose we can use the 5D consistent truncation of 7D $U(1)^2$ gauged supergravity on Σ_g derived in [33–35]. The corresponding 5D theory is written as an $\mathcal{N} = 2$ gauged supergravity with two vector multiplets and one hypermultiplet. At the AdS_5 vacuum one vector multiplet becomes massive through a Higgs mechanism. The theory depends on a parameter z that specifies the twist along Σ_g and the preserved supersymmetry is generically $\mathcal{N} = 2$ corresponding to a dual $\mathcal{N} = 1$ CFT. For the special values $z = \pm 1$ supersymmetry is enhanced to $\mathcal{N} = 4$, and, for $g > 1$, one obtains a dual $\mathcal{N} = 2$ CFT. One can check that, for $z = \pm 1$ and $g > 1$, by setting the hyperscalars to their AdS_5 value and setting to zero the fields in the massive vector multiplet, equations of motion and supersymmetry variations of the theory in [33] coincide with those of the $U(1)^2$ sector of the minimal $\mathcal{N} = 4$ gauged supergravity. Explicitly, in the notations of [33] we need to set $e^{-10B/3} = 2^{-1/3}m^{10/3}\zeta$, $e^{10\lambda_1/9} = e^{-5\lambda_2/3} = 2^{-1/3}\zeta$ and $A^{(0)} = m^{-4/3}2^{-1/6}A$, with $g_2 = 2^{5/6}m^{5/3}$.

We will now show that the entropy (1) matches the prediction of a microscopic computation based on the superconformal index for a generic $\mathcal{N} = 2$ CFT with a holographic dual. The number of BPS states with charges Q_1, Q_2 and spin J_1, J_2 in an $\mathcal{N} = 2$ CFT can be computed (for large charges and spins) by taking the Legendre transform of the superconformal index \mathcal{I} , which is a function of chemical potentials for the electric charges and the angular momenta. The agreement between the gravitational picture and the field theory computation

requires $\log \mathcal{I} = \log \mathcal{Z}$. The general expectation for an $\mathcal{N} = 1$ CFT with a holographic dual is [36]

$$\log \mathcal{I}(\Delta, \omega_i) = -\frac{4\pi i (\omega_1 + \omega_2 \pm 1)^3}{27 \omega_1 \omega_2} a(\Delta), \quad (4)$$

valid at large N . In this formula ω_1 and ω_2 are chemical potentials conjugated to J_1 and J_2 and $(\omega_1 + \omega_2 \pm 1)\Delta/2$ is a set of chemical potentials for the R and flavor symmetries of the theory. They are normalized such that Δ can be interpreted as an assignment of R charges to the fields of the theory with the only constraint that the superpotential has R -charge two. Moreover, $a(\Delta) = \frac{9}{32} \text{Tr} R(\Delta)^3$, where $R(\Delta)$ is the R -symmetry generator and the trace is taken over the fermionic fields, is the trial a charge at large N [37]. For Lagrangian theories with a holographic dual, formula (4) has been derived in the large N limit in [4,10,38,39]. It is also compatible with the Cardy limit performed in [5,8,9,40]. It has not been yet derived in full generality for non-Lagrangian theories. However, the Cardy limit can be also derived by writing the effective theory of the CFT coupled to background fields on $S^3 \times S^1$ in the limit where the circle shrinks [5,9] and this method applies also to non-Lagrangian theories. The two signs in (4) arise in the saddle-point evaluation of the index in different regions in the space of chemical potentials. Consider first an $\mathcal{N} = 2$ Lagrangian CFT with n_V vector multiplets and n_H hypermultiplets. In $\mathcal{N} = 1$ language, the theory can be described by n_V vector multiplets, n_V chiral multiplets ϕ_I , and n_H pairs of chiral multiplets (q_a, \tilde{q}_a) . The trial R symmetry can be written as $R(\Delta) = (\Delta_1 r_1 + \Delta_2 r_2)/2$ with $\Delta_1 + \Delta_2 = 2$, where r_1 is the $U(1)_R$ symmetry assigning charge 2 to ϕ_I and zero to q_a, \tilde{q}_a and r_2 is the Cartan generator of $SU(2)_R$ assigning charge zero to ϕ_I and charge 1 to q_a, \tilde{q}_a . In the gravitational dual, Q_1 is the charge under $r_1/2$ and Q_2 under r_2 . Notice that the exact R symmetry corresponds to $\hat{\Delta}_1 = \frac{2}{3}$, $\hat{\Delta}_2 = \frac{4}{3}$ leading to canonical dimensions for the fields. We easily compute

$$\begin{aligned} \text{Tr} R(\Delta) &= (n_V - n_H)\Delta_1, \\ \text{Tr} R(\Delta)^3 &= \frac{3n_V}{4} \Delta_1 \Delta_2^2 + \frac{n_V - n_H}{4} \Delta_1^3. \end{aligned} \quad (5)$$

Holography requires $a = c$ at large N [28]. Using $16(a - c) = \text{Tr} R$ [41], this implies $n_V = n_H$ at leading order in N . We then find

$$\log \mathcal{I}(\Delta, \omega_i) = -\frac{\pi i a (\omega_1 + \omega_2 \pm 1)^3}{8 \omega_1 \omega_2} \Delta_1 \Delta_2^2, \quad (6)$$

where $a \equiv a(\hat{\Delta}) = (n_V/4)$ denotes the exact central charge of the CFT at large N . We see that, with the redefinitions $X_1 = (\omega_1 + \omega_2 \pm 1)(\Delta_1/2)$ and $X_2 = (\omega_1 + \omega_2 \pm 1)(\Delta_2/4)$, we recover (3) and the

constraint $X_1 + 2X_2 - \omega_1 - \omega_2 = \pm 1$. For non-Lagrangian theories, we just replace n_V and n_H with an effective number of vector multiplets and hypermultiplets, $n_V = 4(2a - c)$, $n_H = 4(5c - 4a)$ [31]. The structure of (5) is completely fixed by $\mathcal{N} = 2$ superconformal invariance [42] and the previous argument still holds.

The universal black string in AdS₅.—We can similarly find a universal black string solution of $\mathcal{N} = 4$ gauged supergravity as a special case of the black strings in $\text{AdS}_5 \times S^5$ found in [43] with angular momentum J , electric charges (Q_1, Q_2, Q_3) , and magnetic charges (p_1, p_2, p_3) when we set $Q_3 = Q_2$ and $p_3 = p_2$. The horizon geometry of this solution has the topology of a warped product $\text{BTZ} \times {}_{\omega}S^2$ and carries an extra conserved charge corresponding to a momentum Q_0 along the Bañados-Teitelboim-Zanelli (BTZ) circle. Upon compactification on the circle we obtain a 4D dyonic black hole with Lifshitz-like asymptotics. In the special case where all of the electric charges Q_i and J are zero, the 5D solution is a domain wall interpolating between AdS_5 and $\text{AdS}_3 \times S^2$. There is a constraint on the magnetic charges $p_1 + 2p_2 = -1$, which corresponds to the fact that the dual field theory is topologically twisted along S^2 , and a further constraint involving the electric charges: $(1 + p_1)Q_2 + (1 + 2p_1)Q_1 = 0$. The entropy of the 4-dimensional black hole reads [43,44]

$$S(J, Q_i, p_i) = 2\pi \sqrt{\frac{c_{\text{CFT}}}{6} \left(Q_0 - \frac{J^2}{2k} - \frac{F^2}{2k_{FF}} \right)}, \quad (7)$$

where $F = Q_1 - Q_2$ and

$$\begin{aligned} c_{\text{CFT}} &= -24a \frac{(1 + p_1)^2}{2 + 3p_1}, \\ k &= -2ap_1(1 + p_1)^2, \quad k_{FF} = -2a(2 + 3p_1). \end{aligned} \quad (8)$$

We expect again that the solution can be embedded in any AdS_5 compactification preserving 16 supercharges. The special case of class \mathcal{S} with no punctures has been constructed in [44], as a part of a more general compactifications of M5-branes on a Riemann surface with 8 supercharges. It is not difficult to check that the solution in [44] becomes that of the universal black string when the theory has 16 supercharges ($\kappa = -1$, $z_1 = -1$, $\mathfrak{g}_1 = 2 - 2\mathfrak{g}$, and $\mathfrak{t}_1 = -2p_1$, in the notation of [44]), and the entropy coincides with (7).

The result is indeed universal from the field theory point of view. The black string is dual to a 2D CFT obtained by compactifying the $\mathcal{N} = 2$ 4D CFT on S^2 with a topological twist. The microscopic entropy is just the number of states of the 2D CFT with $L_0 = Q_0$, electric charges Q_1 and Q_2 under $r_1/2$ and r_2 and charge J under the additional infrared $SU(2)$ symmetry associated with rotation along S^2 . These states are accounted for by the charged Cardy formula. The latter has precisely the form (7), where c_{CFT} is

the central charge of the 2D CFT, k is the level of the rotational $SU(2)$ symmetry, and k_{FF} the level of the flavor symmetry $r_1/2 - r_2$ [44]. All these quantities can be computed with an (equivariant) integration of the 4D anomaly polynomial [44–46] and are universal because, as already noticed, the form of the 4-dimensional anomalies for an $\mathcal{N} = 2$ CFT with a holographic dual is universal and only depends on $a = (n_V/4)$. For details on the anomaly integration leading to (8) see [44].

$\mathcal{N} = 4$ gauged supergravity in 4 dimensions.—Four-dimensional $\mathcal{N} = 4$ $SO(4)$ gauged supergravity can be obtained as the consistent truncation of 11D supergravity on S^7 [47]. The bosonic field content of this theory is the metric $g_{\mu\nu}$, two $SU(2)$ gauge fields $(A_\mu^I, \tilde{A}_\mu^I)$, $I = 1, 2, 3$, a dilaton ϕ , and an axion χ . The fermionic components are four gravitini $\psi_{\mu i}$, $i = 1, 2, 3, 4$, and four spin-1/2 fermions χ_i . These fields form the $\mathcal{N} = 4$ gauged supergravity multiplet $(g_{\mu\nu}, \psi_{\mu i}, A_\mu^I, \tilde{A}_\mu^I, \chi_i, \phi, \chi)$. The bosonic Lagrangian is given by

$$\begin{aligned} \frac{1}{\sqrt{-g}}\mathcal{L} = & R - \frac{1}{2}(\partial^\mu\phi)(\partial_\mu\phi) - \frac{1}{2}e^{2\phi}(\partial^\mu\chi)(\partial_\mu\chi) \\ & + 2g^2[4 + 2\cosh(\phi) + \chi^2 e^\phi] \\ & - \frac{1}{2}e^{-\phi}F_{\mu\nu}^I F^{I\mu\nu} - \frac{1}{2}\frac{e^\phi}{1 + \chi^2 e^{2\phi}}\tilde{F}_{\mu\nu}^I \tilde{F}^{\mu\nu I} \\ & - \frac{\chi}{2\sqrt{-g}}\varepsilon_{\mu\nu\rho\sigma}\left(F^{\mu\nu I}F^{\rho\sigma I} - \frac{e^{2\phi}}{1 + \chi^2 e^{2\phi}}\tilde{F}^{\mu\nu I}\tilde{F}^{\rho\sigma I}\right). \end{aligned}$$

The universal KN black hole in AdS_4 .—The universal 4D KN solution can be obtained by specializing the general KN black holes in $AdS_4 \times S^7$ to have angular momentum J , electric charges (Q_1, Q_2, Q_1, Q_2) , and magnetic charges (p_1, p_2, p_1, p_2) under the Cartan subgroup of $SO(8)$ [48–50]. The magnetic charges are restricted to satisfy $p_1 = -p_2 \equiv p$, which corresponds to the absence of a topological twist, and there is a further constraint among charges

$$J = \frac{Q_1 + Q_2}{2}\left(-1 + \sqrt{1 - 16p^2 + \frac{4\pi^2}{F_{S^3}^2}Q_1 Q_2}\right). \quad (9)$$

The entropy, given by $S(p, Q_1, Q_2, J) = 2F_{S^3} \frac{J}{(Q_1 + Q_2)}$, where $F_{S^3} = \{[\pi\ell_4^2]/[2G_N^{(4)}]\}$ is the S^3 free energy of the dual CFT, can be obtained by extremizing the functional [51,52]

$$S = -F_{S^3} \frac{(\Delta_1 \Delta_2 - 4p^2 \omega^2)}{\omega} + \pi i \sum_{i=1}^2 \Delta_i Q_i + \pi i \omega J, \quad (10)$$

with the constraint $\Delta_1 + \Delta_2 - \omega = 2$. This solution can be embedded in any AdS_4 solution of M-theory or type II string theory with 16 supercharges.

The general expectation for a 3D $\mathcal{N} = 4$ CFT with a holographic dual is that the entropy is the Legendre transform of [52]

$$\log \mathcal{I}(\Delta, \omega) = -\frac{F_{S^3}(\Delta_i - 2\omega p_i)}{2\omega} - \frac{F_{S^3}(\Delta_i + 2\omega p_i)}{2\omega}, \quad (11)$$

where $F_{S^3}(\Delta_i)$ is the S^3 free energy as a function of the trial R symmetry [53] and $\Delta_1 + \Delta_2 = 2$. Δ_1 and Δ_2 are conjugated to the Cartan generators of the $SU(2) \times SU(2)$ R symmetry. This formula follows from gluing gravitational blocks in gravity [52], which is the counterpart of gluing holomorphic blocks in the dual field theory [54,55]. For $p = 0$ and the Aharony, Bergman, Jafferis, Maldacena (ABJM) theory, (11) has been derived in the Cardy limit in [56,57] by factorizing the superconformal index into vortex partition functions. It is expected to hold for more general theories and for $p \neq 0$. We then find a general prediction for the trial free energy of a generic $\mathcal{N} = 4$ CFT with a holographic dual in the large N limit

$$F_{S^3}(\Delta_i) = F_{S^3} \Delta_1 \Delta_2. \quad (12)$$

We can explicitly check this prediction in various examples. Holographic $\mathcal{N} = 4$ CFTs arise as world volume theories of M2-branes probing $\mathbb{C}^2/\Gamma_1 \times \mathbb{C}^2/\Gamma_2$, with Γ_i discrete subgroups of $SU(2)$, where the role of Γ_1 and Γ_2 can be exchanged by mirror symmetry [58]. Consider, for simplicity, the case where $\Gamma_2 = \mathbb{Z}_p$. The world volume theory is based on an $\mathcal{N} = 4$ ADE quiver with gauge groups $U(n_a N)$ corresponding to the nodes of the extended Dynkin diagram of Γ_1 and bifundamental hypermultiplets associated with the links, flavored with the addition of p fundamental hypermultiplets (n_a are the comarks: see [59] for conventions and details). Denote by n_V , n_B , and n_F the total number of vector multiplets, bifundamental and fundamental hypers, respectively. The large N limit of the S^3 partition function can be computed with the methods in [60–62]. In the large N limit the eigenvalue distribution for a group $U(n_a N)$ is given by n_a copies of the segment $\lambda(t) = N^{1/2}t$ with density $\rho(t)$ [$\int dt \rho(t) = 1$]. In the large N limit, using the rules in [61,62], we obtain

$$\begin{aligned} F_{S^3}(\Delta_i) = & \frac{n_B}{N^{1/2}} \frac{\pi^2}{6} \Delta_2 (\Delta_2 - 2)(\Delta_2 - 4) \int \rho(t)^2 dt \\ & + \frac{n_V}{N^{1/2}} \frac{2\pi^2}{3} \Delta_1 (\Delta_1 - 1)(\Delta_1 - 2) \int \rho(t)^2 dt \\ & + \frac{n_F N^{1/2}}{2} (2 - \Delta_2) \int \rho(t) |t| dt, \end{aligned} \quad (13)$$

where we assign charge Δ_1 to the adjoint chiral in the vector multiplet and $\Delta_2/2$ to the chiral fields q_a, \tilde{q}_a in the hypermultiplets, in $\mathcal{N} = 2$ notations. Since the ADE quivers are balanced (the number of hypers for each group

is twice the number of colors), we have $n_V = n_B$ and we find the saddle-point distribution

$$\rho(t) = \frac{\pi\sqrt{2n_F n_V N \Delta_2 - n_F N |t|}}{2\pi^2 n_V \Delta_2^2}, \quad (14)$$

with free energy

$$F_{S^3}(\Delta_i) = \frac{\pi}{3} \sqrt{2n_F n_V} \Delta_1 \Delta_2, \quad (15)$$

which reproduces (12) with $F_{S^3} = (\pi/3)\sqrt{2n_F n_V}$. The previous computation for $\Gamma_1 = \mathbb{Z}_q$ was already done in disguise in [63]. Notice that $F_{S^3} = \mathcal{O}(N^{3/2})$, as expected for M2-brane theories. Formula (15) can be also derived from the identification of the trial free energy with the volume functional of the transverse Calabi-Yau [61,64]. M2-branes probing Abelian hyper-Kähler orbifolds of \mathbb{C}^4 can be also realized in terms of $\mathcal{N} = 4$ circular quivers with nonzero Chern-Simons terms [65]. The simplest example is actually ABJM itself, whose free energy has been computed in [61] and reads $F_{S^3}(\delta_i) = 4F_{S^3} \sqrt{\delta_1 \delta_2 \delta_3 \delta_4}$, where δ_i are conjugated to the Cartan subgroup of $SO(8)$ and satisfy $\sum_{i=1}^4 \delta_i = 2$. This reduces to (12) for $\delta_3 = \delta_1 = \Delta_1/2$ and $\delta_4 = \delta_2 = \Delta_2/2$. Using and extending the results in [63,66], one can also compute the free energy for the more general $\mathcal{N} = 4$ quivers discussed in [65] and check that (12) is valid. Notice that in all these examples the $SU(2) \times SU(2)$ R symmetry acts differently from the case with no Chern-Simons and is fully visible once the theory is written in terms of both standard and twisted hypermultiplets [67–69]. Another large class of $\mathcal{N} = 4$ holographic quivers are the $T_\sigma^p(G)$ theories [70] whose gravitational dual was found in [71,72]. The refined free energy on S^3 for $T(SU(N))$ has been recently computed in the large N limit in [73] and it reads $F_{S^3} = \frac{1}{2} \Delta_1 \Delta_2 N^2 \log N$ which also agrees with (12). The prediction (12) can be also checked for a larger class of $T_\sigma^p(G)$ theories [74].

The universal twisted black hole in AdS_4 .—We can obtain dyonic black holes with a twist (magnetic charge for the R symmetry) and horizon $AdS_2 \times \Sigma_g$, where Σ_g is a Riemann surface of genus g , by specializing the corresponding solution in $AdS_4 \times S^7$ [75–77]. For $g = 0$ we can add an angular momentum J [78]. The case of static solutions of minimal gauged supergravity has been already discussed in [7,18]. Solutions with a generic $\mathcal{N} = 4$ choice of charges are not regular and we will not discuss them further. One can check however that the comparison between the gravity entropy functional and the large N limit of the (refined) topologically twisted index would still agree [since it is also based on (12) [52]], although the extremization leads to a nonphysical value for the entropy. For $J = 0$ one can find an off-shell agreement by considering the Euclidean black saddles discussed in [79].

S. M. H. is supported in part by WPI Initiative, MEXT, Japan at IPMU, the University of Tokyo, JSPS KAKENHI Grant-in-Aid (Wakate-A), No. 17H04837 and JSPS KAKENHI Grant-in-Aid (Early-Career Scientists), No. 20K14462. A.Z. is partially supported by the INFN, the ERC-STG Grant No. 637844-HBQFTNCER, and the MIUR-PRIN Contract No. 2017CC72MK003.

*morteza.hosseini@ipmu.jp

†alberto.zaffaroni@mib.infn.it

- [1] A. Strominger and C. Vafa, *Phys. Lett. B* **379**, 99 (1996).
- [2] J. M. Maldacena, *Int. J. Theor. Phys.* **38**, 1113 (1999).
- [3] F. Benini, K. Hristov, and A. Zaffaroni, *J. High Energy Phys.* **05** (2016) 054.
- [4] F. Benini and P. Milan, *Phys. Rev. X* **10**, 021037 (2020).
- [5] S. Choi, J. Kim, S. Kim, and J. Nahmgoong, [arXiv:1810.12067](https://arxiv.org/abs/1810.12067).
- [6] V. Pestun *et al.*, *J. Phys. A* **50**, 440301 (2017).
- [7] F. Azzurli, N. Bobev, P. M. Cricigno, V. S. Min, and A. Zaffaroni, *J. High Energy Phys.* **02** (2018) 054.
- [8] A. Cabo-Bizet, D. Cassani, D. Martelli, and S. Murthy, *J. High Energy Phys.* **08** (2019) 120.
- [9] J. Kim, S. Kim, and J. Song, *J. High Energy Phys.* **01** (2021) 025.
- [10] F. Benini, E. Colombo, S. Soltani, A. Zaffaroni, and Z. Zhang, *Classical Quantum Gravity* **37**, 215021 (2020).
- [11] C. Couzens, J. P. Gauntlett, D. Martelli, and J. Sparks, *J. High Energy Phys.* **01** (2019) 212.
- [12] J. P. Gauntlett, D. Martelli, and J. Sparks, *J. High Energy Phys.* **11** (2019) 176.
- [13] J. P. Gauntlett and O. Varela, *Phys. Rev. D* **76**, 126007 (2007).
- [14] J. P. Gauntlett and O. Varela, *J. High Energy Phys.* **02** (2008) 083.
- [15] D. Cassani, G. Josse, M. Petrini, and D. Waldram, *J. High Energy Phys.* **11** (2019) 017.
- [16] E. Malek, *Fortschr. Phys.* **65**, 1700061 (2017).
- [17] F. Benini, N. Bobev, and P. M. Cricigno, *J. High Energy Phys.* **07** (2016) 020.
- [18] N. Bobev and P. M. Cricigno, *J. High Energy Phys.* **12** (2017) 065.
- [19] N. Bobev and P. M. Cricigno, *J. High Energy Phys.* **12** (2019) 054.
- [20] L. Romans, *Nucl. Phys.* **B267**, 433 (1986).
- [21] M. Awada and P. Townsend, *Nucl. Phys.* **B255**, 617 (1985).
- [22] J. B. Gutowski and H. S. Reall, *J. High Energy Phys.* **02** (2004) 006.
- [23] J. B. Gutowski and H. S. Reall, *J. High Energy Phys.* **04** (2004) 048.
- [24] Z. W. Chong, M. Cvetič, H. Lu, and C. N. Pope, *Phys. Rev. D* **72**, 041901(R) (2005).
- [25] Z.-W. Chong, M. Cvetič, H. Lu, and C. N. Pope, *Phys. Rev. Lett.* **95**, 161301 (2005).
- [26] H. K. Kunduri, J. Lucietti, and H. S. Reall, *J. High Energy Phys.* **04** (2006) 036.
- [27] S. Kim and K.-M. Lee, *J. High Energy Phys.* **12** (2006) 077.
- [28] M. Henningson and K. Skenderis, *J. High Energy Phys.* **07** (1998) 023.

- [29] S. M. Hosseini, K. Hristov, and A. Zaffaroni, *J. High Energy Phys.* **07** (2017) 106.
- [30] A. Cabo-Bizet, D. Cassani, D. Martelli, and S. Murthy, *J. High Energy Phys.* **10** (2019) 062.
- [31] D. Gaiotto and J. Maldacena, *J. High Energy Phys.* **10** (2012) 189.
- [32] J. M. Maldacena and C. Nunez, *Int. J. Mod. Phys. A* **16**, 822 (2001).
- [33] P. Szepietowski, *J. High Energy Phys.* **12** (2012) 018.
- [34] D. Cassani, G. Josse, M. Petrini, and D. Waldram, *J. High Energy Phys.* **02** (2021) 232.
- [35] E. Malek and V. V. Camell, [arXiv:2012.15601](https://arxiv.org/abs/2012.15601).
- [36] S. M. Hosseini, K. Hristov, and A. Zaffaroni, *J. High Energy Phys.* **05** (2018) 121.
- [37] K. A. Intriligator and B. Wecht, *Nucl. Phys.* **B667**, 183 (2003).
- [38] A. Lanir, A. Nedelin, and O. Sela, *J. High Energy Phys.* **04** (2020) 091.
- [39] A. Cabo-Bizet, D. Cassani, D. Martelli, and S. Murthy, *J. High Energy Phys.* **11** (2020) 150.
- [40] A. Amariti, I. Garozzo, and G. Lo Monaco, [arXiv:1904.10009](https://arxiv.org/abs/1904.10009).
- [41] D. Anselmi, D. Freedman, M. T. Grisaru, and A. Johansen, *Nucl. Phys.* **B526**, 543 (1998).
- [42] A. D. Shapere and Y. Tachikawa, *J. High Energy Phys.* **09** (2008) 109.
- [43] S. M. Hosseini, K. Hristov, and A. Zaffaroni, *J. High Energy Phys.* **11** (2019) 090.
- [44] S. M. Hosseini, K. Hristov, Y. Tachikawa, and A. Zaffaroni, *J. High Energy Phys.* **09** (2020) 167.
- [45] F. Benini and N. Bobev, *J. High Energy Phys.* **06** (2013) 005.
- [46] I. Bah, F. Bonetti, R. Minasian, and E. Nardoni, *J. High Energy Phys.* **01** (2020) 125.
- [47] M. Cvetič, H. Lu, and C. Pope, *Nucl. Phys.* **B574**, 761 (2000).
- [48] V. Kostelecky and M. J. Perry, *Phys. Lett. B* **371**, 191 (1996).
- [49] M. Cvetič, G. Gibbons, H. Lu, and C. Pope, [arXiv:hep-th/0504080](https://arxiv.org/abs/hep-th/0504080).
- [50] K. Hristov, S. Katmadas, and C. Toldo, *Phys. Rev. D* **100**, 066016 (2019).
- [51] S. Choi, C. Hwang, S. Kim, and J. Nahmgoong, *J. Korean Phys. Soc.* **76**, 101 (2020).
- [52] S. M. Hosseini, K. Hristov, and A. Zaffaroni, *J. High Energy Phys.* **12** (2019) 168.
- [53] D. L. Jafferis, *J. High Energy Phys.* **05** (2012) 159.
- [54] S. Pasquetti, *J. High Energy Phys.* **04** (2012) 120.
- [55] C. Beem, T. Dimofte, and S. Pasquetti, *J. High Energy Phys.* **12** (2014) 177.
- [56] S. Choi, C. Hwang, and S. Kim, [arXiv:1908.02470](https://arxiv.org/abs/1908.02470).
- [57] S. Choi and C. Hwang, *J. High Energy Phys.* **03** (2020) 068.
- [58] M. Porrati and A. Zaffaroni, *Nucl. Phys.* **B490**, 107 (1997).
- [59] N. Mekareeya, *J. High Energy Phys.* **12** (2015) 174.
- [60] C. P. Herzog, I. R. Klebanov, S. S. Pufu, and T. Tesileanu, *Phys. Rev. D* **83**, 046001 (2011).
- [61] D. L. Jafferis, I. R. Klebanov, S. S. Pufu, and B. R. Safdi, *J. High Energy Phys.* **06** (2011) 102.
- [62] P. M. Crichigno, C. P. Herzog, and D. Jain, *J. High Energy Phys.* **03** (2013) 039.
- [63] S. M. Hosseini and N. Mekareeya, *J. High Energy Phys.* **08** (2016) 089.
- [64] D. Martelli, J. Sparks, and S.-T. Yau, *Commun. Math. Phys.* **280**, 611 (2008).
- [65] Y. Imamura and K. Kimura, *Prog. Theor. Phys.* **120**, 509 (2008).
- [66] A. Amariti, C. Klare, and M. Siani, *J. High Energy Phys.* **10** (2012) 019.
- [67] D. Gaiotto and E. Witten, *J. High Energy Phys.* **06** (2010) 097.
- [68] Y. Imamura and K. Kimura, *J. High Energy Phys.* **10** (2008) 040.
- [69] K. Hosomichi, K.-M. Lee, S. Lee, S. Lee, and J. Park, *J. High Energy Phys.* **07** (2008) 091.
- [70] D. Gaiotto and E. Witten, *J. Stat. Phys.* **135**, 789 (2009).
- [71] B. Assel, C. Bachas, J. Estes, and J. Gomis, *J. High Energy Phys.* **08** (2011) 087.
- [72] B. Assel, C. Bachas, J. Estes, and J. Gomis, *J. High Energy Phys.* **12** (2012) 044.
- [73] L. Coccia, [arXiv:2006.06578](https://arxiv.org/abs/2006.06578).
- [74] L. Coccia and C. F. Uhlemann, [arXiv:2011.10050](https://arxiv.org/abs/2011.10050).
- [75] S. L. Cacciatori and D. Klemm, *J. High Energy Phys.* **01** (2010) 085.
- [76] S. Katmadas, *J. High Energy Phys.* **09** (2014) 027.
- [77] N. Halmagyi, *J. High Energy Phys.* **03** (2015) 032.
- [78] K. Hristov, S. Katmadas, and C. Toldo, *J. High Energy Phys.* **01** (2019) 199.
- [79] N. Bobev, A. M. Charles, and V. S. Min, *J. High Energy Phys.* **10** (2020) 073.