## **Transport in Strongly Coupled Quark Matter**

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Motivated by the possible presence of deconfined quark matter in neutron stars and their mergers and the important role of transport phenomena in these systems, we perform the first-ever systematic study of different viscosities and conductivities of dense quark matter using the gauge/gravity duality. Using the V-QCD model, we arrive at results that are in qualitative disagreement with the predictions of perturbation theory, which highlights the differing transport properties of the system at weak and strong coupling and calls for caution in the use of the perturbative results in neutron star applications.

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Introduction.—The first recorded observations of binary neutron star (NS) mergers, including both gravitational wave [1,2] and electromagnetic [3] signals, have opened intriguing new avenues for the study of strongly interacting matter at ultrahigh densities. While most of the attention in the field has so far been directed to the macroscopic properties of the stars and correspondingly to the equation of state (EoS) of NS matter (see, e.g., [4-6] and references therein), there exists ample motivation to inspect also transport properties of dense quantum chromodynamics determined by viscosities and conductivities. This is due in particular to the fact that, while near the transition from nuclear to quark matter (QM) the EoSs of the two phases may largely resemble one another, the corresponding transport properties are expected to witness much more dramatic changes, potentially enabling a direct detection of OM either in quiescent NSs or their binary mergers [7,8]. In addition, understanding the relative magnitudes of different transport coefficients may turn out useful for the relativistic hydrodynamic simulations of NS mergers [9,10].

In the QM phase, which is expected to be found in the cores of massive NSs [11] and is very likely created in NS mergers [8,12–14], very few robust results exist for the QCD contribution to even the most central transport coefficients, including the bulk and shear viscosities and the electrical and heat conductivities. In fact, the only first-principles determination of these quantities in unpaired quark

Published by the American Physical Society under the terms of the Creative Commons Attribution 4.0 International license. Further distribution of this work must maintain attribution to the author(s) and the published article's title, journal citation, and DOI. Funded by SCOAP<sup>3</sup>. matter dates back to the early 1990s, amounting to a leadingorder perturbative QCD (pQCD) calculation [15]. Owing to the strongly coupled nature of QM in the density regime relevant for NSs [16,17], these results are, however, of limited predictive value. A nonpertubative analysis at strong coupling would be required for robust predictions, but the well-known problems that lattice Monte Carlo simulations face at nonzero baryon densities presently prohibit the use of this standard tool (see, e.g., [18]).

Another complication in the determination of transport quantities in QM has to do with the fact that the physical phase of QCD realized at moderate densities is at present unknown [19]. While the general expectation is that some type of quark pairing is likely present all the way down to the deconfinement transition at low temperatures [20], it is currently unclear which particular phases are realized in nature. Assuming that the physical moderate-density phase contains at least some nonzero fraction of unpaired quarks, it is, however, often considered a reasonable approximation to first inspect transport coefficients in the somewhat simpler case of unpaired QM [7].

In this Letter, we approach the most important transport properties of dense unpaired QM with the only first-principles machinery currently capable of describing strongly coupled quantum field theories at high baryon density: the gauge/gravity duality or, in short, holography. This correspondence provides a link between classical gravity and strongly coupled quantum field theories, relating their observables in a detailed manner (see, e.g., [17,21,22] for reviews). In the context of heavy-ion physics, many important insights have been gained through the study of questions that are difficult to address with traditional field theory machinery, such as the details of equilibration dynamics

(see, e.g., the recent review [23]). In addition, the conjectured lower limit of the shear viscosity to entropy ratio,  $\eta/s \ge 1/(4\pi)$ , has hinted toward universality in strongly coupled systems, which has had a profound effect on many subfields of theoretical physics [24].

In the context of NS physics, promising progress has been achieved in applying holographic methods to determining the EoS of QCD matter [25–29], but no attempts have been made to analyze transport in strongly coupled dense QM. Here, we shall take the first steps in this direction by using one of the most highly developed "bottom-up" frameworks (i.e., with no known string theory origin) designed to mimic a gravity dual of QCD at finite temperature and density [30]: the Veneziano limit of the improved holographic QCD (IHQCD) model, V-QCD [31–33]. Interestingly, the results obtained for the most important transport coefficients of dense QM are in qualitative disagreement with known perturbative predictions, calling for caution in the application of the latter results to phenomenological studies of NSs.

For comparison and completeness, we shall contrast our V-QCD results not only to pQCD but also to one of the most widely studied top-down holographic models describing deconfined quark matter, i.e., the D3-D7 system in its unbackreacted limit [34,35], implying we neglect the corrections from flavors. This corresponds to the quenched approximation in field theory language [36]. The latter provides reliable results for quantities that can be derived from the free energy [37], which in the present case translates to the shear viscosity, obtainable via the relation  $\eta/s = 1/(4\pi)$ . Other physical quantities, on the other hand, are expected to receive corrections from backreacted (unquenched) quark matter, which we shall indeed witness in our results.

Setup.—As explained above, we approach the description of dense strongly coupled matter via the string-theory-inspired V-QCD model [31–33]. The way we have set up our analysis is more general, however, and can accommodate other holographic models such as the probe-brane limit of the D3-D7 system [35,38]. Both setups have been thoroughly applied to the study of the bulk properties of NS matter, and we refer the interested reader to Refs. [13,14,25,28,33,39].

A significant difference between two holographic models lies in how the effects of the flavor sector are treated. V-QCD has these effects systematically built in, whereas the D3-D7 system treats the flavors as a probe (note, however, that, for massless quarks, Refs. [40–44] have gone beyond the probe approximation). Despite this difference, we can define the two models via the same gravitational action consisting of two terms  $S_{\text{total}} = S_g + S_f$ , where

$$S_g = N_c^2 M_{\rm Pl}^3 \int d^5 x \sqrt{-g} \left[ R - \frac{1}{2} \partial_\rho \phi \partial^\rho \phi - V(\phi) \right] \qquad (1$$

$$S_f = -N_f N_c M_{\rm Pl}^3 \int d^5 x \mathcal{Z}(\phi, \chi) \sqrt{-\det(\Gamma_{\mu\nu})}.$$
 (2)

Here, g is the determinant of the metric  $g_{\mu\nu}$ ,  $M_{\rm Pl}$  denotes the rescaled five-dimensional Planck mass, R is the Ricci scalar, and we have defined

$$\Gamma_{\mu\nu} = g_{\mu\nu} + \kappa(\phi, \chi) \partial_{\mu} \chi \partial_{\nu} \chi + \mathcal{W}(\phi, \chi) F_{\mu\nu}. \tag{3}$$

The potentials and couplings  $V(\phi)$ ,  $\mathcal{Z}(\phi,\chi)$ ,  $\kappa(\phi,\chi)$ , and  $\mathcal{W}(\phi,\chi)$  will be different functions in D3-D7 and in V-OCD, as we shall discuss below.

Of the two parts of the action,  $S_g$  is related to the glue sector of a gauge theory with rank  $N_c$ . The scalar field  $\phi$  is identified with the dilaton, which according to the holographic dictionary is dual to the Yang–Mills running coupling constant. For V-QCD, following the IHQCD model [45,46], the potential V in Eq. (1) is chosen to reproduce the known physics of Yang–Mills theory upon comparison with perturbative results and lattice data [33,47,48]. In the D3-D7 model,  $S_g$  is determined, on the other hand, by the closed string sector of type IIB supergravity and  $V=-12/L^2$ . To fix units, we demand that the asymptotically AdS space have the unit radius L=1.

At the same time, the physics of the  $N_f$  flavors of fundamental quarks is captured by the Dirac–Born–Infeld action  $S_f$ , where the tachyon field  $\chi$  is dual to the chiral condensate  $\bar{q}q$ , whose boundary value is related to the masses of the quarks [49]. We set the quark masses to zero in V-QCD but keep them nonvanishing in the D3-D7 model to achieve the breaking of conformal symmetry. In the D3-D7 model, a quark mass corresponds to the energy necessary to introduce an additional quark over the ground state, implying that it corresponds to a constituent quark mass with a value of the order of  $1/N_c$  times the baryon mass [25]. The field strength  $F_{\mu\nu} = \partial_{\mu}A_{\nu} - \partial_{\nu}A_{\mu}$ , on the other hand, provides the dynamics for the  $U(1)_B$  gauge field  $A_{\mu}$  corresponding to the conserved baryonic charge of the dual field theory.

In the V-QCD model, the functions  $\mathcal{Z}$ ,  $\kappa$ , and  $\mathcal{W}$  are fixed to reproduce the desired features of QCD (such as confinement and asymptotic freedom) for weak [31,45] and strong [29,31,46,50–52] Yang–Mills couplings. Moreover, the potentials were tuned to match with lattice data for the EoS at small chemical potentials [33,53]. In this way, the model in effect extrapolates the lattice results to the regime relevant for NS cores. We also determine  $M_{\rm Pl}$  by using lattice data and set  $N_f/N_c=1$ , and furthermore employ the couplings of the fits 5b, 7a, and 8b given in Appendix A of [33].

For the D3-D7 model, supergravity implies the simple relations

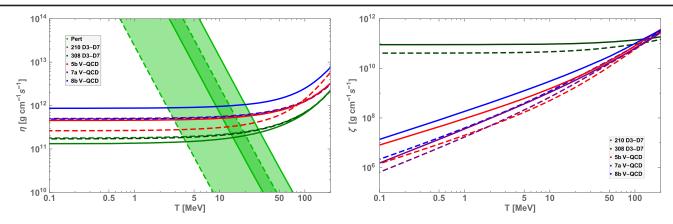


FIG. 1. Shear (left) and bulk (right) viscosities as functions of temperature for  $\mu = 450$  MeV (dashed lines) and  $\mu = 600$  MeV (solid lines). The filled bands on the left correspond to the pQCD results (upper band obtained for  $\mu = 600$  MeV, lower for 450 MeV), and they have been generated by varying the parameter x inside  $\alpha_s$  (see the main text for the definition) inside the interval  $1/2 \le x \le 2$ .

$$M_{\rm Pl}^3 = \frac{1}{8\pi^2}, \qquad \mathcal{Z} = \frac{\lambda_{\rm YM}}{2\pi^2} \cos^3 \chi, \qquad \mathcal{W} = \frac{2\pi}{\sqrt{\lambda_{\rm YM}}},$$
 $\kappa = 1,$ 
(4)

where  $\lambda_{\rm YM}=g_{\rm YM}^2 N_c$  is the 't Hooft coupling (with  $g_{\rm YM}$  the Yang–Mills coupling) of the dual field theory. Following [25], we fix  $\lambda_{\rm YM}\simeq 10.74$  so that the pressure matches the Stefan–Boltzmann value at asymptotically large chemical potentials.

The holographic determination of the thermodynamic quantities and transport coefficients is briefly reviewed in the Supplemental Material [63]. To translate the resulting expressions into numerical results, we need to choose, among other things, the characteristic perturbative energy scale  $\Lambda_{\rm UV}$  for V-QCD. We use the values  $\Lambda_{\rm UV}=226.24$ , 210.76, and 156.68 MeV for the fits 5b, 7a, and 8b in the V-QCD model, following the choices made in [13,33]. The quark mass in the D3-D7 model, on the other hand, is given two values:  $M_q=210.76$  MeV for direct comparison with the V-QCD potential 7a [64], and  $M_q=308.55$  MeV following the logic of Ref. [25]. These values are used to (partially) probe the systematic uncertainties of this model.

An important point to note is that in both of our models, we consistently work with three mass-degenerate quark flavors assuming beta equilibrium, whereby all quark flavors share the same chemical potential  $\mu = \mu_B/3$ . This implies the absence of electrons in the system, which would only change upon taking flavor-dependent masses into account.

Viscosities.—The shear and bulk viscosities of dense QCD matter describe the resistance of the system to deformations. They become relevant in settings where NSs are either strongly deformed or their interiors taken out of thermal equilibrium, both of which occur in different stages of binary NS mergers [10,65]. In addition, viscosities play a role in determining the damping of unstable r modes in rapidly rotating stars [66–71].

Viscosities appear in contributions to stress forces due to an inhomogeneous motion of the fluid. Letting  $v_i$ , i = 1, 2, 3 be the components of the velocity of a fluid moving at low velocities, the resulting stress becomes

$$T_{ij} = -\eta(\partial_i v_j + \partial_j v_i) - \left(\zeta - \frac{2}{3}\eta\right) \delta_{ij} \partial_k v^k, \quad (5)$$

where  $\eta$  and  $\zeta$  are the shear and bulk viscosities, respectively.

The values obtained from the holographic models are plotted in Fig. 1 (left) together with the pQCD result for unpaired quark matter [7,15],

$$\eta \approx 4.4 \times 10^{-3} \frac{\mu^2 m_D^{2/3}}{\alpha_s^2 T^{5/3}}.$$
(6)

In evaluating this expression, we have used the one-loop Debye mass  $m_D^2 = 2\alpha_s[N_f\mu^2 + (2N_c + N_f)\pi^2T^2/3]/\pi$  (see, e.g., [72]) and the two-loop strong coupling  $\alpha_s$  related to the QCD gauge coupling via  $\alpha_s = g_{\rm QCD}^2/(4\pi)$ . Following typical conventions in the literature (see, e.g., [73,74] and the discussion in the beginning of Sec. VII in [75]), the renormalization scale is set to  $\bar{\Lambda} = x\sqrt{(2\pi T)^2 + (2\mu)^2}$ , where x is a parameter that parameterizes the renormalization scale dependence of the pQCD result and is varied between the values 1/2 and 2. Finally, the QCD scale  $\Lambda_{\rm QCD}$  is given the value 378 MeV, which was obtained by demanding that  $\alpha_s(\bar{\Lambda} = 2 \text{ GeV}) = 0.2994$  [76].

We observe that at high temperatures the shear viscosity of the strongly coupled fluid is qualitatively larger than the perturbative result, while at low temperatures it approaches a constant with the crossing of the holographic and pQCD results taking place around  $T \sim 5-50$  MeV. We also note that, in agreement with our naive expectation, both holographic models give comparable results, as the D3-D7 calculation is not hampered by problems related to backreaction in this case.

In addition, we have determined the QCD contribution to the bulk viscosity  $\zeta$ , which, however, represents only a subdominant contribution to r-mode damping [77]. The corresponding results are shown in Fig. 1 (right), where we observe that, in V-QCD,  $\zeta$  is highly suppressed in comparison to  $\eta$  at low temperatures but reaches a value of around 10% of the shear viscosity at high temperatures. The D3-D7 model, on the other hand, leads to a very flat curve in the range of temperatures studied, but the validity of this model is clearly questionable due to the flavor quenching. In leading-order pOCD, the bulk viscosity finally vanishes when the quark masses are negligible compared to the chemical potentials and remains small even at high temperatures [78] (note, however, that nonperturbative effects are expected to increase the value somewhat [79,80]), which explains the lack of a pOCD curve in this figure.

Conductivities.—A newly formed NS undergoes a cooling process, mainly through the emission of neutrinos from its interior. The neutrinos transport heat to the surface, where the energy is emitted as radiation. A closer inspection of this process shows that the thermal evolution of the star depends on several quantities, including the heat conductivity that determines the heat flux to the surface and the electrical conductivity that determines the magnitude of Joule heating through the decay of magnetic fields [81–83]. In the postmerger phase of an NS binary merger, the electrical and thermal conductivities are furthermore relevant for equilibration and the evolution of magnetic fields [10,84]. Finally, these quantities may in principle prove useful in distinguishing between different phases of QCD through the observation of thermal radiation.

In the strongly coupled theories that we study in the present work, matter resides in a state that can be described as a relativistic fluid. This implies that the electrical and thermal conductivities are not independent but are determined by a single coefficient  $\sigma$ , which is defined by the constitutive relation of the current arising from a gradient of the chemical potential,

$$J_{x} = -\sigma \partial_{x} \left( \frac{\mu}{T} \right). \tag{7}$$

The electrical conductivity, defined as the ratio between the current and the electric field  $E_x$ , can be shown to take the form [85]

$$\sigma^{xx} = \frac{J^x}{E_x} = \frac{\varepsilon + p}{Ts} \sigma,\tag{8}$$

where  $\varepsilon$ , p, and s are the energy density, pressure, and entropy density, respectively. Moreover, the thermal conductivity, defined as the ratio between the heat current  $Q^x$  and the temperature gradient, becomes

$$\kappa^{xx} = \frac{Q^x}{-\partial_x T} = \frac{\mu}{T} \frac{\varepsilon + p}{\rho} \sigma = \frac{\mu s}{\rho} \sigma^{xx}, \tag{9}$$

where  $\rho$  is the charge density. It should be stressed that these expressions hold only for a steady state where the gradients of the temperature and chemical potential balance each other such that the transport does not occur via convection.

With the above results established, we see that it suffices to compute the electrical conductivity in the holographic models. Our results for the two conductivities are displayed in Fig. 2 together with the perturbative results for unpaired quark matter [7,15]

$$\sigma^{xx} \approx 0.01 \frac{\mu^2 m_D^{2/3}}{\alpha_s T^{5/3}}, \qquad \kappa^{xx} \approx 0.5 \frac{m_D^2}{\alpha_s^2}. \tag{10}$$

Here, we continue to use the same values for  $m_D$  and  $\alpha_s$  as listed in the previous section, and the electrical conductivity is given in units of  $e^2/(\hbar c)$ .

We observe that, both in the perturbative and D3-D7 calculations, the conductivities either decrease with temperature or are largely independent of it, while in the V-QCD

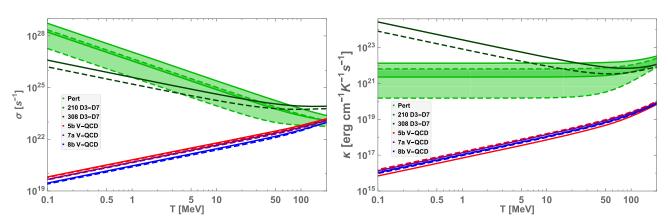


FIG. 2. Electrical (left) and thermal (right) conductivities as functions of temperature for  $\mu = 450$  MeV (dashed lines) and  $\mu = 600$  MeV (solid). The filled bands again correspond to the pQCD results for  $1/2 \le x \le 2$ , with the  $\mu = 600$  MeV bands slightly above the 450 MeV ones.

model they are increasing functions of T. This qualitative discrepancy can be easily understood as a suppression of momentum dissipation of charged particles coming from the assumptions of these models: in the pQCD case, the coupling is assumed small by construction, whereas in the D3-D7 case, only the leading-order effect in an  $N_f/N_c$  expansion is retained [86]. There are strong reasons to expect that properly including backreaction in this holographic model would bring the D3-D7 result into at least qualitative agreement with the V-QCD one (see, e.g., the discussion in [87]).

Discussion.—A first-principles microscopic determination of the fundamental properties of dense QCD matter is a notoriously difficult problem, not least due to the strongly coupled nature of the system at phenomenologically relevant energies. In recent years, the gauge/gravity duality has shown considerable promise as a potential tool, as it allows approaching the problem from an angle complementary to traditional field theory methods; indeed, promising results have been obtained for many bulk thermodynamic quantities, leading to predictions for observables such as the NS mass-radius relation and the phase diagram of the theory [13,25,33].

In this Letter, we have used the holographic machinery to tackle a more challenging class of physical quantities characterizing the response of the medium to external perturbations. In particular, we have studied the behavior of the transport coefficients most relevant for the physics of NSs and their mergers, i.e., the shear and bulk viscosities and the thermal and electrical conductivities. All these quantities have been evaluated in a highly developed bottom-up framework mimicking a gravity dual for QCD, V-QCD, and the corresponding results subsequently compared to those from the D3-D7 probe-brane setup and perturbative QCD [15].

Our main results are depicted in Figs. 1 and 2. Inspecting these plots, an issue that stands out immediately is the stark contrast between the V-QCD and pQCD predictions for all quantities, for which both results are available: in V-QCD, transport coefficients typically increase with temperature, while a qualitatively different behavior is predicted by pQCD. As a result, in the T=0 limit, V-QCD coefficients are comparatively strongly suppressed. In addition, while we observe qualitative agreement between the V-QCD and D3-D7 predictions for the shear viscosity, the same is not true for the other three quantities studied. These observations clearly call for physical interpretation.

As flavors are quenched in the D3-D7 model, the main dissipative effect affecting quark matter is due to drag by the thermal plasma [88,89]. The increasing trend of the electrical conductivity at lower temperatures reflects the decreasing drag force, which, however, is expected to be cut off by radiation effects not captured by the D3-D7 analysis [86,90,91]. Furthermore, at very low temperatures, the quenched approximation is expected to break down

altogether [41,42]. The V-QCD model, on the other hand, does not suffer from these issues, as flavors are unquenched, and consequently the trends exhibited by both conductivities can be expected to reflect the true behavior of these quantities in strongly coupled unpaired quark matter.

For the viscosities, the situation is somewhat different. In both holographic setups, the shear viscosity is proportional to a quantity derived from the free energy, which leads to a fair agreement between the two predictions. On the other hand, in the pQCD calculation, flavor contributions give rise to a strong increase of the quantity in the  $T \to 0$  limit, leading to a stark disagreement with the holographic results. In contrast, for the bulk viscosity, for which no pQCD result is available, we witness a marked sensitivity of the result to the pattern in which conformal invariance is broken (by a running coupling vs quark masses) in our holographic models, which essentially invalidates the prediction of the D3-D7 model for this quantity.

In summary, we have seen dramatic differences arise between the predictions of a strongly coupled unquenched theory and its perturbative or quenched approximations for the transport properties of dense quark matter. This observation clearly calls for significant caution in the application of the perturbative results in any phenomenological study within NS physics and highlights the necessity of further developing the holographic approach to the problem.

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- [1] B. Abbott *et al.* (LIGO Scientific and Virgo Collaborations), Phys. Rev. Lett. **119**, 161101 (2017).
- [2] B. Abbott *et al.* (LIGO Scientific and Virgo Collaborations), Phys. Rev. Lett. **121**, 161101 (2018).
- [3] B. Abbott et al. (LIGO Scientific, Virgo, Fermi GBM, INTEGRAL, IceCube, AstroSat Cadmium Zinc Telluride Imager Team, IPN, Insight-Hxmt, ANTARES, Swift, AGILE Team, 1M2H Team, Dark Energy Camera GW-EM, DES, DLT40, GRAWITA, Fermi-LAT, ATCA, AS-KAP, Las Cumbres Observatory Group, OzGrav, DWF (Deeper Wider Faster Program), AST3, CAASTRO, VINROUGE, MASTER, J-GEM, GROWTH, JAGWAR, CaltechNRAO, TTU-NRAO, NuSTAR, Pan-STARRS, MAXI Team, TZAC Consortium, KU, Nordic Optical Telescope, ePESSTO, GROND, Texas Tech University, SALT Group, TOROS, BOOTES, MWA, CALET, IKI-GW Follow-up, H.E.S.S., LOFAR, LWA, HAWC, Pierre Auger, ALMA, Euro VLBI Team, Pi of Sky, Chandra Team at McGill University, DFN, ATLAS Telescopes, High Time Resolution Universe Survey, RIMAS, RATIR, SKA South Africa/MeerKAT Collaborations), Astrophys. J. 848, L12 (2017).
- [4] E. Annala, T. Gorda, A. Kurkela, and A. Vuorinen, Phys. Rev. Lett. 120, 172703 (2018).
- [5] L. Rezzolla, E. R. Most, and L. R. Weih, Astrophys. J. Lett. 852, L25 (2018).
- [6] C. D. Capano, I. Tews, S. M. Brown, B. Margalit, S. De, S. Kumar, D. A. Brown, B. Krishnan, and S. Reddy, Nat. Astron. 4, 625 (2020).
- [7] A. Schmitt and P. Shternin, Astrophysics and Space Science Library **457**, 455 (2018).
- [8] E. R. Most, L. J. Papenfort, V. Dexheimer, M. Hanauske, S. Schramm, H. Stocker, and L. Rezzolla, Phys. Rev. Lett. 122, 061101 (2019).
- [9] L. Baiotti and L. Rezzolla, Rep. Prog. Phys. 80, 096901 (2017).
- [10] M. G. Alford, L. Bovard, M. Hanauske, L. Rezzolla, and K. Schwenzer, Phys. Rev. Lett. 120, 041101 (2018).
- [11] E. Annala, T. Gorda, A. Kurkela, J. Nättilä, and A. Vuorinen, Nat. Phys., (2020).
- [12] A. Bauswein, N.-U. F. Bastian, D. B. Blaschke, K. Chatziioannou, J. A. Clark, T. Fischer, and M. Oertel, Phys. Rev. Lett. 122, 061102 (2019).
- [13] P. M. Chesler, N. Jokela, A. Loeb, and A. Vuorinen, Phys. Rev. D 100, 066027 (2019).
- [14] C. Ecker, M. Jarvinen, G. Nijs, and W. van der Schee, Phys. Rev. D 101, 103006 (2020).
- [15] H. Heiselberg and C. J. Pethick, Phys. Rev. D 48, 2916 (1993).
- [16] U. Kraemmer and A. Rebhan, Rep. Prog. Phys. **67**, 351 (2004).

- [17] J. Casalderrey-Solana, H. Liu, D. Mateos, K. Rajagopal, and U. A. Wiedemann, https://doi.org/10.1017/CBO978 1139136747, arXiv:1101.0618.
- [18] P. de Forcrand, *Proc. Sci.*, LAT2009 (2009) 010 [arXiv:1005.0539].
- [19] M. G. Alford, A. Schmitt, K. Rajagopal, and T. Schäfer, Rev. Mod. Phys. 80, 1455 (2008).
- [20] K. Rajagopal and F. Wilczek, in *At the Frontier of Particle Physics*. *Handbook of QCD*. *Vol. 1–3*, edited by M. Shifman and B. Ioffe (2000), pp. 2061–2151.
- [21] A. V. Ramallo, Springer Proc. Phys. **161**, 411 (2015).
- [22] N. Brambilla et al., Eur. Phys. J. C 74, 2981 (2014).
- [23] J. Berges, M. P. Heller, A. Mazeliauskas, and R. Venugopalan, arXiv:2005.12299.
- [24] P. Kovtun, D. T. Son, and A. O. Starinets, J. High Energy Phys. 10 (2003) 064.
- [25] C. Hoyos, N. Jokela, D. R. Fernández, and A. Vuorinen, Phys. Rev. Lett. 117, 032501 (2016).
- [26] C. Hoyos, N. Jokela, D. R. Fernández, and A. Vuorinen, Phys. Rev. D 94, 106008 (2016).
- [27] C. Ecker, C. Hoyos, N. Jokela, D. R. Fernndez, and A. Vuorinen, J. High Energy Phys. 11 (2017) 031.
- [28] K. B. Fadafan, J. C. Rojas, and N. Evans, Phys. Rev. D 101, 126005 (2020).
- [29] T. Ishii, M. Jrvinen, and G. Nijs, J. High Energy Phys. 07 (2019) 003.
- [30] Note that at low densities, transport has been studied in the IHQCD and V-QCD frameworks in Refs. [92–94].
- [31] M. Järvinen and E. Kiritsis, J. High Energy Phys. 03 (2012) 002.
- [32] T. Alho, M. Järvinen, K. Kajantie, E. Kiritsis, C. Rosen, and K. Tuominen, J. High Energy Phys. 04 (2014) 124; 02 (2015) 33.
- [33] N. Jokela, M. Järvinen, and J. Remes, J. High Energy Phys. 03 (2019) 041.
- [34] M. Kruczenski, D. Mateos, R. C. Myers, and D. J. Winters, J. High Energy Phys. 07 (2003) 049.
- [35] S. Kobayashi, D. Mateos, S. Matsuura, R. C. Myers, and R. M. Thomson, J. High Energy Phys. 02 (2007) 016.
- [36] C. Nunez, A. Paredes, and A. V. Ramallo, Adv. High Energy Phys. **2010**, 196714 (2010).
- [37] A. Karch, A. O'Bannon, and E. Thompson, J. High Energy Phys. 04 (2009) 021.
- [38] G. Itsios, N. Jokela, and A. V. Ramallo, Nucl. Phys. **B909**, 677 (2016).
- [39] E. Annala, C. Ecker, C. Hoyos, N. Jokela, D. R. Fernández, and A. Vuorinen, J. High Energy Phys. 12 (2018) 078.
- [40] F. Bigazzi, A. L. Cotrone, J. Mas, A. Paredes, A. V. Ramallo, and J. Tarrío, J. High Energy Phys. 11 (2009) 117
- [41] F. Bigazzi, A. L. Cotrone, J. Mas, D. Mayerson, and J. Tarrío, J. High Energy Phys. 04 (2011) 060.
- [42] F. Bigazzi, A. L. Cotrone, and J. Tarrío, J. High Energy Phys. 07 (2013) 074.
- [43] A. F. Faedo, D. Mateos, C. Pantelidou, and J. Tarrio, J. High Energy Phys. 02 (2017) 047.
- [44] A. F. Faedo, D. Mateos, C. Pantelidou, and J. Tarrio, J. High Energy Phys. 10 (2017) 139; 07 (2019) 058.
- [45] U. Gürsoy and E. Kiritsis, J. High Energy Phys. 02 (2008) 032.

- [46] U. Gürsoy, E. Kiritsis, and F. Nitti, J. High Energy Phys. 02 (2008) 019.
- [47] U. Gürsoy, E. Kiritsis, L. Mazzanti, and F. Nitti, Nucl. Phys. **B820**, 148 (2009).
- [48] M. Panero, Phys. Rev. Lett. 103, 232001 (2009).
- [49] For V-QCD, the flavor setup is based on Refs. [95–98].
- [50] D. Areán, I. Iatrakis, M. Järvinen, and E. Kiritsis, J. High Energy Phys. 11 (2013) 068.
- [51] D. Areán, I. Iatrakis, M. Järvinen, and E. Kiritsis, Phys. Rev. D 96, 026001 (2017).
- [52] M. Järvinen, J. High Energy Phys. 07 (2015) 033.
- [53] See Supplemental Material § B at http://link.aps.org/supplemental/10.1103/PhysRevLett.125.241601 for details of the V-QCD model, which includes Refs. [54–62].
- [54] A. Donos and J. P. Gauntlett, J. High Energy Phys. 11 (2014) 081.
- [55] B. Goutéraux, N. Jokela, and A. Pönni, J. High Energy Phys. 07 (2018) 004.
- [56] T. Alho, M. Järvinen, K. Kajantie, E. Kiritsis, and K. Tuominen, Phys. Rev. D 91, 055017 (2015).
- [57] T. Alho, J. Remes, K. Tuominen, and A. Vuorinen, Phys. Rev. D 101, 106025 (2020).
- [58] S. S. Gubser, Adv. Theor. Math. Phys. 4, 679 (2000).
- [59] S. Borsányi, Z. Fodor, C. Hoelbling, S. D. Katz, S. Krieg, and K. K. Szabó, Phys. Lett. B 730, 99 (2014).
- [60] S. Borsanyi, Z. Fodor, S. D. Katz, S. Krieg, C. Ratti, and K. Szabo, J. High Energy Phys. 01 (2012) 138.
- [61] N. Jokela, M. Järvinen, G. Nijs, and J. Remes, arXiv: 2006.01141.
- [62] S. J. Brodsky, G. F. de Téramond, and A. Deur, Phys. Rev. D 81, 096010 (2010).
- [63] § A, including Refs. [99,100].
- [64] Recall that these parameters,  $M_q$  for D3-D7 and  $\Lambda_{\rm UV}$  for V-QCD, both control how the conformal symmetry is broken. Their values must therefore be comparable, even though not necessarily exactly the same.
- [65] S. Fujibayashi, K. Kiuchi, N. Nishimura, Y. Sekiguchi, and M. Shibata, Astrophys. J. 860, 64 (2018).
- [66] N. Andersson, Astrophys. J. 502, 708 (1998).
- [67] J. L. Friedman and S. M. Morsink, Astrophys. J. 502, 714 (1998).
- [68] N. Andersson and K. D. Kokkotas, Int. J. Mod. Phys. D 10, 381 (2001).
- [69] P. Arras, E. E. Flanagan, S. M. Morsink, A. Schenk, S. A. Teukolsky, and I. Wasserman, Astrophys. J. 591, 1129 (2003).
- [70] M. G. Alford, S. Mahmoodifar, and K. Schwenzer, J. Phys. G 37, 125202 (2010).
- [71] B. Haskell, Int. J. Mod. Phys. E 24, 1541007 (2015).
- [72] A. Vuorinen, Phys. Rev. D 68, 054017 (2003).
- [73] A. Kurkela, P. Romatschke, and A. Vuorinen, Phys. Rev. D 81, 105021 (2010).

- [74] A. Kurkela and A. Vuorinen, Phys. Rev. Lett. 117, 042501 (2016).
- [75] J. Ghiglieri, A. Kurkela, M. Strickland, and A. Vuorinen, Phys. Rep. 880, 1 (2020).
- [76] C. Amsler *et al.* (Particle Data Group), Phys. Lett. B 667, 1 (2008).
- [77] The dominant one originates from chemical equilibration through weak processes,  $\zeta \sim 10^{24} 10^{29} \text{ g cm}^{-1} \text{ s}^{-1}$  [7,70]. A quantitative analysis of this effect is unfortunately outside the scope of our work.
- [78] P. B. Arnold, C. Dogan, and G. D. Moore, Phys. Rev. D 74, 085021 (2006).
- [79] D. Kharzeev and K. Tuchin, J. High Energy Phys. 09 (2008) 093.
- [80] M. M. Wang, Y. Jiang, B. Wang, W. M. Sun, and H.-S. Zong, Mod. Phys. Lett. A 26, 1797 (2011).
- [81] D. G. Yakovlev and C. J. Pethick, Annu. Rev. Astron. Astrophys. 42, 169 (2004).
- [82] D. Page, U. Geppert, and F. Weber, Nucl. Phys. **A777**, 497 (2006).
- [83] D. Viganò, N. Rea, J. A. Pons, R. Perna, D. N. Aguilera, and J. A. Miralles, Mon. Not. R. Astron. Soc. 434, 123 (2013).
- [84] A. Harutyunyan, A. Nathanail, L. Rezzolla, and A. Sedrakian, Eur. Phys. J. A **54**, 191 (2018).
- [85] See Supplemental Material [53] § C.
- [86] A. Karch and A. O'Bannon, J. High Energy Phys. 09 (2007) 024.
- [87] J. Tarrio, J. High Energy Phys. 04 (2014) 042.
- [88] S. S. Gubser, Phys. Rev. D 74, 126005 (2006).
- [89] C. P. Herzog, A. Karch, P. Kovtun, C. Kozcaz, and L. G. Yaffe, J. High Energy Phys. 07 (2006) 013.
- [90] A. Mikhailov, arXiv:hep-th/0305196.
- [91] M. Chernicoff and A. Guijosa, J. High Energy Phys. 06 (2008) 005.
- [92] U. Gürsoy, E. Kiritsis, G. Michalogiorgakis, and F. Nitti, J. High Energy Phys. 12 (2009) 056.
- [93] U. Gürsoy, E. Kiritsis, L. Mazzanti, and F. Nitti, J. High Energy Phys. 12 (2010) 088.
- [94] I. Iatrakis and I. Zahed, J. High Energy Phys. 04 (2015) 080.
- [95] F. Bigazzi, R. Casero, A. Cotrone, E. Kiritsis, and A. Paredes, J. High Energy Phys. 10 (2005) 012.
- [96] R. Casero, E. Kiritsis, and A. Paredes, Nucl. Phys. B787, 98 (2007).
- [97] I. Iatrakis, E. Kiritsis, and A. Paredes, Phys. Rev. D **81**, 115004 (2010).
- [98] I. Iatrakis, E. Kiritsis, and A. Paredes, J. High Energy Phys. 11 (2010) 123.
- [99] D. Mateos, R. C. Myers, and R. M. Thomson, Phys. Rev. Lett. 97, 091601 (2006).
- [100] C. Eling and Y. Oz, J. High Energy Phys. 06 (2011) 007.