Probing the Neutron Skin with Ultrarelativistic Isobaric Collisions

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Particle production in ultrarelativistic heavy ion collisions depends on the details of the nucleon density distributions in the colliding nuclei. We demonstrate that the charged hadron multiplicity distributions in isobaric collisions at ultrarelativistic energies provide a novel approach to determine the poorly known neutron density distributions and thus the neutron skin thickness in finite nuclei, which can in turn put stringent constraints on the nuclear symmetry energy.

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Introduction.-Nuclei are bound states of protons and neutrons by the overall attractive nuclear force. The nuclear force is short ranged, and is surpassed by Coulomb repulsion among protons at long distances. This is compensated by more neutrons to keep heavy nuclei bound. With more neutrons comes the penalty symmetry energy associated with the asymmetry between the proton and neutron numbers. The symmetry energy influences the proton and neutron density distributions, and in particular, the neutron skin thickness in nuclei (difference between the rms radii of the neutron and proton distributions, $\Delta r_{np} \equiv r_n - r_p$ [1]. The symmetry energy and its density dependence are crucial to our understanding of the masses and drip lines of neutron-rich nuclei and the equation of state (EOS) of nuclear and neutron star matter [2–10].

Measurements of the neutron density and the Δr_{np} , complemented by state-of-the-art theoretical calculations [11–14], can yield valuable information on the symmetry energy [15–18]. Exact knowledge of nucleon density distributions is also crucial to new physics searches beyond the standard model [19]. Because protons are charged, its density distributions are well measured by electrons scattering off nuclei [20,21]. The neutron density distributions are not as well measured [17]; for example, the Δr_{np} measurements of the benchmark ²⁰⁸Pb nucleus fall in the range of 0.15–0.22 fm with a typical precision of 20%-50% [16,17,22]. One limitation is the inevitable uncertainties in modeling the strong interaction of the reaction mechanisms [23]. Promising ways to measure neutron densities are through electroweak parity-violating scattering processes with electrons [24,25] and neutrinos [26], exploiting the large weak charge of the neutron compared to the diminishing one of the proton. Such measurements, although cleaner to interpret, require large luminosities [16,27]. The current measurement by PREX (Lead Radius Experiment) on the ²⁰⁸Pb Δr_{np} is $0.33^{+0.16}_{-0.18}$ fm [28].

The symmetry energy affects observables in low to intermediate energy heavy ion collisions, such as the isospin diffusion [29,30], the neutron-proton flow difference [31], the isospin-dependent pion production [32], and light cluster formation [33]. Heavy ion collisions at relativistic energies are generally considered insensitive to nuclear structures and the symmetry energy. Recent studies of isobaric ${}^{96}_{44}$ Ru + ${}^{96}_{44}$ Ru and ${}^{96}_{40}$ Zr + ${}^{96}_{40}$ Zr collisions at nucleon-nucleon center-of-mass energy of $\sqrt{s_{\rm NN}} =$ 200 GeV indicate, however, that nuclear density distributions have a noticeable effect on the total charged hadron multiplicity (N_{ch}) [34,35]. Since N_{ch} can be measured precisely, we demonstrate in this work that the $N_{\rm ch}$ distributions in isobaric collisions can be used to determine the Δr_{np} (and hence the symmetry energy) to a precision that may be comparable to or even exceed those achieved by traditional low-energy nuclear experiments.

The symmetry energy and the neutron skin.—The symmetry energy encodes the energy related to neutronproton asymmetry in nuclear matter EOS. It is conventionally defined in the binding energy per nucleon, $E(\rho, \delta) = E_0(\rho) + E_{\text{sym}}(\rho)\delta^2 + \mathcal{O}(\delta^4)$, where $\rho = \rho_n + \rho_p$ is the nucleon number density and $\delta = (\rho_n - \rho_p)/\rho$ is the isospin asymmetry with ρ_p (ρ_n) denoting the proton (neutron) density [4]. The symmetry energy can be obtained as $E_{\text{sym}}(\rho) = \frac{1}{2} [\partial^2 E(\rho, \delta)/\partial\delta^2]|_{\delta=0}$. It can be expanded at ρ_r in $\chi_r = (\rho - \rho_r)/3\rho_r$ as $E_{\text{sym}}(\rho) = E_{\text{sym}}(\rho_r) + L(\rho_r)\chi_r + \mathcal{O}(\chi_r^2)$, where $L(\rho_r) = 3\rho_r [dE_{\text{sym}}(\rho)/d\rho]|_{\rho=\rho_r}$ is the density slope parameter [4]. Especially, for $\rho_r = \rho_0 \approx 0.16 \text{ fm}^{-3}$ (the nuclear saturation density), one has $L \equiv L(\rho_0)$ which characterizes the density dependence of the $E_{\text{sym}}(\rho)$ around ρ_0 . In addition, at a subsaturation cross density $\rho_c = 0.11\rho_0/0.16 \approx 0.11 \text{ fm}^{-3}$, roughly corresponding to the average density of finite nuclei, the $E_{\text{sym}}(\rho_c) = 26.65 \pm 0.20 \text{ MeV}$ [36] is precisely obtained from nuclear binding energies [36]. At this ρ_c , a strong constraint $L(\rho_c) = 47.3 \pm 7.8 \text{ MeV}$ is obtained from the electric dipole polarizability data of ²⁰⁸Pb [37]. Generally, it is found that the $L(\rho_c)$ displays a particularly strong positive correlation with the Δr_{np} of heavy nuclei.

In the present work, we use two different nuclear energy density functionals to describe nuclear matter EOS and the properties of finite nuclei, namely, the standard Skyrme-Hartree-Fock (SHF) model (see, e.g., Ref. [38]) and the extended SHF (eSHF) model [39,40]. These two models have been very successful in describing the structures of finite nuclei, especially global properties such as binding energies and charge radii. Compared to SHF, the eSHF contains additional momentum and density-dependent twobody forces to effectively simulate the momentum dependence of the three-body forces [40]. Fitting to data using the strategy in Ref. [41], we obtain an interaction parameter set (denoted as Lc47) within eSHF by fixing $L(\rho_c) =$ 47.3 MeV [37] with $E_{\text{sym}}(\rho_c) = 26.65$ MeV [36]. We also construct two more parameter sets (Lc20 and Lc70) with $L(\rho_c) = 20$ MeV and 70 MeV, respectively, keeping the same $E_{\text{sym}}(\rho_c)$ [36], to explore the effects of the symmetry energy (and neutron skin) variations. For the SHF calculations, we use the well-known interaction set SLy4 [42,43].

Table I lists the nuclear radii of 96 Zr and 96 Ru, assuming spherical symmetry, from the eSHF calculations using Lc20, Lc47, and Lc70, and the SHF calculation with SLy4, together with the $L(\rho_c)$ and L parameters. It is seen that the four interactions give similar proton rms radius r_p for 96 Zr and 96 Ru since they are experimentally well constrained, but the neutron radius r_n increases with $L(\rho_c)$ and L, leading to a positive correlation between Δr_{np} and $L(\rho_c)$ (and L) as expected. The Δr_{np} of 208 Pb from our calculations are also listed in Table I. We note that those values essentially cover the current uncertainty in the 208 Pb measurements. Shown in Fig. 1 are the corresponding nucleon density distributions of 96 Ru and 96 Zr. In the following, we use these density distributions in heavy ion collision models to examine the effects on N_{ch} .



FIG. 1. Proton and neutron density distributions of (a) 96 Ru and (b) 96 Zr nuclei from eSHF (parameter sets Lc20, Lc47, Lc70) and SHF (SLy4).

Heavy ion collision models.—We use four typical, well developed, commonly used models for relativistic heavy ion collisions. The HIJING (Heavy ion jet interaction generator, v1.411) [44,45] is an event generator of heavy ion collisions using binary nucleon-nucleon (NN) collisions based on the GLAUBER theory, incorporating nuclear shadowing effect and partonic energy loss in medium. Each NN collision is described by multiple minijet production inspired by perturbative quantum chromodynamics (QCD), with the LUND [46] string fragmentation. The AMPT (A Multi-Phase Transport) is a dynamical model [47]. Its default version (AMPT-def, v1.26) uses HIJING but subjects the minijet partons to partonic scatterings via ZPC (Zhang's paton cascasde) [48] and, after fragmentation, hadronic scatterings via ART (a relativistic transport) [49]. The string melting version of AMPT (AMPT-sm, v2.26) [50] converts all hadrons from HIJING to partons to undergo partonic scatterings, and uses a simple coalescence for hadronization, followed by hadronic rescatterings. The UrOMD (Ultra Relativistic Quantum Molecular Dynamics, v3.4) [51,52] is a microscopic transport model with covariant propagation of hadrons on classical trajectories, combined with stochastic binary scatterings, color string formation, and resonance decays. Except for the input neutron and proton density distributions, all parameters are set to default. About 3×10^7 events within the impact parameter (b) range [0, 20] fm are simulated in each model for each set of

TABLE I. The rms radii (in fm) for neutron (r_n) and proton (r_p) distributions and the neutron skin thickness ($\Delta r_{np} \equiv r_n - r_p$) of ⁹⁶Ru and ⁹⁶Zr, and the symmetry energy slope parameters $L(\rho_c)$ and L (in MeV) from eSHF (parameter sets Lc20, Lc47, Lc70) and SHF (SLy4) calculations. The Δr_{np} values for ²⁰⁸Pb are also listed for comparison.

	$L(ho_c)$	L	⁹⁶ Zr			⁹⁶ Ru			²⁰⁸ Pb
			r_n	r_p	Δr_{np}	r_N	r _P	Δr_{np}	Δr_{np}
Lc20	20	13.1	4.386	4.27	0.115	4.327	4.316	0.011	0.109
Lc47	47.3	55.7	4.449	4.267	0.183	4.360	4.319	0.042	0.190
Lc70	70	90.0	4.494	4.262	0.232	4.385	4.32	0.066	0.264
SLy4	42.7	46.0	4.432	4.271	0.161	4.356	4.327	0.030	0.160



FIG. 2. (a) Charged hadron multiplicity (N_{ch}) distributions from AMPT-sm, UrQMD, HIJING, and AMPT-def for density set Lc47. The results for the other density sets are similar. (b) The relative widths *w* of the N_{ch} tails for four density sets in four models.

the nuclear densities for Ru + Ru and Zr + Zr collisions at $\sqrt{s_{\text{NN}}} = 200$ GeV.

Model results and discussions.—Charged hadrons are counted with transverse momentum $p_T > 0.2 \text{ GeV}/c$ and pseudorapidity $|\eta| < 0.5$. Figure 2(a) shows the N_{ch} distributions in Zr + Zr collisions calculated by the four models using the nuclear density set Lc47. The distributions are similar except at large N_{ch} . The absolute N_{ch} values are subject to large model dependence because particle production in heavy ion collisions is generally hard to model precisely. The shape of the N_{ch} distribution is, on the other hand, more robust. It is primarily determined by the interaction cross section as a function of *b*. While the tail falloff shapes are similar among AMPT-sm, AMPT-def, and UrQMD, that of HUING is distinct. To quantify the shape, we fit the tail distributions by

$$dP/dN_{\rm ch} \propto -\text{Erf}[-(N_{\rm ch}/N_{1/2}-1)/w] + 1,$$
 (1)

where $N_{1/2}$ is the N_{ch} value at half height and *w* is the width of the tail relative to $N_{1/2}$. The fitted curves are superimposed in Fig. 2(a). Figure 2(b) depicts the fit *w* values. The HIJING model has a factor of ~2 narrower tail than the other three transport models which are similar. This feature can be readily used to distinguish models once data are available, though not the main goal of this work.

The main goal of this work is to identify which density set would best describe data and hence to determine the neutron skin thickness and the symmetry energy. In a given model, at a given b, the numbers of participants (N_{part}) and binary nucleon-nucleon collisions (N_{bin}) slightly differ for different nuclear densities. Since N_{ch} is generally considered to depend on N_{part} and perhaps N_{bin} , those differences could produce an effect on N_{ch} . The effect is understandably small, hardly observable in a plot of the N_{ch} distributions themselves, but can be magnified by the ratio of the N_{ch} distribution in Ru + Ru to that in Zr + Zr [35]. These ratios using the four sets of densities, in AMPT-sm as



FIG. 3. Ratio of the N_{ch} distribution in Ru + Ru to that in Zr + Zr for various densities in AMPT-sm. The other models are similar.

an example, are shown in Fig. 3. The splittings of the $N_{\rm ch}$ tails are clear.

The ratios in Fig. 3 are illustrative to highlight the differences but are cumbersome to quantify. As seen from Fig. 2(b), the tail widths are equal among the densities in a given model, so the splittings are mostly due to the slight shifts in $N_{1/2}$, or differences in the average N_{ch} values. The $N_{1/2}$ value is sensitive to the chosen fit range. We thus use the relative $\langle N_{ch} \rangle$ difference between Ru + Ru and Zr + Zr,

$$R = 2 \frac{\langle N_{\rm ch} \rangle_{\rm RuRu} - \langle N_{\rm ch} \rangle_{\rm ZrZr}}{\langle N_{\rm ch} \rangle_{\rm RuRu} + \langle N_{\rm ch} \rangle_{\rm ZrZr}},$$
(2)

to quantify the splitting of the $N_{\rm ch}$ tails. Experimental measurements of $N_{\rm ch}$ are affected by tracking inefficiency, usually multiplicity dependent [53]. While this effect is mostly canceled in R, it is better to use only central collisions, say top 5%, where the tracking efficiency is constant to a good degree. To experimentally determine the centrality percentage, the peripheral collisions that are not recorded because of online trigger inefficiency should be taken into account. This trigger inefficiency can be experimentally corrected, and is equal between the isobar systems as a function of $N_{\rm ch}$. Even without correction, taking a conservative trigger efficiency of 95% [53], the less than 2% difference in very peripheral collisions [35] would yield only 10^{-3} mismatch in the centralities between the two systems. This would give a negligible uncertainty on R on the order of 2×10^{-4} . In short, since R is a relative measure between Ru + Ru and Zr + Zr collisions, much of the experimental effects cancel.

The *R* in each model must depend on how much the Ru and Zr nuclear density distributions differ, which can be characterized by Δr_{np} of the Zr (or Ru) nucleus. We therefore plot in Fig. 4 the *R* in the top 5% centrality against Δr_{np} of the Zr nucleus from the eSHF (SHF) calculations with Lc20, Lc47, and Lc70 (SLy4). It is found that *R*



FIG. 4. The relative $\langle N_{\rm ch} \rangle$ ratio *R* as a function of the Zr neutron skin thickness. The four sets of data points in order from left to right are from Lc20, SLy4, Lc47, Lc70 densities. The results from AMPT-sm, UrQMD, HIJING, and AMPT-def are connected by solid lines. The results from GLAUBER and TRENTO models are connected by dashed lines.

monotonically increases with Δr_{np} . This is because, with increasing Δr_{np} , the difference between Ru and Zr densities increases. This results in an increasing difference in N_{ch} between Ru + Ru and Zr + Zr collisions.

Figure 4 further shows that the value of R has a relatively weak model dependence. This includes even HJJING whose tail distribution is significantly narrower than the other models (cf. Fig. 2), which can already be discriminated by data measurements as aforementioned. In what follows, we further decipher what R entails by using nondynamical but intuitive geometrical models, namely, the GLAUBER and TRENTO models.

In a GLAUBER model [54,55], it is postulated that $N_{\rm ch} \propto [(1-x)N_{\rm part}/2 + xN_{\rm bin}]$, where x is the so-called hard-component fraction. The N_{part} and N_{bin} differ between the two isobar collisions, in a manner sensitive to the nuclear density parameters. This yields a nonvanishing Rdependence of Δr_{np} , as shown by the dynamical model results. In addition, the different contribution weights from N_{part} and N_{bin} to N_{ch} , characterized by the x parameter, also affect the R. The four models giving similar R as a function of Δr_{np} may be indicative of their similar x values. This is not surprising because all these models can approximately describe the centrality dependence of N_{ch} observed by the PHOBOS experiment, which suggested an x value of 0.1 [56]. In order to investigate the sensitivity of R to the xvalue, we show in Fig. 4 results from the GLAUBER model [57–59] with x = 0.1 and a significantly larger value of x = 0.2 (which can be taken as an extreme). Indeed, the x = 0.1 results fall within the range of the four dynamical models. The x = 0.2 results have a stronger sensitivity to Δr_{np} , indicating that the N_{bin} is more sensitive to the nuclear density distribution.

On the other hand, the two-component particle production model has drawbacks. The centrality dependences of $N_{\rm ch}$ at $\sqrt{s_{\rm NN}} = 200$ GeV and 2.76 TeV have essentially the same shape [60], whereas the hard-scattering cross section, hence the x value, should vary with energy. The PHENIX experiment found that a simple wounded quark model can successfully describe the centrality dependence of $N_{\rm ch}$ in Au + Au collisions over 7.7-200 GeV [61] and in several small systems at 200 GeV [62]. Recent azimuthal anisotropy data by STAR in ultracentral U + U collisions cannot be described by the GLAUBER approach [63] and the TRENTO model was proposed [64]. In the TRENTO model, particle production is only related to the reduced thickness, $N_{\rm ch} \propto T_R(p; T_A, T_B) \equiv [(T_A^p + T_B^p)/2]^{1/p}$ [64,65]. We use the parameter p = 0 (i.e., $N_{ch} \propto \sqrt{T_A T_B}$), a gamma fluctuation parameter k = 1.4, and a Gaussian nucleon size of 0.6 fm, which were found to well describe the multiplicity data in heavy ion collisions [64,65]. The R calculated by the TRENTO model is shown in Fig. 4. The Δr_{np} dependence is weaker because only N_{part} contributes to the N_{ch} . Considering all model results, the overall spread in R is wider at larger Δr_{np} . This is because the nuclear density difference between the Ru and Zr nuclei increases with increasing Δr_{np} , so does the model dependence.

Experimentally, the $N_{\rm ch}$ can be measured exquisitely precisely. The relative $\langle N_{\rm ch} \rangle$ difference in central collisions is immune to many experimental uncertainties. Figure 4 thus strongly suggests that the isobar data may determine Δr_{np} relatively accurately. This is afforded by the rather weak dependence of the R observable to the details of QCD physics on particle production. This is in contrast to the hadronic observables in previous lowenergy studies, where strong model dependences prevent a more precise determination [17,22]. The current experimental range of the ²⁰⁸Pb Δr_{NP} is indicated by the band on the top of Fig. 4. The ²⁰⁸Pb Δr_{np} calculated by the eSHF (SHF) to accommodate those low-energy measurements are written above the band. Our results in Fig. 4 indicate that with a given measurement of R, the precision in the derived Δr_{np} of ⁹⁶Zr can be as good as 0.05 fm covered by the four dynamical models (AMPT-sm, UrQMD, HIJING, AMPT-def) and the two static models (TRENTO and GLAUBER), as illustrated by the lower band (taking hypothetically R = 0.006). This would be an appreciable improvement over the current constraint from ²⁰⁸Pb. Our results shall thus provide a significant input to help constrain the symmetry energy, bearing important implications to nuclear matter and neutron star EOS.

We have assumed spherical nuclei in our calculations. The main idea of our work is still valid with deformed nuclei. There are a number of promising ways to determine the nuclear deformity from heavy ion collisions [66–69]. We postpone such a study to a future work.

We focused on central collisions only, in part because of the technical advantages aforementioned. However, one would naively expect the neutron skin effect to be the strongest in peripheral collisions. This would be true if one could uniquely determine and sort events in b. Experimentally, the collision centrality is usually determined by final-state particle multiplicity. In peripheral low multiplicity events, a wide range of b is mixed due to large fluctuations so that the nuclear density difference is mostly smeared out.

We note that the final state $N_{\rm ch}$ is sensitive only to the overall nucleon density at relativistic energies, and therefore indirectly sensitive to the neutron density (and neutron skin) given that the proton density is well determined. In low-energy nuclear reactions, isospin-sensitive observables [70–73] were extensively studied where the isospin-dependent interactions are important. Note that we have studied isobaric collisions at fairly high energy of $\sqrt{s_{\rm NN}} = 200$ GeV primarily because those data have already been taken so that our method can be readily applied.

Conclusions.—We have calculated nuclear densities by energy density functional theory using several symmetry energy parameters. We show, using four dynamical heavy ion collision models and two static models, that the charged hadron multiplicity difference between isobar ${}^{96}_{44}$ Ru + ${}^{96}_{42}$ Ru and ${}^{96}_{40}$ Zr + ${}^{96}_{40}$ Zr collisions has a strong sensitivity to the neutron skin and the symmetry energy, with weak model dependence. Because the charged hadron multiplicity can be precisely measured and because the systematic uncertainties largely cancel, our findings suggest potentially significant improvement to neutron skin and symmetry energy determination using relativistic isobar collision data collected in 2018 at the Relativistic Heavy Ion Collider.

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