Comment on "Einstein-Gauss-Bonnet Gravity in Four-Dimensional Spacetime"

In this Comment, we elaborate on several points raised in Ref. [1]. The authors claimed to have found a fourdimensional gravitational theory which fulfills the assumptions of the Lovelock theorem [2] though not its implications. To that end, they employed a regularization procedure already outlined in Ref. [3]. This procedure consists of rescaling the coupling constant of the Gauss-Bonnet (GB) term by 1/(D-4) and taking the $D \rightarrow 4$ *limit* after varying the Einstein-Gauss-Bonnet (EGB) action. The authors of Ref. [1] claim that the variation of the GB term is proportional to (D-4), canceling the 1/(D-4) factor and hence yielding a nonvanishing contribution to the field equations in D = 4. This claim does not stand a thorough analysis given that the variation of the kth order Lovelock Lagrangian can be decomposed as [4,5]

$$\frac{1}{\sqrt{|g|}} \frac{\delta S^{(k)}}{\delta g^{\mu\nu}} = (D - 2k) A^{(k)}_{\mu\nu} + W^{(k)}_{\mu\nu}, \tag{1}$$

where $W_{\mu\nu}^{(k)}$ is a tensor from which no (D-2k) can be extracted and which does not vanish for a general D>2k. The GB term is given by k = 2, and the field equations proposed in Ref. [1] for arbitrary dimension are

$$G_{\mu\nu} + \frac{1}{M_P^2} \Lambda_0 g_{\mu\nu} + \frac{2\alpha}{M_P^2} \left(A_{\mu\nu}^{(2)} + \frac{W_{\mu\nu}^{(2)}}{D - 4} \right) = 0,$$
 (2)

where the explicit form of $A_{\mu\nu}^{(2)}$ and $W_{\mu\nu}^{(2)}$ can be found in Ref. [5]. Although the 1/(D-4) rescaling compensates the factor multiplying the $A^{(2)}_{\mu\nu}$ term, the same procedure renders the $W^{(2)}_{\mu\nu}$ term indeterminate. This owes to the fact that, although $W_{\mu\nu}^{(2)}$ vanishes in D=4, it does so due to algebraic reasons [4,5] and not because it is proportional to some power of (D-4). Hence, the indeterminate term $W_{\mu\nu}^{(2)}/(D-4)$ renders the field equations (2) ill-defined in four dimensions. Indeed, the first problem to address would be to make sense of the limit of a tensor field [4–8], since these objects are defined for integer values of D only.

Despite the above discussion, the field equations (2) are well defined in D = 4 when constrained to particular geometries in which $W_{\mu\nu}^{(2)}$ vanishes in arbitrary dimensions. This is the case, for instance, for all conformally flat geometries, including the Friedmann-Lemaitre-Robertson-Walker or maximally symmetric solutions found in Ref. [1]. However, metric perturbations will be sensible to the ill-defined terms, hence rendering these solutions unphysical. Indeed, though linear perturbations around a maximally symmetric background are oblivious to these pathologies [1], they enter at second order through terms proportional to 1/(D-4) [5]. In this direction, other works also showed that an infinitely strongly coupled new scalar degree of freedom appears beyond linear order [9]. These results strongly suggest that Eq. (2) is generally ill-defined in four dimensions.

A possible way to circumvent the pathologies of the above field equations (2) would be to get rid of the $W_{\mu\nu}^{(2)}$ term. However, $A_{\mu\nu}^{(2)}$ is not divergenceless [5]. Hence, by virtue of the Bianchi identity under diffeomorphisms, we conclude that there is no diffeomorphism-invariant action whose variation gives Eq. (2) without the $W^{(2)}_{\mu\nu}$ term. Regarding spherically symmetric metrics of the form

$$ds^{2} = B(r)dt^{2} - B^{-1}(r)dr^{2} - r^{2}d\Omega_{D-2}^{2},$$
 (3)

for them to have a vanishing $W_{\mu\nu}^{(2)}$ in arbitrary dimension, there are conditions that B(r) must fulfill [5]. The spherically symmetric geometries presented in Ref. [1] do not satisfy these requirements and, therefore, cannot be solutions of Eq. (2), since this is not a well-defined set of equations in this case. Indeed, as proven in Ref. [5], they are neither solutions of the truncated equations (2) without the $W_{\mu\nu}^{(2)}$ term. This is not surprising, since the authors of Ref. [1] derived these geometries by taking D = 4 in the $D \ge 5$ solutions for the EGB theory found in Ref. [10] and then rescaling the GB coupling by 1/(D-4), instead of finding a solution to the rescaled field equations (2) in D=4. Furthermore, the authors of Ref. [1] state that the central curvature singularity of these geometries cannot be reached by an observer. Nevertheless, as shown in Ref. [5], radial freely falling observers do reach the singularity in finite proper time, contradicting this claim.

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