

Anomalous Supersymmetry

Georgios Katsianis,^{1,2} Ioannis Papadimitriou,³ Kostas Skenderis,^{1,2} and Marika Taylor^{1,2}
¹*STAG Research Centre, Highfield, University of Southampton, SO17 1BJ Southampton, United Kingdom*
²*Mathematical Sciences, Highfield, University of Southampton, SO17 1BJ Southampton, United Kingdom*
³*School of Physics, Korea Institute for Advanced Study, Seoul 02455, Korea*

 (Received 4 April 2019; published 14 June 2019)

We show that supersymmetry is anomalous in $\mathcal{N} = 1$ superconformal quantum field theories (SCFTs) with an anomalous R symmetry. This anomaly was originally found in holographic SCFTs at strong coupling. Here we show that this anomaly is present, in general, and demonstrate it for the massless superconformal Wess-Zumino model via a one-loop computation. The anomaly appears first in four-point functions of two supercurrents either with two R currents or with an R current and an energy-momentum tensor. In fact, the Wess-Zumino consistency conditions together with the standard R -symmetry anomaly imply the existence of the anomaly. We outline the implications of this anomaly.

DOI: [10.1103/PhysRevLett.122.231602](https://doi.org/10.1103/PhysRevLett.122.231602)

Anomalies of symmetries play an important role in quantum field theories. If a global symmetry is anomalous, classical selection rules are not respected in the quantum theory, and classically forbidden processes may occur. This is a feature of the theory, and it is linked with observable effects. For example, the axial anomaly explains the π^0 decay and leads to the resolution of the $U(1)$ problem in QCD [1,2]. On the other hand, anomalies in local (gauge) symmetries lead to inconsistencies, such as lack of unitarity, and they must be canceled. An important corollary is that anomalous global symmetries cannot be consistently coupled to corresponding local symmetries. Reviews on anomalies in quantum field theories may be found in Refs. [3,4].

Anomalies in supersymmetric theories.—In this Letter, we discuss a new anomaly in four-dimensional supersymmetric quantum field theories with an anomalous R symmetry: Global supersymmetry itself is anomalous. This anomaly was discovered in the context of superconformal theories that can be realized holographically [5]. Here, we show that the same anomaly arises in the perturbation theory in the simplest supersymmetric model: the free superconformal Wess-Zumino (WZ) model.

An anomaly may be detected either by putting the theory on a nontrivial background or by computing correlation functions on a flat background and checking whether the Ward identities are satisfied. The latter method was the one that led to the original discovery of anomalies via one-loop

triangle diagrams [1,2]. Here, we will carry out the analogous computation for the supersymmetry anomaly. The anomaly is associated, in particular, with anomalous one-loop contributions to four-point correlation functions between two supersymmetry currents and two R currents or an R current and an energy-momentum tensor. We will discuss the former in the free superconformal WZ model, but analogous contributions would arise in any supersymmetric theory with a (softly broken) anomalous R symmetry. Actually, as will be sketched below and is shown in detail in the companion paper [6], the WZ consistency conditions [7] together with the standard triangle anomalies imply that supersymmetry must be anomalous.

Discussion of anomalies in 4d (super)conformal quantum field theory (QFT) has a long history. It has been known since the 1970s [8,9] that the trace of the stress tensor \mathcal{T}_μ^μ is anomalous in the presence of a curved background metric $g_{\mu\nu}$ and background source A_μ for a chiral current \mathcal{J}_μ , and the R current is similarly anomalous. Moreover, there are generally mixed anomalies involving two energy-momentum tensors and a chiral current [10,11]. It has also been known since Ref. [12] that the currents sit in a supermultiplet, as do the anomalies. In particular, the trace anomaly and the R -current anomaly are in the same multiplet as the gamma trace of the supercurrent, $\gamma^\mu \mathcal{Q}_\mu$. The latter is an anomaly in the conservation of the special supersymmetry current, $x^\nu \gamma_\nu \mathcal{Q}_\mu$. It follows that special supersymmetry (sometimes also called S supersymmetry) is anomalous. It was believed, however, that supersymmetry itself (sometimes called Q supersymmetry) is preserved; i.e., the conservation of \mathcal{Q}_μ is nonanomalous.

There have been extensive studies in the past regarding anomalies in supersymmetry. It was realized early on [13–18] that one cannot maintain at the quantum level simultaneously $\partial^\mu \mathcal{Q}_\mu = 0$ and $\gamma^\mu \mathcal{Q}_\mu = 0$ and, if the model

Published by the American Physical Society under the terms of the [Creative Commons Attribution 4.0 International license](https://creativecommons.org/licenses/by/4.0/). Further distribution of this work must maintain attribution to the author(s) and the published article's title, journal citation, and DOI. Funded by SCOAP³.

is a gauge theory, gauge invariance: One of the three conditions must be relaxed, and the standard choice is to have a superconformal anomaly. This is the standard superconformal anomaly mentioned above and is distinct from the anomaly discussed here. Also distinct is the Konishi anomaly [19,20], which is a superspace version of the chiral anomaly in supersymmetric gauge theories.

Another set of studies, reviewed in Ref. [21], considers the effective action for elementary fields and examines whether it is invariant under supersymmetry including loop effects; it investigates the conservation of the supercurrent inside correlators of elementary fields and/or solves the WZ consistency conditions relevant for this setup and finds no supersymmetry anomaly. This does not contradict the results we present below: To find the anomaly, one should either put the theory on a nontrivial background or consider correlation functions of (classically) conserved currents. (To illustrate this point, consider a free fermion in a complex representation in flat spacetime. This theory has a standard axial anomaly originating from the three-point function of the axial current. However, if one looks only at correlators of elementary fields, these are nonanomalous, and the axial current inside such correlators is conserved.) Studies involving correlators of currents have also appeared, but typically only discuss three-point functions of bosonic currents. As mentioned above, the supersymmetry anomaly appears first in four-point functions involving two supercurrents and two bosonic currents, and to our knowledge these have not been computed before.

Anomalies associated with correlation functions of conserved currents can be analyzed by coupling the currents to external sources, which in our case form an $\mathcal{N} = 1$ superconformal multiplet. As such, the anomaly we discuss here could be related to existing superspace results on anomaly candidates for $D = 4$, $\mathcal{N} = 1$ supergravity theories [22–26] (in particular, in type II anomalies in Ref. [25]), though we emphasize that in our case the supergravity fields are external and thus nondynamical (off shell).

A supersymmetry anomaly appears in super Yang-Mills (SYM) theory in the WZ gauge when there are gauge anomalies [27] (see also [28–30]). This anomaly is easy to understand: In the WZ gauge, supersymmetry transformations require a compensating gauge transformation, and this transfers the anomaly from the gauge sector to supersymmetry. When the SYM theory is consistent at the quantum level (i.e., the gauge anomalies cancel), then supersymmetry is also nonanomalous. A supersymmetry anomaly appears in theories with gravitational anomalies [31–33], as one may anticipate based on the fact that the energy-momentum tensor and the supercurrent are part of the same supermultiplet. Indeed, this supersymmetry anomaly sits in the same multiplet as the gravitational anomaly.

Here, we will discuss a supersymmetry anomaly in consistent QFTs (no gauge anomalies) which have a

conserved energy-momentum tensor. We also emphasize that we are concerned with local anomalies, not with beta functions.

Holographic anomalies.—The anomaly we discuss here was first computed holographically [5]. In holography, given a bulk action, one can use holographic renormalization [34,35] to compute the Ward identities and anomalies of the dual QFT. AdS/CFT relates the $\mathcal{N} = 1$ superconformal quantum field theory (SCFT) in four dimensions to $\mathcal{N} = 2$ gauged supergravity in five dimensions. Starting from gauged supergravity in an asymptotically locally AdS₅ spacetime and turning on sources for all superconformal currents, one can compute the complete set of superconformal anomalies. This computation is available for holographic CFTs, which, in particular, means that the central charges should satisfy $a = c$ as $N \rightarrow \infty$ [34].

Early attempts to compute the supertrace Ward identity can be found in Refs. [36,37], but these missed contributions to the anomaly involving the R -symmetry current and the Ricci tensor. Following the work of Pestun [38], there was renewed interest in supersymmetric theories on curved spacetimes and their holographic duals. The holographic anomalies for bosonic currents were computed in Ref. [39], reproducing (and correcting) known field theory results [40]. The full superconformal anomalies for the $\mathcal{N} = 1$ current multiplet were computed holographically in Ref. [5], while Ref. [41] obtained the superconformal anomalies in the presence of local supersymmetric scalar couplings. An analogous holographic computation relevant to two-dimensional SCFTs was reported in Ref. [42].

The holographic results leave open the possibility that the anomaly is special to holographic theories at strong coupling. In this Letter, we show that this is not the case. One could have anticipated the anomaly based on the structure of the supersymmetric variation of the supercurrent, which is of the schematic form $\delta Q^\mu \sim \gamma_\nu T^{\mu\nu} \epsilon + C^{\mu\nu\rho} \partial_\nu \mathcal{J}_\rho \epsilon$, where $C^{\mu\nu\rho}$ is a tensor constructed from gamma matrices and the metric. The Ward identity for the four-point function involving two supercurrents and two R currents would then involve terms of the form

$$\begin{aligned} & \partial_\mu^{x_1} \langle Q^\mu(x_1) \bar{Q}^\nu(x_2) \mathcal{J}^\kappa(x_3) \mathcal{J}^\lambda(x_4) \rangle \\ & \sim \delta(x_1 - x_2) \langle \delta \bar{Q}^\nu(x_2) \mathcal{J}^\kappa(x_3) \mathcal{J}^\lambda(x_4) \rangle + \dots, \quad (1) \end{aligned}$$

where the dots denote additional terms [the exact Ward identity is given (9)]. Using the variation of the supercurrent, we find that the rhs contains the three-point function of three R currents, which is anomalous, and correspondingly one may anticipate (1) will be anomalous. Similarly, the same four-point function but with one of the R currents replaced by an energy-momentum tensor is expected to be anomalous, since $\langle \mathcal{J} T T \rangle$ is anomalous. To determine whether an anomaly appears or not, we need to carry out the computation explicitly. Before we turn to this,

we discuss the consistency condition that the anomalies must satisfy.

Wess-Zumino consistency.—Let e_μ^a , A_μ , and ψ_μ denote the sources (vierbein, gauge field, and gravitino) that couple to the superconformal currents and $\mathcal{W}[e, A, \psi]$ be the generating functional of connected graphs. We define the currents in the presence of sources (as usual) by

$$\mathcal{T}_a^\mu = e^{-1} \frac{\delta \mathcal{W}}{\delta e_\mu^a}, \quad \mathcal{J}^\mu = e^{-1} \frac{\delta \mathcal{W}}{\delta A_\mu}, \quad \mathcal{Q}^\mu = e^{-1} \frac{\delta \mathcal{W}}{\delta \bar{\psi}_\mu}, \quad (2)$$

where $e \equiv \det(e_\mu^a)$. In the presence of anomalies

$$\delta_i \mathcal{W} = \int d^4x e \epsilon_i \mathcal{A}_i, \quad (3)$$

where δ_i denotes the superconformal transformations, ϵ_i are the (local) parameters of the transformations, and \mathcal{A}_i are the corresponding anomalies. The variations form an algebra, $[\delta_i, \delta_j] = f_{ij}^k \delta_k$, and using this in (3) we obtain the WZ consistency condition

$$\int d^4x [\delta_i(e \epsilon_j \mathcal{A}_j) - \delta_j(e \epsilon_i \mathcal{A}_i) - f_{ij}^k e \epsilon_k \mathcal{A}_k] = 0. \quad (4)$$

The transformation rules and the local algebra they satisfy are derived in Ref. [6] and are given in Table I.

Assuming the R -symmetry current has the standard triangle anomalies (i.e., assuming the form of \mathcal{A}_R in Table II), the WZ consistency conditions (4) may be viewed as equations to determine the remaining anomalies. This computation is presented in Ref. [6], and the results are summarized in Table II. Note, in particular, that all anomalies are given in terms of the central charges a and c . The anomalies of the bosonic currents are in agreement with the results derived in Refs. [39,40]. The supersymmetry anomaly \mathcal{A}_Q that we discuss here is related to the R -symmetry anomaly \mathcal{A}_R through the same descent equation that relates the supersymmetry anomaly discussed in Ref. [27] to the corresponding gauge anomaly. However, as noted earlier, there are important differences in the physics (in Ref. [27], the gauge anomalies must vanish for consistency of the model, while this is not so for the R

anomalies relevant for us), as well as in the context (the WZ conditions discussed in Ref. [27] are for a vector multiplet in flat space, while the anomalies in Table II are those of $\mathcal{N} = 1$ conformal supergravity [6]).

Here we discuss only one of the WZ equations: the one obtained by considering the commutator of R symmetry (with parameter θ) with Q supersymmetry (with parameter ϵ):

$$\int d^4x [\delta_\epsilon(e \theta \mathcal{A}_R) - \delta_\theta(e \epsilon \mathcal{A}_Q)] = 0. \quad (5)$$

Using the explicit form of \mathcal{A}_R , it is easy to see that $\delta_\epsilon \mathcal{A}_R \neq 0$ and the WZ consistency condition requires that $\mathcal{A}_Q \neq 0$. This argument does not rely on the theory having conformal invariance, and thus we expect any 4d supersymmetric theory with an R -symmetry anomaly to have a corresponding anomaly in the conservation of the supercurrent. (This expectation has been verified in the followup paper [43].)

One may wonder whether this anomaly can be removed by adding a local counterterm \mathcal{W}_{ct} to the action such that $\mathcal{W}_{\text{ren}} = \mathcal{W} + \mathcal{W}_{\text{ct}}$ is nonanomalous, i.e., $\delta_\epsilon \mathcal{W}_{\text{ren}} = 0$. Using the commutator of two supersymmetry variations, $[\delta_\epsilon, \delta_{\epsilon'}]$, given in Table I, we find

$$(\delta_\xi + \delta_\lambda + \delta_\theta) \mathcal{W}_{\text{ren}} = 0 \Rightarrow (\delta_\xi + \delta_\lambda) \mathcal{W}_{\text{ren}} \neq 0, \quad (6)$$

since $\delta_\theta \mathcal{W}_{\text{ren}} = \mathcal{A}_R \neq 0$. It follows that if one wishes to preserve supersymmetry, \mathcal{W}_{ct} must break diffeomorphisms and/or local Lorentz transformations. [Note that, since \mathcal{A}_R is a genuine anomaly, it is not possible to set the rhs of the second equation in (6) to zero using a local counterterm. This implies that there are no further local counterterms that can restore diffeomorphisms and local Lorentz invariance.] Next, we calculate this anomaly by one-loop computations within a specific model.

Model.—Consider the massless Wess-Zumino action with one complex bosonic field ϕ and one Majorana fermionic field χ :

$$S = - \int d^4x \left(\partial_\mu \phi \partial^\mu \phi^* + \frac{1}{2} \bar{\chi} \not{\partial} \chi \right). \quad (7)$$

The conserved currents are given in Table III. We have included improvement terms so that classically $\mathcal{T}_\mu^\mu = 0$, $\gamma^\mu \mathcal{Q}_\mu = 0$ and we are dealing with an $\mathcal{N} = 1$ SCFT.

TABLE I. Transformation rules of the current sources and their algebra, to leading order in the gravitino. All other commutators vanish, except for that of two diffeomorphisms and two local Lorentz transformations, which take a standard form.

$$\begin{aligned} \delta e_\mu^a &= \xi^\lambda \partial_\lambda e_\mu^a + e_\mu^a \partial_\lambda \xi^\lambda - \lambda_b^a e_\mu^b + \sigma e_\mu^a - \frac{1}{2} \bar{\psi}_\mu \gamma^a \epsilon, \quad \delta \psi_\mu = \xi^\lambda \partial_\lambda \psi_\mu + \psi_\lambda \partial_\mu \xi^\lambda - \frac{1}{4} \lambda_{ab} \gamma^{ab} \psi_\mu + \frac{1}{2} \sigma \psi_\mu + D_\mu \epsilon - \gamma_\mu \eta - i \gamma^5 \theta \psi_\mu, \\ \delta A_\mu &= \xi^\lambda \partial_\lambda A_\mu + A_\lambda \partial_\mu \xi^\lambda + (3i/4) \bar{\phi}_\mu \gamma^5 \epsilon - (3i/4) \bar{\psi}_\mu \gamma^5 \eta + \partial_\mu \theta, \quad \phi_\mu \equiv \frac{1}{3} \gamma^\nu (D_\nu \psi_\mu - D_\mu \psi_\nu - (i/2) \gamma^5 \epsilon_{\nu\mu}^\rho D_\rho \psi_\sigma) \\ [\delta_\epsilon, \delta_{\epsilon'}] &= \delta_\xi + \delta_\lambda + \delta_\theta, \quad \xi^\mu = \frac{1}{2} \bar{\epsilon}' \gamma^\mu \epsilon, \quad \lambda_b^a = -\frac{1}{2} (\bar{\epsilon}' \gamma^\nu \epsilon) \omega_{\nu b}^a, \quad \theta = -\frac{1}{2} (\bar{\epsilon}' \gamma^\nu \epsilon) A_\nu \\ [\delta_\epsilon, \delta_\eta] &= \delta_\sigma + \delta_\lambda + \delta_\theta, \quad \sigma = \frac{1}{2} \bar{\epsilon} \eta, \quad \lambda_b^a = -\frac{1}{2} \bar{\epsilon} \gamma_b^a \eta, \quad \theta = -\frac{3i}{4} \bar{\epsilon} \gamma^5 \eta \end{aligned}$$

TABLE II. Anomalous Ward identities and corresponding anomalies [6]. ($D_\mu \psi_\nu \equiv [\partial_\mu + \frac{1}{4} \omega_\mu^{ab}(e, \psi) \gamma_{ab} + i\gamma^5 A_\mu] \psi_\nu - \Gamma_{\mu\nu}^\rho \psi_\rho$ with $\omega_\mu^{ab}(e, \psi) \equiv \omega_\mu^{ab}(e) + \frac{1}{4}(\bar{\psi}_a \gamma_\mu \psi_b + \bar{\psi}_\mu \gamma_a \psi_b - \bar{\psi}_\mu \gamma_b \psi_a)$; ∇_μ is the Levi-Civita connection; ϕ_μ is defined in Table I.)

$e_\mu^a \mathcal{T}_a^\mu + \frac{1}{2} \bar{\psi}_\mu \mathcal{Q}^\mu = \mathcal{A}_W$, $\nabla_\mu \mathcal{T}^\mu + i \bar{\psi}_\mu \gamma^5 \mathcal{Q}^\mu = \mathcal{A}_R$	Weyl square: $W^2 \equiv W_{\mu\nu\rho\sigma} W^{\mu\nu\rho\sigma}$
$D_\mu \mathcal{Q}^\mu - \frac{1}{2} \gamma^a \psi_\mu \mathcal{T}_a^\mu - \frac{3i}{4} \gamma^5 \phi_\mu \mathcal{T}^\mu = \mathcal{A}_Q$, $\gamma_\mu \mathcal{Q}^\mu - \frac{3i}{4} \gamma^5 \psi_\mu \mathcal{T}^\mu = \mathcal{A}_S$	Euler density: $E = R_{\mu\nu\rho\sigma} R^{\mu\nu\rho\sigma} - 4R_{\mu\nu} R^{\mu\nu} + R^2$
$\mathcal{A}_W = (c/16\pi^2)(W^2 - \frac{8}{3}F^2) - (a/16\pi^2)E + \mathcal{O}(\psi^2)$, $\mathcal{A}_R = [(5a-3c)/27\pi^2] \tilde{F}F + [(c-a)/24\pi^2] \mathcal{P}$	Pontryagin density: $\mathcal{P} \equiv \tilde{R}^{\mu\nu\rho\sigma} R_{\mu\nu\rho\sigma}$ $\tilde{R}_{\mu\nu\rho\sigma} \equiv \frac{1}{2} \epsilon_{\mu\nu}^{\kappa\lambda} R_{\kappa\lambda\rho\sigma}$
$\mathcal{A}_Q = -[(5a-3c)i/9\pi^2] \tilde{F}^{\mu\nu} A_\mu \gamma^5 \phi_\nu + [(a-c)/6\pi^2] (\nabla_\mu (A_\rho \tilde{R}^{\rho\sigma\mu\nu}) \gamma_{(\nu} \psi_{\sigma)}) - \frac{1}{4} F_{\mu\nu} \tilde{R}^{\mu\nu\rho\sigma} \gamma_\rho \psi_\sigma + \mathcal{O}(\psi^3)$	Schouten tensor: $P_{\mu\nu} \equiv \frac{1}{2} (R_{\mu\nu} - \frac{1}{6} R g_{\mu\nu})$
$\mathcal{A}_S = [(5a-3c)/6\pi^2] \tilde{F}^{\mu\nu} [D_\mu - (2i/3)A_\mu \gamma^5] \psi_\nu + (ic/6\pi^2) F^{\mu\nu} (\gamma_\mu^{\rho\sigma} \delta_\nu^\rho - \delta_\mu^{\rho\sigma} \delta_\nu^\rho) \gamma^5 D_\rho \psi_\sigma + [3(2a-c)/4\pi^2] P_{\mu\nu} g^{\mu[\nu} \gamma^{\rho\sigma]} D_\rho \psi_\sigma + [(a-c)/8\pi^2] \times (R^{\mu\nu\rho\sigma} \gamma_{\mu\nu} - \frac{1}{2} R g_{\mu\nu} g^{\mu[\nu} \gamma^{\rho\sigma]}) D_\rho \psi_\sigma + \mathcal{O}(\psi^3)$	$U(1)_R$ field strengths: $\tilde{F}_{\mu\nu} \equiv \frac{1}{2} \epsilon_{\mu\nu}^{\rho\sigma} F_{\rho\sigma}$ $F^2 \equiv F_{\mu\nu} F^{\mu\nu}$ $F\tilde{F} \equiv F_{\mu\nu} \tilde{F}^{\mu\nu}$

From the form of the anomaly \mathcal{A}_Q in Table II follows that the first anomalous contribution in flat space correlators appears in four-point functions involving two supercurrents and either two R currents or an R current and an energy-momentum tensor. Here we discuss the former, referring to Ref. [44] for a detailed account of both cases.

Since we seek to investigate the possibility of a supersymmetry anomaly, we should not assume the existence of a supersymmetric regulator: The one-loop computation should not be done in superspace. (On the other hand, the form of anomalies respects the symmetries they break, and, thus, one may use superspace to analyze possible anomaly candidates.) We will instead do the computation in components and use the same regulator as in the original triangle anomaly computation, namely, momentum cutoff [1,2]. We will consider the four-point correlation function

$$\langle \mathcal{Q}^\mu(x_1) \bar{\mathcal{Q}}^\nu(x_2) \mathcal{T}^\kappa(x_3) \mathcal{T}^\lambda(x_4) \rangle. \quad (8)$$

Standard path integral manipulations show that this correlator classically satisfies the following Ward identity:

$$\begin{aligned} & -i \partial_\mu^{x_1} \langle \mathcal{Q}_1^\mu \bar{\mathcal{Q}}_2^\nu \mathcal{T}_3^\kappa \mathcal{T}_4^\lambda \rangle \\ & = \delta^{(4)}(x_{12}) \langle \delta \bar{\mathcal{Q}}_1^\nu \mathcal{T}_3^\kappa \mathcal{T}_4^\lambda \rangle + \{ \delta^{(4)}(x_{13}) \langle \delta \mathcal{T}_1^\kappa \bar{\mathcal{Q}}_2^\nu \mathcal{T}_4^\lambda \rangle \\ & \quad - \partial_\rho^{x_1} [\delta^{(4)}(x_{13}) \langle \delta \mathcal{T}_1^{\rho\kappa} \bar{\mathcal{Q}}_2^\nu \mathcal{T}_4^\lambda \rangle] + (3, \kappa) \leftrightarrow (4, \lambda) \} \\ & \quad - \partial_\rho^{x_1} [\delta^{(4)}(x_{12}) \langle \delta \bar{\mathcal{Q}}_1^{\nu\rho} \mathcal{T}_3^\kappa \mathcal{T}_4^\lambda \rangle], \end{aligned} \quad (9)$$

where we have used the shorthand notation $\mathcal{Q}^\mu(x_i) \equiv \mathcal{Q}_i^\mu$, etc., $x_{ij} \equiv x_i - x_j$, and the contributions on the rhs are

expressed in terms of the supersymmetry variations of the currents: $\delta_\epsilon \mathcal{Q}^\mu = \epsilon \delta \mathcal{Q}^\mu + \partial_\nu \epsilon \delta \mathcal{Q}^{\mu\nu}$ and idem for \mathcal{T}^μ . A similar Ward identity follows from R invariance.

One-loop computation.—We now compute (9). Since the theory is free, the complete computation is one loop. The four-point function receives contributions from three classes of Feynman box diagrams, shown in Fig. 1; this computation is straightforward but tedious.

One may verify that (9), as well as the corresponding R -symmetry Ward identity, is (naively) satisfied by a simple shift of the loop momentum, much the same way as the triangle Ward identity is naively satisfied. Again in parallel with the triangle anomaly, (part of) the one-loop contributions to the four-point function are superficially linearly divergent. This implies that there is a momentum routing ambiguity when using a momentum cutoff regulator (see, for example, Jackiw's lectures in Ref. [3]).

We proceed by taking the $\partial_\mu^{x_3}$ of (9) and subtracting from it the $\partial_\mu^{x_1}$ derivative of the corresponding R -symmetry Ward identity. By construction, the four-point functions cancel, and one is left with an identity involving three-point functions only (namely, the terms appearing on the rhs of the Ward identities). Had these three-point functions been nonanomalous, this would be an identity. However, the three-point functions involve the anomalous $\langle \mathcal{T} \mathcal{T} \mathcal{T} \rangle$ correlator, and this implies that either (9) or the corresponding R -symmetry Ward identity should be anomalous. Assuming the form of the bosonic Ward identities is standard (i.e., given by the expressions in Table II), the

 TABLE III. The (on-shell) energy-momentum tensor T_a^μ , the R -symmetry current \mathcal{T}^μ , and the supersymmetry current \mathcal{Q}^μ , for the massless superconformal WZ model in flat space.

$$\begin{aligned} T_a^\mu &= (\eta^{\mu\rho} \eta_a^\sigma + \eta^{\mu\sigma} \eta_a^\rho - \eta_a^\mu \eta^{\rho\sigma}) \partial_\rho \phi^* \partial_\sigma \phi - \frac{1}{3} (\partial^\mu \partial_a - \eta_a^\mu \partial^2) (\phi^* \phi) + \frac{1}{4} \bar{\chi} (\gamma^\mu \partial_a + \gamma_a \partial^\mu) \chi \\ \mathcal{T}^\mu &= (2i/3) (\phi^* \partial^\mu \phi - \phi \partial^\mu \phi^* + \frac{1}{4} \bar{\chi} \gamma^\mu \gamma^5 \chi) \\ \mathcal{Q}^\mu &= (1/\sqrt{2}) (\not{\partial} \phi \gamma^\mu \chi_R + \not{\partial} \phi^* \gamma^\mu \chi_L) + (\sqrt{2}/3) \gamma^{\mu\nu} \partial_\nu (\phi \chi_R + \phi^* \chi_L), \chi_R \equiv \frac{1}{2} (1 - \gamma^5) \chi. \end{aligned}$$

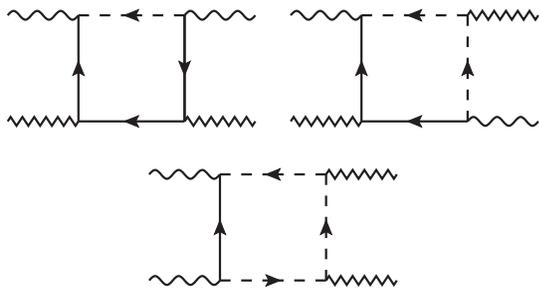


FIG. 1. Box diagrams contributing to the four-point correlation function (8). Zigzag lines denote R currents, wavy lines denote supersymmetry currents, straight lines denote fermionic propagators, and dashed lines denote bosonic propagators.

R -symmetry four-point function Ward identity is not anomalous, and, therefore, the supersymmetry Ward identity is anomalous. This computation is the counterpart of (5) but now in terms of Feynman diagrams.

One can then show that there is a momentum routing such that (i) the triangle R -symmetry anomaly is reproduced, (ii) the four-point R -symmetry Ward identity is nonanomalous, and (iii) the supersymmetry Ward identity is anomalous, with the anomaly given in Table II and with $c = 2a = 1/24$, which are the values in our model. In addition, upon taking the gamma trace of the same four-point function, $\gamma_\mu \langle Q^\mu \bar{Q}^\nu \mathcal{J}^\kappa \mathcal{J}^\lambda \rangle$, one automatically reproduces the \mathcal{A}_S anomaly given in Table II.

In general, changing the momentum routing, one may move the anomaly from one conserved current to another. This would be equivalent to adding local finite counterterms, and as argued earlier there is no choice of such counterterms that would remove the supersymmetry anomaly while preserving diffeomorphisms and local Lorentz transformations.

It is also straightforward to check that the same anomaly is present in the massive WZ model as well. As in the case of standard triangle anomalies, adding a mass term modifies the Ward identities, but the anomaly remains the same. This is as expected, since the anomaly arises from the UV behavior of Feynman diagrams and the parts of the loop computation that give rise to the anomaly remain the same.

Implications of the anomaly.—Let us conclude with a few comments about the implications of this anomaly. As mentioned earlier, an important consequence is that a SQFT with such a supersymmetry anomaly cannot be coupled to dynamical supergravity. (The anomalous R symmetry alone implies that coupling to a supergravity that gauges the R symmetry is inconsistent. Here, we see that couplings to supergravity that do not gauge the R symmetry are also inconsistent.) In the context of supersymmetric model building, one does not usually work with theories with an R symmetry, anomalous or nonanomalous; nonanomalous R symmetry is not compatible with gaugino masses (see [45]). More generally, one does not expect a theory with continuous symmetry to emerge from a consistent

quantum theory of gravity, such as string theory. However, such models may be considered in bottom-up approaches (see [46] for a recent example). Similar comments apply to bottom-up string cosmology models. This anomaly also affects supersymmetric localization computations, as has already been noted in Refs. [5,6,41,42]. However, it is possible that a suitable noncovariant local counterterm (for theories with $a = c$, such a counterterm evaluated on supersymmetric backgrounds of the form $S^1 \times M_3$, with M_3 a Seifert manifold, should agree with the counterterm used in Ref. [47]) may cancel the rigid supersymmetry anomaly at the expense of breaking certain diffeomorphisms on a given supersymmetric background. From a more formal perspective, it would be interesting to explore how the supersymmetry anomaly is captured in index theorems. It would also be interesting to investigate the existence of such an anomaly in other dimensions and/or extended supersymmetry.

We thank Benjamin Assel, Roberto Auzzi, Friedmann Brandt, Lorian Bonora, Davide Cassani, Cyril Closset, Camillo Imbimbo, Manthos Karydas, Heeyeon Kim, Zohar Komargodski, Dario Martelli, Sunil Mukhi, Sameer Murthy, Parameswaran Nair, Dario Rosa, Stanislav Schmidt, Ashoke Sen, and Peter West for illuminating discussions and email correspondence. K. S. and M. T. are supported in part by the Science and Technology Facilities Council (Consolidated Grant “Exploring the Limits of the Standard Model and Beyond”). This research was supported in part by the National Science Foundation under Grant No. NSF PHY-1748958, and this project has received funding or support from the European Union’s Horizon 2020 research and innovation program under the Marie Skłodowska-Curie Grant Agreement No. 690575. I. P. thanks the University of Southampton, King’s College London, and the International Center for Theoretical Physics in Trieste for hospitality and partial financial support during the completion of this work. M. T. thanks the Kavli Institute for the Physics and Mathematics of the Universe for hospitality during the completion of this work.

Note added.—Recently, a related work [48] appeared on the arXiv.

-
- [1] S. L. Adler, *Phys. Rev.* **177**, 2426 (1969).
 - [2] J. S. Bell and R. Jackiw, *Nuovo Cimento A* **60**, 47 (1969).
 - [3] S. B. Treiman, E. Witten, R. Jackiw, and B. Zumino, *Current Algebra and Anomalies* (World Scientific, Singapore, 1986).
 - [4] K. Fujikawa and H. Suzuki, *Path Integrals and Quantum Anomalies* (Clarendon, Oxford, 2004).
 - [5] I. Papadimitriou, *J. High Energy Phys.* **07** (2017) 038.
 - [6] I. Papadimitriou, *J. High Energy Phys.* **04** (2019) 040.
 - [7] J. Wess and B. Zumino, *Phys. Lett.* **37B**, 95 (1971).

- [8] D. M. Capper and M. J. Duff, *Nuovo Cimento A* **23**, 173 (1974).
- [9] S. Deser, M. J. Duff, and C. J. Isham, *Nucl. Phys.* **B111**, 45 (1976).
- [10] R. Delbourgo and A. Salam, *Phys. Lett.* **40B**, 381 (1972).
- [11] L. Alvarez-Gaume and E. Witten, *Nucl. Phys.* **B234**, 269 (1984).
- [12] S. Ferrara and B. Zumino, *Nucl. Phys.* **B87**, 207 (1975).
- [13] B. de Wit and D. Z. Freedman, *Phys. Rev. D* **12**, 2286 (1975).
- [14] L. F. Abbott, M. T. Grisaru, and H. J. Schnitzer, *Phys. Rev. D* **16**, 2995 (1977).
- [15] L. F. Abbott, M. T. Grisaru, and H. J. Schnitzer, *Phys. Lett.* **71B**, 161 (1977).
- [16] L. F. Abbott, M. T. Grisaru, and H. J. Schnitzer, *Phys. Lett.* **73B**, 71 (1978).
- [17] K. Hieda, A. Kasai, H. Makino, and H. Suzuki, *Prog. Theor. Exp. Phys.* **2017**, 063B03 (2017).
- [18] Y. R. Batista, B. Hiller, A. Cherchiglia, and M. Sampaio, *Phys. Rev. D* **98**, 025018 (2018).
- [19] K. Konishi, *Phys. Lett.* **135B**, 439 (1984).
- [20] K.-i. Konishi and K.-i. Shizuya, *Nuovo Cimento A* **90**, 111 (1985).
- [21] O. Piguet and K. Sibold, *Renormalized supersymmetry. The perturbation theory of $N = 1$ supersymmetric theories in flat space-time* (Birkhauser, Boston, 1986), Vol. 12.
- [22] L. Bonora, P. Pasti, and M. Tonin, *Nucl. Phys.* **B252**, 458 (1985).
- [23] I. L. Buchbinder and S. M. Kuzenko, *Nucl. Phys.* **B274**, 653 (1986).
- [24] F. Brandt, *Classical Quantum Gravity* **11**, 849 (1994).
- [25] F. Brandt, *Ann. Phys. (N.Y.)* **259**, 357 (1997).
- [26] L. Bonora and S. Giaccari, *J. High Energy Phys.* **08** (2013) 116.
- [27] H. Itoyama, V. P. Nair, and H.-c. Ren, *Nucl. Phys.* **B262**, 317 (1985).
- [28] O. Piguet and K. Sibold, *Nucl. Phys.* **B247**, 484 (1984).
- [29] E. Guadagnini and M. Mintchev, *Nucl. Phys.* **B269**, 543 (1986).
- [30] B. Zumino, in *Symposium on Anomalies, Geometry, Topology* Argonne, Illinois, 1985, edited by W. A. Bardeen and A. R. White (World Scientific, Singapore, 1985).
- [31] P. S. Howe and P. C. West, *Phys. Lett.* **156B**, 335 (1985).
- [32] Y. Tani, *Nucl. Phys.* **B259**, 677 (1985).
- [33] H. Itoyama, V. P. Nair, and H.-c. Ren, *Phys. Lett.* **168B**, 78 (1986).
- [34] M. Henningson and K. Skenderis, *J. High Energy Phys.* **07** (1998) 023.
- [35] S. de Haro, S. N. Solodukhin, and K. Skenderis, *Commun. Math. Phys.* **217**, 595 (2001).
- [36] M. Chaichian and W. F. Chen, *Nucl. Phys.* **B678**, 317 (2004).
- [37] M. Chaichian and W. F. Chen, in *Symmetries in Gravity and Field Theory*, edited by V. Aldaya and J. M. Cervero (Ediciones Universidad de Salamanca, Salamanca, 2003), pp. 449–472.
- [38] V. Pestun, *Commun. Math. Phys.* **313**, 71 (2012).
- [39] D. Cassani and D. Martelli, *J. High Energy Phys.* **10** (2013) 025.
- [40] D. Anselmi, D. Z. Freedman, M. T. Grisaru, and A. A. Johansen, *Nucl. Phys.* **B526**, 543 (1998).
- [41] O. S. An, *J. High Energy Phys.* **12** (2017) 107.
- [42] O. S. An, Y. H. Ko, and S.-H. Won, *Phys. Rev. D* **99**, 106007 (2019).
- [43] I. Papadimitriou, [arXiv:1904.00347](https://arxiv.org/abs/1904.00347).
- [44] G. Katsianis, I. Papadimitriou, K. Skenderis, and M. Taylor (to be published).
- [45] M. Drees, R. Godbole, and P. Roy, *Theory and Phenomenology of Sparticles: An Account of Four-Dimensional $N = 1$ Supersymmetry in High Energy Physics* (World Scientific, Hackensack, 2004).
- [46] C. Pallis, [arXiv:1812.10284](https://arxiv.org/abs/1812.10284).
- [47] P. Benetti Genolini, D. Cassani, D. Martelli, and J. Sparks, *J. High Energy Phys.* **02** (2017) 132.
- [48] O. S. An, J. U. Kang, J. C. Kim, and Y. H. Ko, *J. High Energy Phys.* **05** (2019) 146.