Spontaneous Breaking of Lepton Number and the Cosmological Domain Wall Problem

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We show that if global lepton number symmetry is spontaneously broken in a postinflation epoch, then it can lead to the formation of cosmological domain walls. This happens in the well-known "Majoron paradigm" for neutrino mass generation. We propose some realistic examples that allow spontaneous lepton number breaking to be safe from such domain walls.

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Introduction.—Topological defects such as monopoles, strings, and domain walls [1,2] can arise in many gauge theories including grand unification. In addition, there can appear (hybrid) configurations such as monopoles connected via strings or walls bounded by strings. Two well-known examples of the latter arise in SO(10) [3] and axion models [4]. Stable or sufficiently long lived domain walls, associated with symmetry breaking scales comparable to or larger than in the standard model (SM) will sconer or later become the dominant energy component of the early Universe. As a consequence, such domain walls pose a serious challenge in cosmology and should therefore be avoided in realistic model building (some possibilities were recently discussed in Ref. [5]).

Domain walls are well known to appear associated with the spontaneous breaking of the Peccei-Quinn symmetry [6]. Here we note that the "weak" $SU(2)_L$ may be associated with the presence of domain walls. This may happen in the context of spontaneous violation of lepton number symmetry. Indeed, such models in which the lepton number is violated by a gauge singlet Higgs vacuum expectation value (vev) [7,8] provides an attractive way to generate Majorana masses for neutrinos [9], as needed to account for current neutrino data [10]. In addition, it implies the existence of a physical Nambu-Goldstone boson, called Majoron. The latter may pick up a mass from explicit symmetry breaking by gravity effects [11–14]. Under such circumstances the Majoron may provide a good dark matter candidate [15–20] (since gravitational effects are not calculable in a reliable way, here we prefer not to invoke their existence).

The origin of the domain wall problem in this case stems from the existence of an unbroken residual subgroup Z_2 arising from the spontaneous lepton number violation, which clashes with the unbroken Z_3 from the nonperturbative instanton effects associated with the weak $SU(2)_{I}$. This implies that the domain wall problem associated with the weak $SU(2)_L$ exists in a broad class of Majoron models of neutrino mass generation. A standard mechanism for evading the domain wall problem is to invoke a suitable inflationary phase during their formation such that the walls are inflated away. In this Letter, we propose a more direct resolution of the domain wall problem that does not rely on inflation. We present various possible mechanisms for having realistic Majoron models, with and without supersymmetry, which allow spontaneous lepton number violation to occur without encountering a domain wall problem.

Global lepton number and domain wall problem.—Apart from the gauge symmetries, it is well know that in the SM there are two "accidental" global U(1) symmetries, namely the baryon number $U(1)_B$ and the lepton number $U(1)_L$ symmetries. Although, accidental within SM, these symmetries nonetheless play a very important role. The baryon number symmetry $U(1)_{B}$ is responsible for the stability of the proton and the lepton number symmetry plays a key role in neutrino mass generation and in determining the Dirac or Majorana nature of neutrinos. The lepton number in the SM is conserved at the Lagrangian level to all orders in perturbation theory. However, the lepton number is an anomalous symmetry; hence it is explicitly broken by nonperturbative effects [21]. (Note that both baryon and lepton numbers are anomalous symmetries; however a particular combination $U(1)_{B-L}$ is anomaly free. The other orthogonal combination $U(1)_{B+L}$ remains anomalous and

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hence it is explicitly broken by nonperturbative effects [21]). In particular, owing to the $[SU(2)_L]^2 \times U(1)_L$ anomaly, the nonperturbative instantons will explicitly break the initial lepton number symmetry $U(1)_L$ down to the discrete Z_N subgroup, with

$$N = \sum_{R} N(R) \times L(R) \times T(R) = 3 \times 1 \times 1 = 3, \quad (1)$$

where N(R) is the number of copies of a given fermion in representation R, L(R) is the lepton number of the fermion and T(R) is the $SU(2)_L$ Dynkin multiplicity index. For the $SU(2)_L$ group, the index T(R) for the lowest representations, singlets, doublets, and triplets are respectively, T(1) = 0, T(2) = 1, and T(3) = 4.

It is clear, from Eq. (1), that the nonperturbative instantons associated with the weak $SU(2)_L$ break $U(1)_L \rightarrow Z_3$. Notice also that the threefold family replication in the SM plays a crucial role in dictating the breaking $U(1)_L \rightarrow Z_3$. The residual Z_3 symmetry is exact at the classical and quantum level, implying the existence of degenerate vacua in our theory. Notice also that, in contrast to the case of axions, where the anomaly is related to the U(1) Peccei-Quinn charge assignments, in the case of the lepton number there is an anomaly intrinsically associated with the chiral nature of weak $SU(2)_L$.

Although the tunneling rate from one vacuum to another due to instantons is extremely small at zero temperature (this rate is proportional to $\exp(-2\pi/\alpha_W)$ and thus unimportant for our discussion [22]), sphaleron induced transitions between the vacua become relevant at higher temperatures [23]. Moreover, it is argued that frequent transitions between the vacua occur even above the critical temperature T_c for the electroweak transition (our argument below involves only temperatures between around 200 GeV and T_c , where the sphalerons operate, and the transitions from one vacuum to another are very frequent [24]). Thus, the *B* and *L* violating reactions at high temperatures are fast, so that the $U(1)_L$, $U(1)_B$ are explicitly broken by nonperturbative effects down to discrete Z_3 symmetries.

If the SM is the final gauge theory, the nonperturbative breaking of the lepton number won't be a serious issue. However, a dynamical understanding of the smallness of neutrinos mass often requires that lepton number is further broken down either explicitly or spontaneously by the new physics associated to neutrino mass generation. A popular and well studied scenario is the case of spontaneous breaking of lepton number [7,8]. This is specially attractive scenario that not only leads to Majorana masses, but also implies the existence of a Nambu-Goldstone boson, called Majoron. It breaks the global $U(1)_L$ lepton number symmetry down to a Z_2 subgroup through the vev of a $SU(3)_c \otimes SU(2)_L \otimes U(1)_Y$ singlet scalar carrying two units of lepton number. However, we notice the mismatch between the unbroken residual subgroup Z_2 arising from the spontaneous lepton number violation and the subgroup Z_3 that is left unbroken by the nonperturbative effects. Owing to this mismatch the domain walls will appear.

For temperatures between 200 GeV and the electroweak critical temperature T_c the tunneling rate between the vacua connected by the Z_3 subgroup, which remains explicitly unbroken by instantons, is very frequent. The barrier separating different vacua, related by the Z_3 , has static energy $E_{\rm sph}(T)$, the sphaleron mass, and the width is of order m_W^{-1} . This is the size of the "restricted instanton" that minimizes the height of the barrier and corresponds to the sphaleron (see Ref. [25] and references therein). Since the wall thickness is much smaller than the horizon size at these temperatures, the walls are expected to be present. The mass per unit surface is $\simeq v^2 m_W$, where v is the order parameter, i.e., the vev that breaks $U(1)_L$. Even one such wall per horizon would provide an energy density $\simeq 3v^2 m_W/4t$ (t is the cosmic time). This exceeds the radiation energy density $\rho_r = (\pi^2/30)g_*^{1/2}T^4$ ($g_* = 106.75$ is the effective number of degrees of freedom) at a cosmic temperature T > 200 GeV if the vev (v) is larger than about $(8\pi/3)^{1/2}(g_*/10)^{1/4}(m_W m_P)^{1/2} \simeq 8.2 \times 10^{10} \text{ GeV} \ (m_P =$ 2.44×10^{18} GeV is the reduced Planck mass). Such values for v are very reasonable if they are to generate, say, the right handed neutrino masses within a type-I seesaw mechanism [27–31]. Right after the wall domination, the Schwarzschild radius corresponding to the mass within each horizon becomes bigger than the horizon itself and the system becomes unstable and collapses into black holes, leading to a cosmological catastrophe [26]. Therefore, unless a suitable remedy is provided, we expect the standard high-scale type-I seesaw Majoron model of neutrino mass generation to be cosmologically inconsistent due to the existence of such domain walls. Note however that lowscale scenarios, such as the inverse seesaw Majoron schemes [32–35] constitute a potential way out. This is because, in that case, the lepton number violating order parameter can lie much below the electroweak scale, where sphaleron effects are negligible and $U(1)_L$ can be regarded as an exact continuous symmetry.

The above spontaneous breaking of $U(1)_L \rightarrow Z_2$ by the vev of a field carrying two units of the lepton number can be connected to neutrino masses in full generality at the operator level. Consider the $U(1)_L$ invariant effective operator

$$\frac{1}{\Lambda^2}\bar{L}^c H H \sigma L.$$
 (2)

In Eq. (2) the field *L* is the $SU(2)_L$ lepton doublet, *H* is the Higgs doublet and σ is a SM gauge singlet scalar field charged under the $U(1)_L$ symmetry. Also, Λ is the cutoff scale for the effective operator above which the full ultraviolet complete theory should be specified. This operator is $U(1)_L$ invariant if σ has charge -2 under the $U(1)_L$ symmetry. After σ develops a nonzero vev, $\langle \sigma \rangle$, $U(1)_L$ is broken down to Z_2 and the expression in Eq. (2) reduces to

the famous Weinberg operator [36]. Again the *CP* odd part of σ will be a Nambu-Goldstone boson, the Majoron. Here *CP* denotes the combined action of charge conjugation (*C*) and parity (*P*).

Solutions to the domain wall problem.—In this section we consider alternative solutions to the domain wall problem that arises from the spontaneous breaking of $U(1)_L$ by the vev of a lepton-number-carrying scalar field. We focus attention on Majoron-type models characterized by the spontaneous breaking of lepton number at a high scale. The examples in the following subsections involve only the SM gauge structure. On the other hand the model considered in the family symmetry subsection requires an extension of the SM with a gauge family symmetry.

Majoron with singlet-triplet seesaw: The simplest solution of the domain wall problem in the Majoron model uses only the usual SM gauge framework. It requires, in addition to the SM fields, the following new ones with their $SU(3)_c \otimes SU(2)_L \otimes U(1)_Y$ quantum numbers indicated in parenthesis and subscripts denoting their charges under $U(1)_L$:

$$\nu_R = (1,1,0)_{+1}, \quad \Sigma_R = (1,3,0)_{+1}, \quad \sigma = (1,1,0)_{-2}, \quad (3)$$

where the first field (ν_R) is a gauge singlet right-handed neutrino present in seesaw schemes [27–31] (with arbitrary multiplicity, which we take equal to one for simplicity, given that this is sufficient to account for the current neutrino oscillation data). The second field (Σ_R) is a $SU(2)_L$ triplet right-handed fermion and, the last field is the complex scalar whose vev $\langle \sigma \rangle$ is responsible for the spontaneous lepton number breaking. The Lagrangian will now contain the following new couplings:

$$\mathcal{L}_{\text{new}} = y_{\nu_R}^{Di} \bar{L}^i \tilde{H} \nu_R + y_{\Sigma}^{Di} \bar{L}^i \tilde{H} \Sigma_R + y_{\Sigma}^M \sigma \bar{\Sigma}_R^c \Sigma_R + y_{\nu_R}^M \sigma \bar{\nu}_R^c \nu_R,$$
(4)

where $\tilde{H} = i\tau_2 H^*$ with τ_2 denoting the second Pauli matrix.

After electroweak symmetry breaking the Higgs field will get a vev $\langle H \rangle = v$ and we will have a seesaw-like mechanism for light neutrinos with mass matrix $m_{\nu} = M_D^T M_R^{-1} M_D$ where

$$M_D = \begin{pmatrix} v y_{\nu}^{D1} & v y_{\nu}^{D2} & v y_{\nu}^{D3} \\ v y_{\Sigma}^{D1} & v y_{\Sigma}^{D2} & v y_{\Sigma}^{D3} \end{pmatrix},$$

$$M_R = \begin{pmatrix} y_{\nu_R}^M \langle \sigma \rangle & 0 \\ 0 & y_{\Sigma}^M \langle \sigma \rangle \end{pmatrix}.$$
 (5)

The resulting matrix, m_{ν} , has rank 2, leaving one light neutrino massless. Note that, since Σ_R has nontrivial $SU(2)_L$ quantum numbers, it produces a significant change in the $[SU(2)_L]^2 \times U(1)_L$ anomaly, which is now given by

$$N = \sum_{R} N(R) \times L(R) \times T(R)$$

= 3 × 1 × 1 - 1 × 1 × 4 = -1. (6)

By computing the anomaly factor one sees that the domain wall problem is absent in this extension. Therefore, the heavy triplet Σ_R acts as an auxiliary Majorana field to address the domain wall issue. Moreover, it also acts as a heavy messenger for small neutrino mass generation through the seesaw mechanism.

Majoron seesaw within supersymmetry: The simple solution illustrated in the previous section can be generalized within a supersymmetric context. We present here a simple supersymmetric model that also addresses the domain wall problem. The particle content and charges of the superfields are as shown in Table I.

In addition to the usual minimal supersymmetric standard model (MSSM) superfields and the right-handed neutrinos (ν^c), one adds the $SU(2)_L$ triplet superfields T, \overline{T} and the gauge singlet superfields S, ϕ , $\overline{\phi}$ with charges as listed in Table I. The superpotential of our model is given by

$$\mathcal{W} = \kappa S(\phi\phi - M^2) + y_{ij}^u H_u Q_i u_j^c + y_{ij}^d H_d Q_i d_j^c$$

+ $y_{ij}^v H_u L_i \nu_j^c + y_{ij}^e H_d L_i e_j^c + \lambda S H_u H_d + y_i^T T L_i H_d$
+ $y'^T \bar{\phi} T \bar{T} + y_{ij}^\phi \frac{\bar{\phi}^2 \nu_i^c \nu_j^c}{m_P},$ (7)

where i, j = 1, 2, 3 are generation indices.

Owing to the presence of the triplet superfield T, the $[SU(2)_L]^2 \times U(1)_L$ anomaly is again found to be

$$N = \sum_{R} N(R) \times L(R) \times T(R)$$

= 3 × 1 × 1 - 1 × 1 × 4 = -1. (8)

TABLE I. Particle content and charges. $U(1)_R$ is an *R* symmetry under which the superpotential W has an *R* charge of 2 units.

Superfields	$SU(3)_c$	$SU(2)_L$	$U(1)_Y$	$U(1)_B$	$U(1)_L$	$U(1)_R$
$\overline{Q_i}$	3	2	1/6	1/3	0	1
u_i^c	3	1	-2/3	-1/3	0	1
d_i^c	3	1	1/3	-1/3	0	1
L_i	1	2	-1/2	0	1	1
e_i^c	1	1	1	0	-1	1
ν_i^c	1	1	0	0	-1	1
Ť	1	3	1	0	-1	1
\overline{T}	1	3	-1	0	0	1
H_{u}	1	2	1/2	0	0	0
H_d	1	2	-1/2	0	0	0
S	1	1	0	0	0	2
ϕ	1	1	0	0	-1	0
$\bar{\phi}$	1	1	0	0	1	0

Thus, unlike the usual Majoron models, here the instanton effects will break $U(1)_L \rightarrow Z_1$, avoiding the domain wall problem. As in the previous case, this holds irrespective of the number of right-handed neutrino superfields, the minimal realistic model has just one.

Notice that this solution differs from the standard seesaw mechanism in that the Majoron coming from the imaginary parts of the $\phi, \bar{\phi}$ scalars carry one unit of lepton number, instead of two. Moreover, our model has other attractive features which make it quite appealing. Apart from solving the domain wall problem, it automatically addresses the so-called μ problem of the MSSM [37]. In addition we also have a *R* symmetry that contains the usual *R* parity of the MSSM, forbidding all the potentially dangerous terms in the superpotential [Eq. (7)]. Finally, right-handed neutrino masses arise through the nonrenormalizable term $\bar{\phi}^2 \nu^c \nu^c / m_P$, where we take the high scale as m_P .

 $SU(3)_{lep}$ family symmetry for leptons: Consider now a $SU(3)_{lep}$ gauge extension of the SM scenario. Let quarks be singlets under this group, while leptons transform under it in a vector-like way (note that this differs from the usual $SU(3)_{lep}$ family symmetry used to address the observed fermion mass hierarchy [38,39]),

$$L = (1, 2, -1/2, 3),$$

$$e_R = (1, 1, -1, 3),$$

$$\nu_R = (1, 1, 0, 3),$$
(9)

with the first three entries in parenthesis indicating the standard model charges and the last entry the $SU(3)_{\text{lep}}$ representation. This extension has several consequences. First of all, right-handed neutrinos cannot have a bare mass term. Their masses must be generated through the spontaneous violation of $U(1)_{\text{L}}$. This is related with the breaking of $SU(3)_{\text{lep}}$ and is achieved by the vev of a flavor sextet scalar field σ with lepton number -2 via the coupling

$$\sigma \bar{\nu}_R^c \nu_R. \tag{10}$$

The second and more important implication is that this extension automatically solves the domain wall problem. The reason is that the center of $SU(3)_{\text{lep}}$, which is Z_3 , exactly coincides with the discrete Z_3 subgroup of $U(1)_L$ left unbroken by the anomaly. Since this accidental subgroup can be embedded in the continuous gauge group $SU(3)_{\text{lep}}$, the degenerate minima are now connected by a gauge transformation, so that any difference among them becomes unphysical. In this way, the domain wall problem is solved. This is a Majoron variant of the domain wall axion solution given in the context of Grand Unified Theory (GUT) in Ref. [40,41].

Diracon solution: Another possible solution to the domain wall problem is obtained by enforcing that the spontaneous lepton number breaking is such that $U(1)_L \rightarrow Z_3$ instead of Z_2 . In this case, there is no mismatch between the residual subgroup preserved by the anomaly and that preserved by the spontaneous lepton number violation due to $\langle \sigma \rangle$, so the domain wall problem will be automatically solved. Clearly, the $U(1)_L \rightarrow Z_3$ spontaneous breaking cannot be accomplished within the framework of the canonical Majoron model. In fact, if Z_3 is the residual unbroken symmetry then neutrinos cannot be Majorana particles. However, we note that for Dirac neutrinos the $U(1)_L \rightarrow Z_3$ breaking is viable, and will lead to a solution of the domain wall problem within a variant of the "Diracon models" [42,43].

To see this Diracon solution, the first thing is to realize that the lepton number of right-handed neutrinos ν_R need not be the same as that of the left-handed neutrinos [44,45]. In fact, a nonconventional lepton number assignment of (4,4,-5) for the three generations of $\nu_{i,R}$; i = 1, 2, 3, proposed in Refs. [46,47] is equally acceptable.

If the $\nu_{i,R}$ transform with such nonconventional charges under $U(1)_L$, then one cannot write down the tree level Dirac term $\tilde{L} \tilde{H} \nu_{i,R}$ nor the Majoron Weinberg operator of Eq. (2). However, one can still write down the following $U(1)_L$ invariant operators

$$\frac{1}{\Lambda} \bar{L} \tilde{H} \chi \nu_{i,R}, \qquad \frac{1}{\Lambda^2} \bar{L} \tilde{H} \chi^* \chi^* \nu_{3,R}, \qquad (11)$$

where $\nu_{i,R}$, i = 1, 2 are the two right-handed neutrinos carrying charge 4 units under $U(1)_L$, and $\nu_{3,R}$ has $U(1)_L$ charge of -5. Also, the field χ has charge of -3 under $U(1)_L$. It can be easily seen that the vev of the χ field will spontaneously break $U(1)_L \rightarrow Z_3$ with the resulting neutrinos being Dirac in nature. Furthermore, the *CP* odd part of χ will be a Nambu-Goldstone boson, which we call Diracon, and is associated with the Dirac mass generation of the neutrinos. Now, since the $U(1)_L$ in this case is spontaneously broken to the same residual subgroup Z_3 as that preserved by the nonperturbative $SU(2)_L$ instantons, there is no mismatch and hence the problem of domain walls is automatically avoided.

Conclusions.—We have shown that if the global lepton number symmetry is broken spontaneously in a post inflationary epoch, then it can lead to the formation of cosmological domain walls. Since the presence of these domain walls may spoil the standard picture of cosmological evolution, we have studied the conditions to prevent their formation as a result of spontaneous symmetry breaking. We have shown that the simplest seesaw Majoron models of neutrino masses have, in principle, a domain wall problem associated with the chiral $SU(2)_L$ gauge group describing the weak interaction. We have also provided some explicit and realistic solutions that allow a safe spontaneous breaking of lepton number, free of domain walls. Some of these models involve new particles that could potentially lead to phenomenological implications.

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- Ya. B. Zeldovich, I. Yu. Kobzarev, and L. B. Okun, Zh. Eksp. Teor. Fiz. 67, 3 (1974) [Sov. Phys. JETP 40, 1 (1974)].
- [2] A. Vilenkin and A. E. Everett, Phys. Rev. Lett. 48, 1867 (1982).
- [3] T. W. B. Kibble, G. Lazarides, and Q. Shafi, Phys. Lett. 113B, 237 (1982).
- [4] T. W. B. Kibble, G. Lazarides, and Q. Shafi, Phys. Rev. D 26, 435 (1982).
- [5] R. Sato, F. Takahashi, and M. Yamada, Phys. Rev. D 98, 043535 (2018).
- [6] P. Sikivie, Phys. Rev. Lett. 48, 1156 (1982).
- [7] Y. Chikashige, R. N. Mohapatra, and R. D. Peccei, Phys. Lett. 98B, 265 (1981).
- [8] J. Schechter and J. W. F. Valle, Phys. Rev. D 25, 774 (1982).
- [9] J. W. F. Valle and J. C. Romao, *Neutrinos in High Energy* and Astroparticle Physics (John Wiley & Sons, New York, 2015).
- [10] P. F. de Salas, D. V. Forero, C. A. Ternes, M. Tórtola, and J. W. F. Valle, Phys. Lett. B 782, 633 (2018).
- [11] I. Z. Rothstein, K. S. Babu, and D. Seckel, Nucl. Phys. B403, 725 (1993).
- [12] E. K. Akhmedov, Z. G. Berezhiani, G. Senjanovic, and Z.-j. Tao, Phys. Rev. D 47, 3245 (1993).
- [13] E. K. Akhmedov, Z. G. Berezhiani, R. N. Mohapatra, and G. Senjanovic, Phys. Lett. B 299, 90 (1993).
- [14] S. F. King and P. O. Ludl, J. High Energy Phys. 03 (2017) 174.
- [15] V. Berezinsky and J. W. F. Valle, Phys. Lett. B 318, 360 (1993).
- [16] M. Lattanzi and J. W. F. Valle, Phys. Rev. Lett. 99, 121301 (2007).

- [17] F. Bazzocchi, M. Lattanzi, S. Riemer-Sørensen, and J. W. F. Valle, J. Cosmol. Astropart. Phys. 08 (2008) 013.
- [18] M. Lattanzi, S. Riemer-Sørensen, M. Tórtola, and J. W. F. Valle, Phys. Rev. D 88, 063528 (2013).
- [19] M. Lattanzi, R. A. Lineros, and M. Taoso, New J. Phys. 16, 125012 (2014).
- [20] J.-L. Kuo, M. Lattanzi, K. Cheung, and J. W. F. Valle, J. Cosmol. Astropart. Phys. 12 (2018) 026.
- [21] G. 't Hooft, Phys. Rev. Lett. 37, 8 (1976).
- [22] D. J. Gross, R. D. Pisarski, and L. G. Yaffe, Rev. Mod. Phys. 53, 43 (1981).
- [23] V. A. Kuzmin, V. A. Rubakov, and M. E. Shaposhnikov, Phys. Lett. **155B**, 36 (1985).
- [24] P.B. Arnold and L. D. McLerran, Phys. Rev. D 37, 1020 (1988).
- [25] G. Lazarides, Lect. Notes Phys. 592, 351 (2002).
- [26] A. Vilenkin, Phys. Rev. D 23, 852 (1981).
- [27] P. Minkowski, Phys. Lett. 67B, 421 (1977).
- [28] M. Gell-Mann, P. Ramond, and R. Slansky, Conf. Proc. C790927, 315 (1979).
- [29] T. Yanagida, Conf. Proc. C7902131, 95 (1979).
- [30] J. Schechter and J. W. F. Valle, Phys. Rev. D 22, 2227 (1980).
- [31] R. N. Mohapatra and G. Senjanovic, Phys. Rev. Lett. 44, 912 (1980).
- [32] R. N. Mohapatra and J. W. F. Valle, Phys. Rev. D 34, 1642 (1986).
- [33] M. C. Gonzalez-Garcia and J. W. F. Valle, Phys. Lett. B 216, 360 (1989).
- [34] A. S. Joshipura and J. W. F. Valle, Nucl. Phys. B397, 105 (1993).
- [35] Z. G. Berezhiani, A. Yu. Smirnov, and J. W. F. Valle, Phys. Lett. B 291, 99 (1992).
- [36] S. Weinberg, Phys. Rev. D 22, 1694 (1980).
- [37] G. R. Dvali, G. Lazarides, and Q. Shafi, Phys. Lett. B 424, 259 (1998).
- [38] Z. G. Berezhiani, Phys. Lett. 129B, 99 (1983).
- [39] Z. G. Berezhiani and M. Yu. Khlopov, Yad. Fiz. 51, 1157 (1990) [Sov. J. Nucl. Phys. 51, 739 (1990)].
- [40] G. Lazarides and Q. Shafi, Phys. Lett. 115B, 21 (1982).
- [41] S. M. Barr, D. B. Reiss, and A. Zee, Phys. Lett. 116B, 227 (1982).
- [42] C. Bonilla and J. W. F. Valle, Phys. Lett. B 762, 162 (2016).
- [43] C. Bonilla, E. Ma, E. Peinado, and J. W. F. Valle, Phys. Lett. B 762, 214 (2016).
- [44] L. Bento and J. W. F. Valle, Phys. Lett. B 264, 373 (1991).
- [45] J. Peltoniemi, D. Tommasini, and J. W. F. Valle, Phys. Lett. B 298, 383 (1993).
- [46] E. Ma and R. Srivastava, Phys. Lett. B 741, 217 (2015).
- [47] E. Ma, N. Pollard, R. Srivastava, and M. Zakeri, Phys. Lett. B 750, 135 (2015).